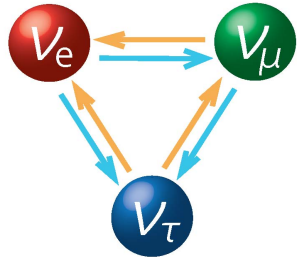
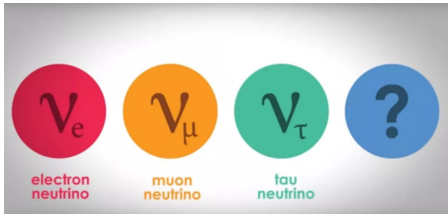
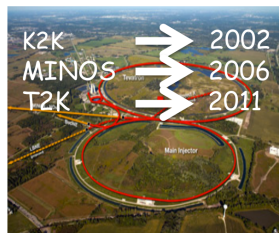
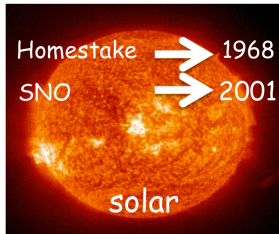
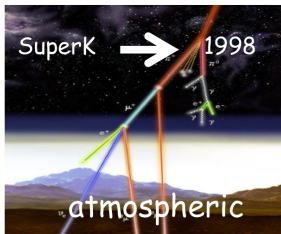


Present Status of Neutrino Physics from a Theory Perspective



Vedran Brdar (CERN-TH)

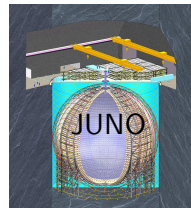
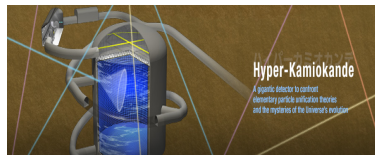
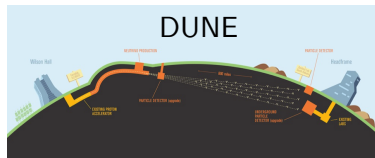
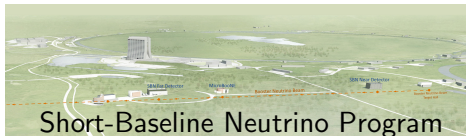
Neutrino Oscillations



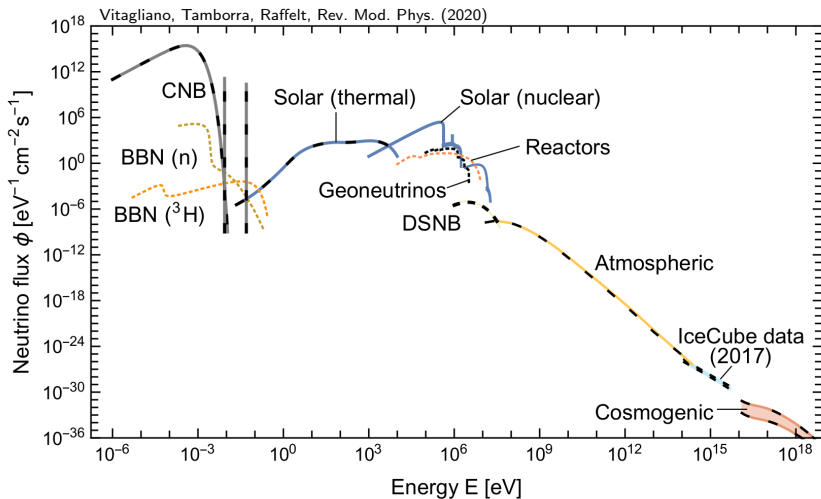
$$P_{\alpha\beta} = \sin^2 2\theta \sin^2\left(\frac{\Delta m^2 L}{4E}\right)$$

(Some of the) Open Questions

- ▶ What is the origin of neutrino mass?
- ▶ CP violation in neutrino sector?
- ▶ Ordering of neutrino masses?
- ▶ Is the neutrino its own antiparticle?
- ▶ Absolute neutrino mass scale?
- ▶ Sterile neutrinos?

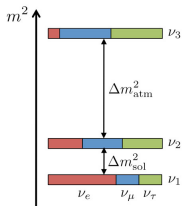


Grand Unified Neutrino Spectrum at Earth

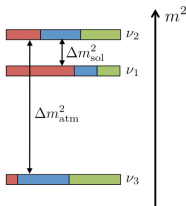


Neutrino Mass Ordering

normal hierarchy (NH)



inverted hierarchy (IH)



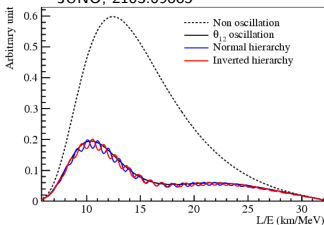
NuFIT, 2007.14792

	Normal Ordering (best fit)		Inverted Ordering ($\Delta\chi^2 = 2.7$)	
	bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range
$\sin^2 \theta_{12}$	$0.304^{+0.013}_{-0.012}$	0.269 \rightarrow 0.343	$0.304^{+0.013}_{-0.012}$	0.269 \rightarrow 0.343
$\theta_{12}/^\circ$	$33.44^{+0.78}_{-0.75}$	31.27 \rightarrow 35.86	$33.45^{+0.78}_{-0.75}$	31.27 \rightarrow 35.87
$\sin^2 \theta_{23}$	$0.570^{+0.018}_{-0.024}$	0.407 \rightarrow 0.618	$0.575^{+0.017}_{-0.021}$	0.411 \rightarrow 0.621
$\theta_{23}/^\circ$	$49.0^{+1.1}_{-1.4}$	39.6 \rightarrow 51.8	$49.3^{+1.0}_{-1.2}$	39.9 \rightarrow 52.0
$\sin^2 \theta_{13}$	$0.02221^{+0.00068}_{-0.00062}$	0.02034 \rightarrow 0.02430	$0.02240^{+0.00062}_{-0.00062}$	0.02053 \rightarrow 0.02436
$\theta_{13}/^\circ$	$8.57^{+0.13}_{-0.12}$	8.20 \rightarrow 8.97	$8.61^{+0.12}_{-0.12}$	8.24 \rightarrow 8.98
$\delta_{CP}/^\circ$	195^{+51}_{-25}	107 \rightarrow 403	286^{+27}_{-32}	192 \rightarrow 360
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.42^{+0.21}_{-0.20}$	6.82 \rightarrow 8.04	$7.42^{+0.21}_{-0.20}$	6.82 \rightarrow 8.04
$\frac{\Delta m_{3l}^2}{10^{-3} \text{ eV}^2}$	$+2.514^{+0.028}_{-0.027}$	$+2.431 \rightarrow +2.598$	$-2.497^{+0.028}_{-0.028}$	$-2.583 \rightarrow -2.412$

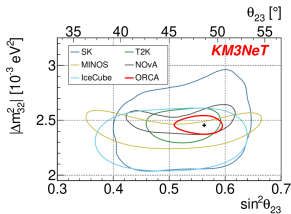
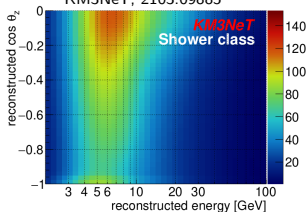
► vacuum, $E_\nu \sim \text{MeV}$

► matter effects, $E_\nu \gtrsim \text{GeV}$

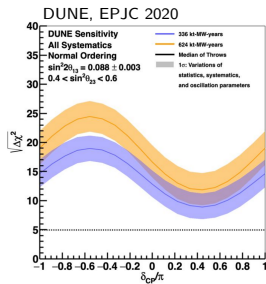
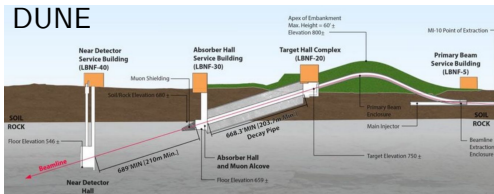
JUNO, 2103.09885



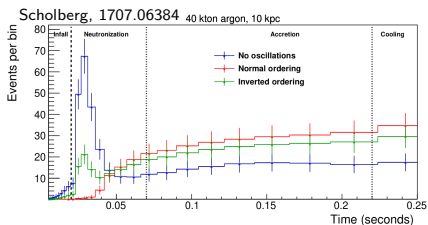
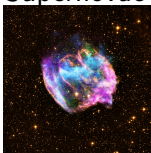
KM3NeT, 2103.09885



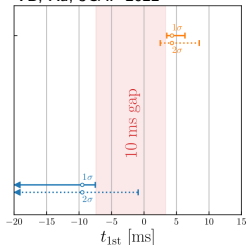
Neutrino Mass Ordering



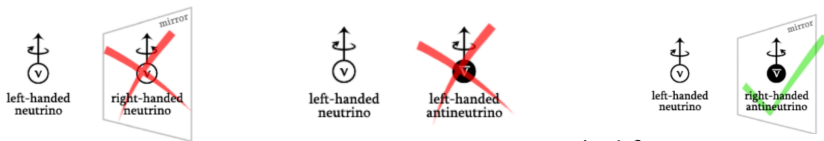
Supernovae



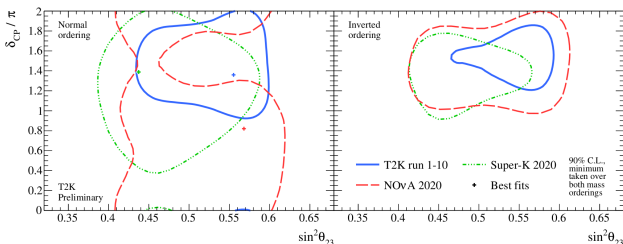
VB, Xu, JCAP 2022



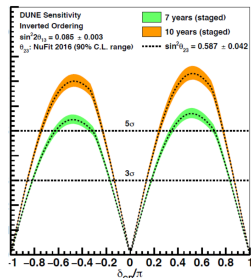
CP Violation in Lepton Sector



$$P(\nu_\alpha \rightarrow \nu_\beta) - P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) = \sum_{j \neq k} 2i \overbrace{\text{Im}(U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^*)}^{J \sim \sin \delta} e^{-i \frac{(m_j^2 - m_k^2)L}{2E}}$$

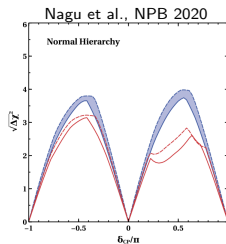
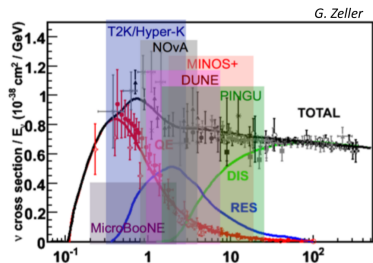


Pan in parallel sessions

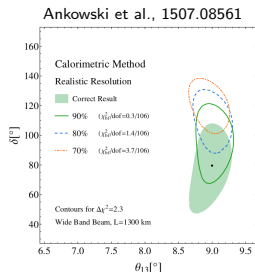
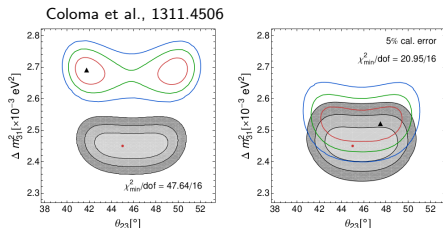


Prerequisites for the Discovery

- ▶ Reduction of ν -nucleus cross section uncertainties



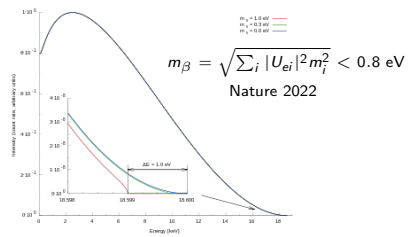
- ▶ Reconstruction of neutrino energy



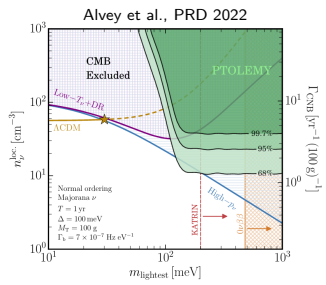
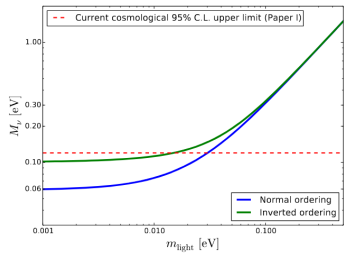
Neutrino Mass Scale?



Köllnberger in parallel sessions

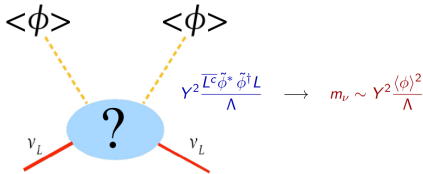


► Testing non-standard cosmology



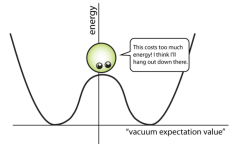
Is the Neutrino its own Antiparticle?

Majorana



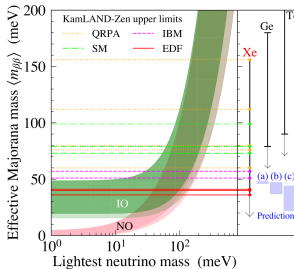
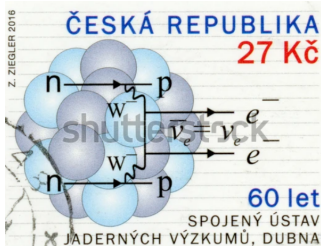
Dirac

	U	C	T
Quarks	u (up)	c (charm)	t (top)
Leptons	e (electron)	μ (muon)	τ (tau)



$$y \bar{\psi}_L \phi \psi_R \Rightarrow m_\nu = y \langle \phi \rangle \Rightarrow y \sim 10^{-12}$$

$0\nu 2\beta$ Decay



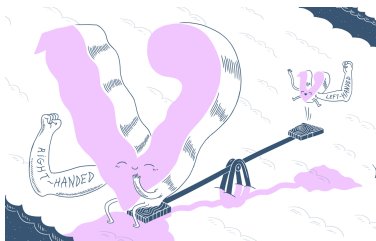
Beretta in parallel sessions

$$\Gamma_{0\nu 2\beta} \propto G_F^4 |\tilde{M}_{0\nu 2\beta}|^2 \left| \sum U_{ej}^2 m_j \right|^2 p_e^2$$

Theory Challenge:
Matrix Elements

Neutrino Mass

Type-I Seesaw

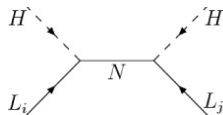


Scotogenic Model

	$SU(2)_L$	$U(1)_Y$	Z_2
Σ	2	1/2	-
N_i	1	0	-
Φ	2	1/2	+
L	2	-1/2	+

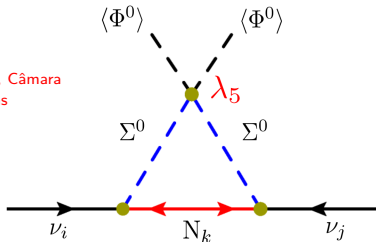
Costa, Escribano, Câmara
in parallel sessions

Minkowski, Mohapatra, Senjanović,
Gell-Mann, Ramond, Slansky, Yanagida

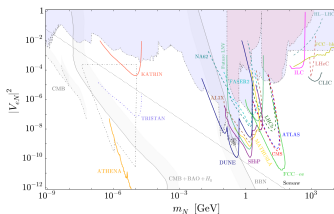
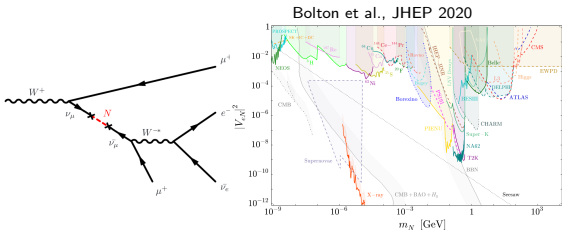


$$m_\nu = -M_D M_R^{-1} M_D^T$$

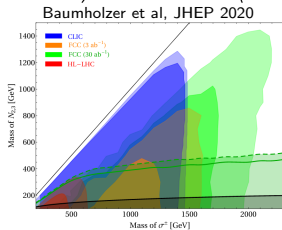
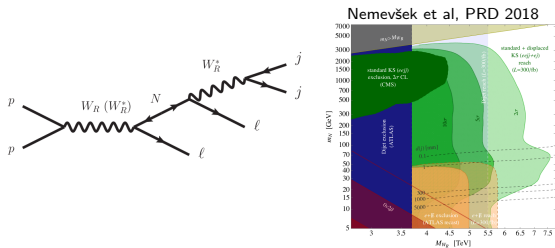
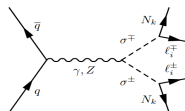
Ma (2006)



Probing the Low Scale Origin of Neutrino Mass

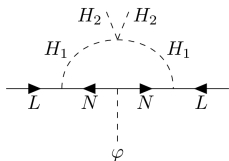


- ▶ for type-I Seesaw, $V_{\alpha N} \sim 10^{-6} \sqrt{100 \text{ GeV}/M_R} \Rightarrow$ not testable
- ▶ **remedy**: inverse seesaw, new interactions of N ...



Energy-Dependent Mixing Parameters

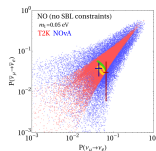
Babu, VB, de Gouvêa, Machado, PRD 2022



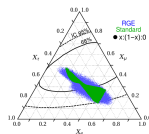
$$\Rightarrow U(Q_p^2) \neq U(Q_d^2) \Rightarrow$$

Model containing
light new particles

Mismatch between PMNS
matrix at Q_p^2 and Q_d^2



T2K
NOvA

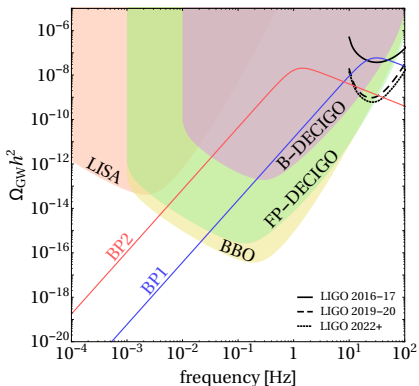


IceCube

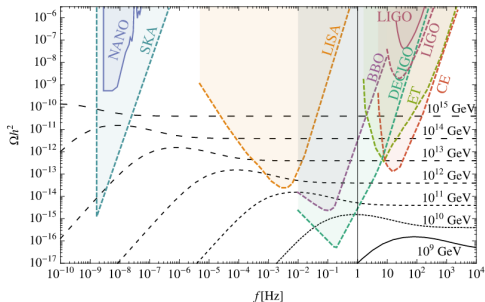
Probing the High Scale Origin of Neutrino Mass

- ▶ **GW detectors** as a window to unexplored seesaw scales
- ▶ testing $M_N \lesssim 10^8$ GeV in models featuring first-order phase transition and $M_N \lesssim$ GUT from topological defects, e.g. cosmic strings

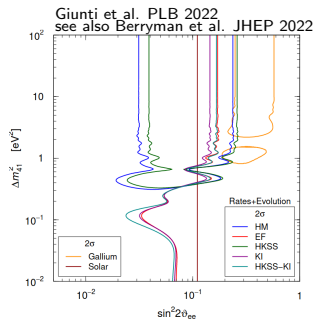
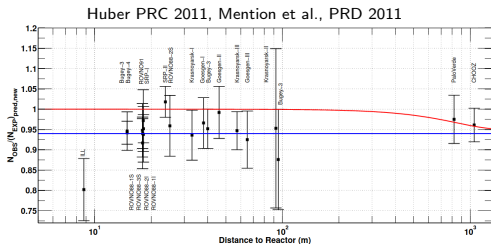
VB, Helmboldt, Kubo, JCAP 2019



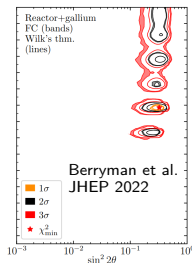
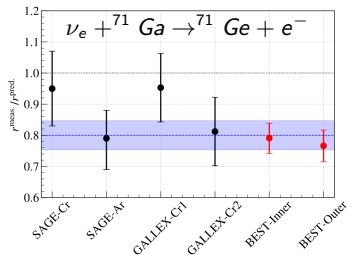
Dror et al., PRL 2020



Anomalies: Reactor and Gallium



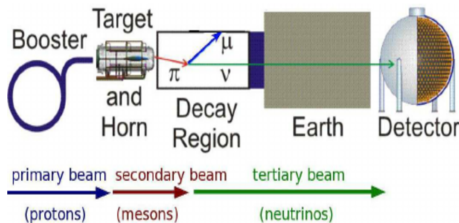
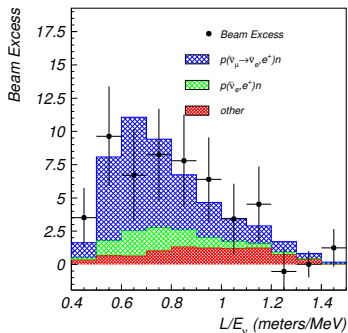
▶ reactor anomaly is fading away



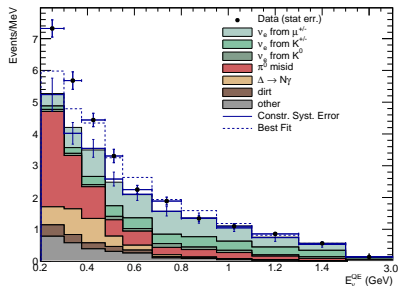
from B. Littlejohn



Anomalies: LSND and MiniBooNE



- ▶ **LSND**: $\bar{\nu}_e$ in $\bar{\nu}_\mu$ beam from stopped pion source ($> 3\sigma$) at $L/E \sim 1\text{km GeV}^{-1}$
- ▶ **MiniBooNE**: reports electron-like event excess (4.8σ)
- ▶ in combination with LSND 6.1σ



Standard Model Explanation?

arXiv > hep-ph > arXiv:2109.08157

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High Energy Physics - Phenomenology

[Submitted on 16 Sep 2021 (v1), last revised 17 Jun 2022 (this version, v2)]

An Altarelli Cocktail for the MiniBooNE Anomaly?

Vedran Brdar (Fermilab and Northwestern U), Joachim Kopp (CERN and JGU Mainz)

We critically examine a number of theoretical uncertainties affecting the MiniBooNE short-baseline neutrino oscillation experiment in an attempt to better understand the observed excess of electron-like events. We re-examine the impact of fake charged current quasi-elastic (CCQE) events, the background due to neutral current π^0 production, and the single-photon background. For all processes, we compare the predictions of different event generators (GENIE, GIBUU, NUANCE, and NuWro) and, for GENIE, of different tunes. Where MiniBooNE uses data-driven background predictions, we discuss the uncertainties affecting the relation between the signal sample and the control sample. In the case of the single-photon background, we emphasize the uncertainties in the radiative branching ratios of heavy hadronic resonances. We find that not even a combination of uncertainties in different channels adding up unfavorably (an "Altarelli cocktail") appears to be sufficient to resolve the MiniBooNE anomaly. We finally investigate how modified background predictions affect the fit of a $3 + 1$ sterile neutrino scenario. We carefully account for full four-flavor oscillations not only in the signal, but also in the background and control samples. We emphasize that because of the strong correlation between MiniBooNE's ν_e and ν_μ samples, a sterile neutrino mixing only with ν_μ is sufficient to explain the anomaly, even though the well-known tension with external constraints on ν_μ disappearance persists.

arXiv > hep-ph > arXiv:2210.08021

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High Energy Physics - Phenomenology

[Submitted on 14 Oct 2022]

More Ingredients for an Altarelli Cocktail at MiniBooNE

Kevin J. Kelly, Joachim Kopp

The MiniBooNE excess of low-energy electron-neutrino-like events persists as one of the most significant puzzles in particle physics today and remains the subject of intense experimental and theoretical focus. A key feature of this excess is the inability of the detector to discriminate between electron-like signals and backgrounds due to photons. Besides single photons, this includes also pairs of overlapping and/or highly-asymmetric photons. In this work, we study photon backgrounds in MiniBooNE in depth, first considering a novel single-photon background arising from multi-nucleon scattering with coherently enhanced initial or final state radiation. This class of processes, which we dub "2p2 γ " (two-particle-two-hole + photon) can explain ~ 40 of the ~ 560 excess events observed by MiniBooNE in neutrino mode (positive horn polarity). Second, we consider the background from neutral-current single- π^0 production, where the two photons from $\pi^0 \rightarrow \gamma\gamma$ decay are mis-identified as an electron-like shower. We find two significant effects when comparing our results to those of MiniBooNE: first, the impact of limited Monte Carlo statistics may be larger than naively expected due to highly non-Gaussian behavior of the fluctuations. This is an effect that needs to be understood with detailed experimental studies, not least because it may affect future experiments as well. Second, in spite of data-driven background estimation techniques, there is a residual dependence on the Monte Carlo generator used. In tandem, we estimate that when combining all effects the significance of the MiniBooNE neutrino-mode low-energy excess may be reduced from 4.7σ to around 3σ .

► reducing the significance to $\simeq 3\sigma$

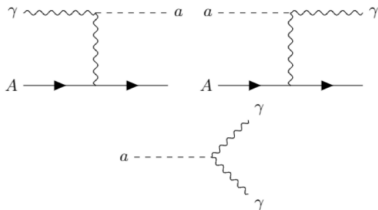
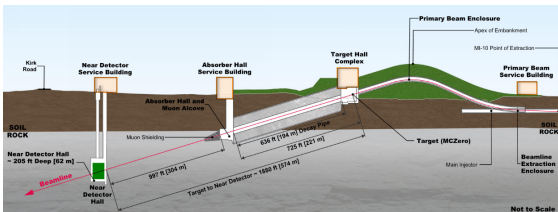
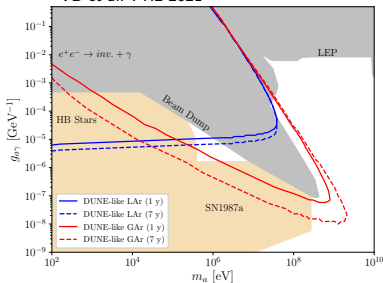
Beyond Neutrino Mass, Beyond Anomalies

from Jae Yu (Snowmass 2021)

Pulido in parallel sessions

Process	Signatures	Background
ALP	Scattering: $\gamma+e/\gamma+N$ (n) Decay in flight: $\gamma\gamma$	ν coherent, NC w/ π^0 , ν_e CC w/ π^0 , etc
LDM	$\chi e \rightarrow \chi e^*$, $\chi N \rightarrow N^* n$	NC w/ π^0 , ν_e CC, QE, RES
mCP	Multiple e^- scatterings	ν_e CC w/ π^0
Dark Photon	$A \rightarrow e^+e^*$, $\mu^+\mu^*$	ν CC + mis-ID π , Accidental overlap of CC
HNL	$N \rightarrow \nu e^+e^*$, $\nu\mu^+\mu^*$, $\nu e\mu$, $\nu\pi^0$, $e\pi$, $\mu\pi$	ν CC + mis-ID π , ν_e CC w/ π^0
ν trident	$\nu \rightarrow \nu e^+e^*$, $\nu\mu^+\mu^*$, $\nu e\mu$	$\nu_\mu N \rightarrow \nu_\mu \pi N$ (ν CC)
BDM/IBDM	$\chi N \rightarrow e N$	ν coherent, NC w/ π^0 , ν_e CC

VB et al. PRL 2021



BACKUP

Direct Detection of Cosmic Neutrino Background

Giachero (2020)

Nuclear Reactors

$E_\nu = 1 - 10$ MeV
Detected ✓



Sun

$E_\nu = 10.4$ MeV
Detected ✓

Accelerators

E_ν up to 12 GeV
Detected ✓



Supernovae (SN 1987A)

$E_\nu = 10$ MeV
Detected ✓

Atmosphere (Cosmic Rays)

E_ν up to 1 GeV
Detected ✓



Astrophysical accelerators

$E_\nu \sim \text{TeV} - \text{PeV}$
Detected ✓

Terrestrial radioactivity

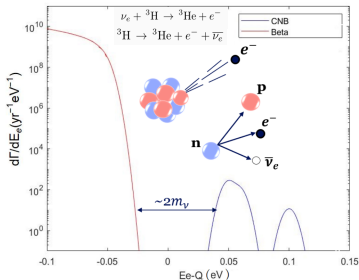
E_ν up to 1 MeV
Detected ✓



Early Universe

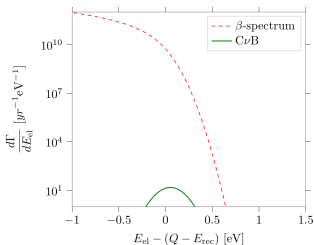
$E_\nu \sim 10^{-4}$ eV
Detected ✗ → Indirect evidence

Tan, Cheianov, 2202.07406



▶ Graphene substrates to hold ^3H at PTOLEMY

▶ Intrinsic uncertainty in the e^- energy (Cheipesh et al., 2101.10069)



▶ Heavier nuclei?

VB, Plestid, Rocco, PRC 2022

