

Tunka-Rex: Radio Detection of Air Showers

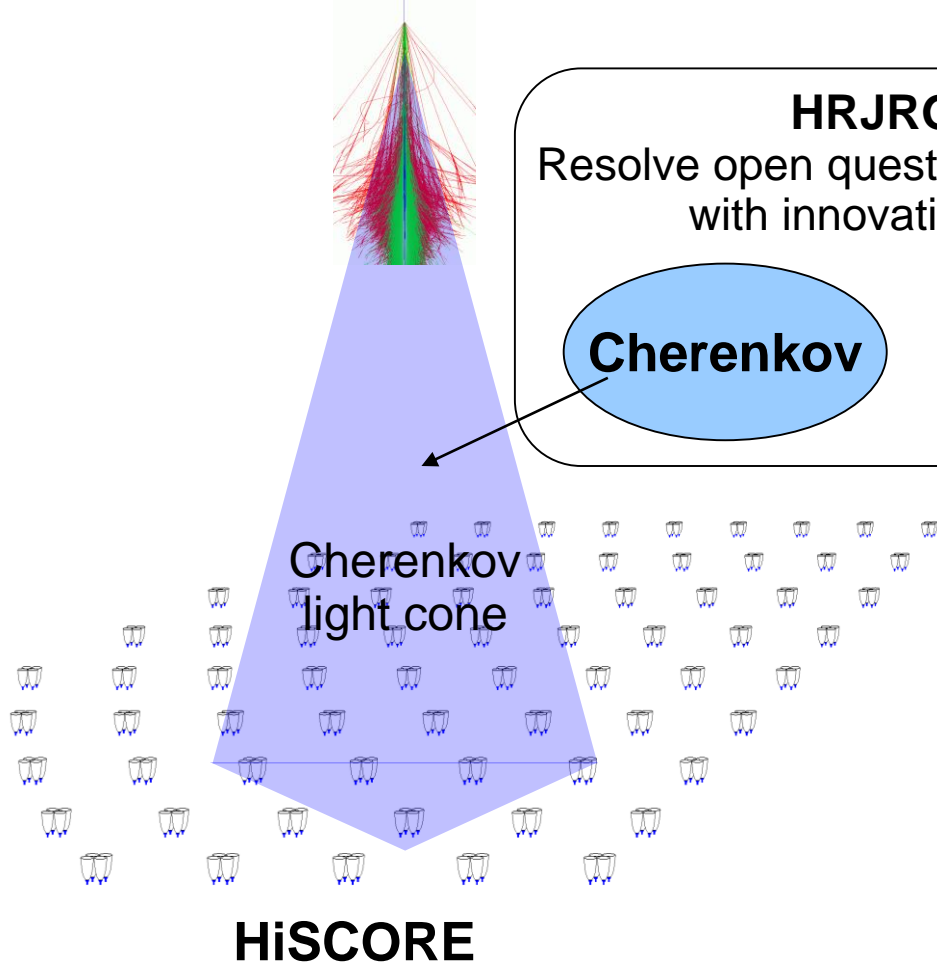
Frank G. Schröder
for the Tunka-Rex Collaboration

Karlsruhe Institute of Technology (KIT), Institut für Kernphysik, Karlsruhe, Germany



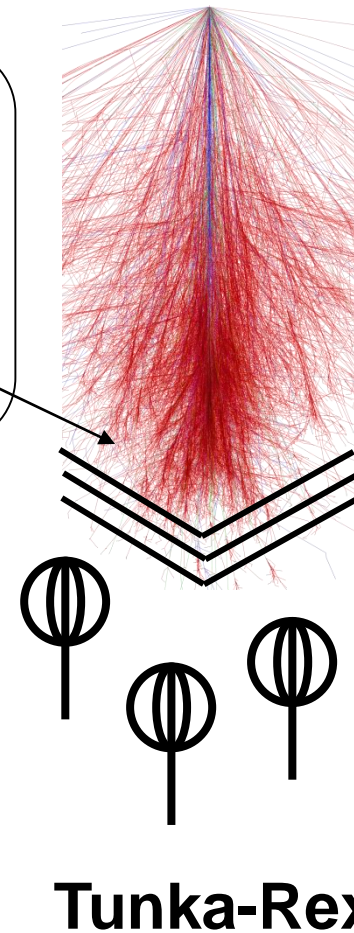
High energy γ (GeV / TeV / PeV):

Ultra-high energy CR (PeV - EeV):



HRJRG-303:
Resolve open questions of cosmic rays with innovative methods

Cherenkov **Radio**



- Funding
 - 2012 –15:
HRJRG
 - 2016 –17:
DFG + Russia
 - HAP *iProgress*

Tunka-Rex Collaboration



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Measurement of cosmic-ray air showers with the Tunka Radio Extension (Tunka-Rex)

P.A. Bezyazeev et al. - Tunka-Rex Collaboration,
Nucl. Instr. and Meth. A 802 (2015) 89-96

Reconstruction of air-shower parameters for large-scale radio detectors using the lateral distribution

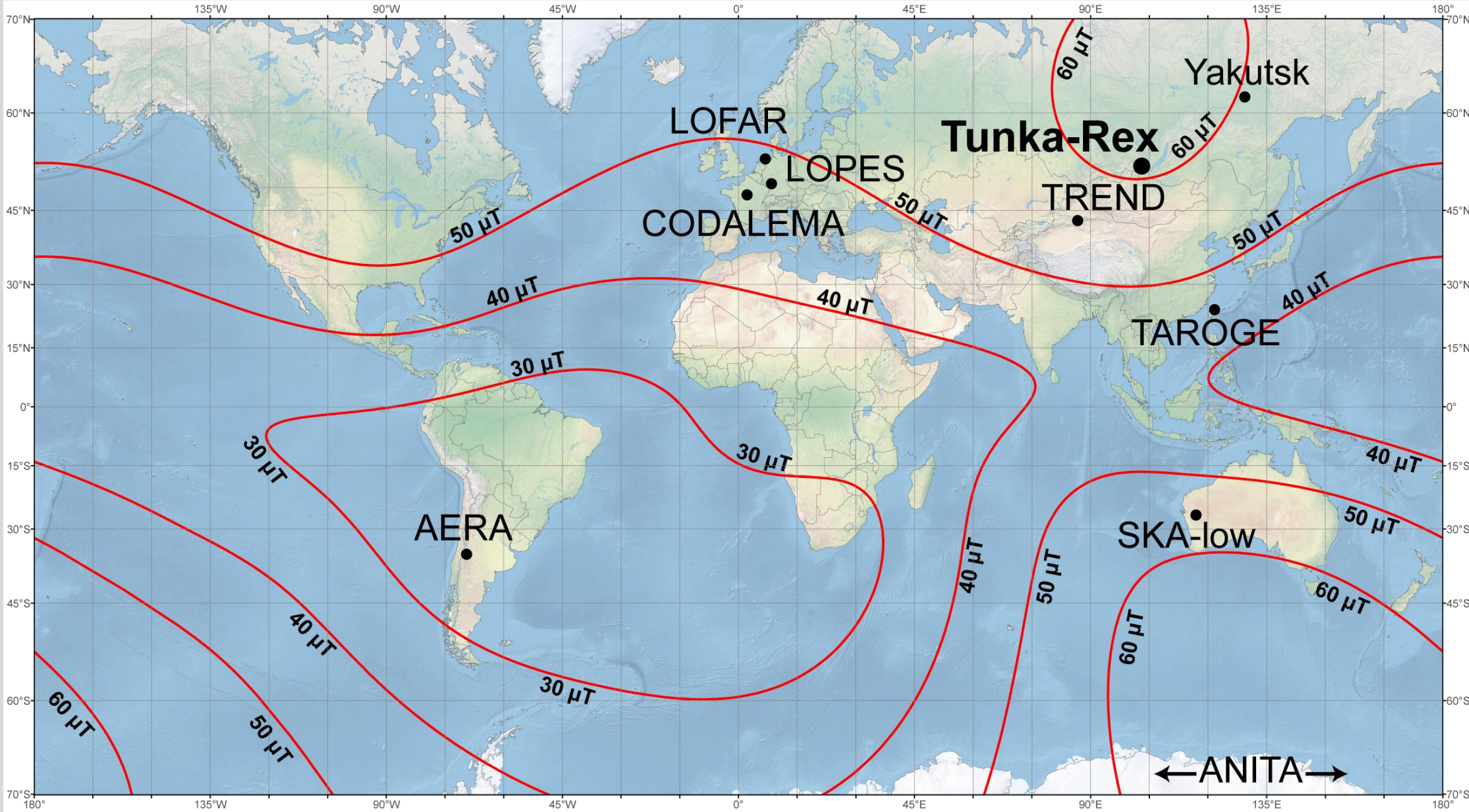
D. Kostunin, et. al (short authorlist),
Astropart.Phys. 74 (2016) 79-86

Radio measurements of the energy and depth of maximum of cosmic-ray air showers by Tunka-Rex

P.A. Bezyazeev et al. - Tunka-Rex Collaboration,
JCAP 01 (2016) 052 – open access

A comparison of the cosmic-ray energy scales of Tunka-133 and KASCADE-Grande via their radio extensions Tunka-Rex and LOPES

Tunka-Rex + LOPES Collaborations,
submitted to PLB



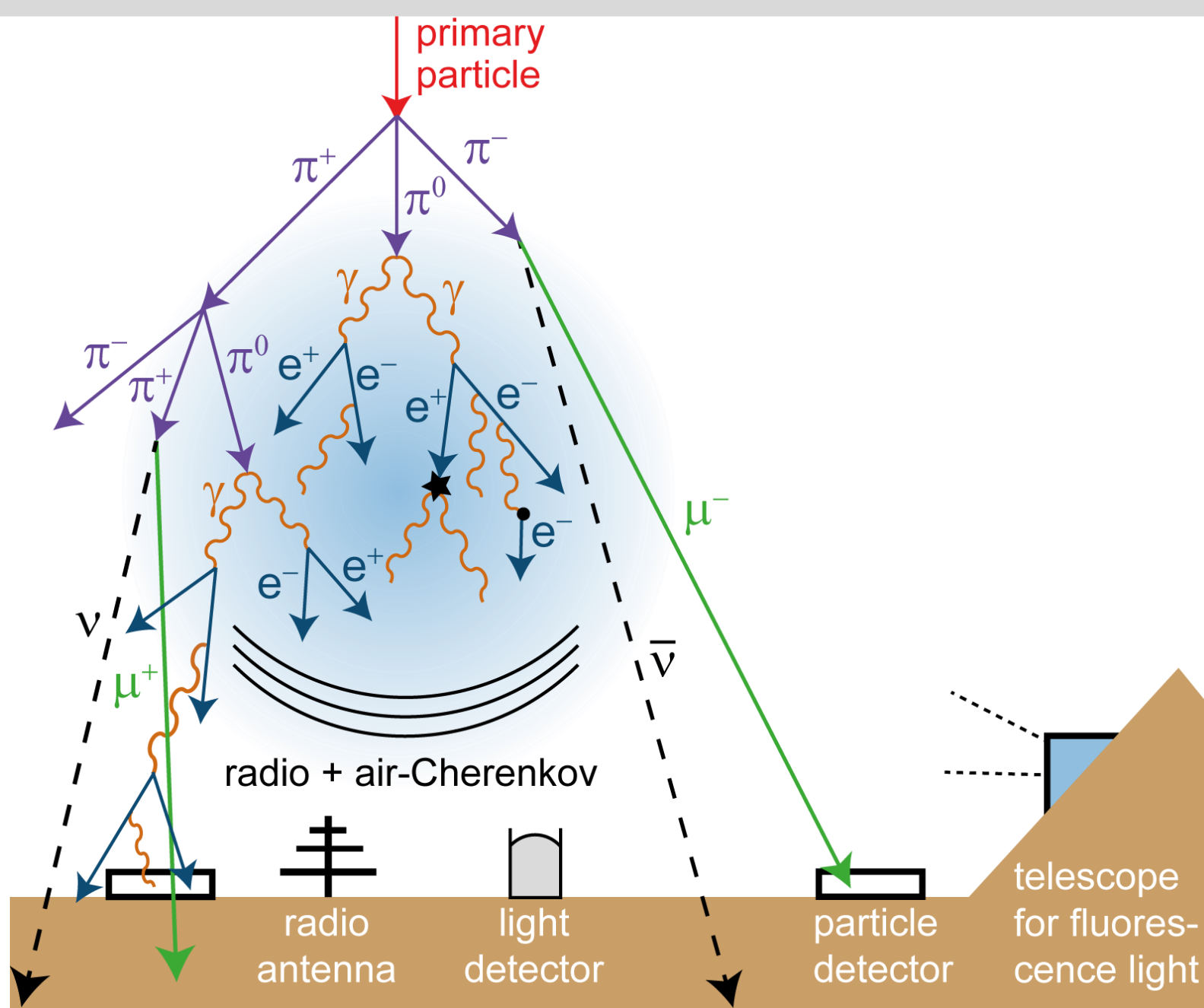
Underlying map (Mercator projection):
Main Geomagnetic Field Total Intensity with contour intervals of 1000 nT
according to US/UK World Magnetic Model - Epoch 2015.0

developed by NOAA/NGDC & CIRES
<http://ngdc.noaa.gov/geomag/WMM>

Map reviewed by NGA and BGS
Published December 2014

Overlaid: Location of radio experiments for cosmic-ray air showers
added on underlying map by Frank G. Schröder
Karlsruhe Institute of Technology (KIT), Germany

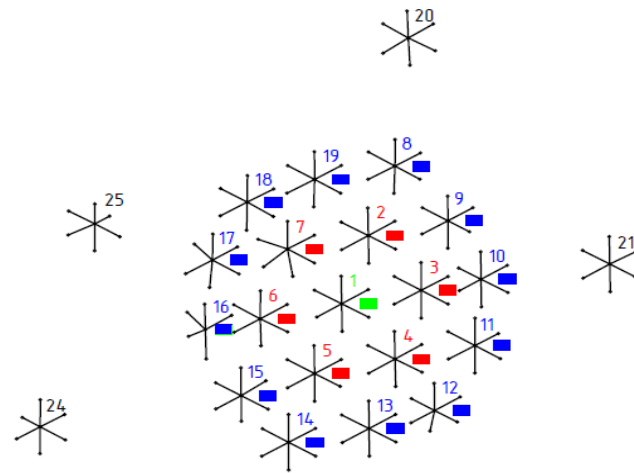
arXiv:
1607.08781



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1607.08781



Tunka-133 and Tunka-Grande at TAIGA

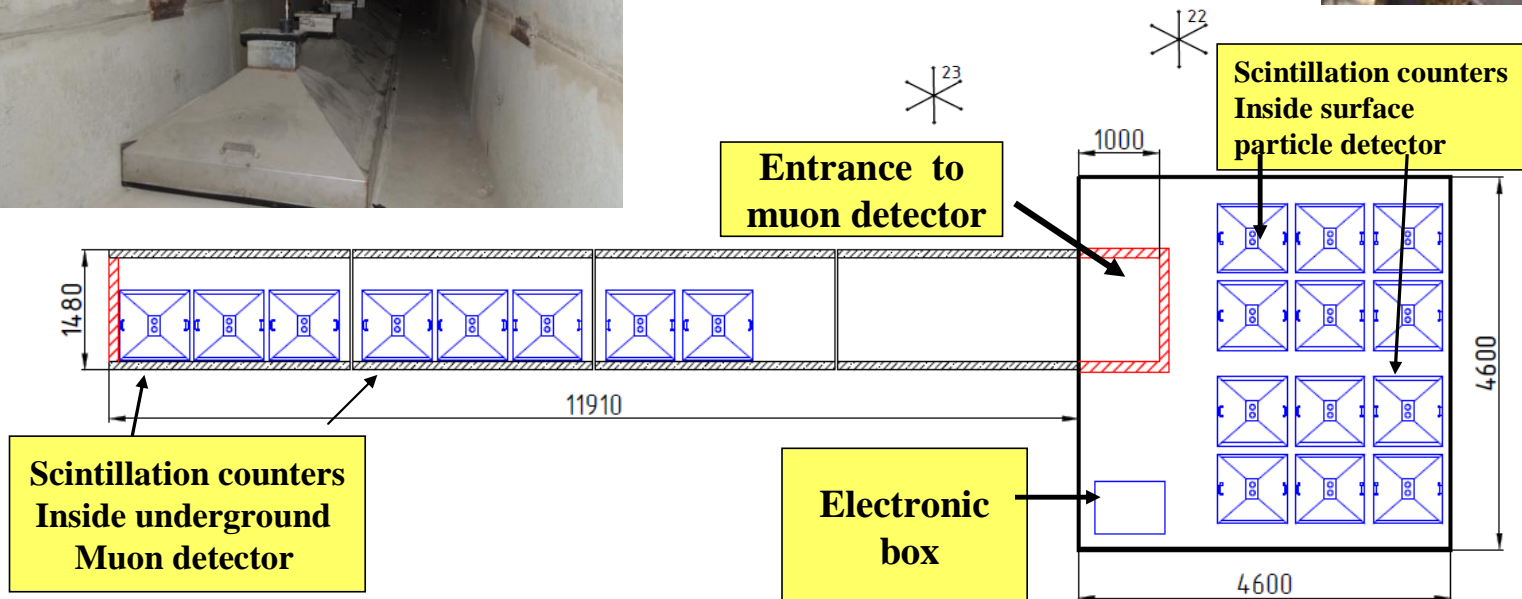


■ Cosmic Rays

- Tunka-133
- Tunka-Grande
- Tunka-Rex

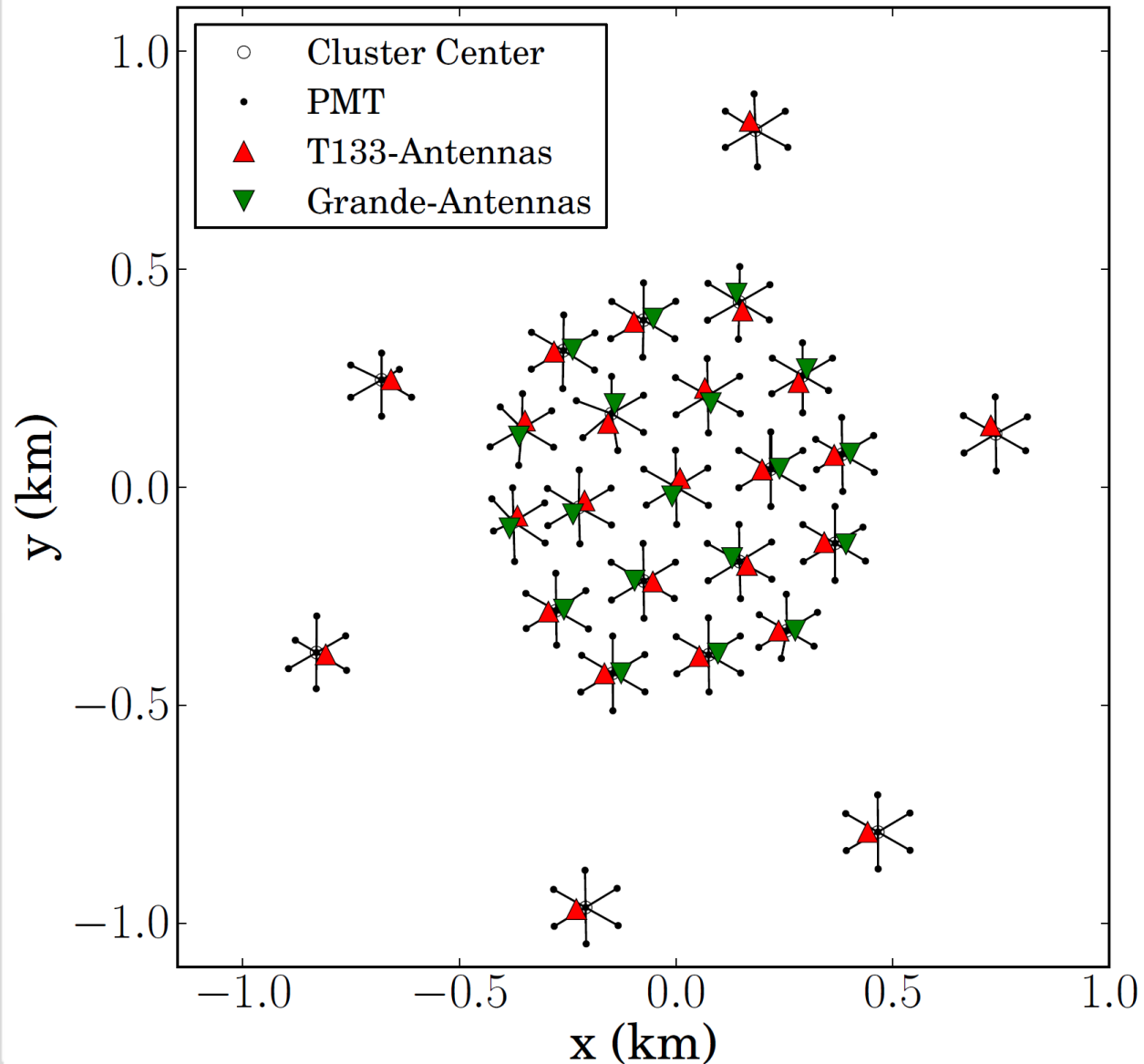
■ Gamma Rays

- HiSCORE
- IACT
- Muon detectors as veto



ECRS 2014
Tunka Coll.

Layout of Tunka-Rex



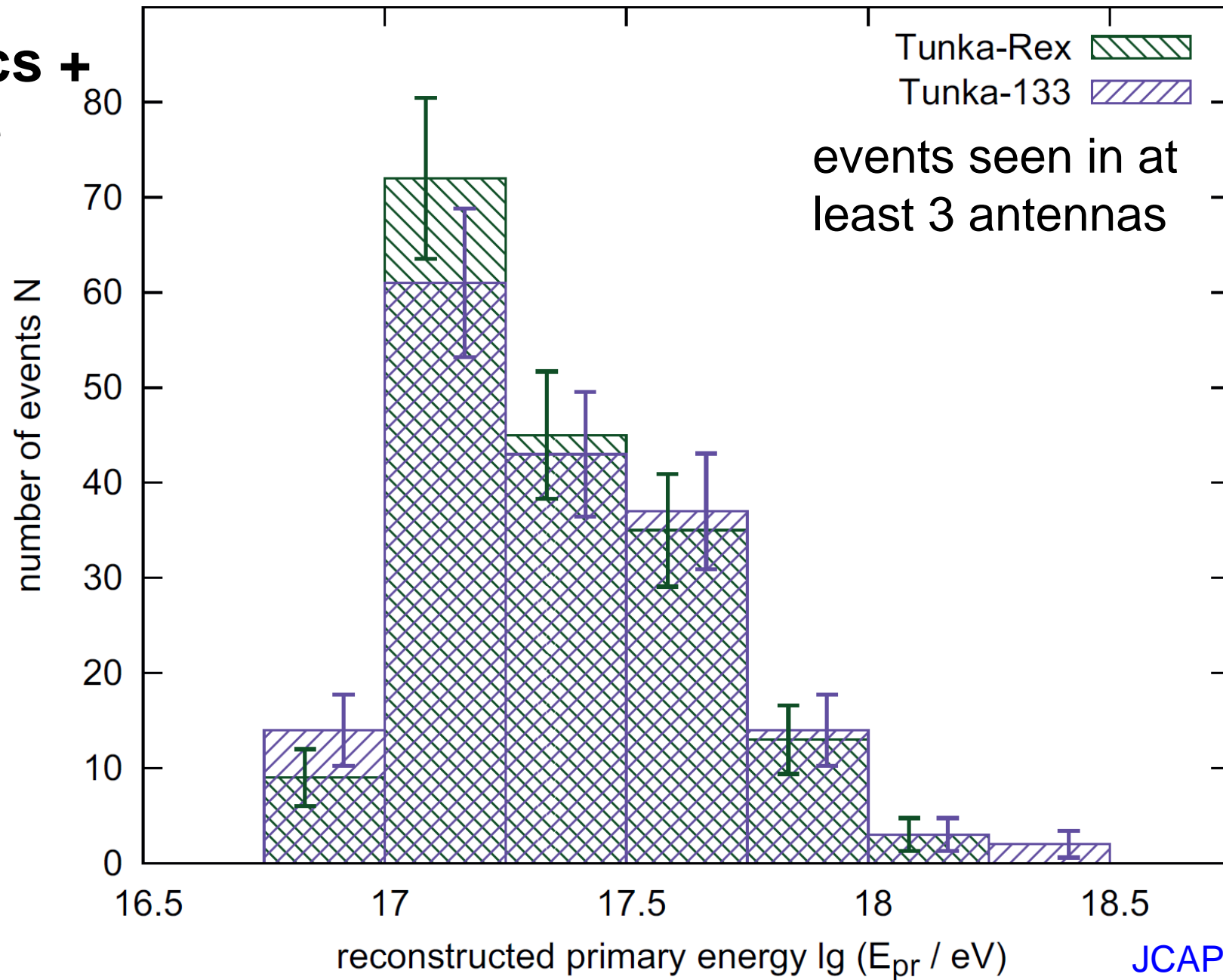
■ Array layout

■ 200 m spacing in 1 km² inner area

■ Fully analyzed: Oct 2012 – Apr 2014

| Year | Stations | Trigger |
|------|----------------------|-----------------------|
| 2012 | 19 | Tunka-133 |
| 2013 | 25 | Tunka-133 |
| 2014 | timing problem | |
| 2015 | 44 | Tunka-133 + Grande |
| 2016 | 44 + 19 in autumn | Tunka-133 + Grande |

Event Statistics + Energy Range

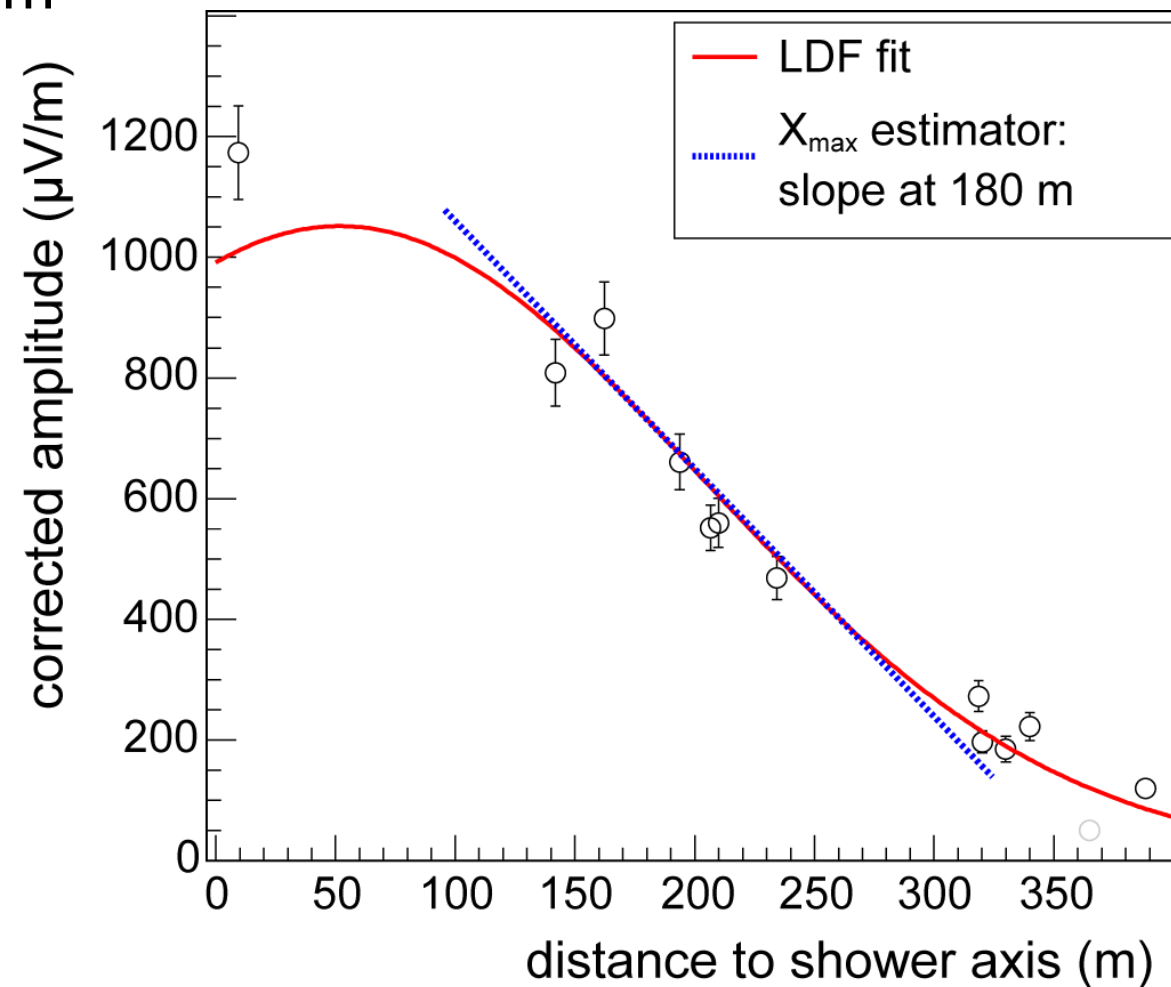
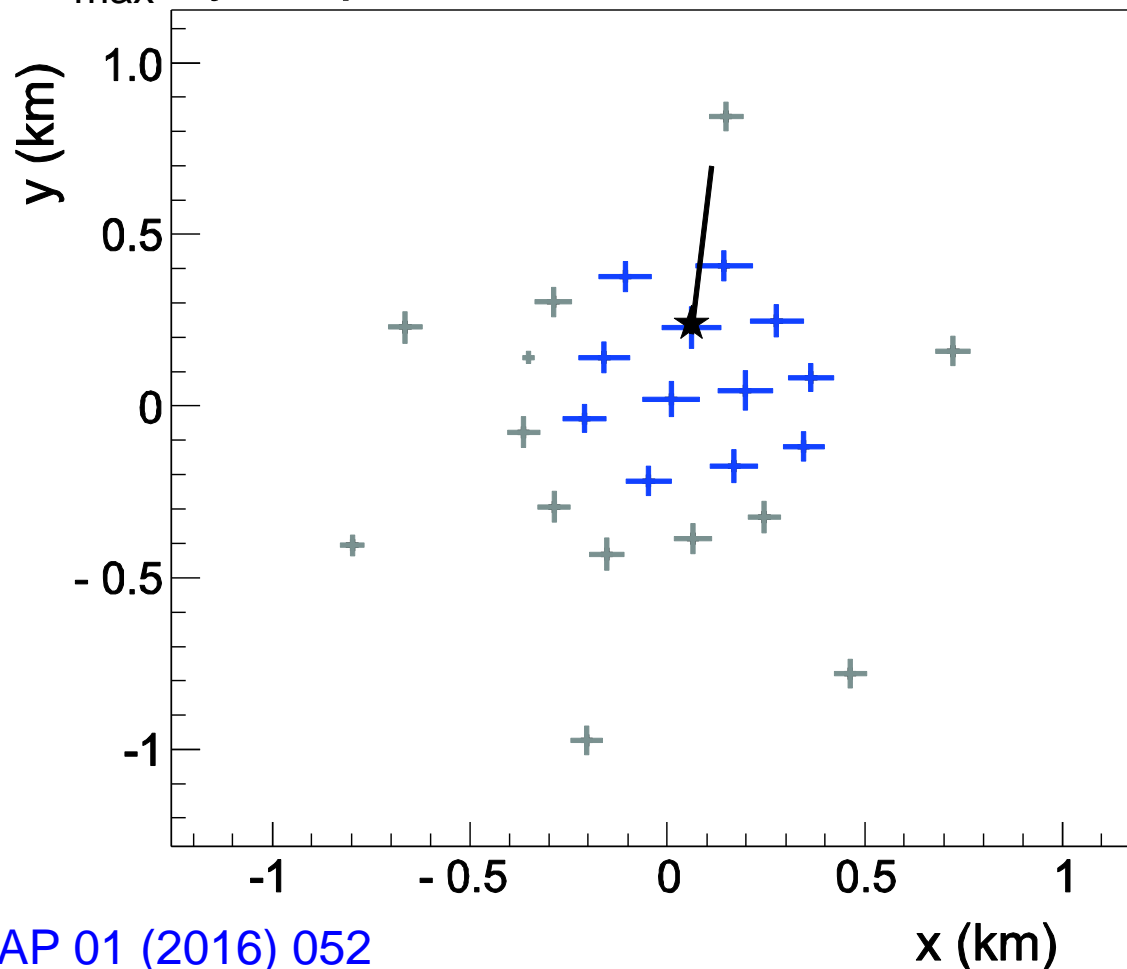


Data analysis with Auger Offline software framework.

JCAP 01 (2016) 052

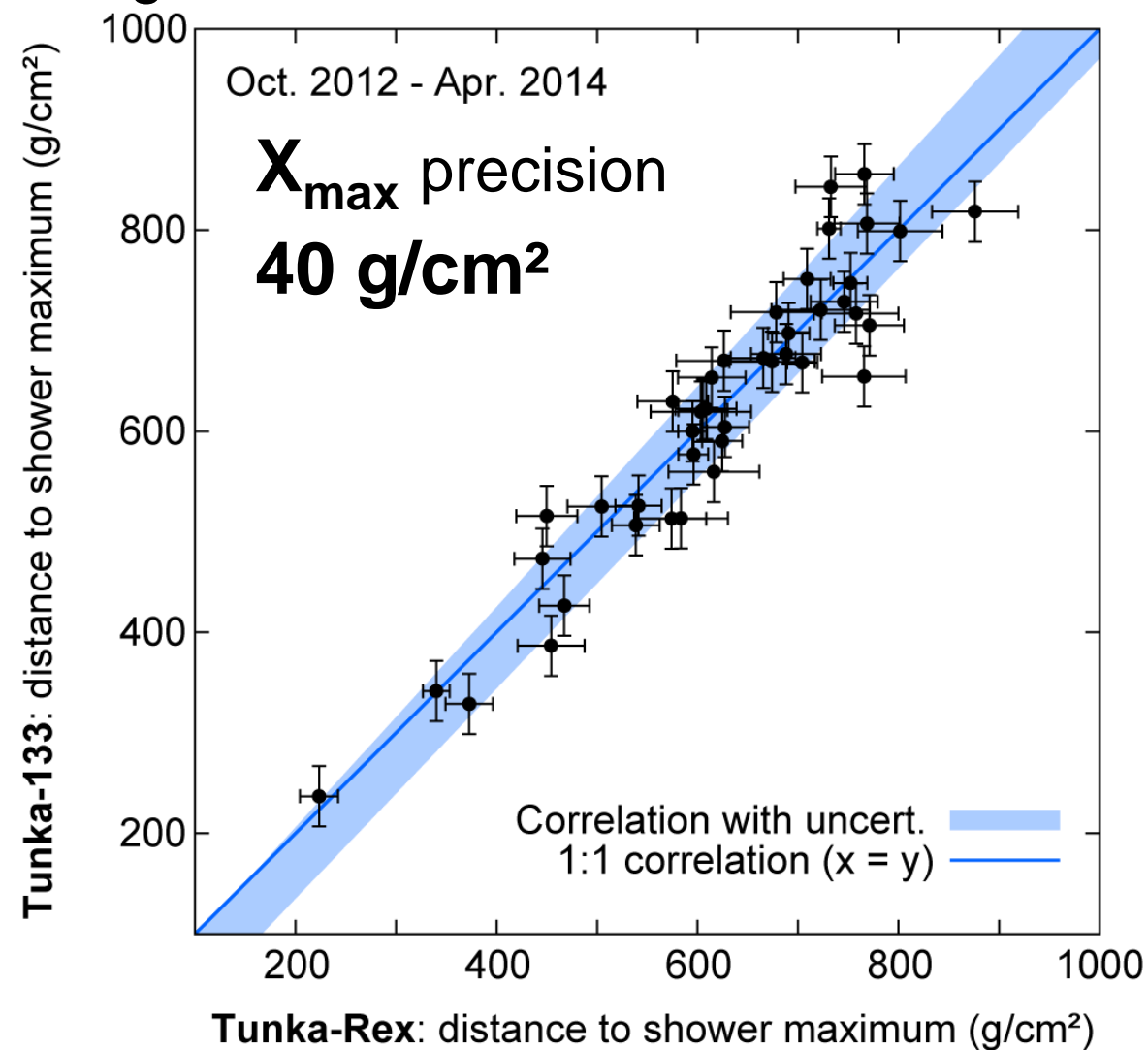
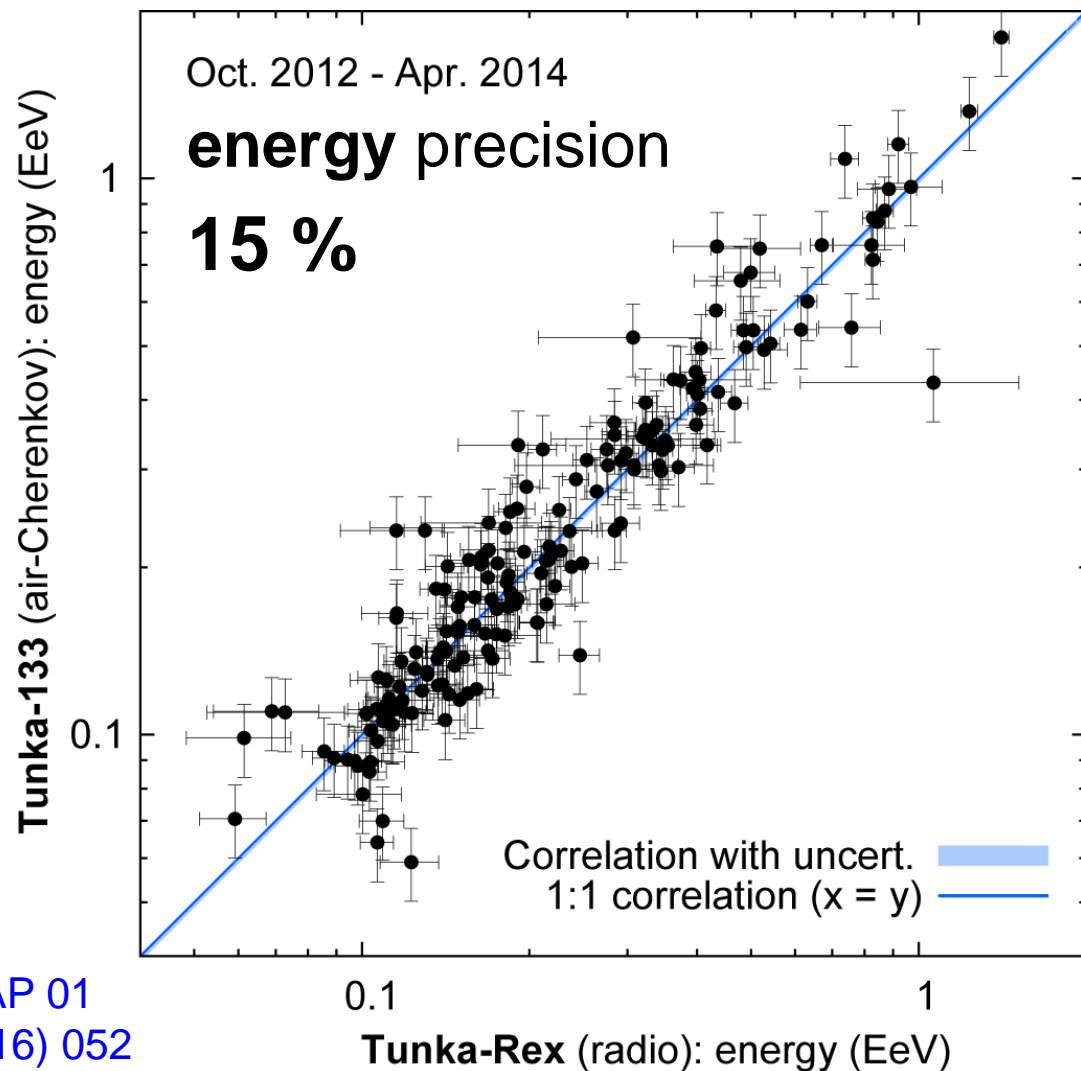
Example Event: Footprint + Lateral Distribution

- Amplitude at 120 m used as energy estimator
- X_{\max} by slope of lateral distribution at 180 m



Tunka-Rex vs. Tunka-133

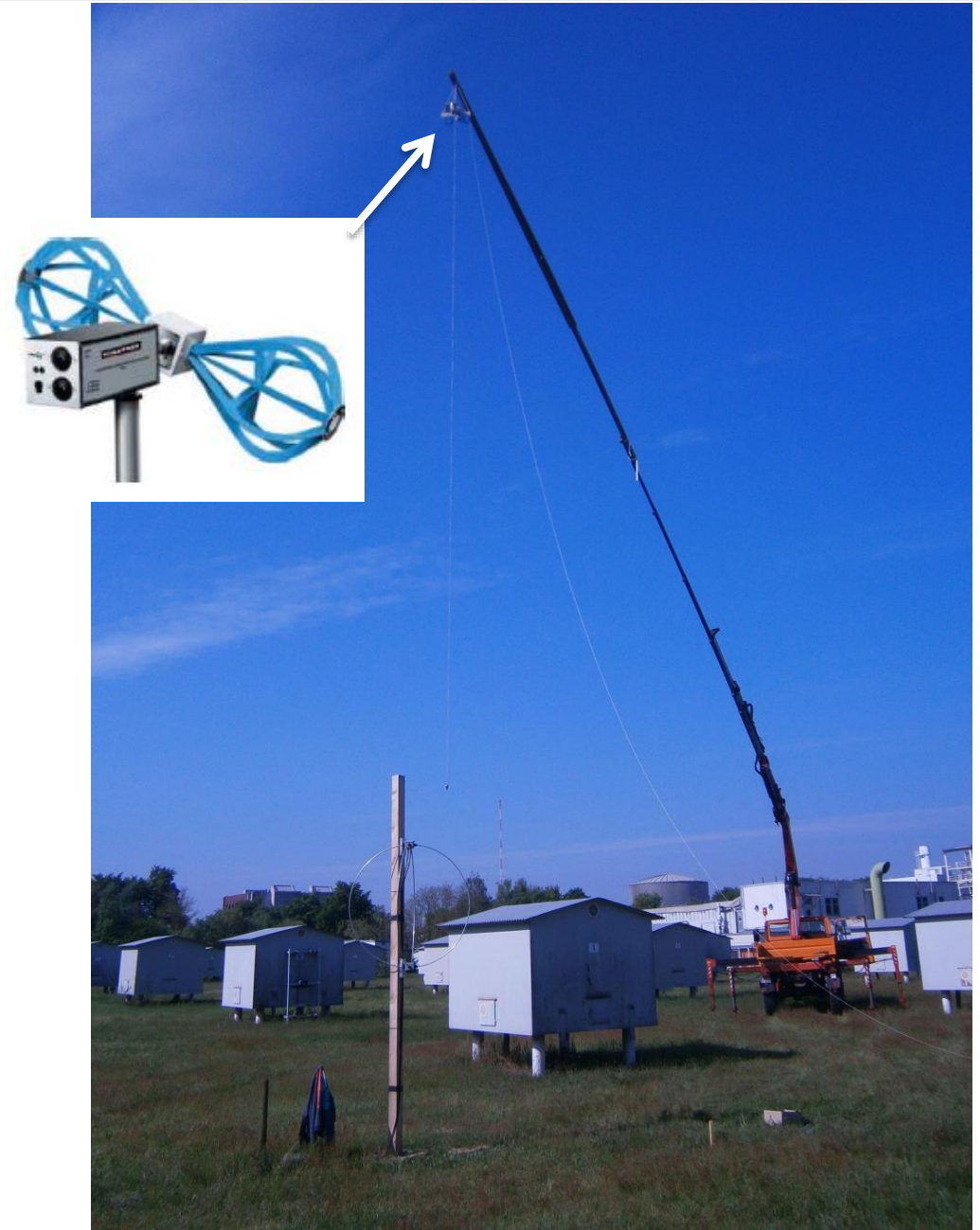
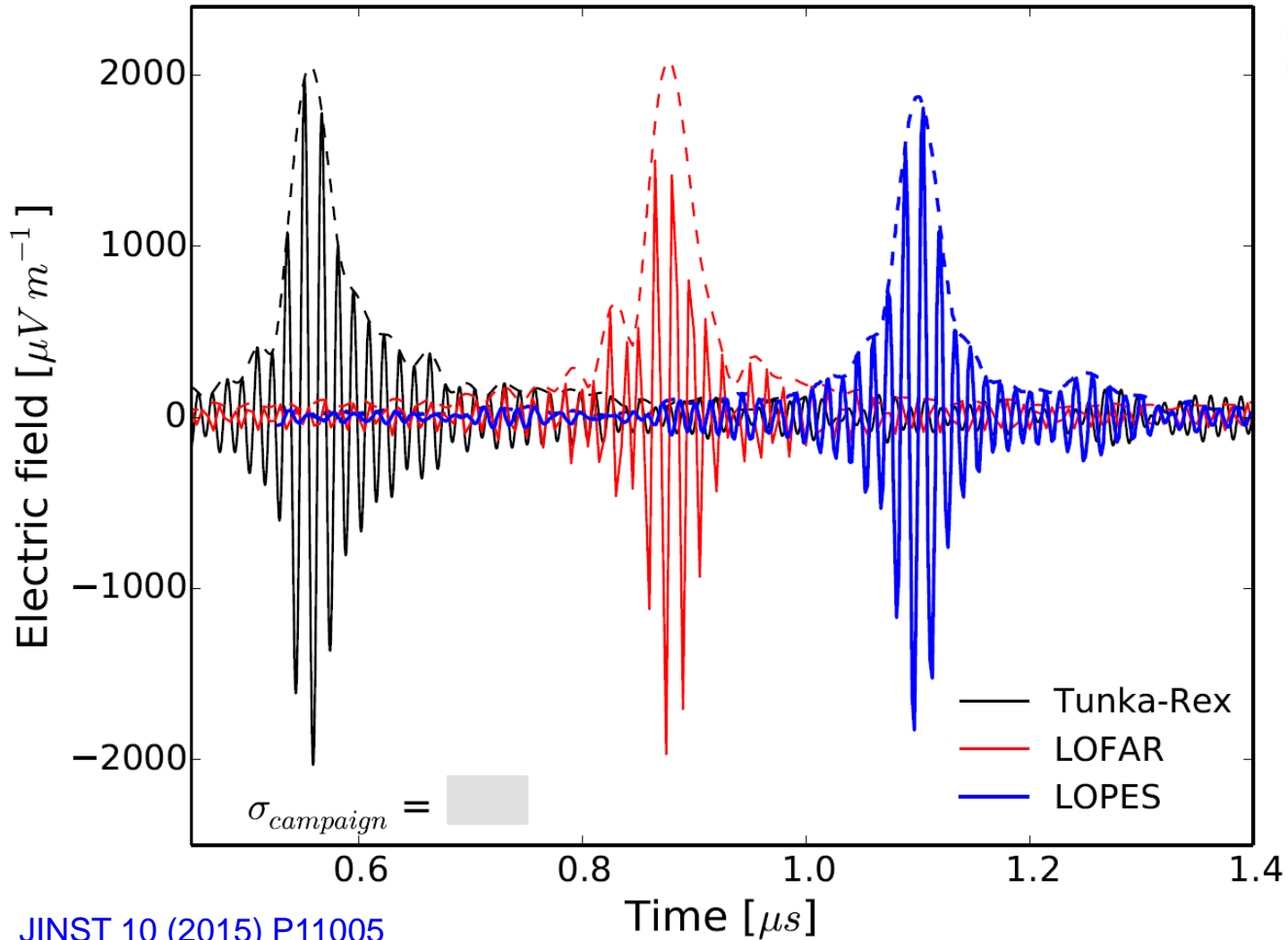
- Radio calibrated on CoREAS simulations agrees with air-Cherenkov



JCAP 01
(2016) 052

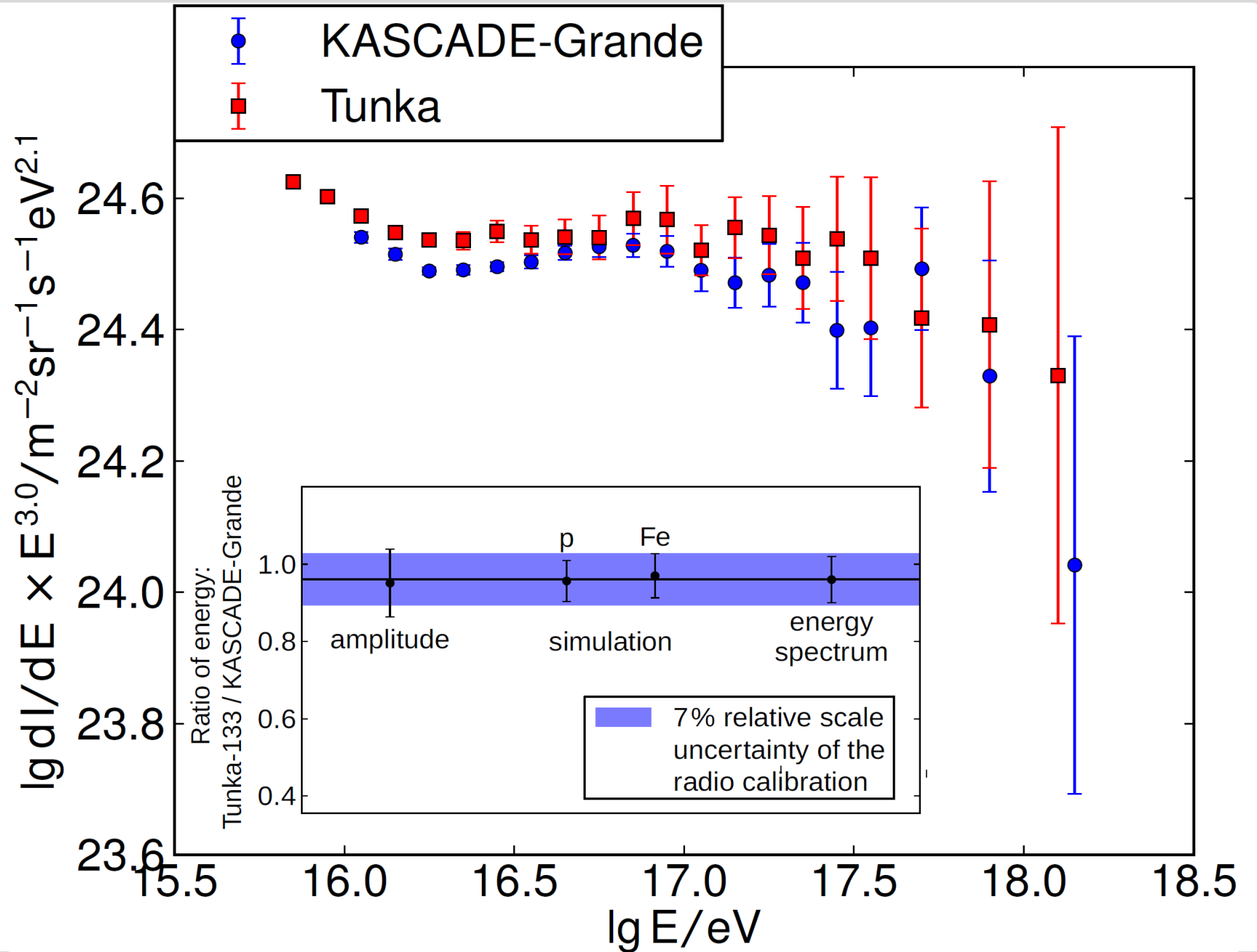
Amplitude Calibration

- Same reference as for LOPES and LOFAR



JINST 10 (2015) P11005

Comparing energy scales via radio



Tunka-Rex + LOPES Colls.,
submitted to PLB

Conclusion

■ Tunka-Rex successful and in time

- Project, antenna deployment, and physics measurement all started in 2012
- Main scientific goal X_{max} *cross-calibration* achieved in 2015 (as planned)

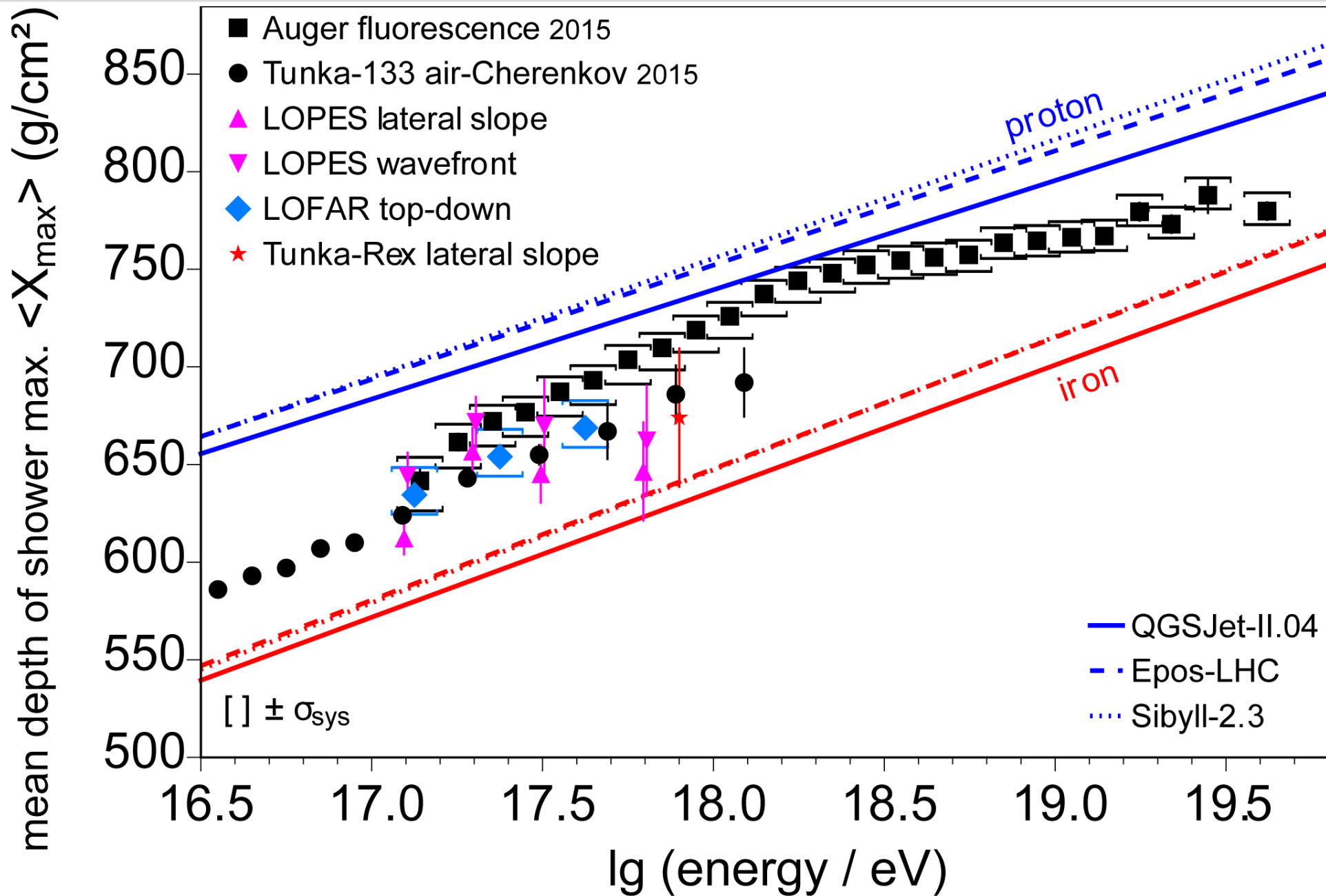
■ Main results

- Energy precision of 15 % with 20 % absolute scale accuracy
- X_{max} precision of 40 g/cm² - improvements under study

■ Outlook

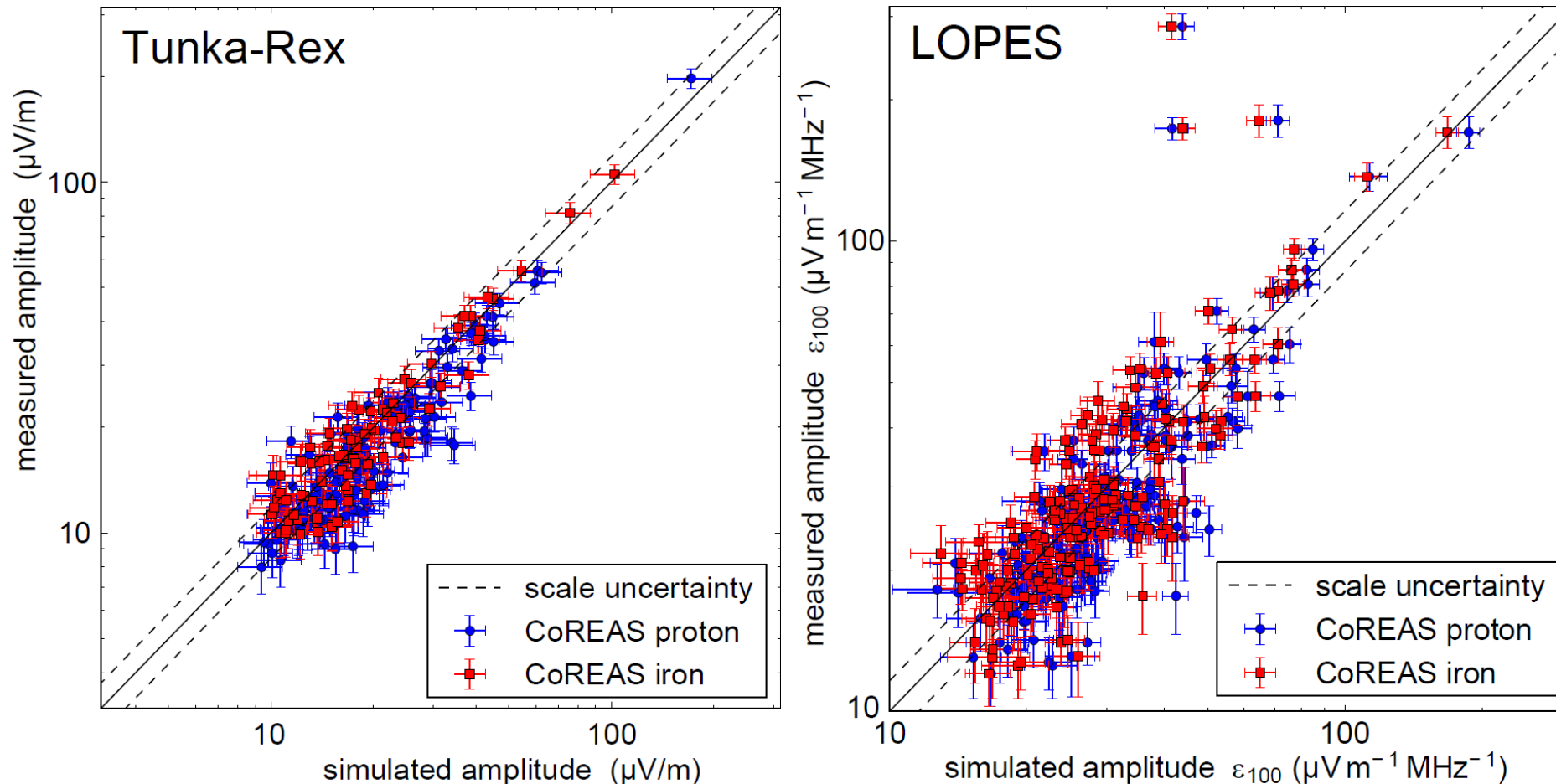
- Extension to 3 antennas per cluster will be finished next month
 - Trigger by Tunka-Grande scintillators boosts the event statistics
- Mass composition between 10^{17} and 10^{18} eV

Backup

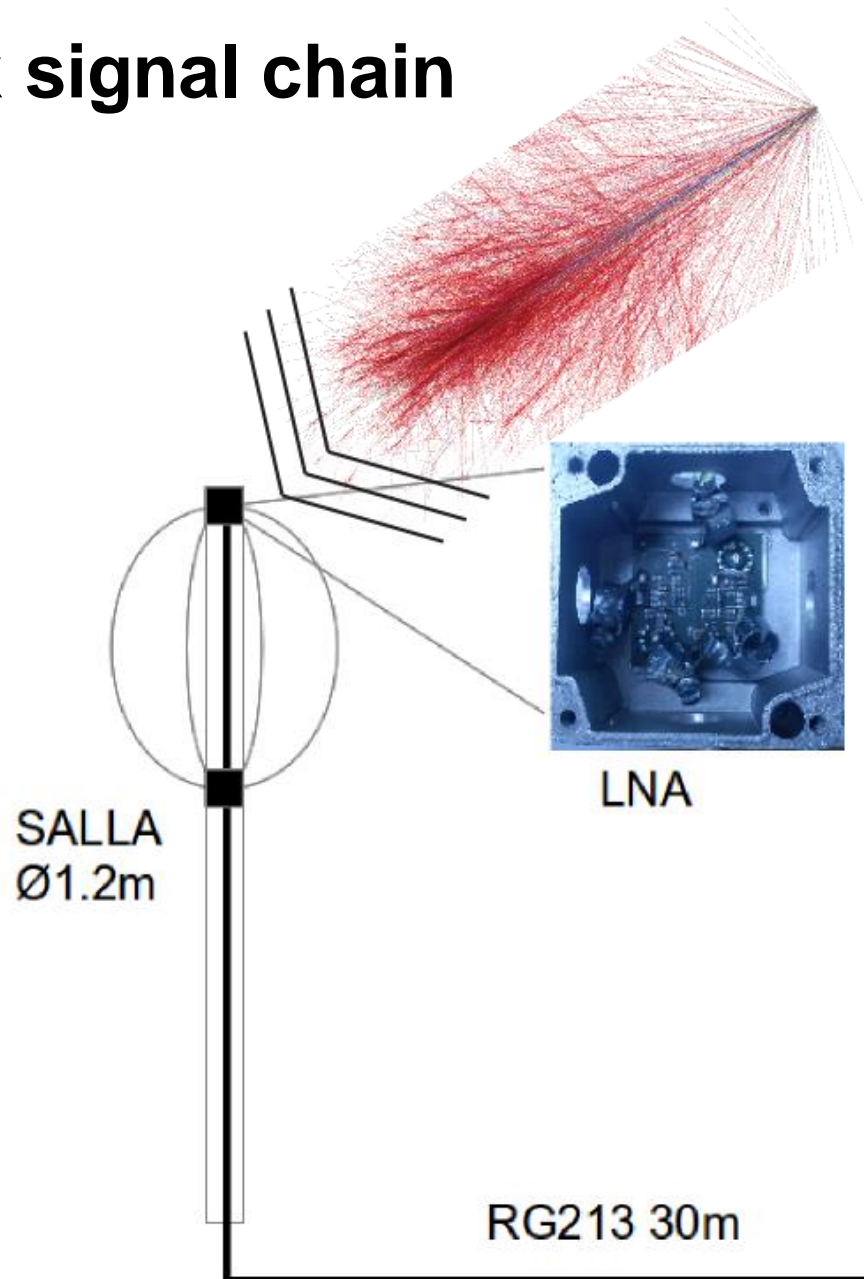


Comparison of energy scales

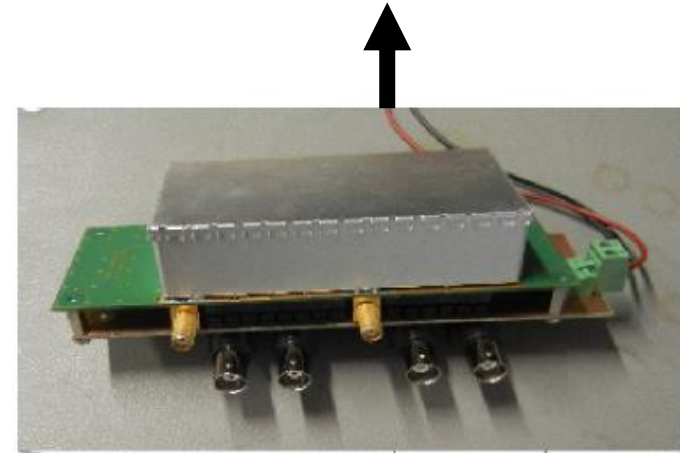
- Simulations with energy reconstructed by host experiments



Tunka-Rex signal chain



Using data acquisition of
Tunka-133 / Tunka-Grande



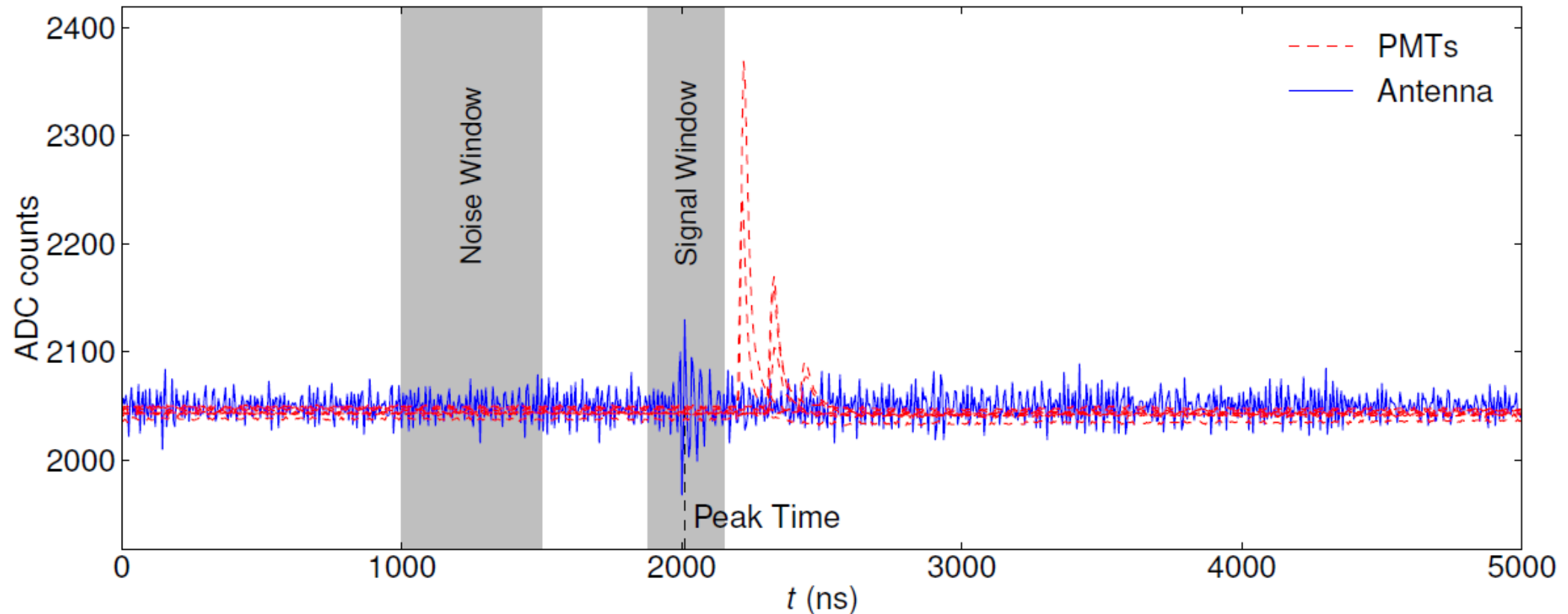
Filter
30-80 MHz

Cluster
Box

ADCs

Reconstruction of Tunka-Rex event

- Signal-to-noise ratio (in power) > 10
 - 5% chance probability of false-positive detection in a single antenna
- 3 antennas + direction agreement with Tunka-133 required



Efficiency of Tunka-Rex standard reconstruction

