

REAL-TIME ALERTS FROM ICECUBE

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FOR THE ICECUBE COLLABORATION

HAP WORKSHOP - THE NON-THERMAL UNIVERSE
2016/09/22, ERLANGEN

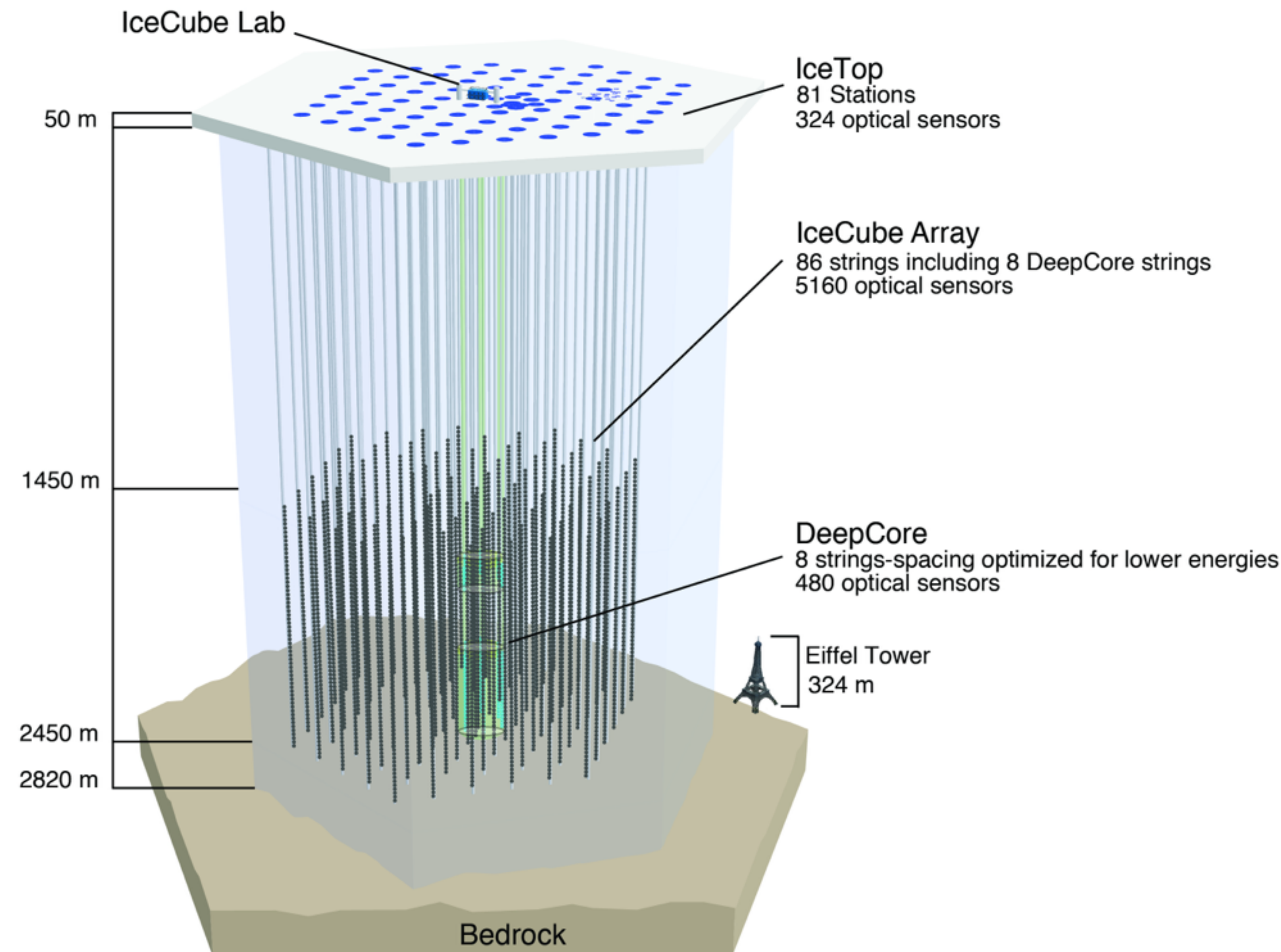


OUTLINE

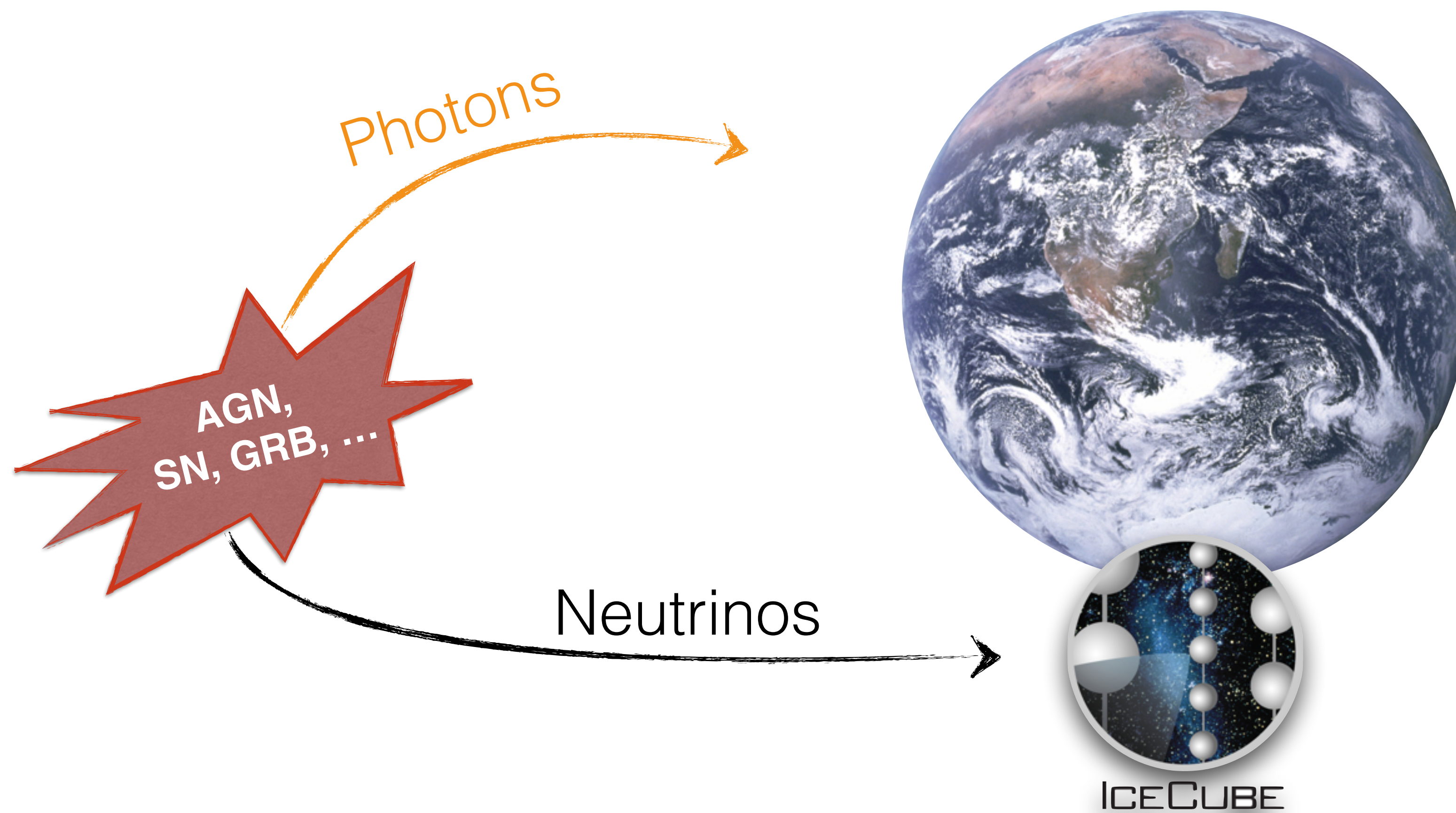
- ▶ IceCube Detector
- ▶ Real-Time Infrastructure
- ▶ Follow-Up / Alert Programs
 - ▶ Gamma-Ray Follow-Up, Optical Follow-Up
 - ▶ HESE and EHE Alerts
- ▶ Summary

THE ICECUBE DETECTOR

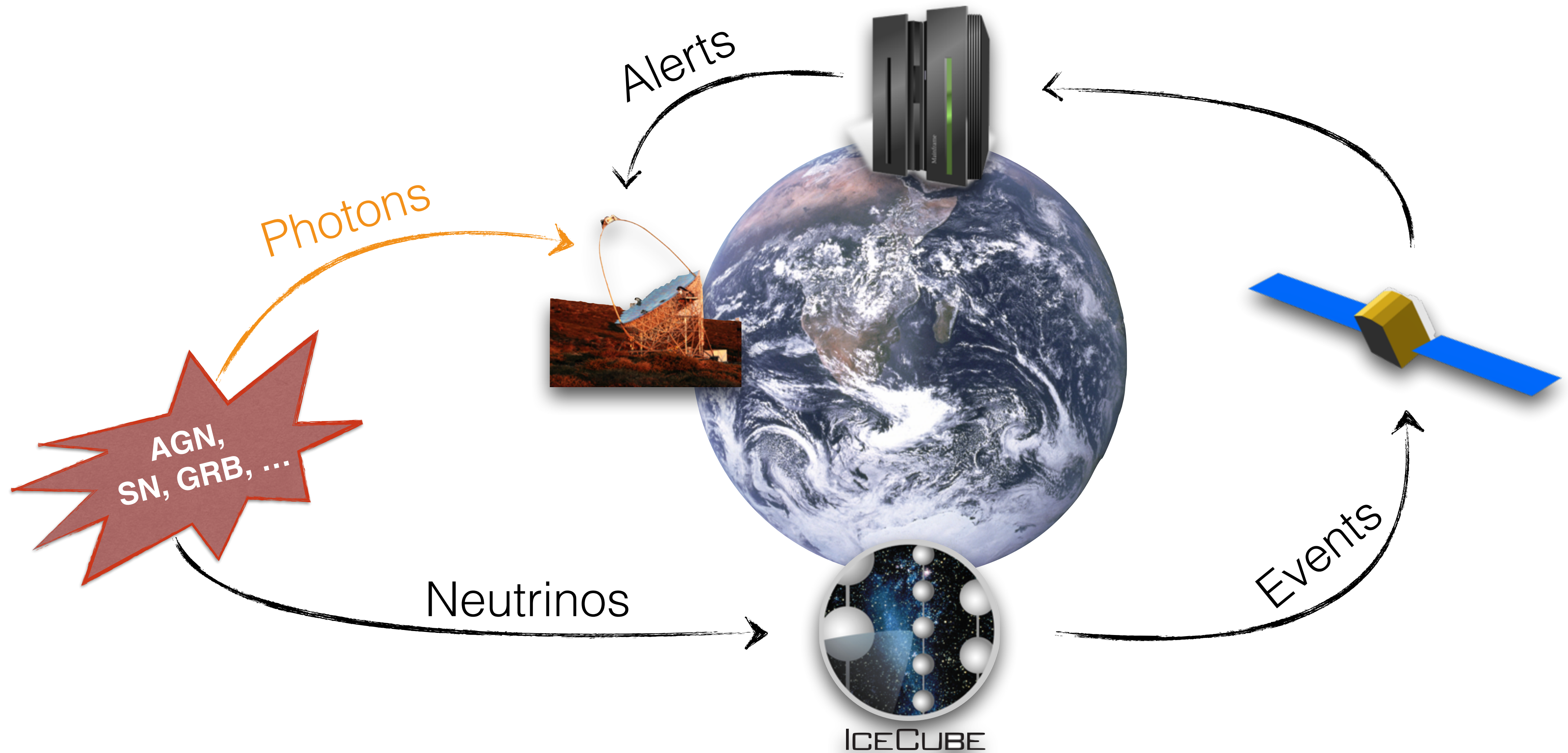
- ▶ 1 km³ detector for Cherenkov light at the South Pole
- ▶ 5160 optical modules on 86 strings
- ▶ Scope of this talk:
Muons from CC ν_μ interaction
- ▶ Angular resolution:
typically better than 1 deg



REAL-TIME ALERT SYSTEM



REAL-TIME ALERT SYSTEM



REAL-TIME OPERATIONS



- ▶ Limited computing resources at the South Pole

- ▶ Limited connectivity
Iridium: low latency, low bandwidth
TDRSS: high latency, high bandwidth

- ▶ Limited detector monitoring
(i.e. no „Good Run List“)

- ▶ Fast/specialized event selections
- ▶ Basic event information first for quick generation of alerts
- ▶ Full event later for improved reconstruction
- ▶ Conservative online monitoring



REAL-TIME ANALYSES IN ICECUBE

Clustering
Searches

2008: Optical Follow-Up (OFU)

2012: Gamma-Ray Follow-Up (GFU)

Individual
Events

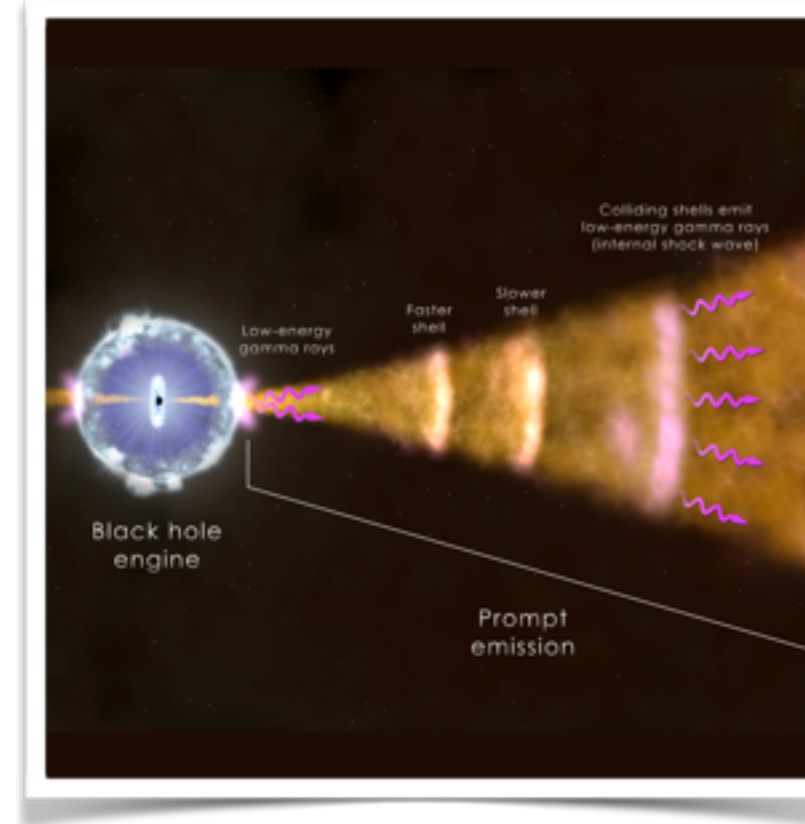
2015: High-Energy Starting Events (HESE)

2016: Extreme High Energy Events (EHE)

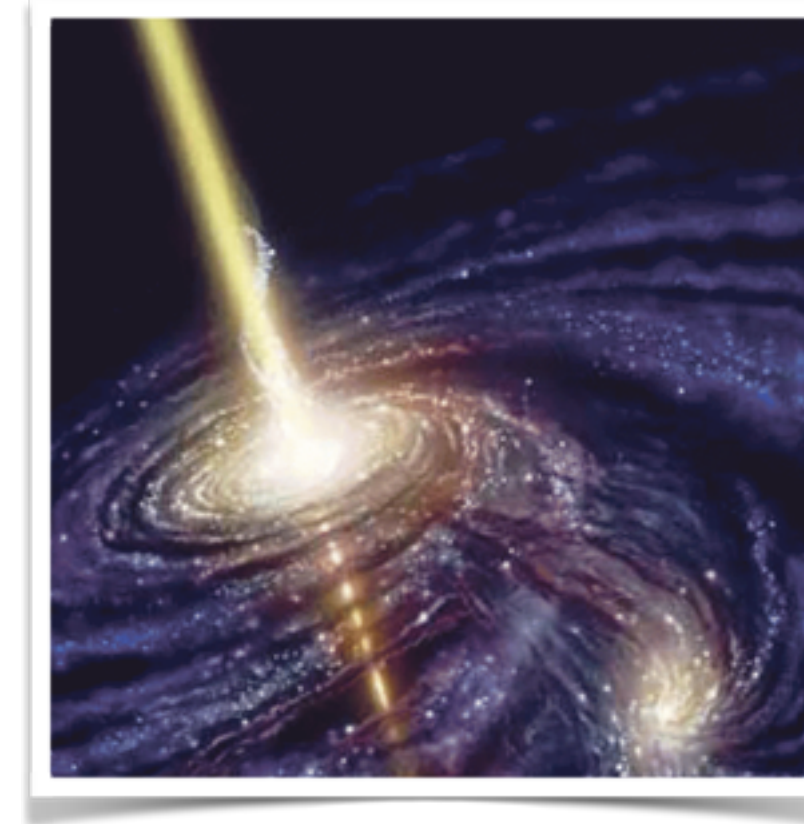
GAMMA, OPTICAL AND X-RAY FOLLOW-UP

Target:

transient and variable sources



GRB: 10–100 s



AGN: ~ 10 d



SN: ~ 100 d

Time-dependent analyses:

- Optical/X-ray Follow-Up: Multiplets within 100 seconds
- Gamma-Ray Follow-Up: Cluster search for up to 21 days

GAMMA-RAY FOLLOW-UP

Event Selection

- ▶ Fast, BDT-based
- ▶ Picking well-reconstructed, through-going muon tracks

Sending to the North

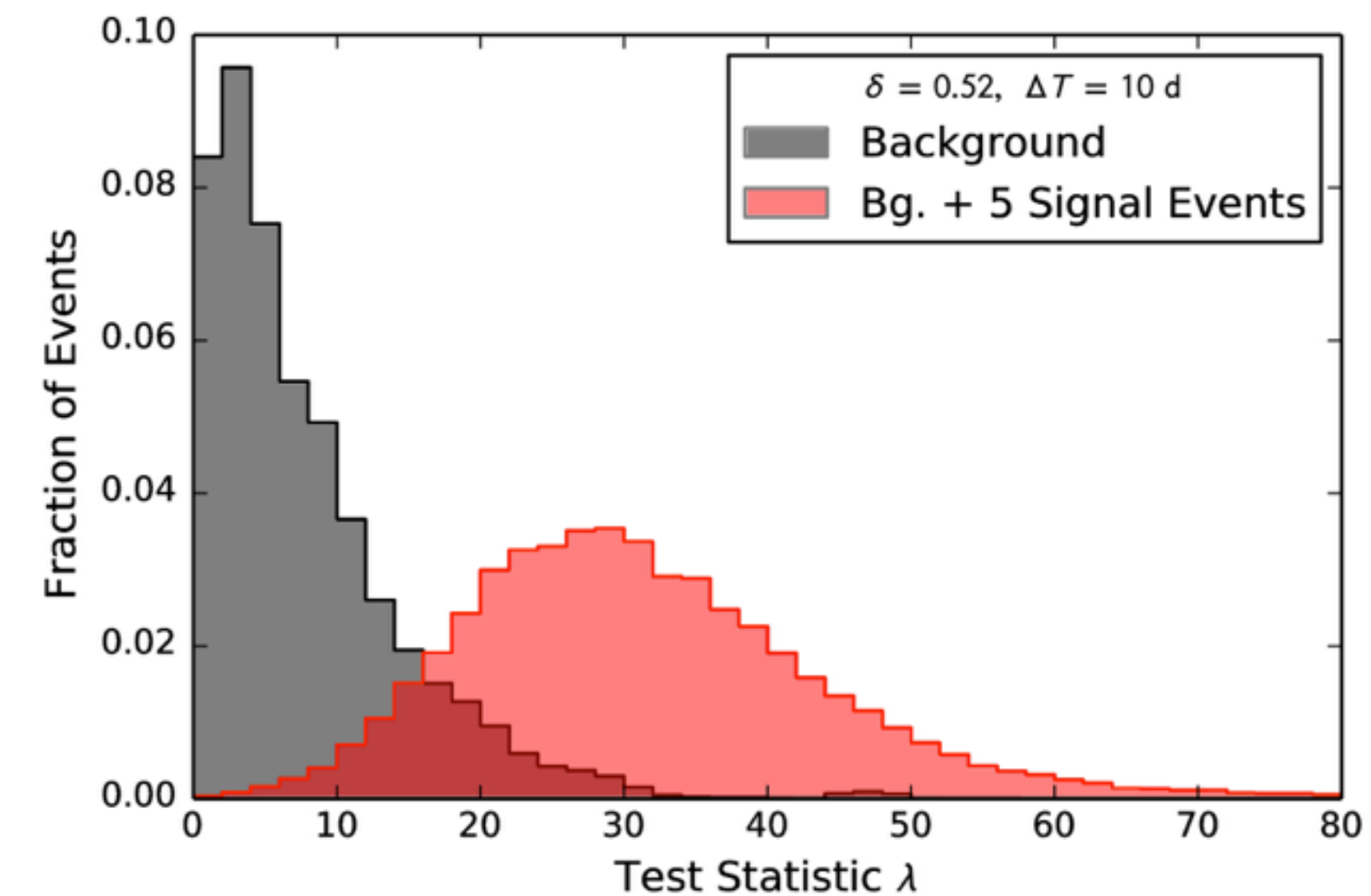
- ▶ Basic information: direction, energy, uncertainty

Clustering Analysis

- ▶ Likelihood analysis:
21 day point-source search
- ▶ Pre-agreed upon list of sources

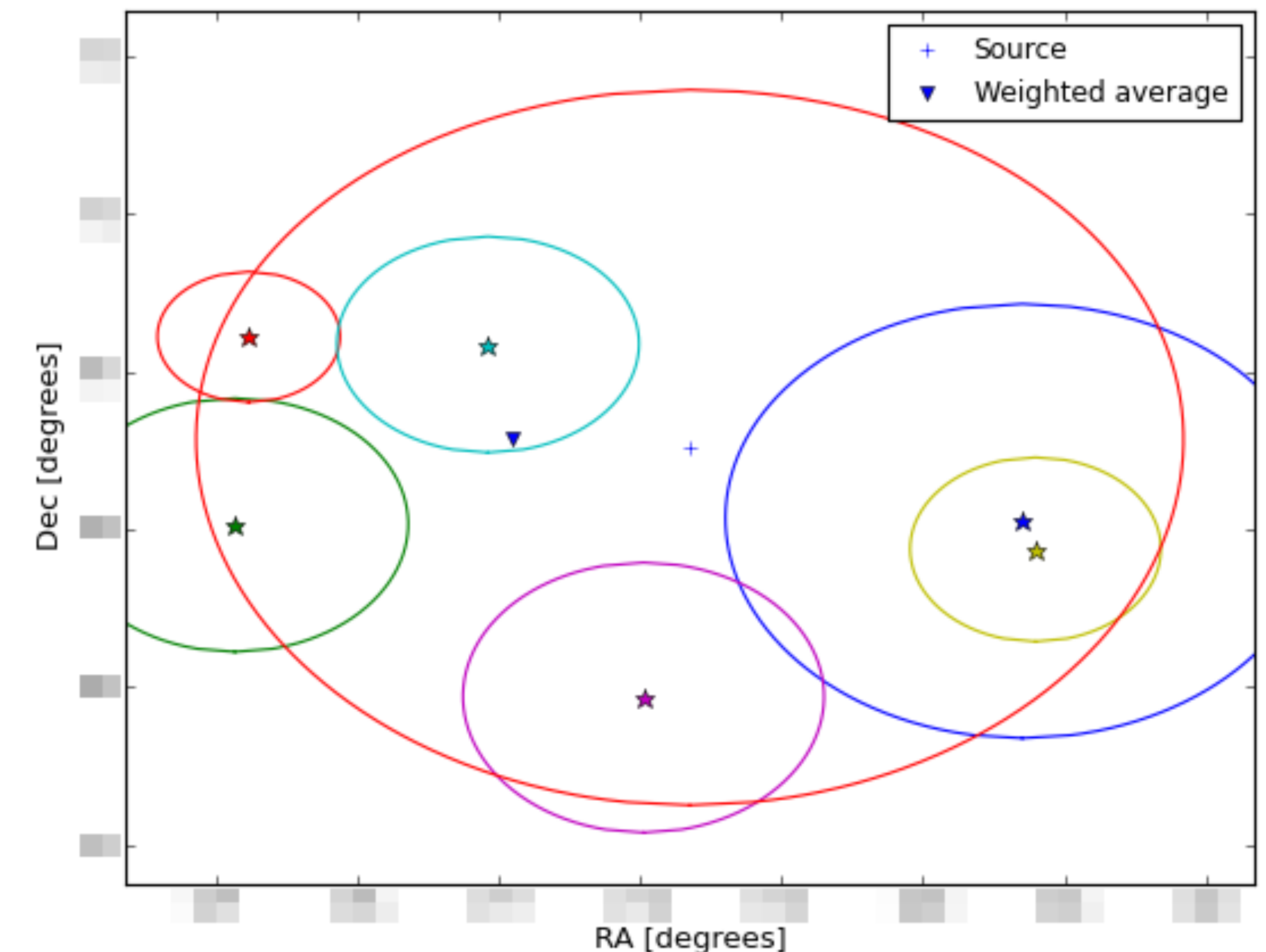
Alerts

- ▶ 2 alerts/year to MAGIC and VERITAS



RESULTS FROM THE GFU PROGRAM

- ▶ **Most significant alert: 2012/11/09**
 - 6 events within 4.2 days
 - p-Value: 0.002% (pre-trial)
- ▶ Alert forwarded to VERITAS
- ▶ No significant evidence for gamma-ray emission seen

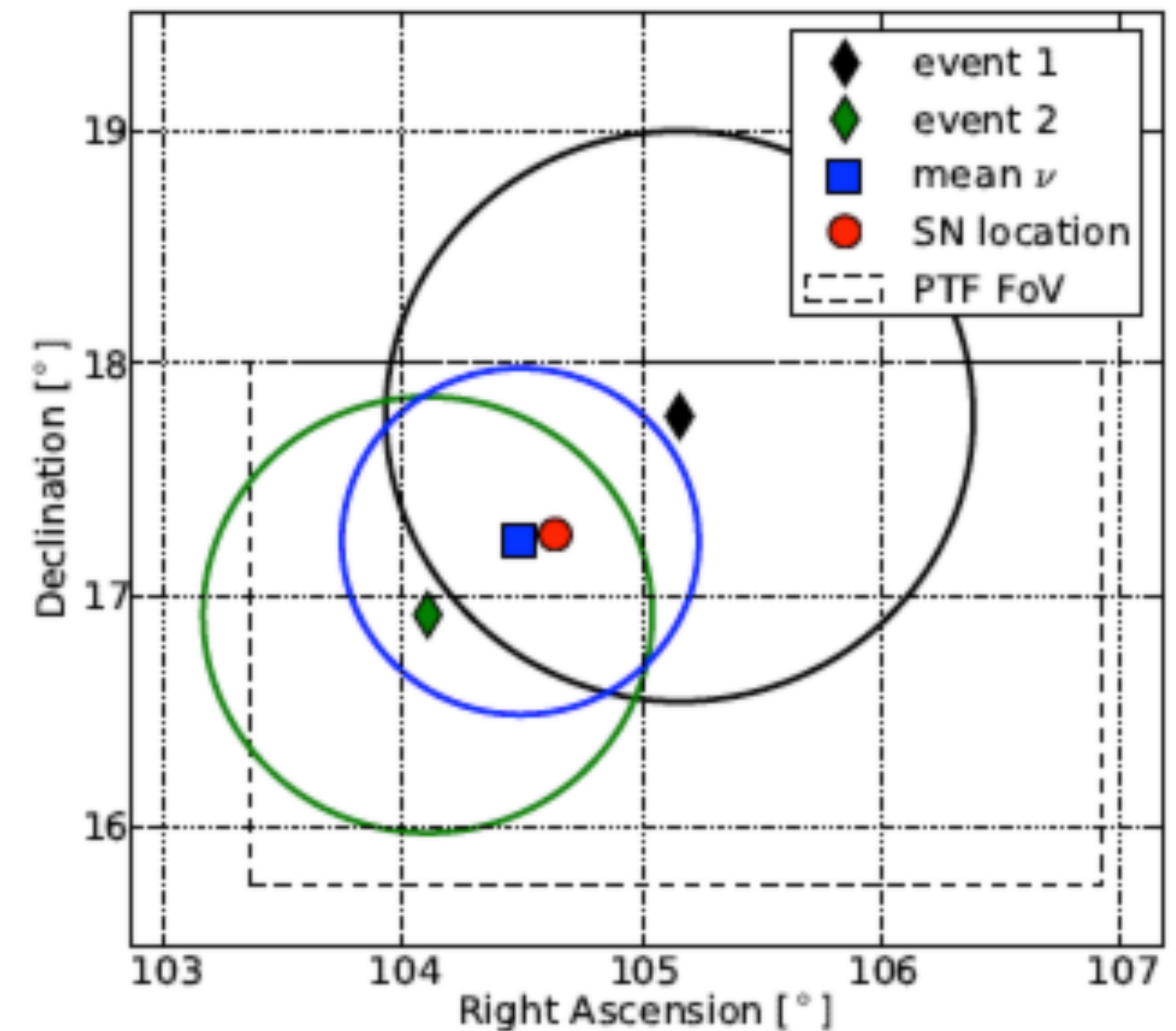


Future plan: Addition of a generic clustering search on the whole sky

RESULTS FROM THE OFU PROGRAM

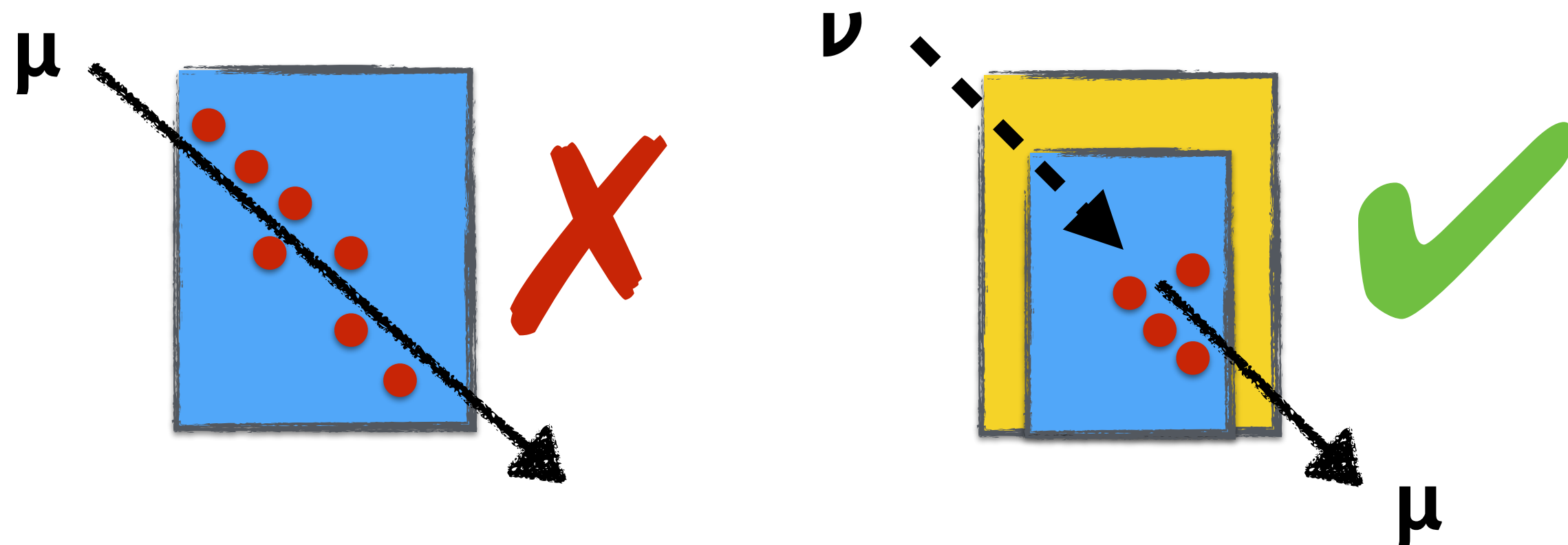
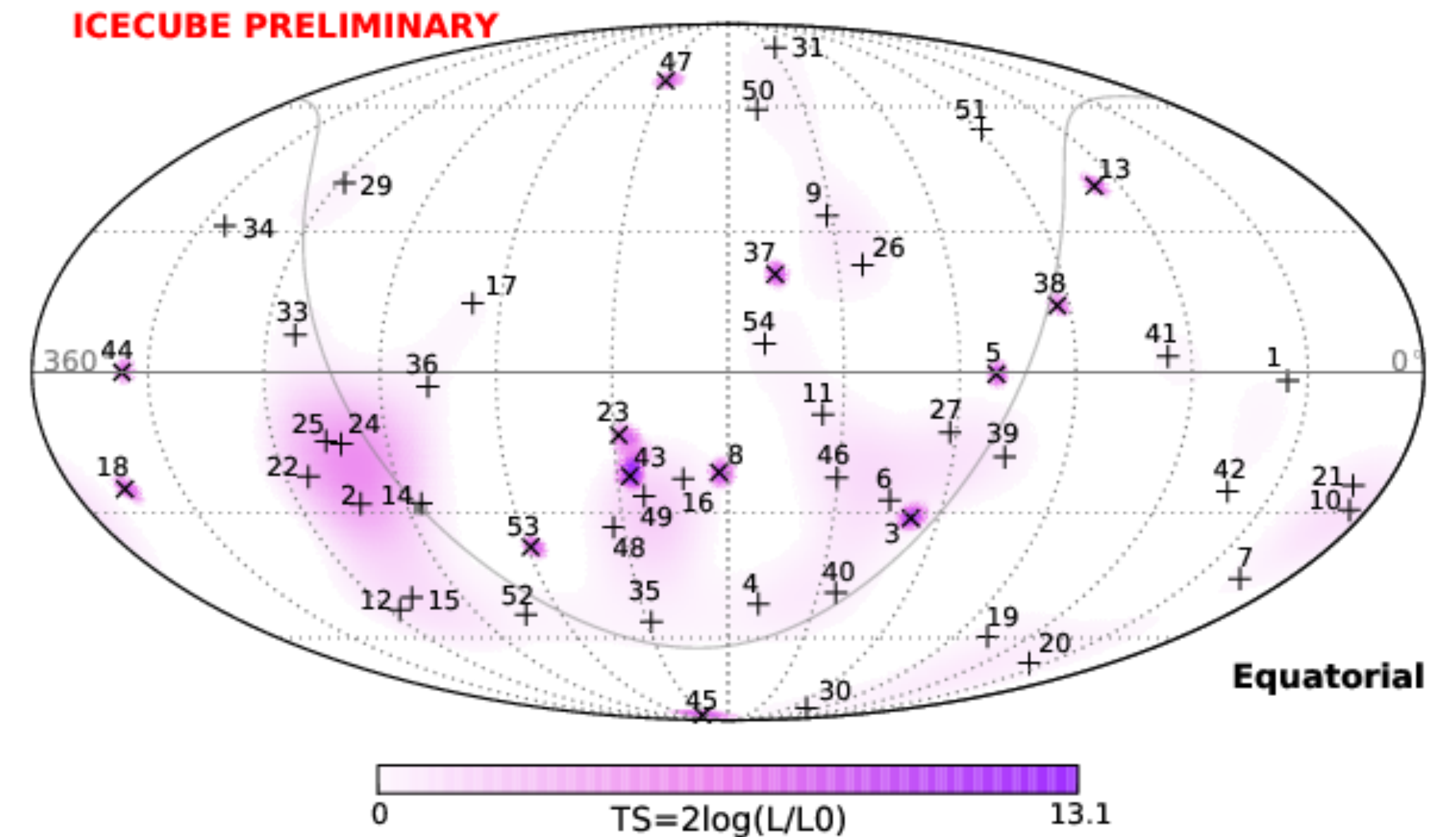
- ▶ **Straight cuts for multiplets**
 - ... in space (3.5 deg)
 - ... and time (100 s)
- ▶ **Most significant doublet: 2012/03/30**
 - two neutrinos within 1.7 s
 - 0.14 deg away from CCSN PTF12csy
 - p-Value: 0.014
 - neutrinos likely unrelated to SN
- ▶ **Growing list of follow-up observatories:**

~~ROTSE~~, PTF, MASTER, SWIFT, (ASAS-SN, LCOGT)
in preparation



HIGH ENERGY STARTING EVENTS

- ▶ Veto against atmospheric muons by outer detector layer
- ▶ Starting tracks with $Q > 6000$ pe, but short lever arm
- ▶ Expensive reconstruction (→revised alerts)

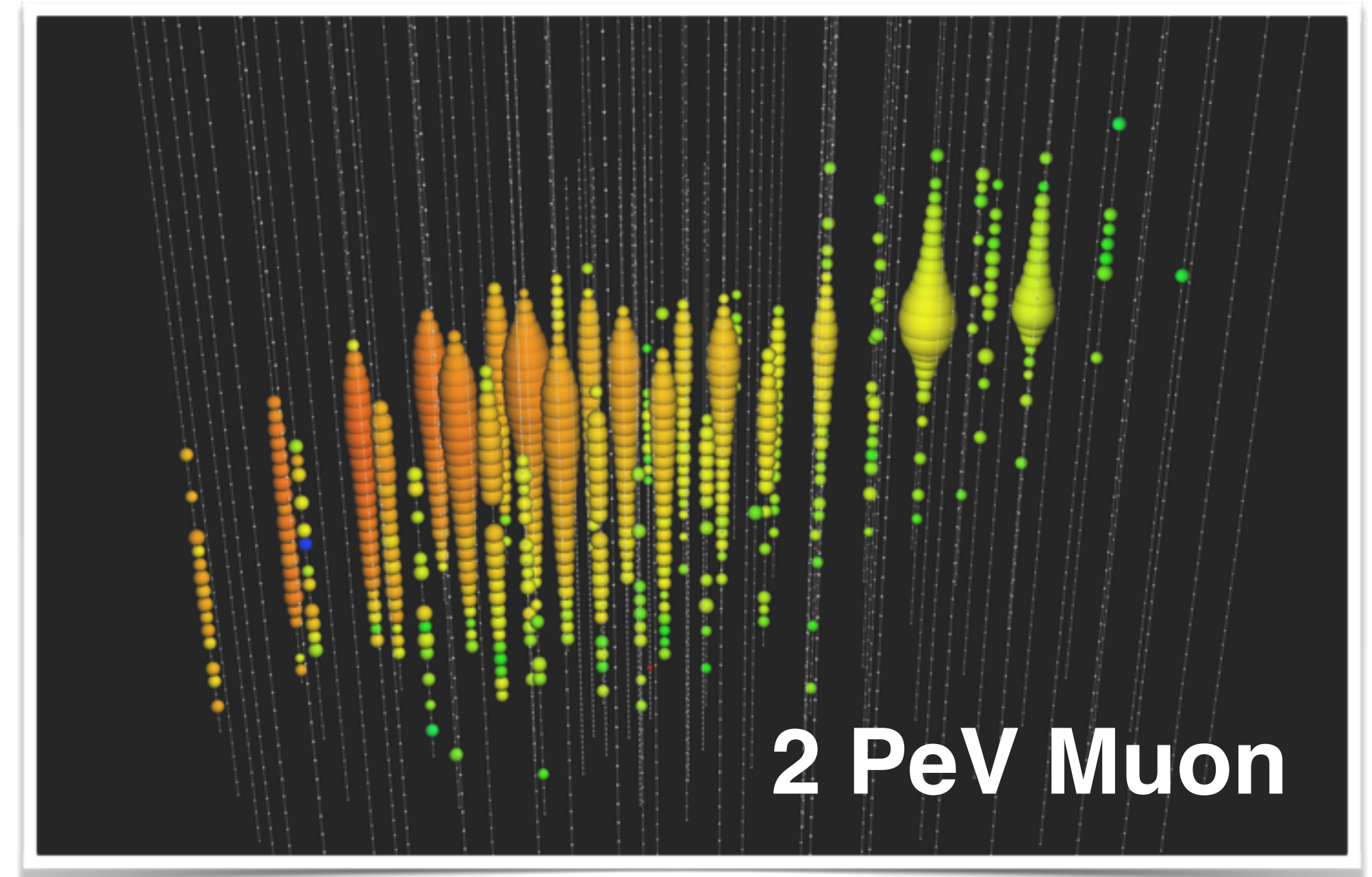


Public events (GCN)
~ 4 alerts / year

EXTREME HIGH ENERGY EVENTS

- ▶ High-energy, throughgoing events
- ▶ $N_{pe} > 3000$
- ▶ Very good resolution (<0.2 deg)
- ▶ Public alerts (GCN)

- ▶ Expected yield (S+B):
 - $4+2$ events/year (E^{-2})
 - $2+2$ events/year ($E^{-2.5}$)



FIRST HESE/EHE ALERTS

Date	Type	RA	Dec	50% Error
2016/04/27	HESE	240.6 deg	9.3 deg	0.6 deg
2016/07/31	EHE + HESE	214.5 deg	-0.3 deg	0.35 deg
2016/08/06	EHE	122.8 deg	-0.7 deg	0.11 deg
2016/08/14	HESE	200.3 deg	-32.4 deg	0.48 deg

- ▶ all alerts published as GCN (via AMON)
- ▶ responses from observers also via GCN

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Gamma-Rays

Optical

Observer	Result
iPTF	3 transients, all AGN
MASTER	no detection
PanSTARRS	7 SN candidates

Observer	Result
IPN	no detection
Fermi-LAT	5 unrelated blazars
Fermi-GBM	no detection
FACT	no detection
VERITAS	no detection
HAWC	no detection
MAGIC	no detection?

SUMMARY

- ▶ **Long-standing experience in realtime analyses**
 - Clustering searches
 - Single events
- ▶ **Successful exchange of alerts with**
MAGIC, VERITAS, HESS, ROTSE, PTF, SWIFT, MASTER, ...
- ▶ **Alerts are shared...**
 - ... directly through MoU with partners
 - ... through AMON / GCN