



# Particle Production after a first order phase transition: What happens when bubbles collide?

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# First Order Phase Transitions

- A scalar field  $\phi$  with a temperature-dependent effective potential

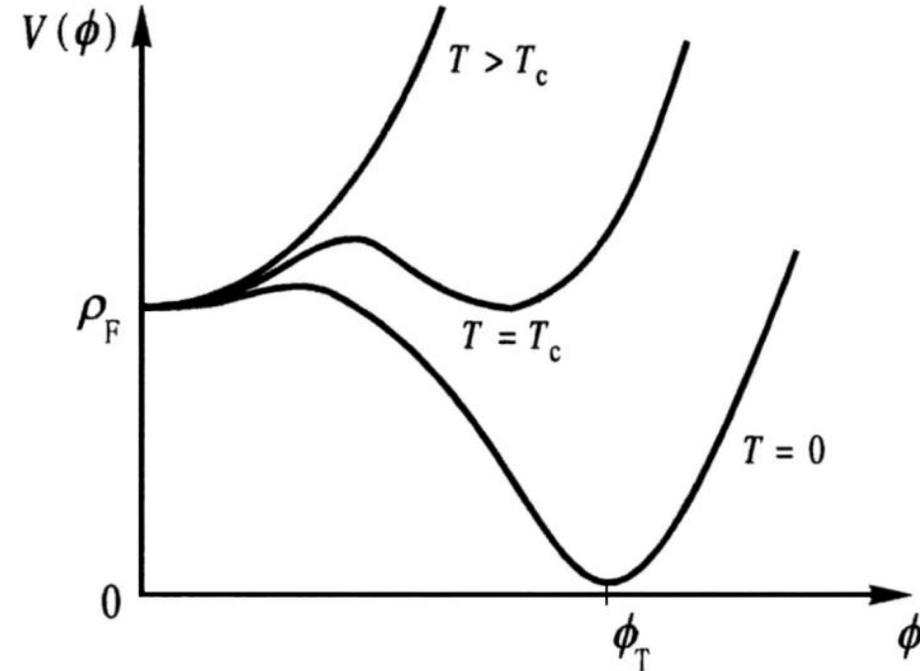
$$V_{\text{eff}}(\phi, T) = V_{\text{tree}}(\phi) + \Delta V(\phi, T)$$

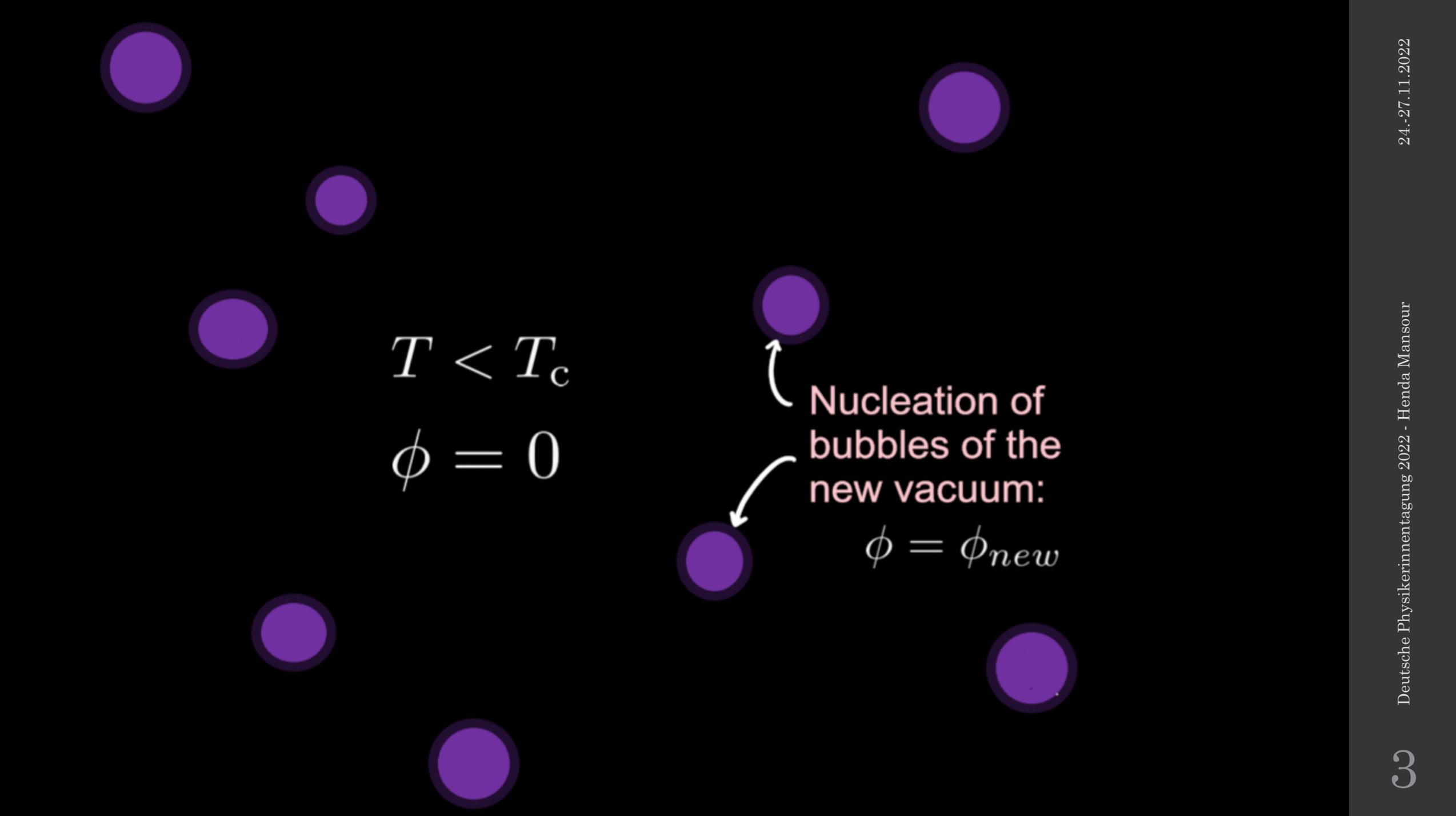
- Example: HT approximation of Higgs 1-loop effective potential

$$V(\phi, T) = \alpha(T^2 - T_0^2)\phi^2 - \beta T\phi^3 + \frac{\lambda(T)}{2}\phi^4$$

- New local minimum with decreasing temperature and at  $T = T_c$  both minima are degenerate.
- Strength of the transition  $\alpha$ , amount of energy released compared to the energy density of the plasma

$$\alpha = \frac{\rho_v}{\rho_{\text{plasma}}} \propto \frac{\Delta V}{T^4}$$

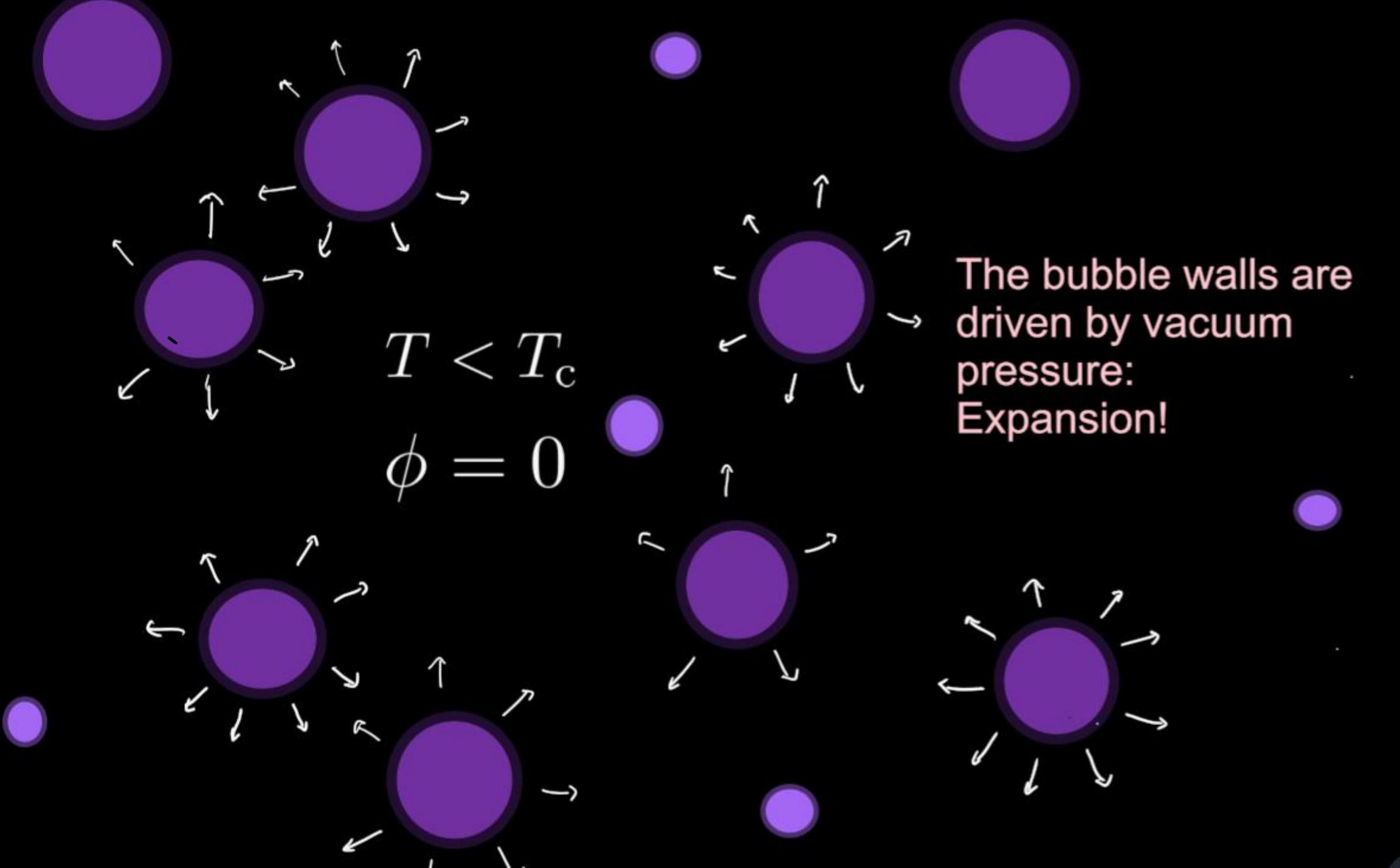



$$T < T_c$$

$$\phi = 0$$

Nucleation of  
bubbles of the  
new vacuum:

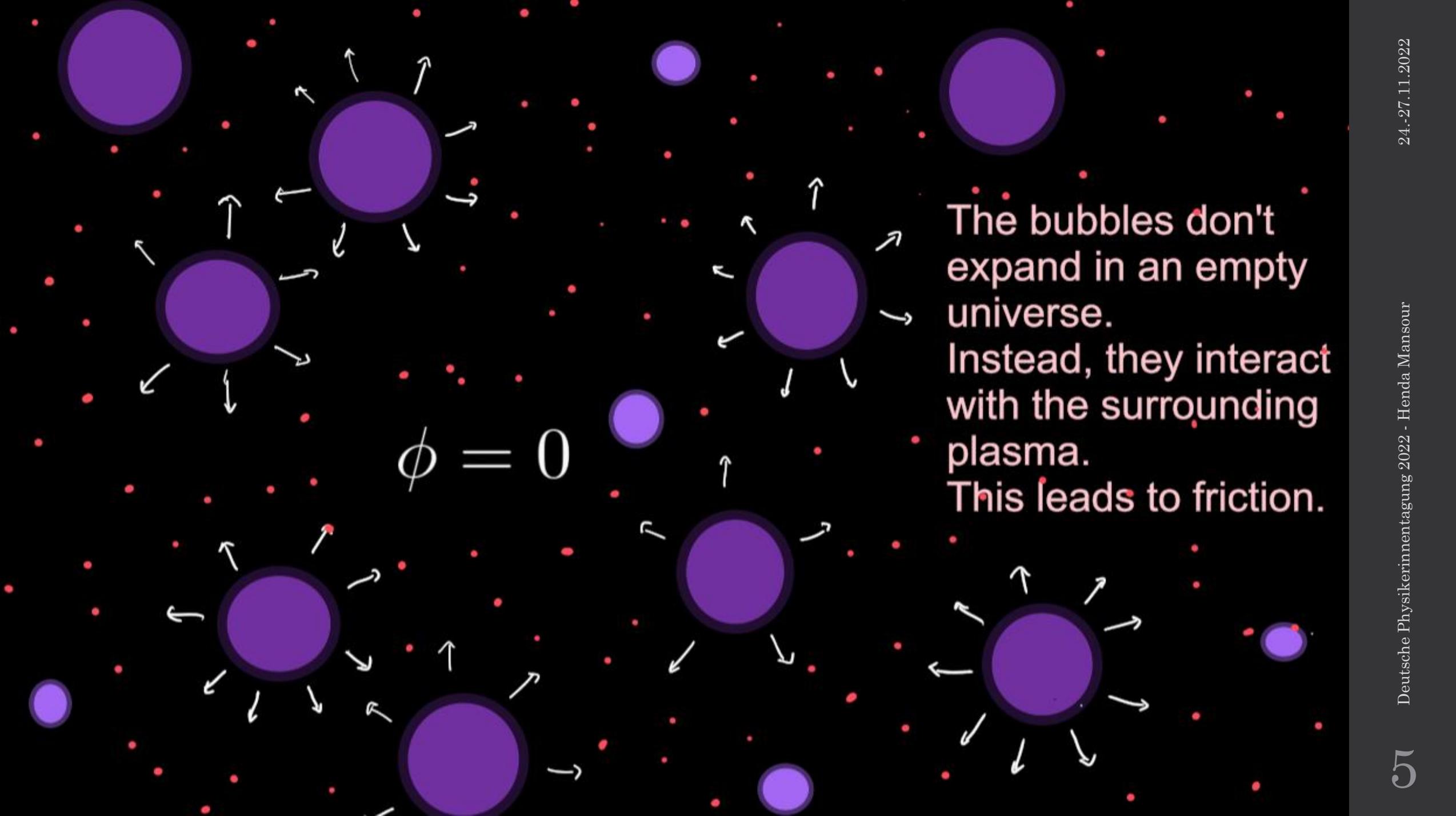
$$\phi = \phi_{new}$$



$$T < T_c$$

$$\phi = 0$$

The bubble walls are driven by vacuum pressure: Expansion!



$$\phi = 0$$

The bubbles don't expand in an empty universe. Instead, they interact with the surrounding plasma. This leads to friction.

## Bubble Collisions

In a strong first-order phase transition, the walls could reach very high lorentz factors  $\gamma_w$

The energy per area in the walls

$$\frac{E_w}{A} = \frac{2}{3} v^2 \frac{\gamma_w}{l_w}$$

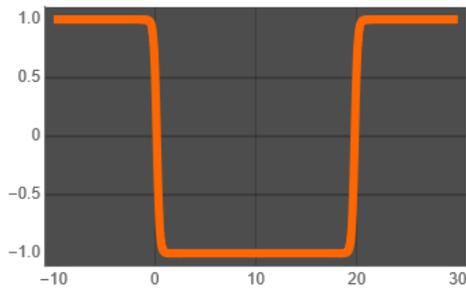
# Investigating The Dynamics

To study the dynamics, we consider planar waves and the following potential

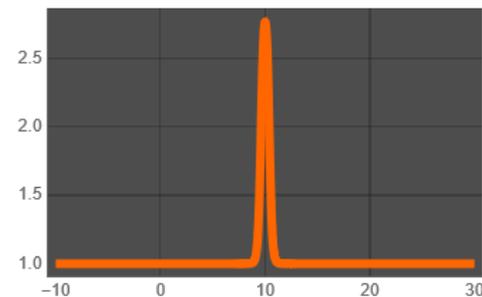
$$V(\phi) = av^2\phi^2 - (2a + 4)v\phi^3 + (a + 3)\phi^4 \quad \text{with a coupling} \quad \mathcal{L}_\chi = \frac{1}{2}m_\chi^2\chi^2 + \lambda\phi^2\chi^2$$

In the ultra-relativistic limit, the equation of motion can be simplified to:

$$\partial_s^2\phi + \frac{1}{s}\partial_s\phi + \frac{dV}{d\phi} = 0 \quad s = \sqrt{t^2 - x^2}$$

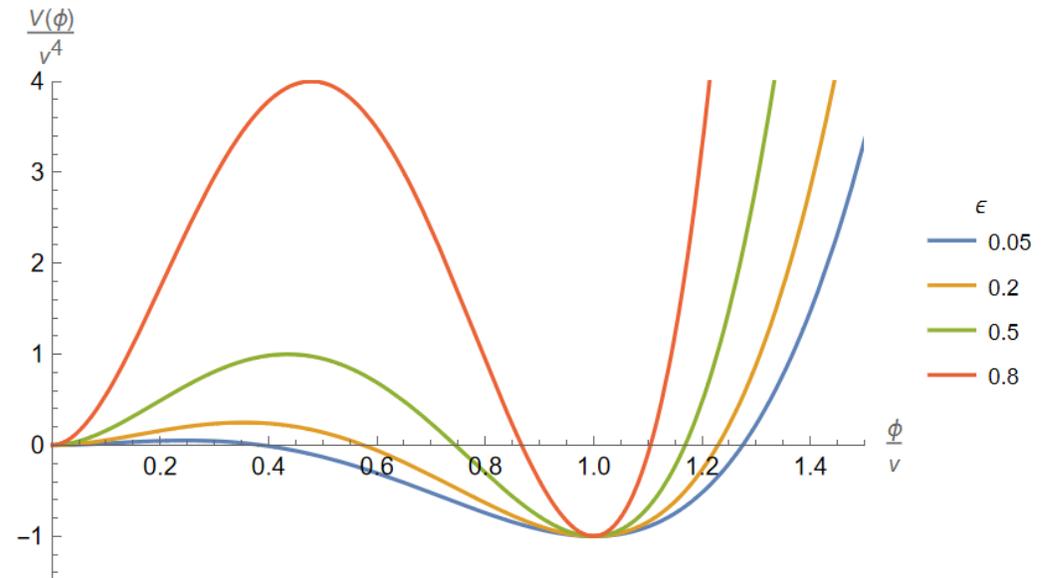


**Before the collision:**  
Two planar walls  
approaching each other.



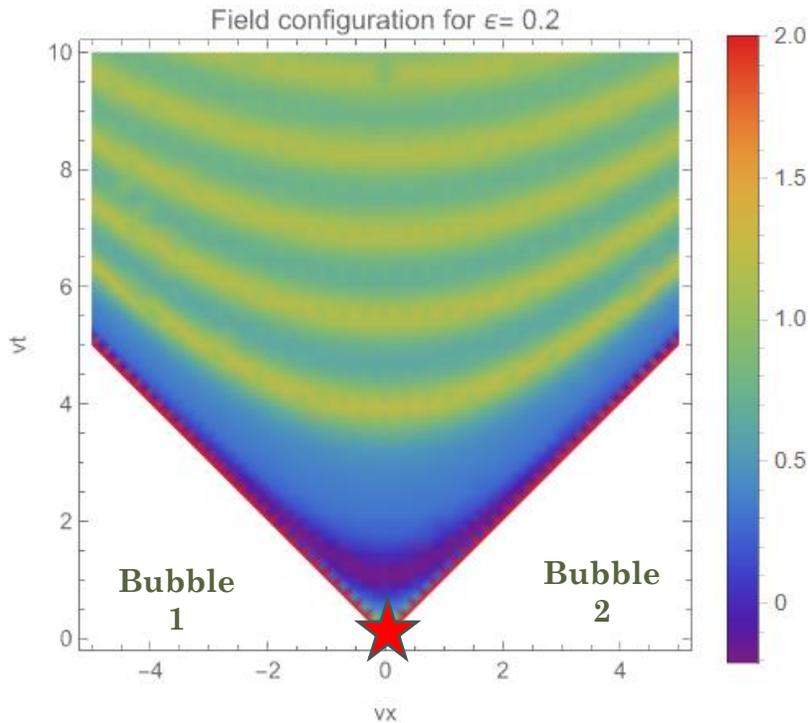
**At collision:**  
The two walls superpose and the  
field experiences a jump.

$$\phi_{\text{after}} = 2\phi_{\text{inside}} - \phi_{\text{outside}}$$



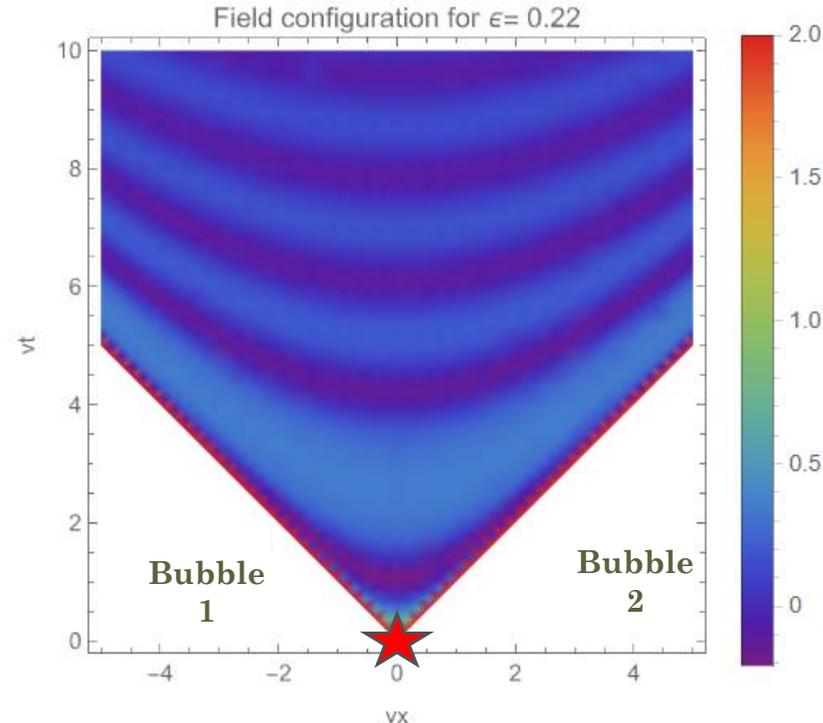
$$\epsilon = \frac{V_m - V_-}{V_m - V_+} = \frac{a^3(a+4)}{a^3(a+4) + 16(a+3)^3}$$

# Field Configuration After Collision



**Inelastic collision: the field oscillates around the new minimum at  $\varphi = 1$ .**

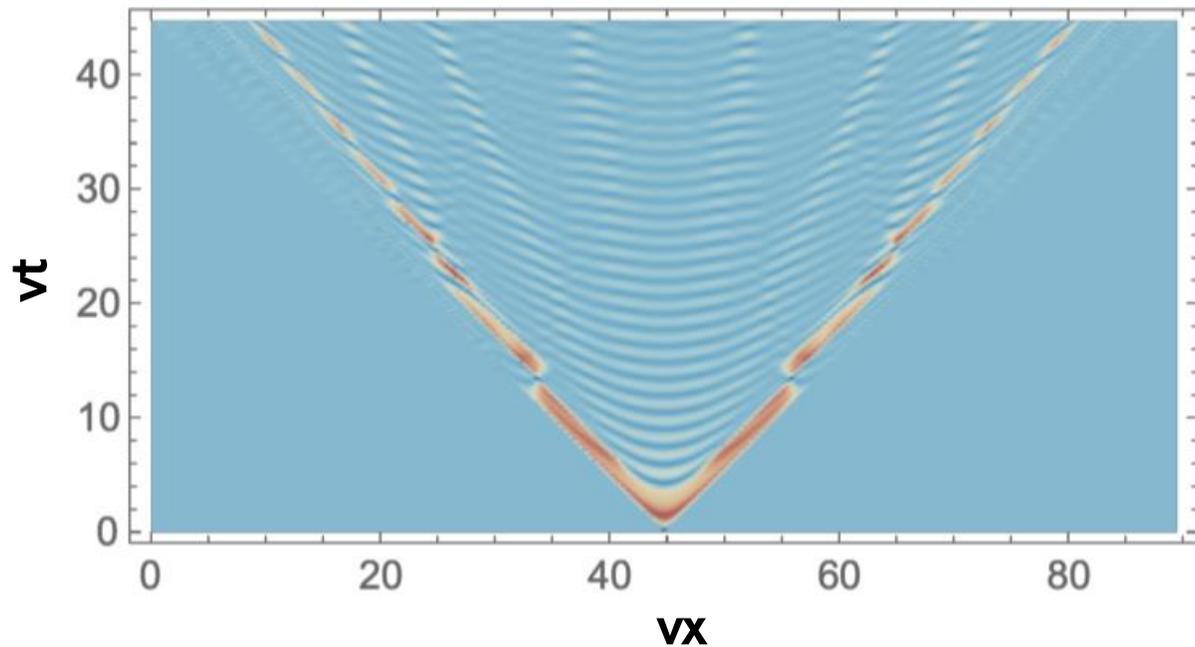
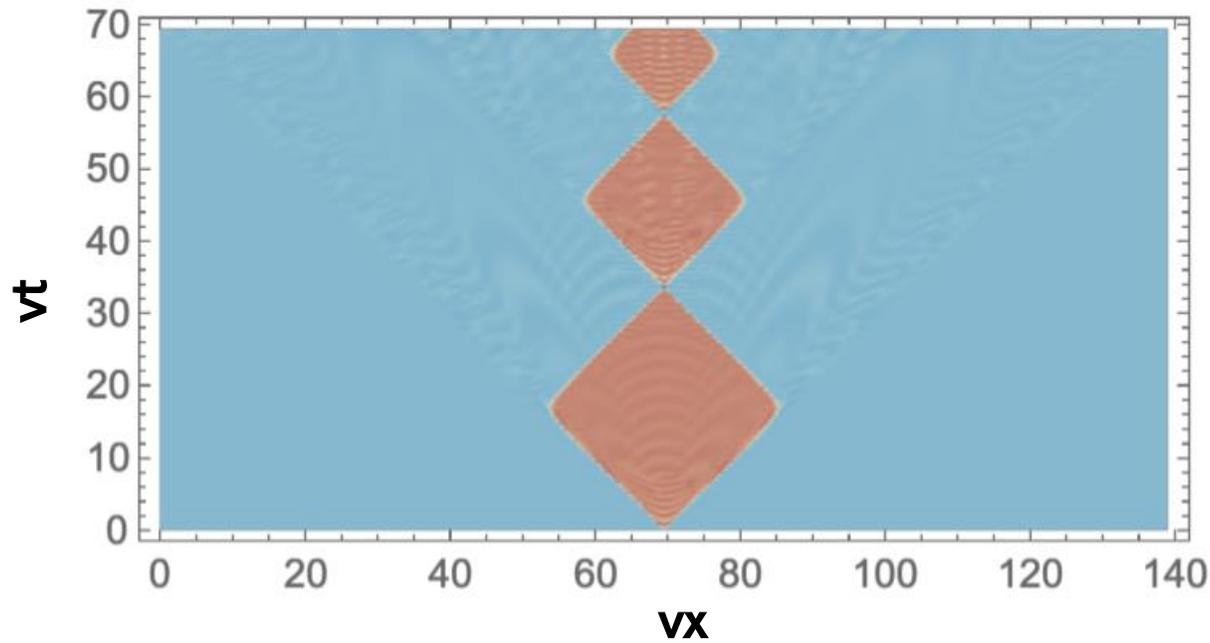
★ Marks the collision point.



**Elastic collision: the field oscillates back around the old minimum at  $\varphi = 0$ .**

**The behaviour of the field strongly depends on the potential shape!**

[Thomas Konstandin and Géraldine Servant. Apr 2011, ArXiv:1104.4793]



## Results from simulations of bubble collisions

### (almost) elastic collisions:

Multiple cycles of collisions before establishment of the true vacuum. Re-establishment of the false vacuum in the regions between the walls.

### Inelastic collisions:

The true vacuum is established after one collision

[Ryusuke Jinno, Thomas Konstandin, and Masahiro Takimoto. sep 2019 , ArXiv: 1906.02588]

# Particle Production: A First Approach

Treat  $\varphi(x, t)$  as an external background field to which the other quantum fields couple

$$\frac{N}{A} = 2 \int \frac{dp_z dw}{(2\pi)^2} \underbrace{|\tilde{\phi}(p_z, w)|^2}_{\text{Fourier transformation of scalar field configuration}} \underbrace{\text{Im} \left( \tilde{\Gamma}^{(2)}(w^2 - p_z^2) \right)}_{\text{Fourier transformation of the effective action}}$$

Decompose the scalar configuration in Fourier modes and interpret each mode as a particle.

The integral over all „particles“ multiplied with the decay rate gives the number of particles produced.

[R. Watkins, L. Widrow. Aspects of reheating in first order inflation. Nucl. Phys. B, 374:446–468, 1992]

# Other non-perturbative effects

- Other mechanisms could play a role, so the Fourier decomposition approach might not be enough.
- Parametric resonance: exponential growth due to bose enhancement [1]

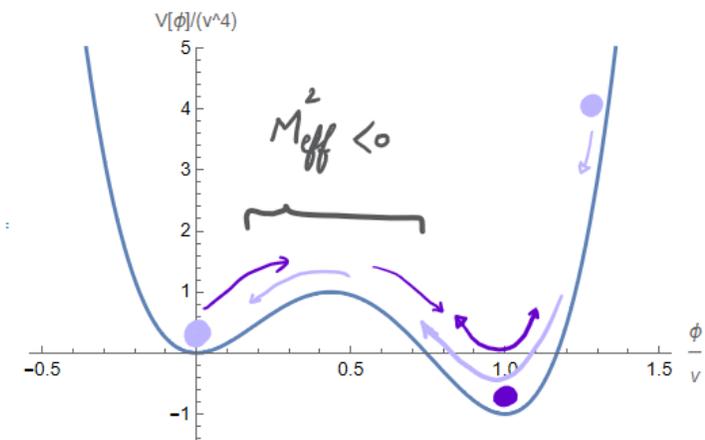
$$\ddot{\chi}_k + (k^2 + m_\chi^2(0) + 2\lambda\phi(t)^2) \chi_k = 0$$

$$\ddot{\chi}_k + \omega_k^2(t) \chi_k = 0$$

- Tachyonic preheating: when the  $\phi$  field is over the potential barrier, the effective mass squared is negative, this leads to exponential growth of modes with  $k < m_\phi$

$$\phi_k(t) \propto \exp\left(it\sqrt{k^2 + M_\phi^2}\right) \rightarrow \phi_k(t) \propto \exp\left(t\sqrt{M_\phi^2 - k^2}\right)$$

- Since **the field is spatially inhomogeneous** in the case of Bubble collisions, estimates of these effects need to be Computed more carefully.



[1] Y. Shtanov, J. Traschen, and R. Brandenberger. Universe reheating after inflation. Physical Review D, 51(10):5438–5455, may 1995  
 [2] Lev Kofman. Tachyonic preheating, 2001

# Conclusions and further questions

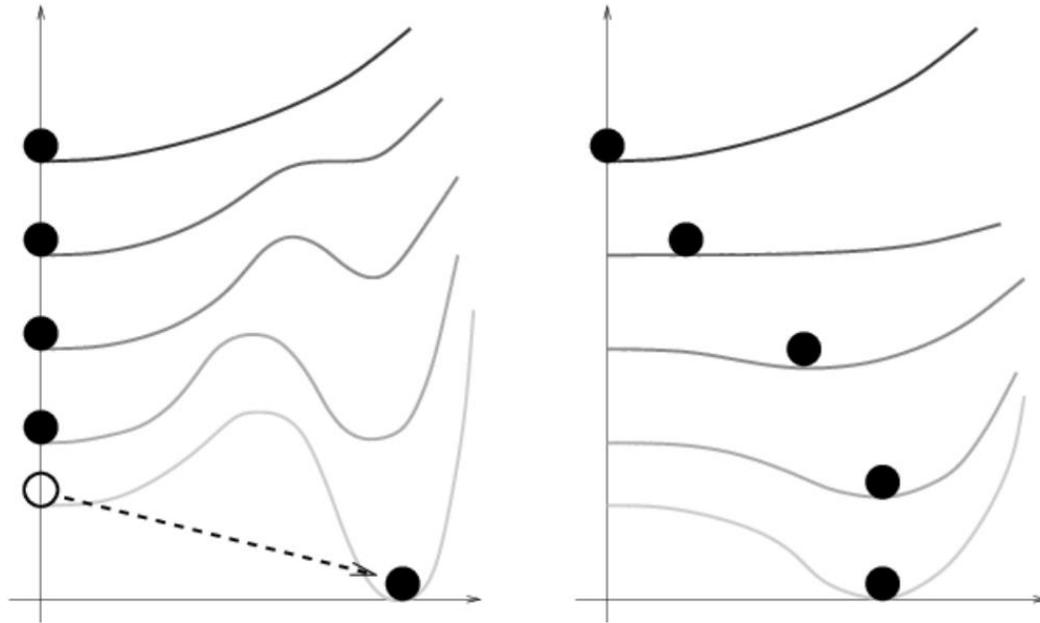
- Bubbles are cool!
- Research on the topic has until now over-simplified aspects of the field dynamics as well as other possibly very relevant mechanisms of particle production.
- To get the full picture in such scenarios, we will need to further investigate the interplay of different mechanisms.
- The importance of each mechanism depends on the dynamics of the field (i.e. Elastic vs. Inelastic collisions) , but also on the wall velocity.



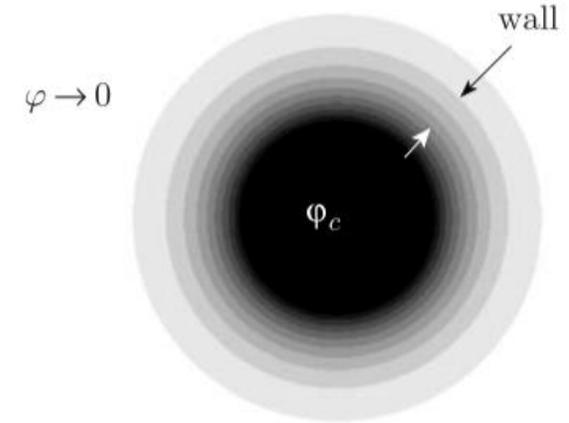
**Thank you for your  
attention!**

Lisa Website.  
Credit: D. Weir,  
University of  
Helsinki

# First-Order Vs. Second-Order PT



Effective potential as function of the scalar field for different temperatures. **Right:** The field rolls down continuously to the new vacuum (second order phase transition). **Left:** Abrupt change of the field value (first order phase transition)



Bubble nucleation in the case of first-order PT.