

Monte Carlo Pathfinding in Radio Astronomy

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SKA Data Challenges

□ Planning the German National SKA Data Centre

■ Cost of computing

- Strong impact on construction costs
- Determined mostly by processing power and growth of archive

$$\frac{2}{3} C_{\text{tot}} \sim \alpha C_{\text{CPU}} + (1-\alpha) C_{\text{arch}}$$

■ Impact

- Only a small fraction of all “raw data” taken by the antennas can be stored in long-term archives

□ Reducing the data volume

- Developing algorithms for weeding out data of interest in near-realtime
- Each key science project (EoR, pulsar, ...) demands its own solution

□ Data analytics

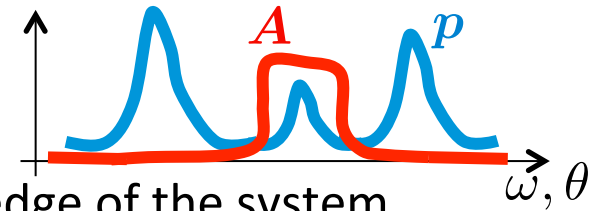
- Simply copying software from LOFAR not effective
- SKA much more complex than LHC
 - Growth of archive / “raw data rate”: **SKA \gtrsim 100 LHC**



Monte Carlo Simulation

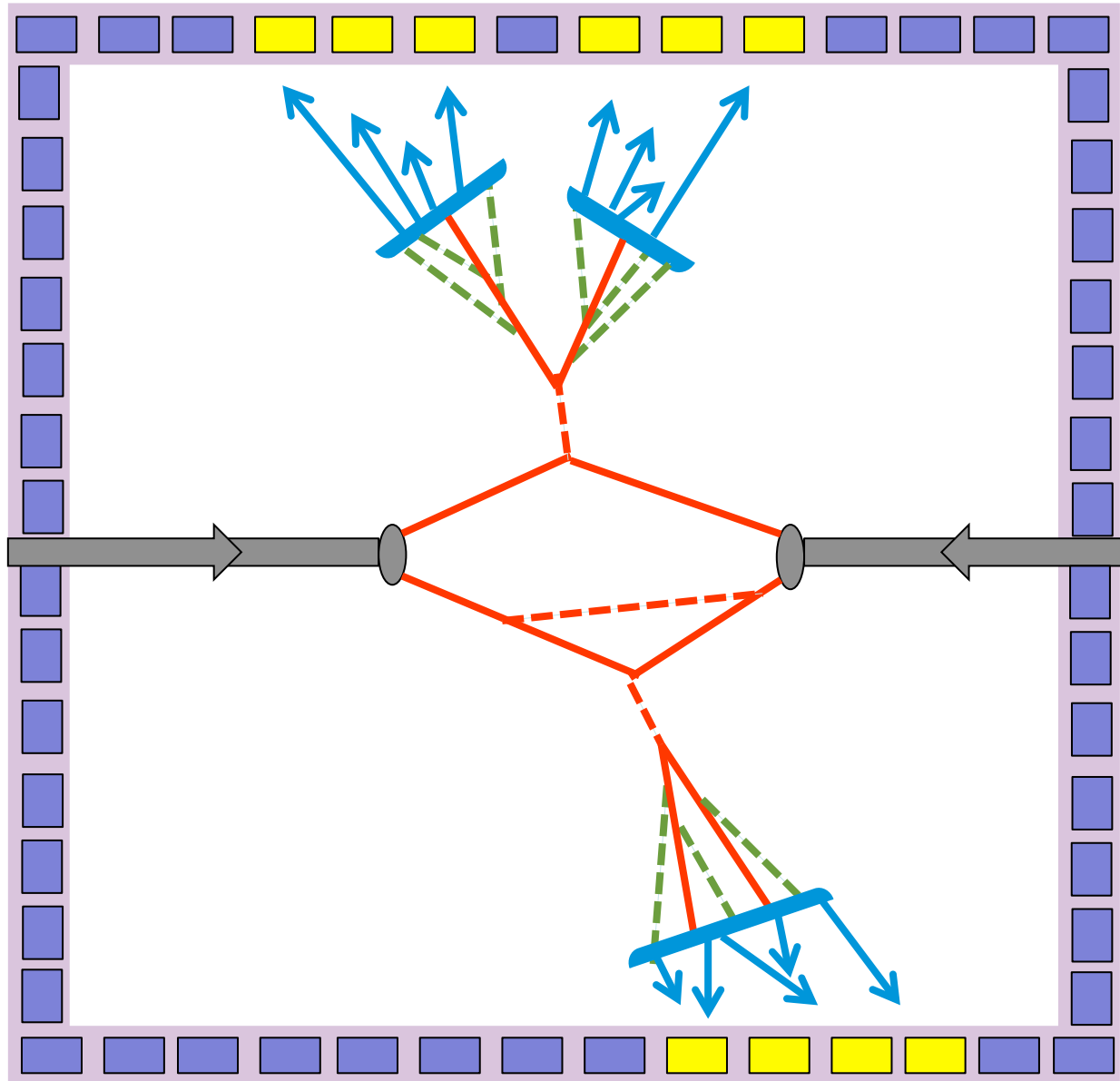
- Outcome of measurements is compared to expectation values of observables A

$$\langle A \rangle = \int A(\omega, \theta) p(\omega, \theta) d\omega d\theta$$



- The **probability distribution** p characterizes the knowledge of the system
 - Thermal equilibrium: $p \sim e^{-H/kT}$
 - Fitting models to data: $p(\omega, \theta) = p(\omega|\theta) p(\theta)$
 - $p(\omega|\theta)$ is the likelihood function of some model distribution
 - The prior $p(\theta)$ describes the initial knowledge of the parameter set θ
- Importance of Monte Carlo methods for evaluating expectation values
 - Scaling of error $\sim 1/\sqrt{N}$, N = number of **random points**: $(\omega_1, \theta_1), \dots, (\omega_N, \theta_N)$
 - Independent of dimension of integral
 - Parallel computing
 - Many Monte Carlo algorithms are parallelizable
 - Finding “rare events” in large search spaces
 - Importance sampling

Monte Carlo Event Generators at LHC



“MCTruth” data

- Hard interaction
- Parton shower
- Hadronization

⇒ Probabilistic simulation of the final state

Simulated data

- Detector simulation
- ⇒ Comparison with experimental data



Monte Carlo Event Generators at LHC

- Theory
 - Predicting final state
 - Simulating background / noise

- Experiment
 - Optimizing trigger algorithms
 - ⇒ reducing data volume
 - Determining influence of detector (e.g. dead material)
 - ⇒ estimating acceptance corrections
 - Unfolding measured data to “hadron level”
 - ⇒ publishing “true” results

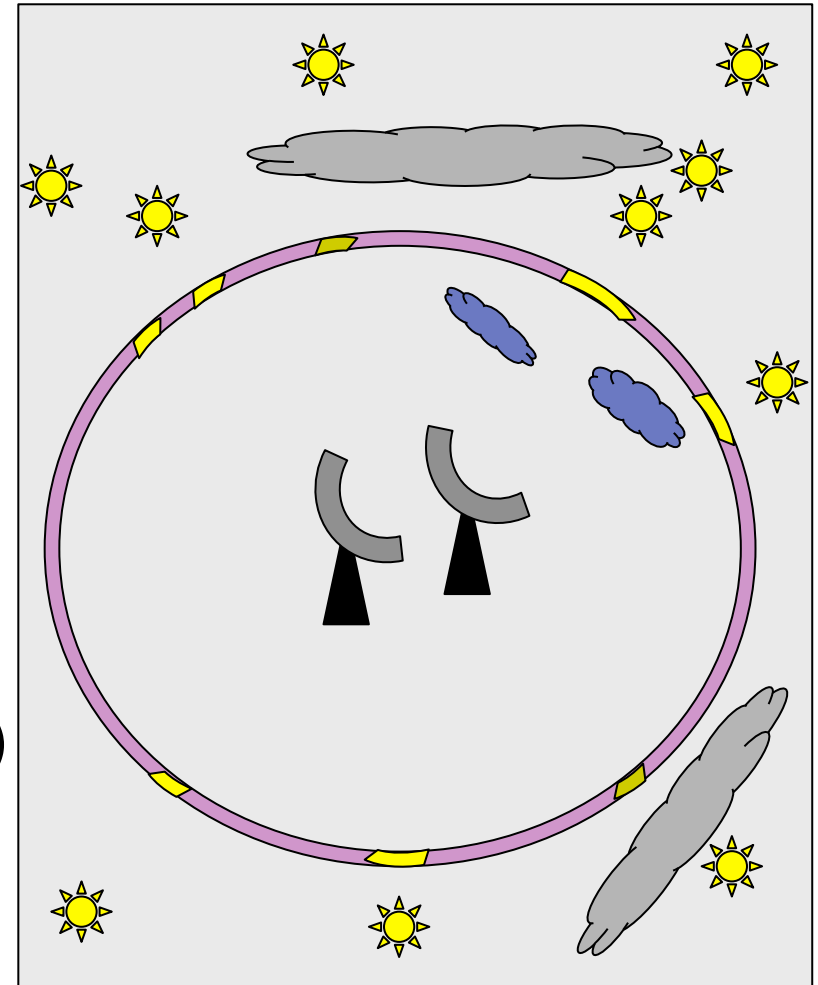
Monte Carlo Simulation at SKA

□ LHC

- Single source
 - Inside the detector
- Theory: QCD + QED + EW
 - Initial state known
- MC simulation of collision events
 - Several MC event generators

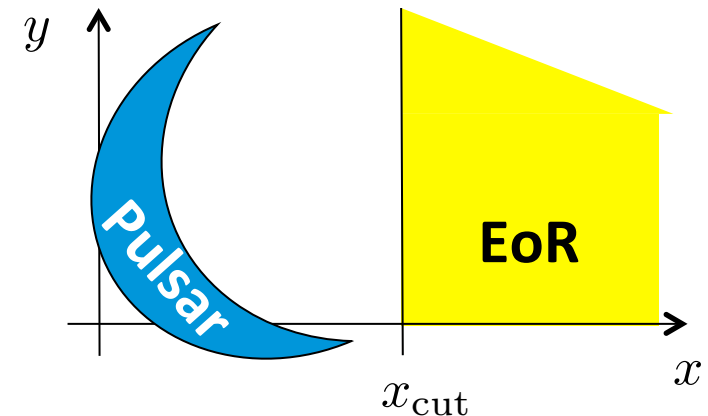
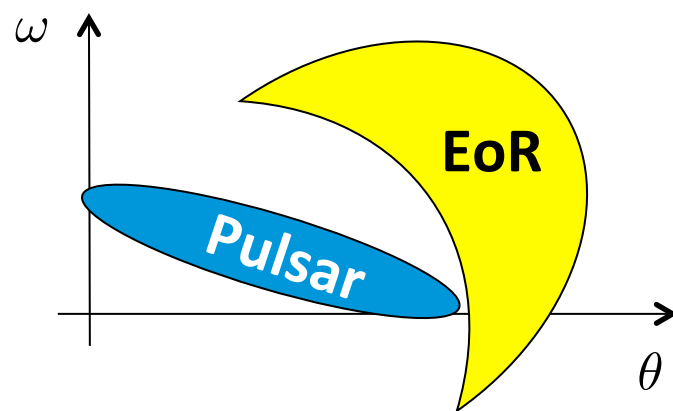
□ SKA

- ✧ Huge number of sources
 - ✧ Outside the celestial sphere (detectors)
- Theory: Maxwell equation with sources
 - + GRT
 - Boundary value problem
- Goal: MC simulation of the cosmos



Monte Carlo Simulation at SKA

- Probability distribution $p(\omega, \theta)$ on the phase space
 - Specify “subregions of interest” for every key science project



- Determine deformation map $(\theta, \omega) \rightarrow (x, y)$ for applying “trigger cuts” on data
 - LHC: e.g. cut on transverse energy of calorimeter cells, $\sum_i E_i \sin \theta_i \geq E_T$
 - SKA: cuts on functions of visibility data, e.g. $f_{\text{EoR}}(u, v, w, V(u, v, w)) \geq x_{\text{cut}}$
- Monte Carlo Simulation
 - New physics by simulating the background in dedicated regions of phase space
 - Reducing data volume in near-realtime based on simulation of “trigger cuts”
 - Key science projects obtain a fixed share of long-term data archive