

# Precise time synchronisation of autonomous radio stations at the Pierre Auger Observatory

Qader Dorosti for the Pierre Auger Observatory

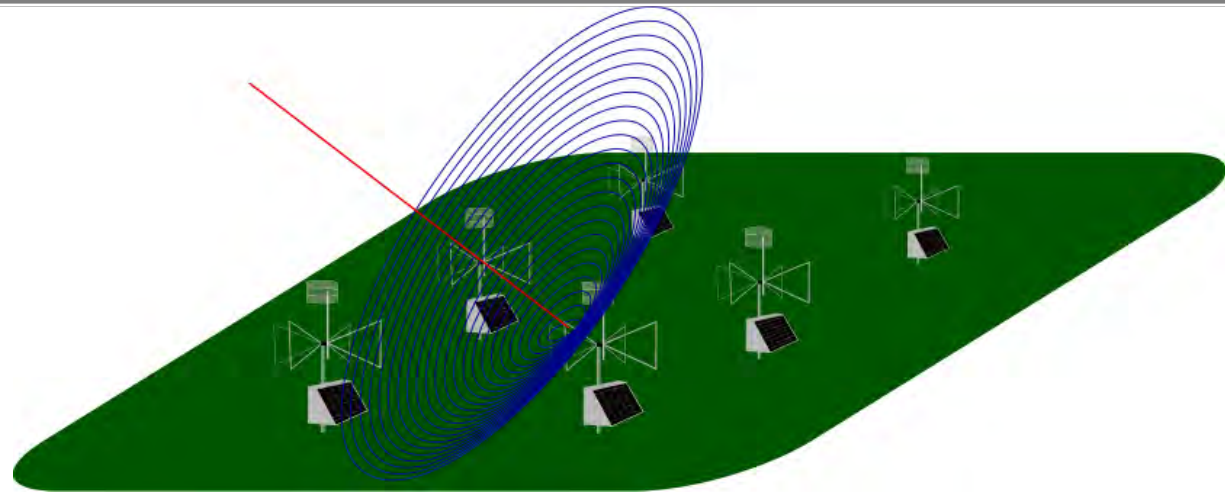
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## Outline:

- Motivation
- Pierre Auger Observatory
- Time Calibration
- Results

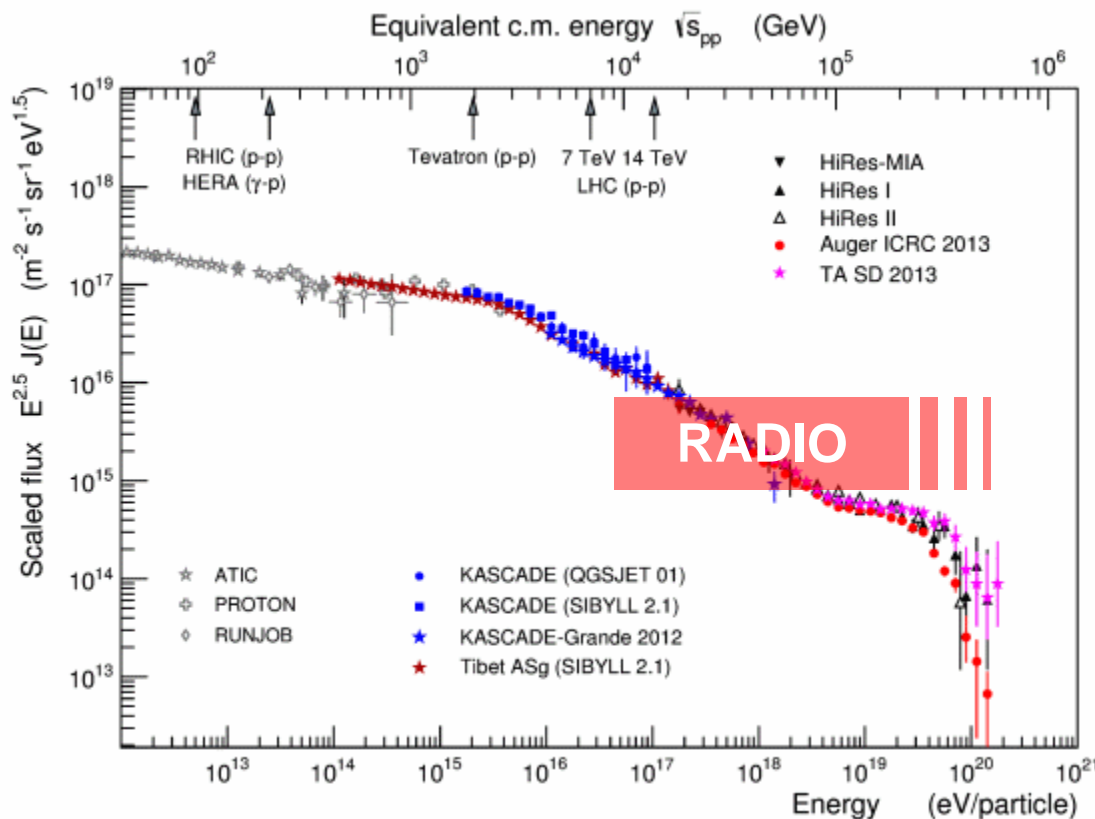
Take home message:

Time resolution < **2 ns**



# Cosmic-ray physics

## Energy spectrum of cosmic rays



Plot by R. Engel & T. Huege

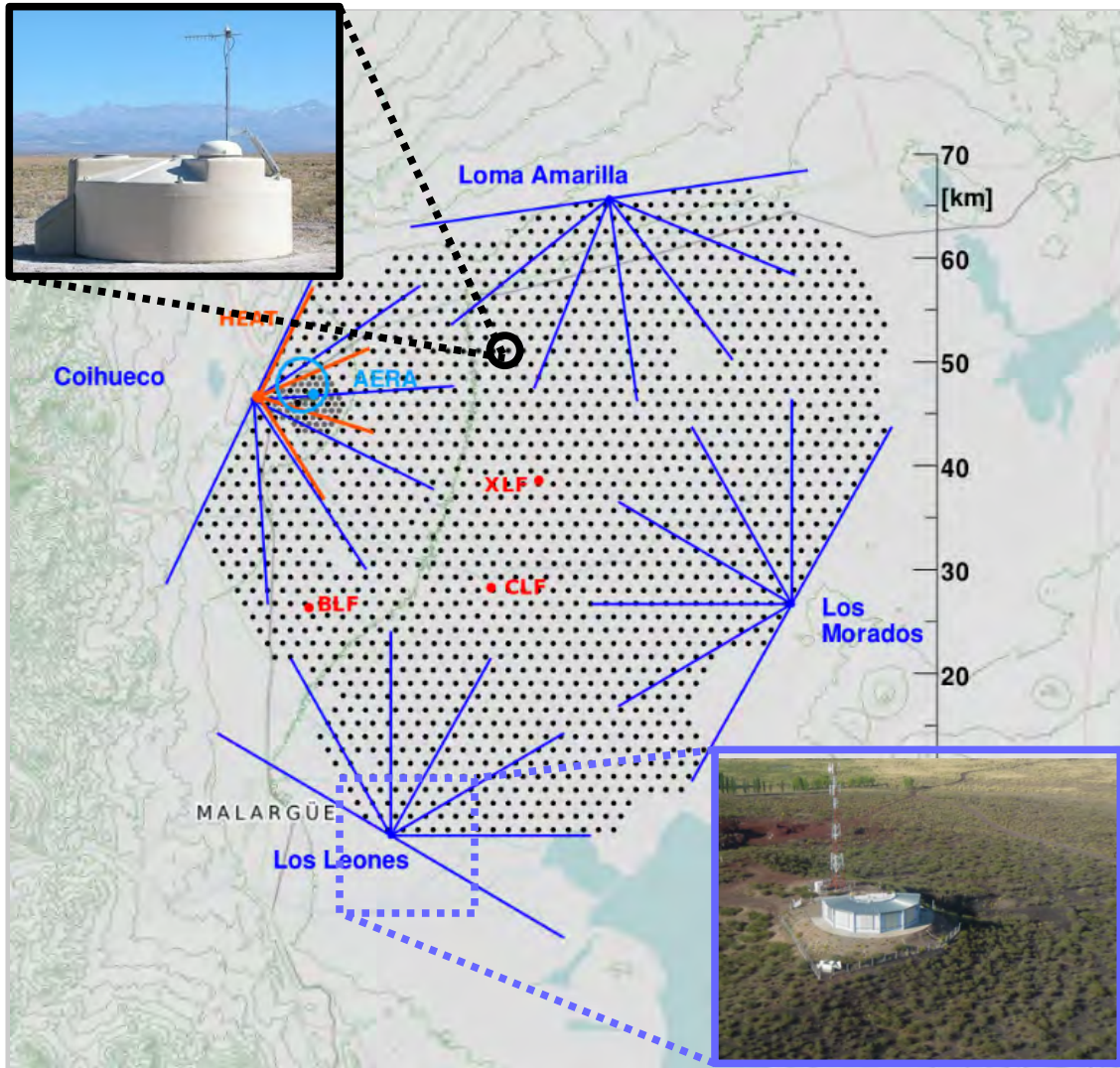
- Intriguing science questions:
  - Origin of cosmic rays  
→  **$10^{20}$  eV ?**
  - Acceleration mechanism?

- Answer requires measurements:
  - Direction
  - Energy
  - Composition

- We need large exposure
- We need precise calibration

# Pierre Auger Observatory

## Layout of the Pierre Auger Observatory



**Location:** Argentina,  
Mendoza, Malargue



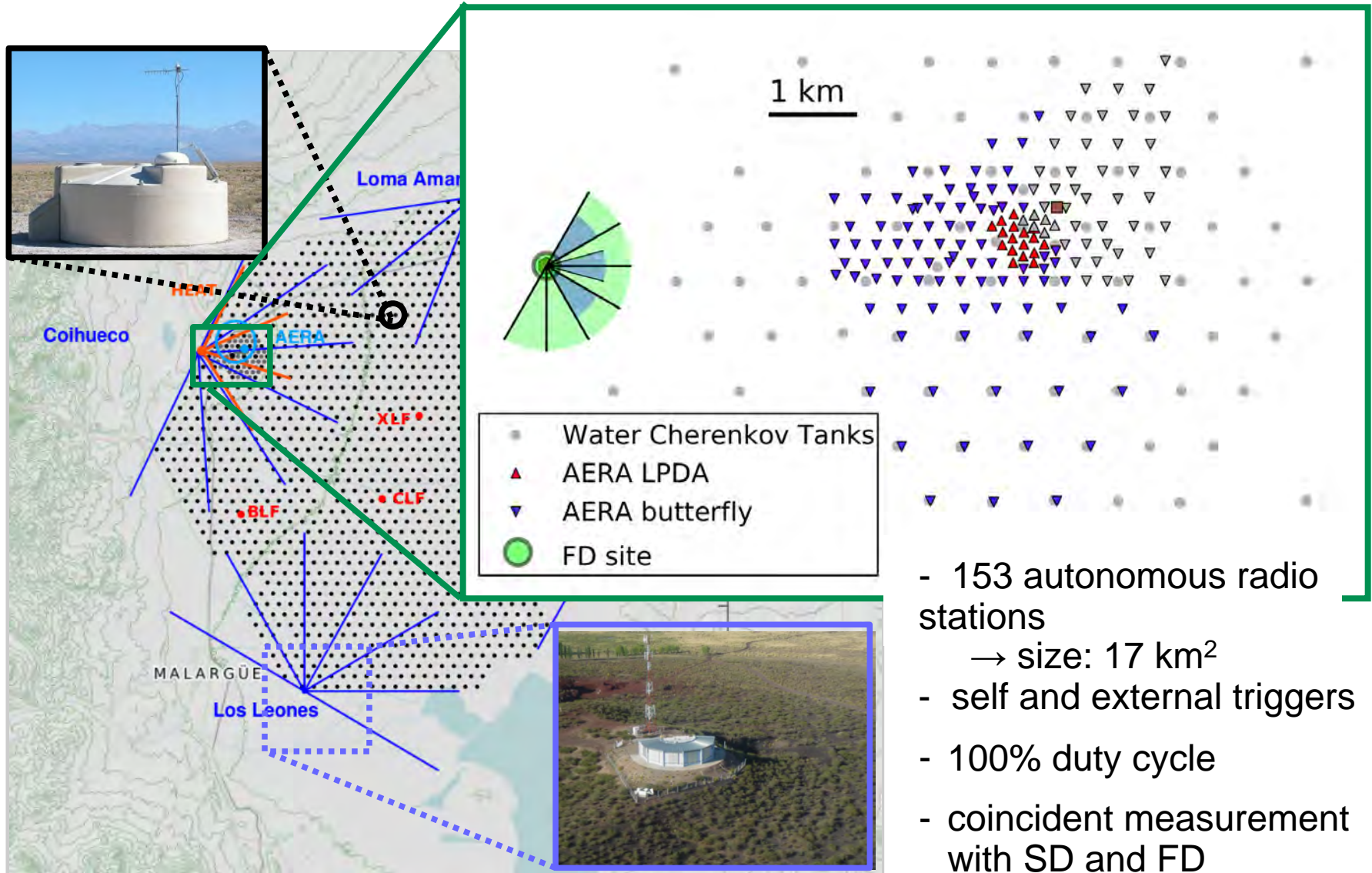
### Surface Detectors (SD)

- 1660 Cherenkov tanks
- 100% duty cycle
- High angular resolution

### Fluorescence Detector (FD)

- 27 telescopes
- 15% duty cycle
- Composition measurement

# Auger Engineering Radio Array (AERA)



# AERA stations

Log-periodic dipole antenna (LPDA)



Butterfly antenna



- 1 GPS clock / station for timing
- Wireless communication antenna
- frequency range: 30-80 MHz
- Differential GPS: position accuracy better than 10 cm

# HAP supports of radio detection of cosmic rays

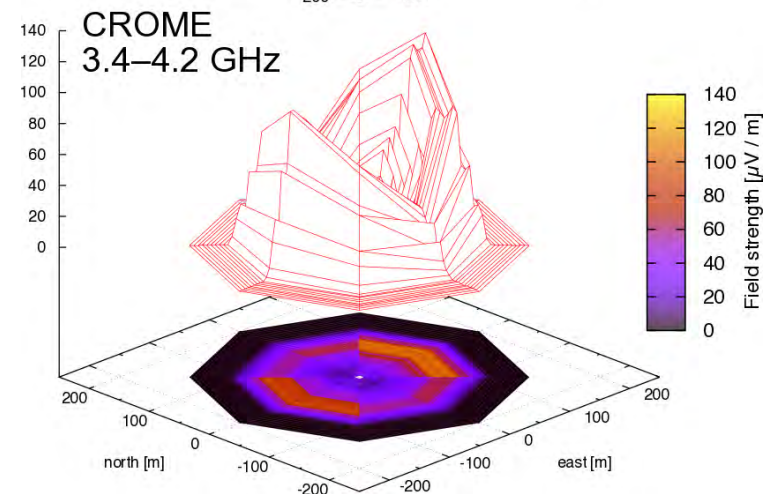
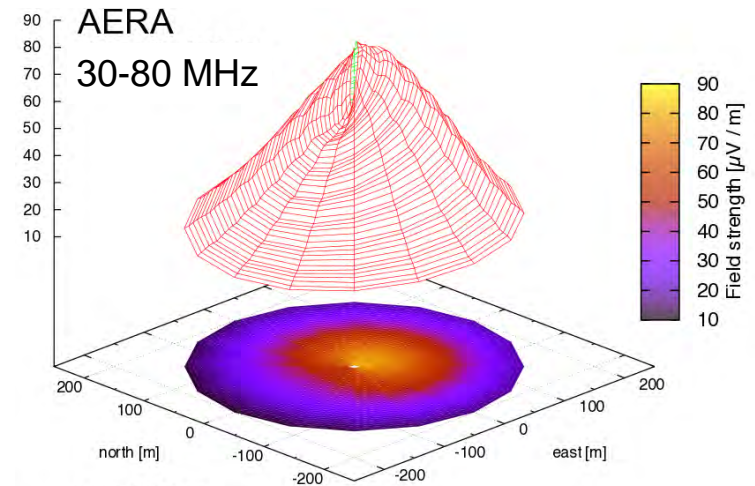
Tuna-Rex antenna



CROME antenna



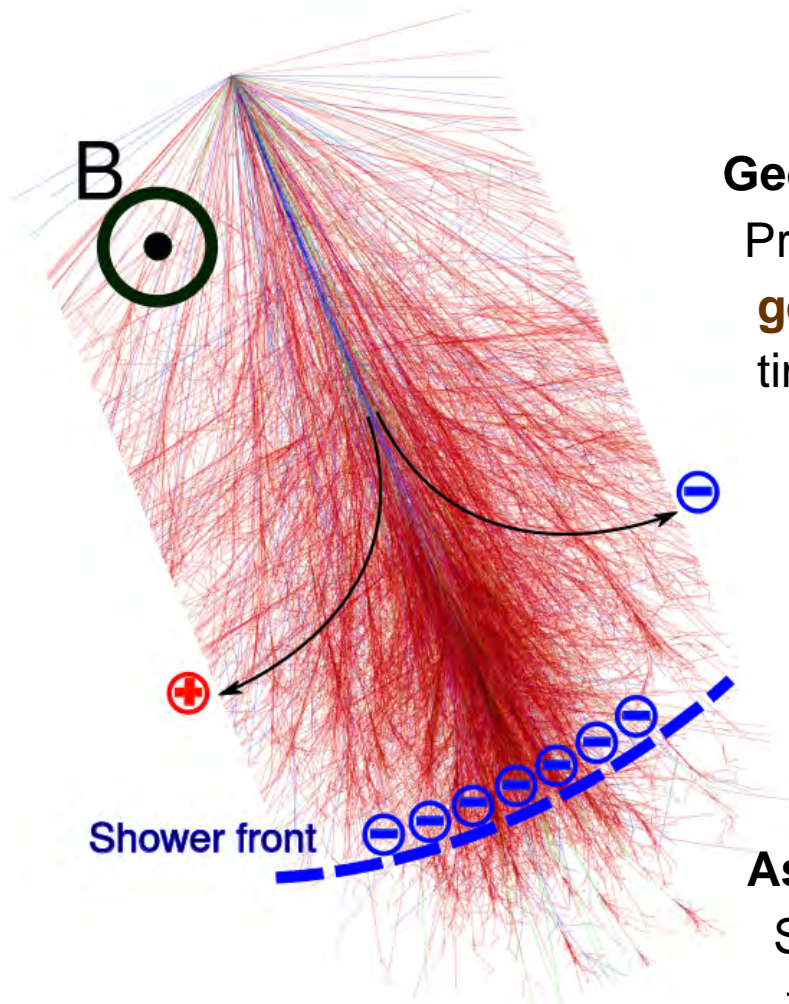
simulated radio emission for different frequency ranges



## Objectives:

- Scalability of cosmic-ray radio detectors to large arrays
- Investigating radio detection in different frequency ranges
- Initial investigation for radio antenna at South Pole (RASTA)

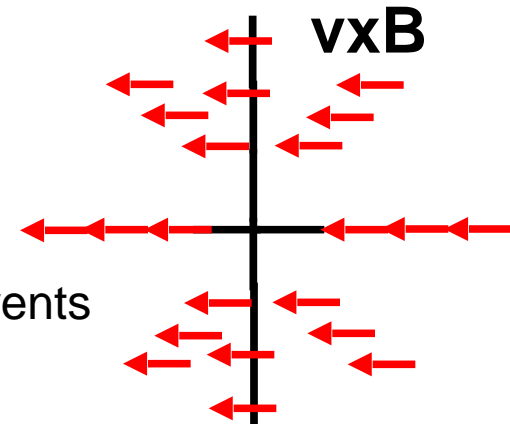
# Radio emission mechanism



## Geomagnetic effect

Primary effect:

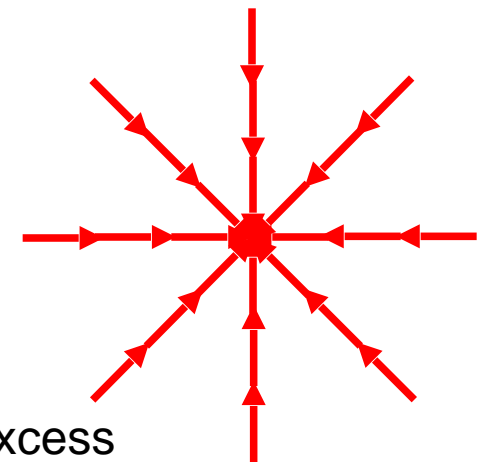
**geomagnetic field** induces  
time-varying transverse currents



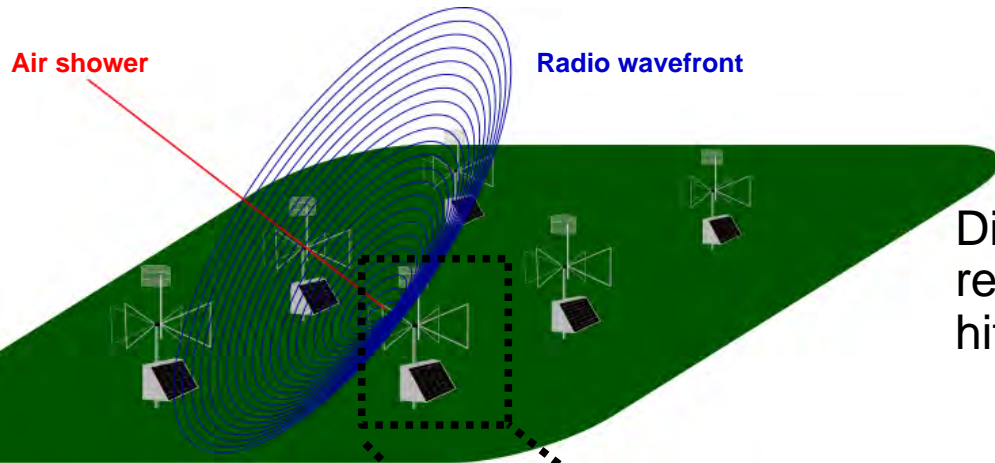
## Askaryan effect

Secondary effect:

time-varying **net charge** excess

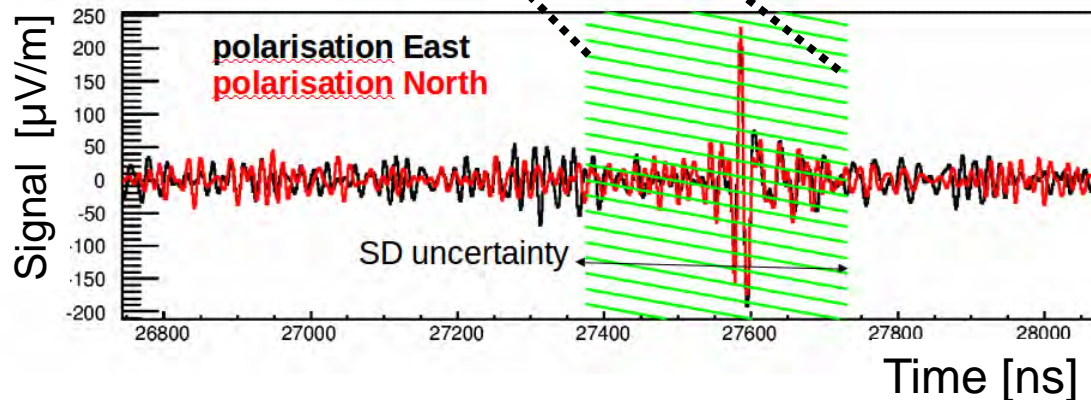


# Radio detection principle



Direction of **air shower** is reconstructed from **radio wavefront** hitting 2D array of **radio stations**

Typical signal trace recorded by radio stations



- Aims to measure:
  - In self and external trigger modes
  - Arrival direction
  - Energy
  - Mass composition
  - Radio interferometry

To exploit the full potential of radio measurement timing of **1 ns** is ideal

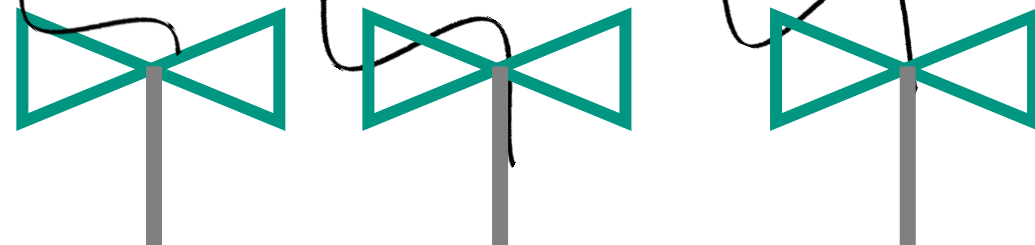


# Time calibration of AERA

## Beacon transmitter



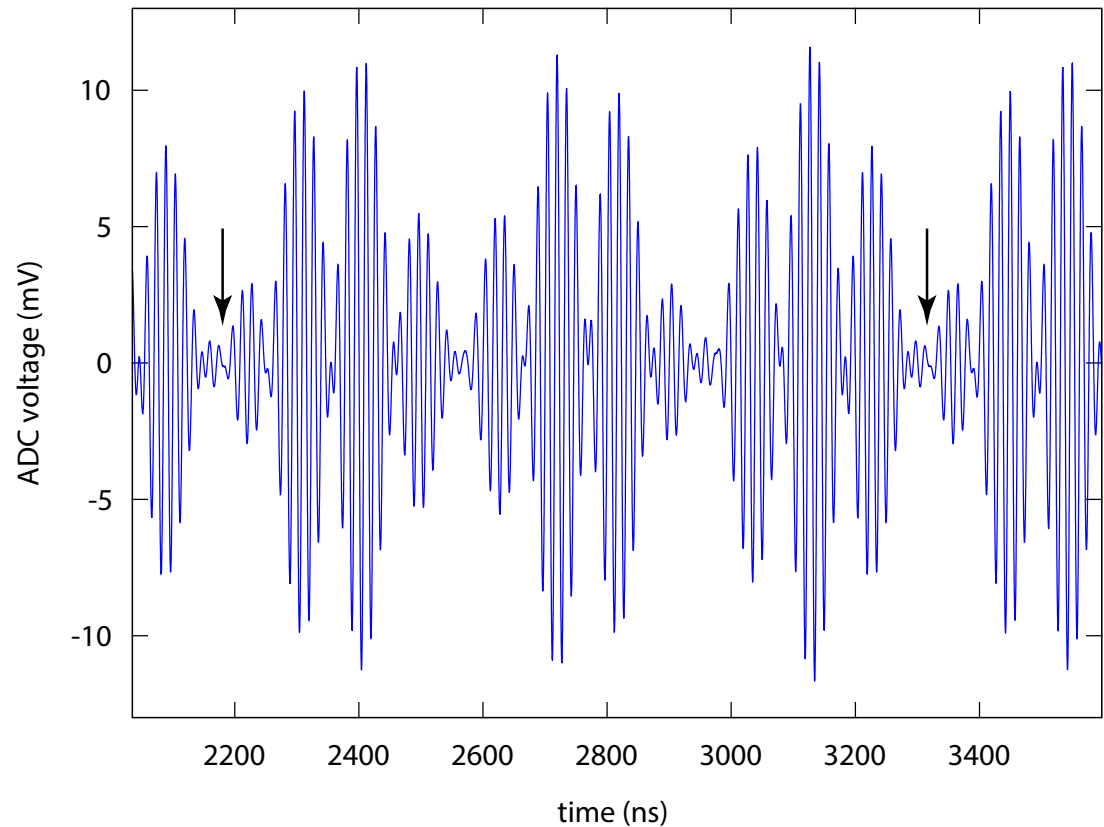
- deployed 3 km west of AERA
- consists of a passive emitting antenna
- generates constant wave emission:  
4 frequencies in **30-80 MHz** range  
→ inside the measurement band of AERA



# Principle of time calibration

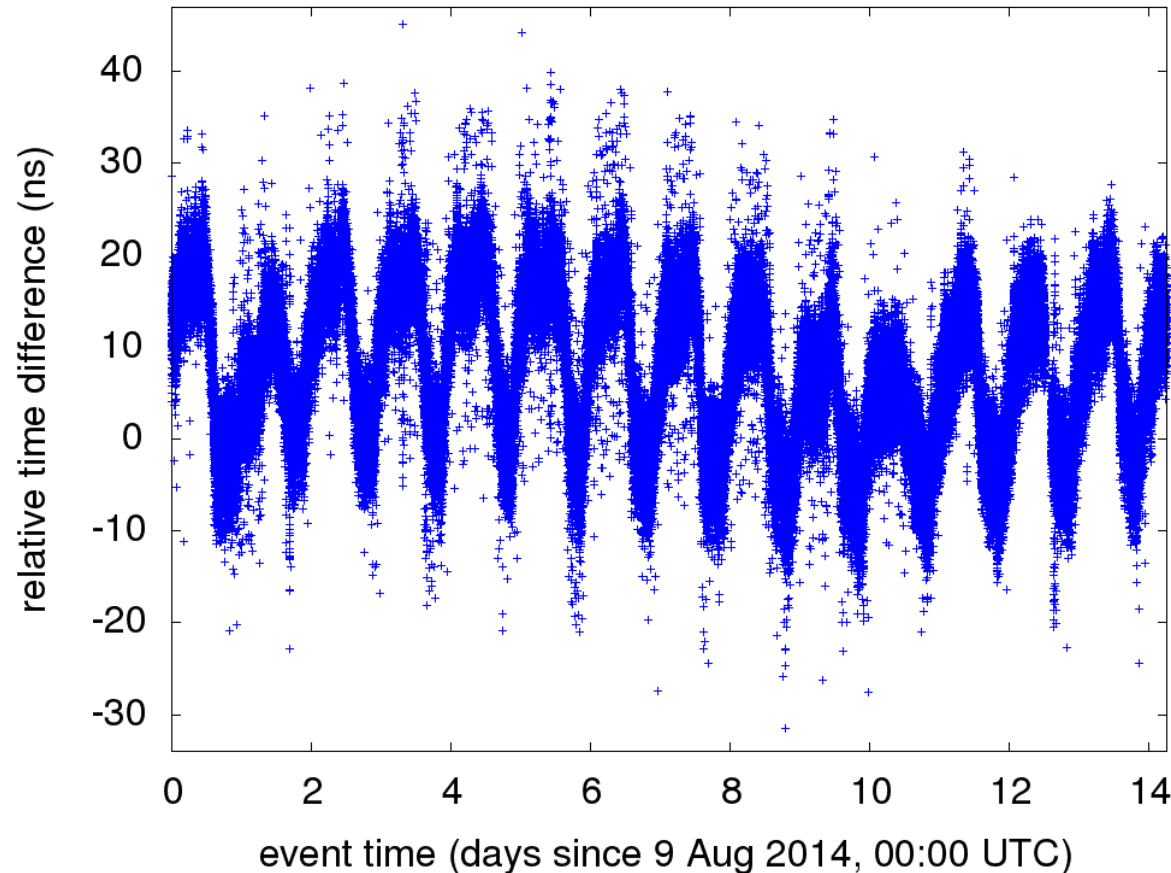
## Measured time trace of a typical beacon signal

- 4 sine waves beat repeats every  $1.1 \mu\text{s}$  < AERA time trace length, i.e.  $10 \mu\text{s}$
- compare arrival time of beat with expected propagation time
- calculate relative time difference



# Performance of the time calibration

Typical time difference between radio stations

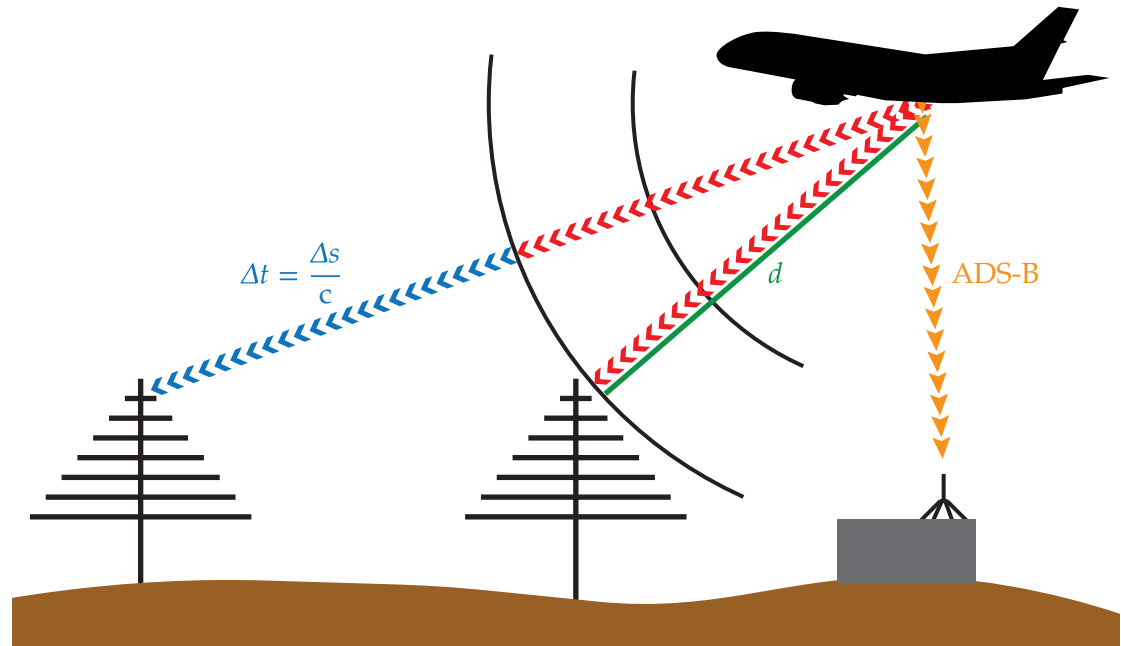


Strong fluctuations of a few 10 ns between stations: due to **GPS time offsets**?  
 Can beacon simply correct for this?

→ we need independent cross check

# Airplane: independent method

- Some commercial airplanes send radio pulses which can be detected by AERA
- Airplanes also send ADS-B (Automatic Dependent Surveillance - Broadcast) data which can be used to determine their positions

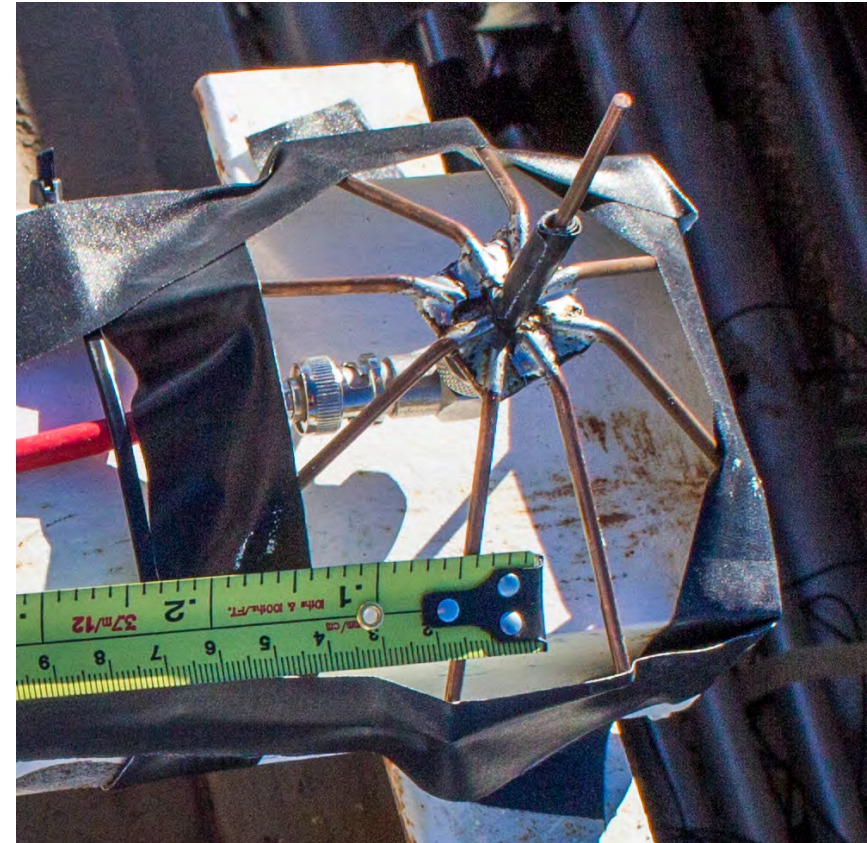


- estimate the pulse arrival time
- cross check the time calibration of beacon method

# Determination of airplane position

- ADS-B transmitted at 1090 MHz at rate of 0.5 – 1 Hz
- contains information, e.g. latitude, longitude, altitude, heading and speed of the airplane
- can be received with equipments costing less than 20 USD

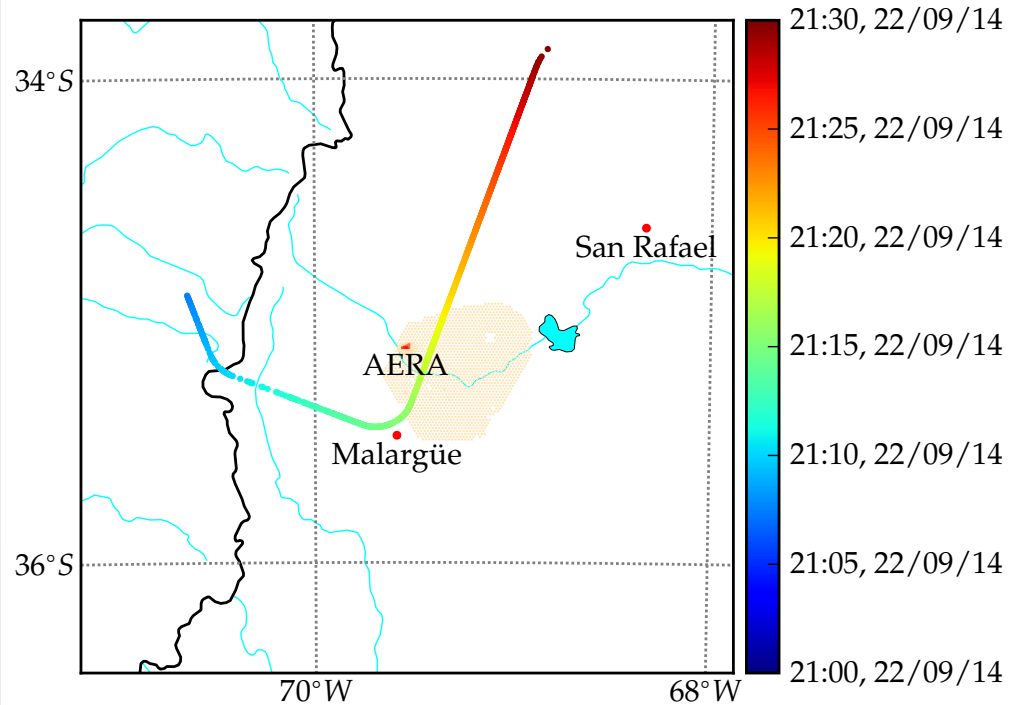
ADS-B receiver antenna



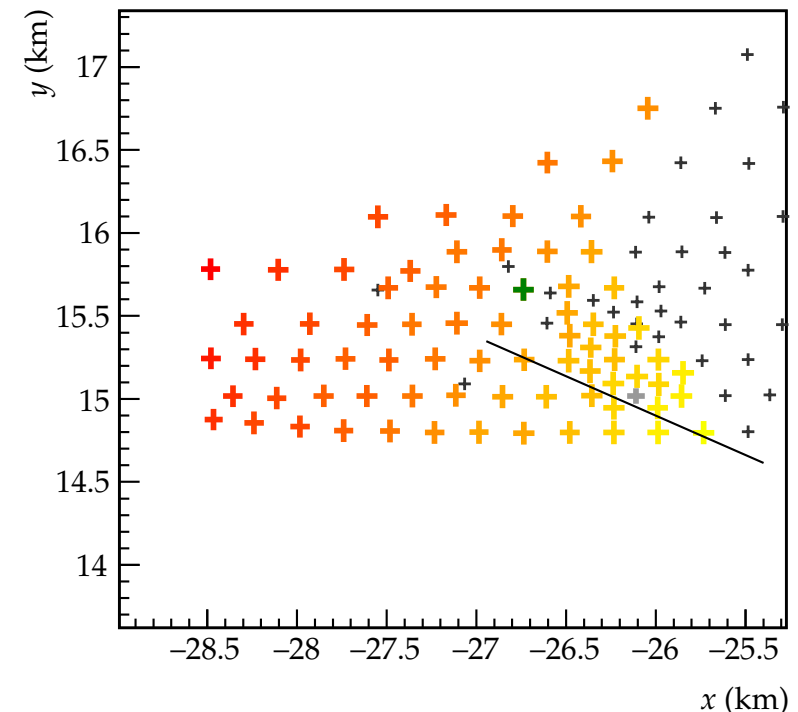
We deployed an ADS-B receiver antenna at the AERA field  
→ range of detection 400 km

# An example event

An airplane trajectory reconstructed from ADS-B data



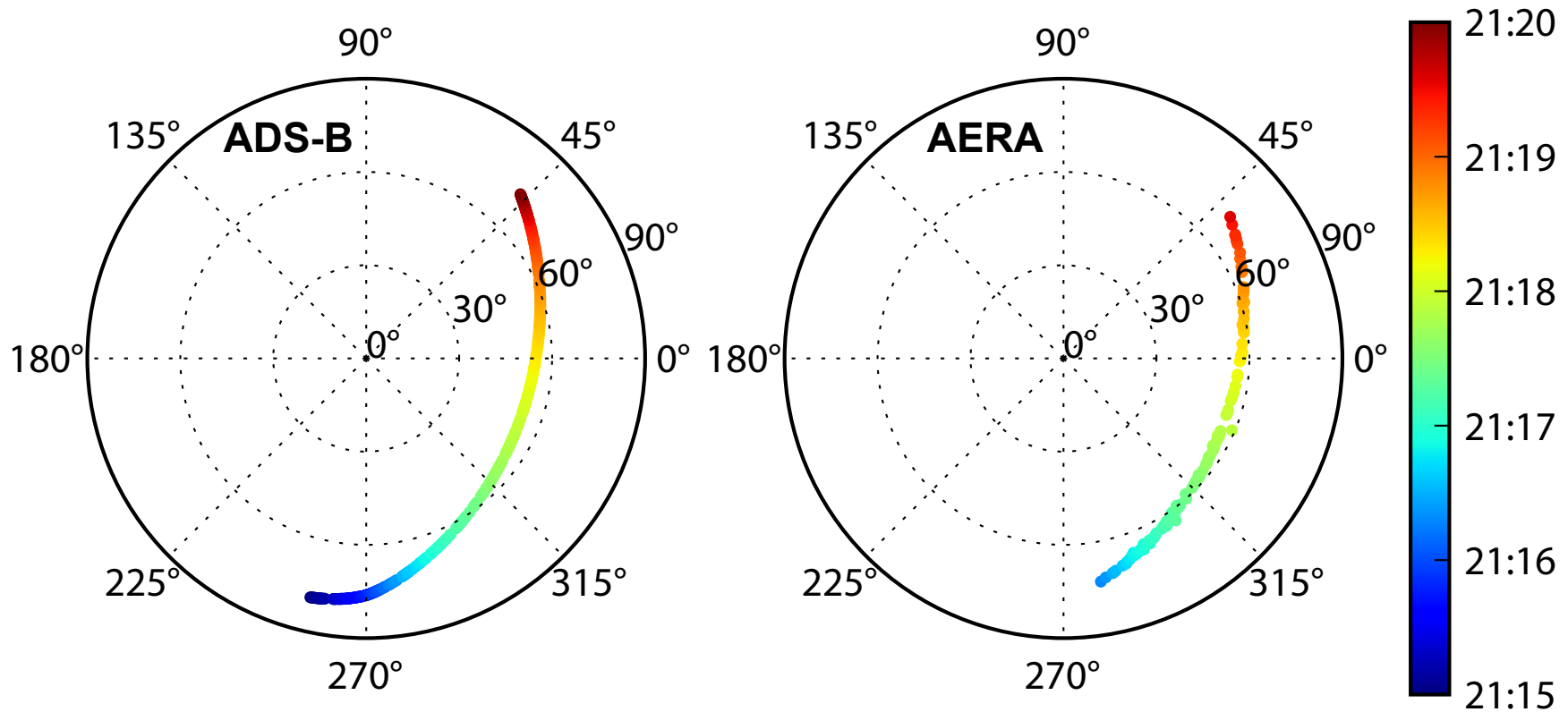
The airplane pulse measured by AERA



By combining the real-time position information from ADS-B  
and the radio pulses emitted by airplanes

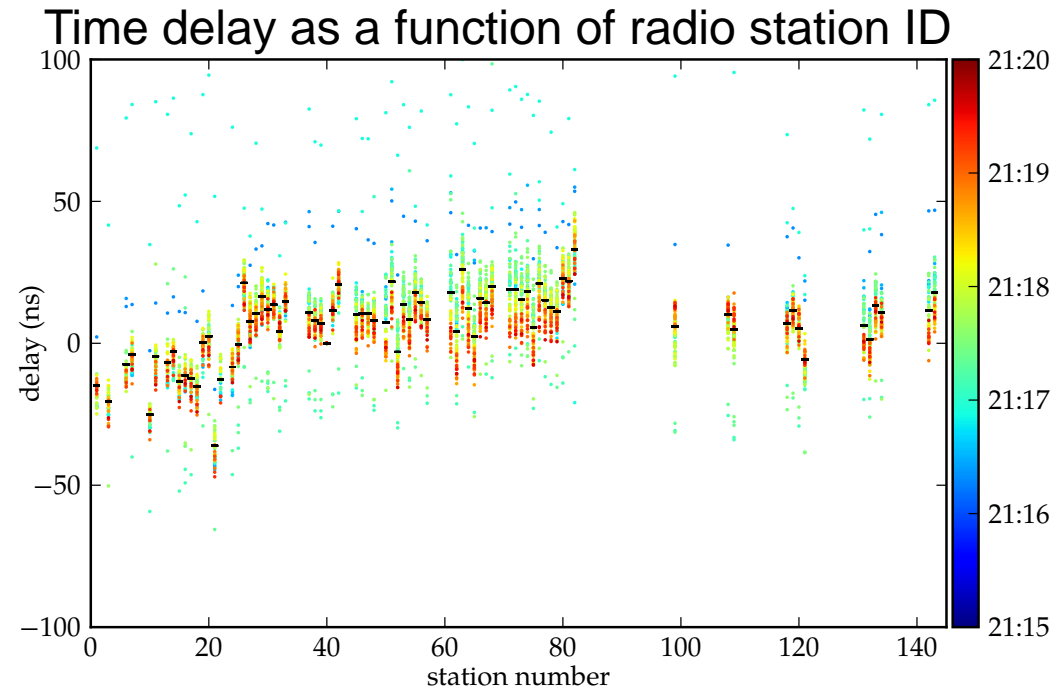
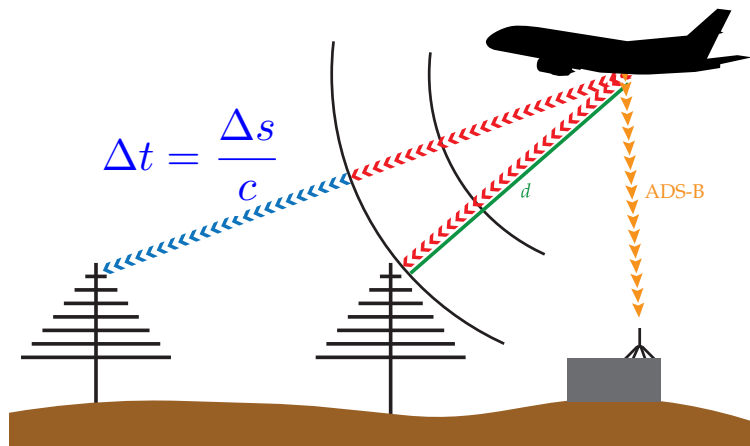
→ **time offsets** between GPS clocks can be calculated

# Comparison between ADS-B and AERA data



ADS-B and AERA reconstructed airplanes are highly correlated

# Performance of the airplane time calibration



Time delay:

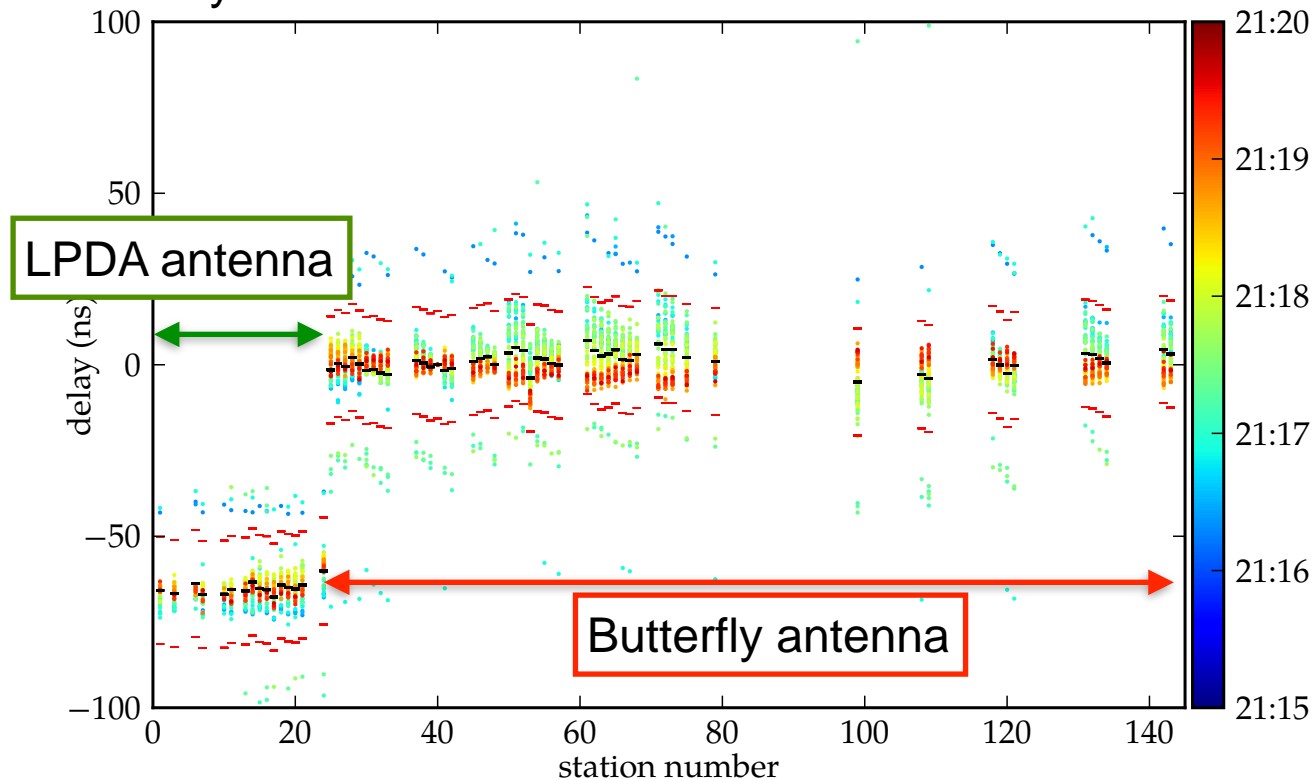
Time difference between a reference station (40)  
and other stations, corrected for expected propagation time  $\Delta t$

Again strong time fluctuations between stations: due to **GPS time offsets**??  
→ apply beacon correction



# Combination of beacon calibration and airplane measurement

Time delay as a function of radio station ID after beacon correction



Different antenna response → different group delay

**Mean time offsets** over the course of several months are consistent within **2 ns**

# Reconstruction improvement

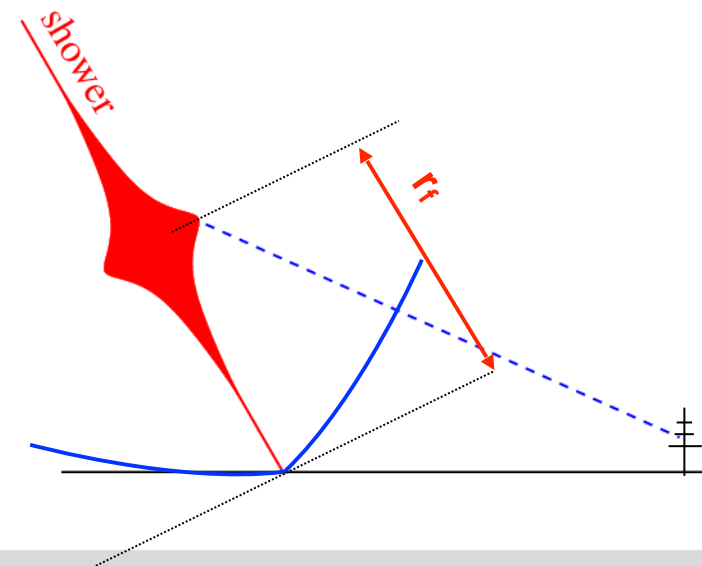
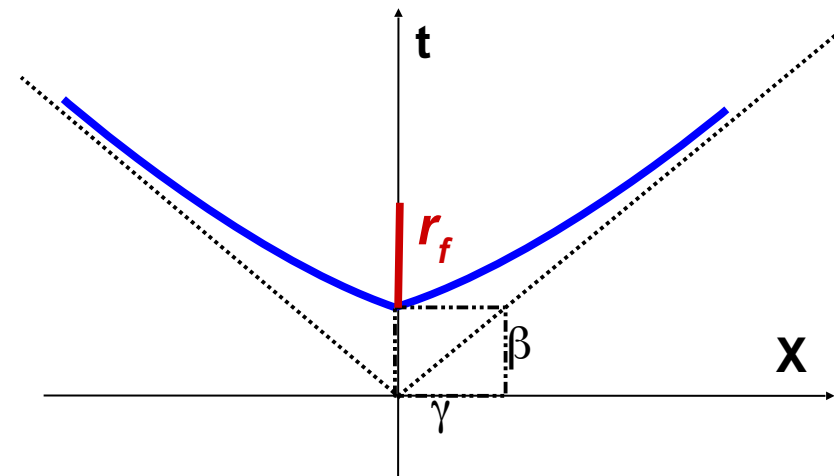
- The wavefront of radio emission is known to be of hyperbolic shape

- We fit a hyperbolic function
 
$$t = \beta \left( \sqrt{1 + x^2/\gamma^2} - 1 \right)$$

to extrapolate radio pulse time measured by stations as a function of its distance to the shower axis

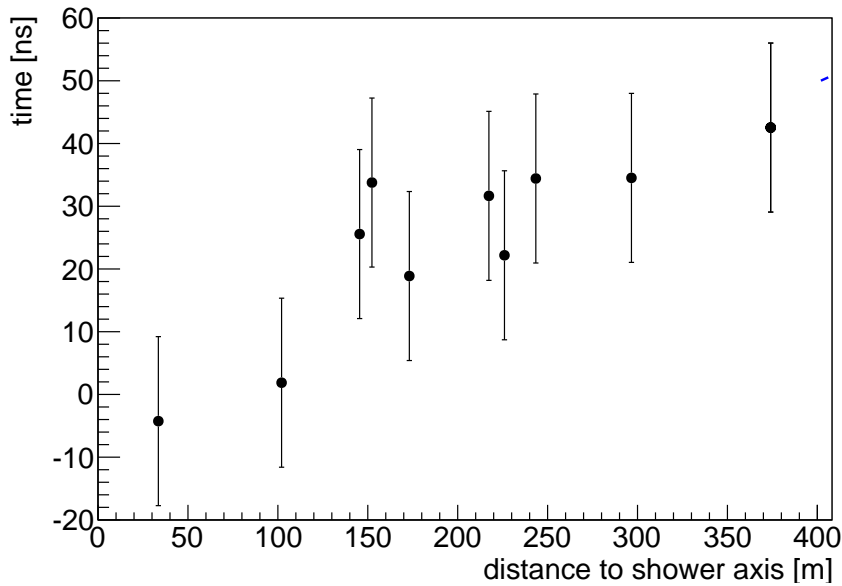
- Distance  $r_f$  to focus is correlated to depth of shower maximum
   
 → composition measurement

Hypothetical hyperbolic wavefront

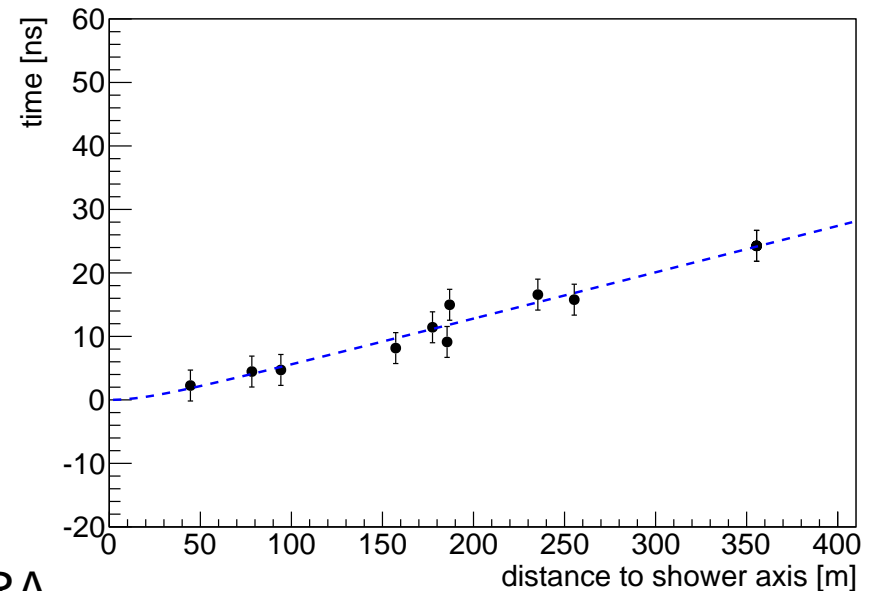


# Measurement of radio wavefront

measured wavefront **before** beacon correction



measured wavefront **after** beacon correction



This event imply both antenna type of AERA

After beacon correction the spread between the antenna types are **suppressed**

# Summary

- Radio detection of cosmic rays provides complementary information on air shower physics
- Beacon provides a high precision timing calibration for autonomous radio detector
- We have cross-checked the beacon timing calibration with a novel method using the signal emitted from commercial airplanes:
  - 2 ns time precision
- The time calibration is already included in analysis

# Outlook

- Collect more airplane events

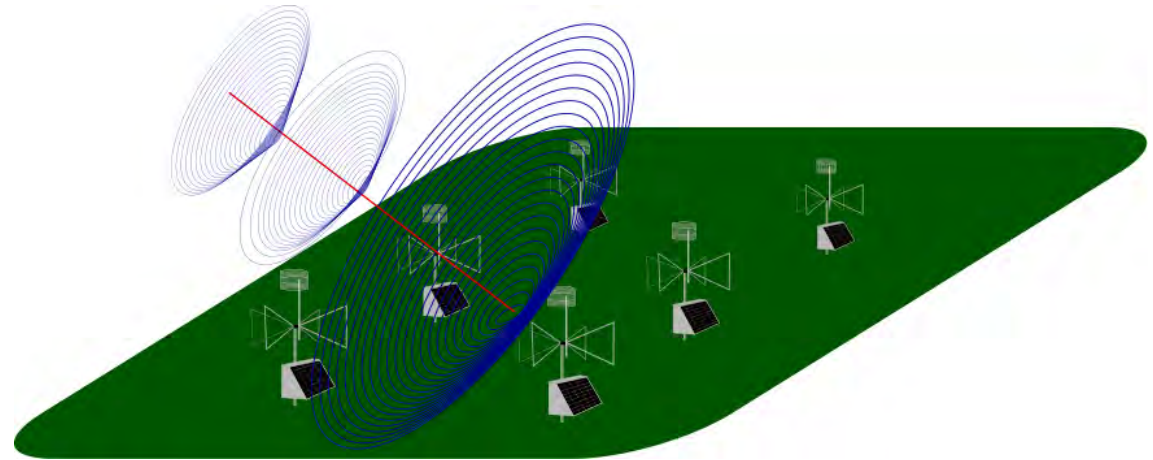
JINST 11 (2016) P01018  
doi: 10.1088/1748-0221/11/01/P01018

## Spare slides

# Physics objective of AERA

Artist's impression of radio wavefront measured by AERA

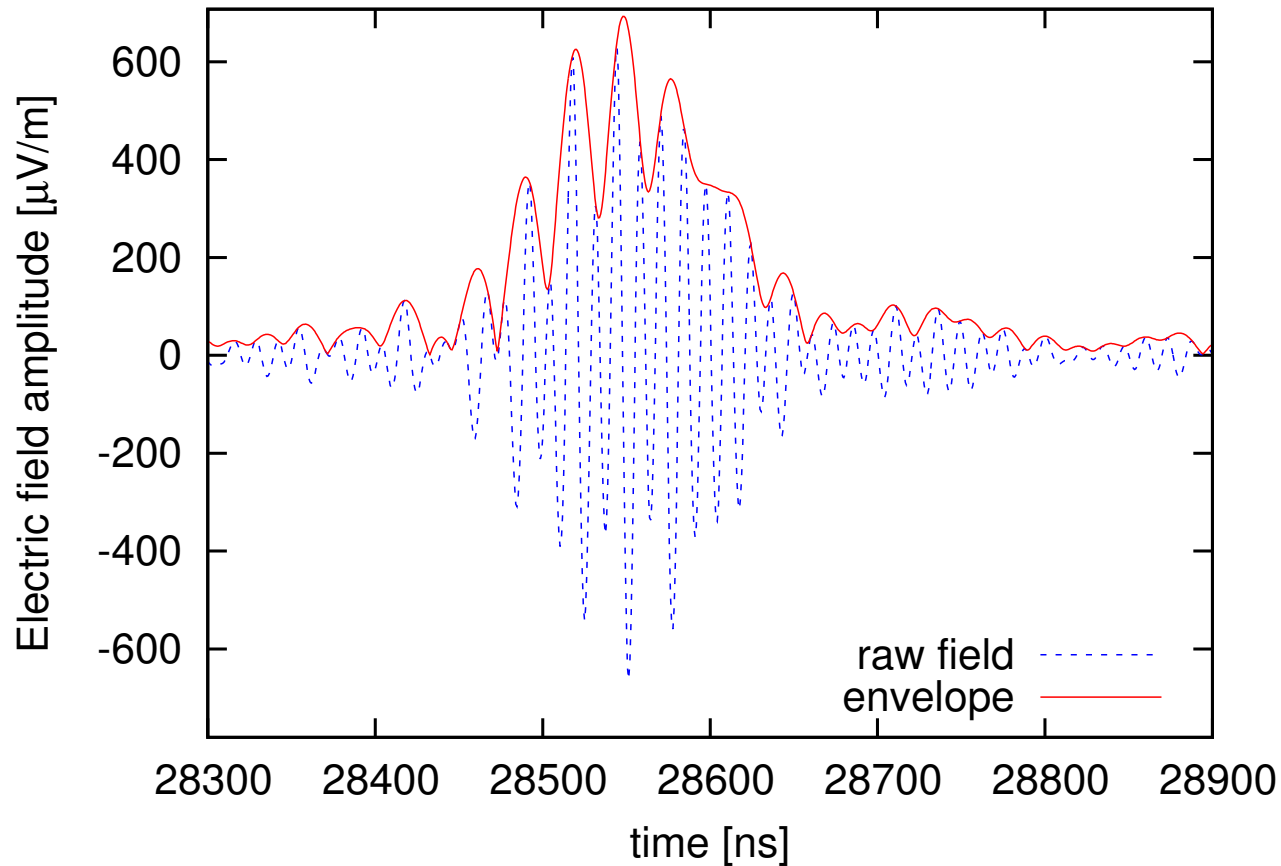
- Accurate measurements of:
  - Arrival direction
  - Energy
  - Mass composition
  - Radio interferometry



Shape of the wavefront measured on ground relates to the composition of cosmic rays

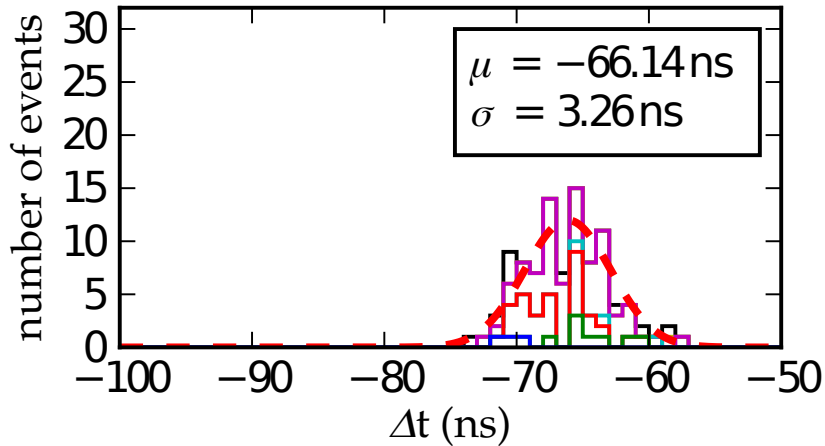
To exploit the full potential of radio measurement timing of 1 ns is needed

# Typical airplane radio pulse measured by AERA stations

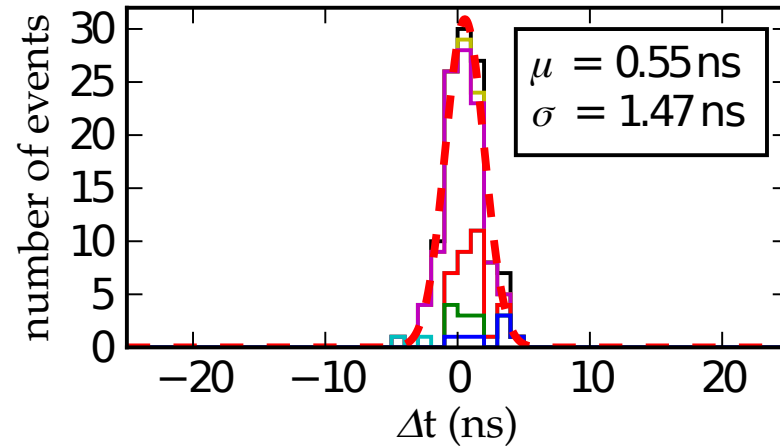


# Time delay measured with AERA station

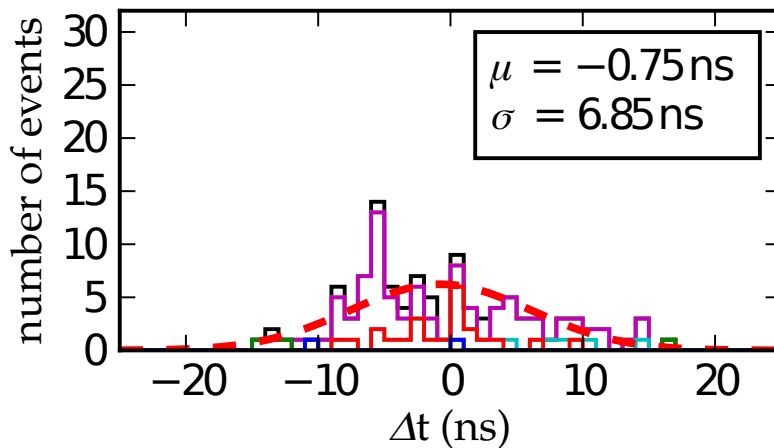
station 18



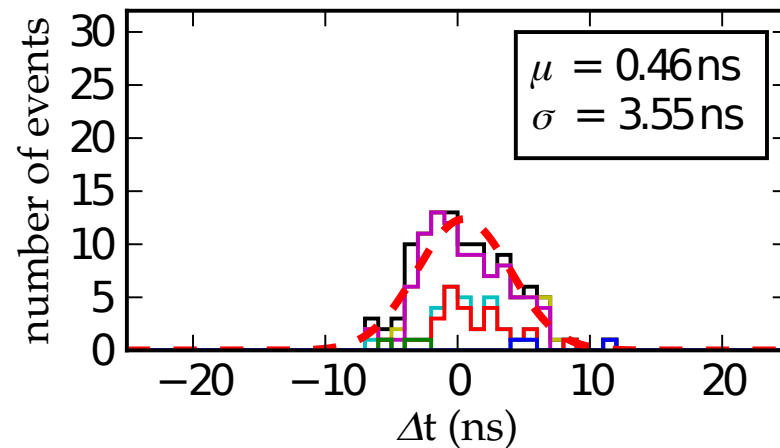
station 31



station 72



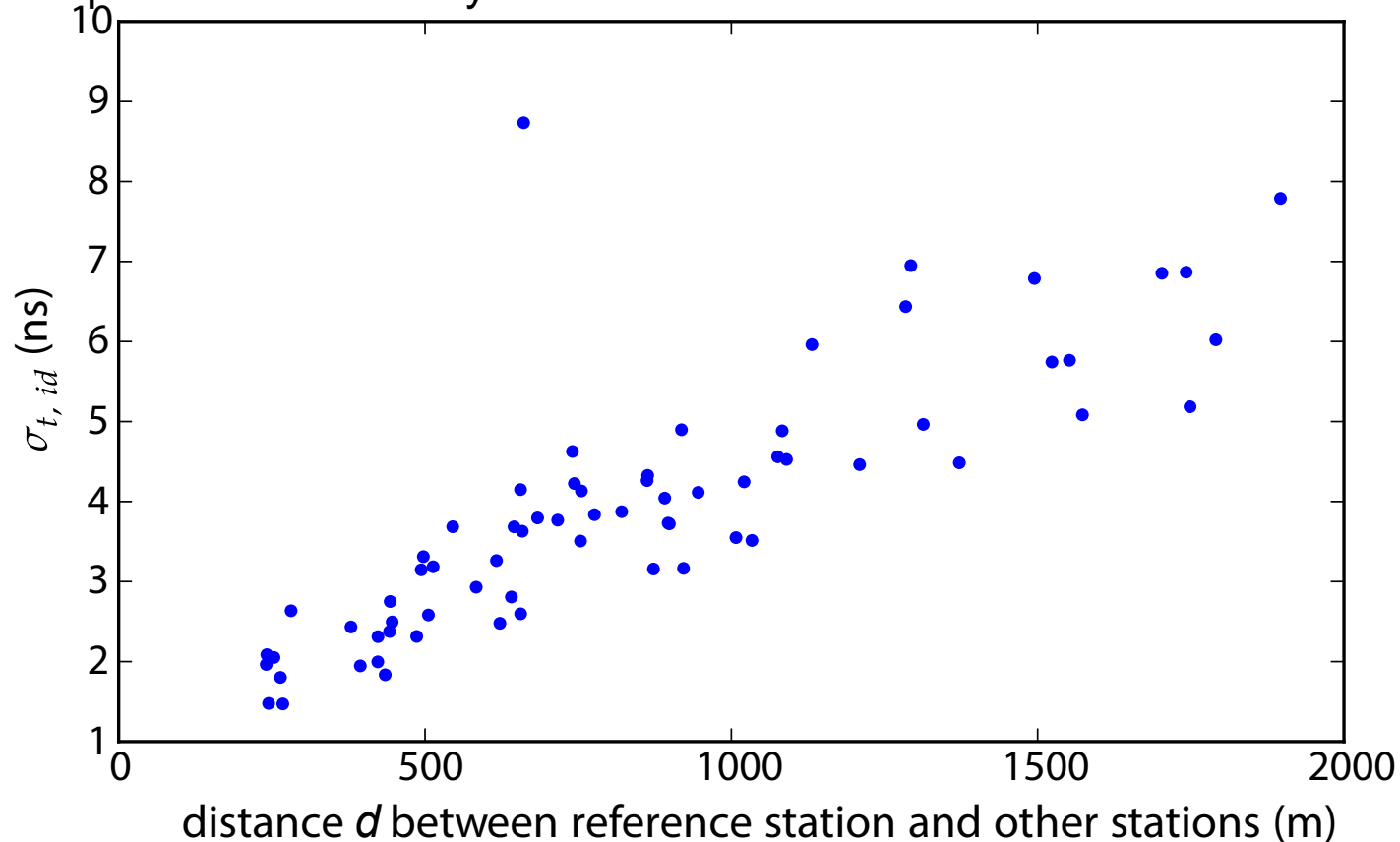
station 133





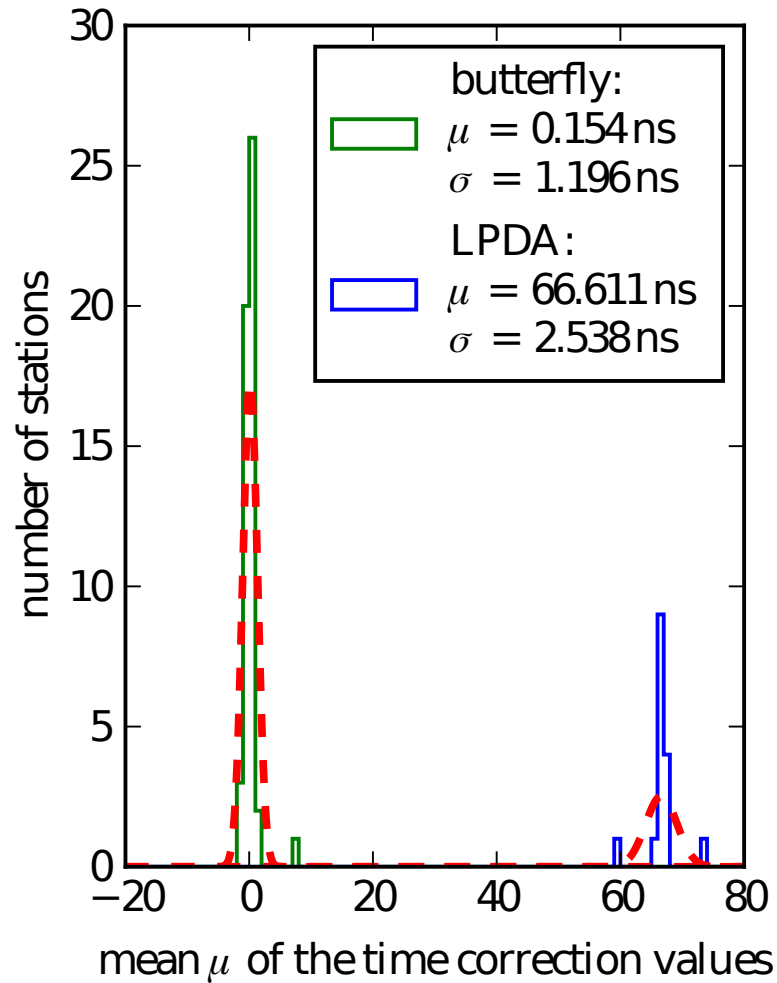
# Systematic uncertainties of airplane measurement

Spread of time delay as function of distance to reference antenna



The further from the reference station, the more precisely airplane position should be known

# Airplane timing after beacon correction



- quadratic sum of the spread of the mean of time delay: **2 ns**