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The Aachen Muon Detector and the FAMOUS telescope

Christine Peters for the AMD and FAMOUS groups



HAP workshop Non-thermal Universe September 2016



Outline

Motivation

The muon detector AMD

The fluorescence telescope FAMOUS

AACHEN UNIVERSITY

Aachen Muon Detector prototype



First Auger MPPC camera for the Observation of Ultra-high-energy air Showers



Cosmic ray air showers components



Fluorescence light

- measurement of energy
- measurement of depth of shower maximum
 - observable sensitive to mass composition

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Muonic component

- muons contain information about primary type
- separating electromagnetic & muonic component by hybrid detection
- separation of different primaries
 entangle hadronic
 interaction scenarios
 - understand flux suppression
- anisotropy studies with protons
 searches for point sources

Cosmic ray air showers observables

Combination of mass sensitive observables allows for an improved determination of the chemical composition of cosmic rays



Future detection principles





Aachen Muon Detector



First measurements - light yield of scintillator tiles

Coincidence measurements of two stacked prototype tiles read out by SiPMs



We measure muons!

SiPMs excellent choice for the detection of muons due to single p.e. resolution!!

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Impact of knowledge - MiniAMD

Calibration device for SSD?

- Number of tiles + SiPMs: 8 each
- Size SiPM: 1.3 mm x 1.3 mm
- SiPMs directly connected to tile
- Weight: ~ 40 kg







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courtesy: Michael Eichler

The fluorescence telescope FAMOUS



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549.7 mm

First field test with full 61 pixel camera

- Measurement of star trails
- Check of the pixel to channel assignment
- Measurement of the mean current over 1 h
- One measurement each 5 sec

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- Photograph taken with CMOS camera
- Automatic image analysis for star detection

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AMD PoS(ICRC2015)596

- Steel support structure and first 16 scintillator tiles are built
- Electronics ready
- First measurements show: Performance is promising Use of SiPMs also option for SSD upgrade
- Next step: Deployment of MiniAMD at Auger site

FAMOUS PoS(ICRC2015)605 PoS(ICRC2015)649

- Successful commissioning of the new 61 pixel focal plane
- Measurement of star trails

Deployment of IceAct Cherenkov telescope at South pole based on FAMOUS design

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New 61 pixel focal plane and power supply unit

61 pixels

- 3 blind pixels for noise monitoring
- Winston cones polished by hand and glued into base plate

- PSU: stable power-supply individually for all pixels
- Automatical temperature compensation

Live monitoring of all channels

Primary Particle Nuclear Interaction Fluorescence K+,K⁰ ◀ with Air Nucleus Light π+,π $\pi^{\pm} - K^+, K^0 \rightarrow \pi^0$ Hadronic Cascade ► μ⁻,μ+ e- e+ e+ e-____μ μ^+ μ^{\pm} υμ $e^{-} \gamma \gamma e^{+} e^{+} \gamma \gamma e^{-}$ p, n, π[±], K[±] μ **Nuclear Fragments** hadronic electromagnetic muonic

Air showers have surprisingly high number of muons (not yet understood)

study of extensive air showers and hadronic multiparticle production
 exploration of fundamental particle physics at energies well beyond those accessible at terrestrial accelerators

Muons are important - Auger Upgrade needed

The Aachen Muon Detector

First measurements - light yield of scintillator tiles

Subtract dark noise spectrum, signal remains.

Each peak corresponds to a certain number of photon equivalents.

SiPMs excellent choice for the detection of muons due to single p.e. resolution!!