



Massive quark form factors at three loops

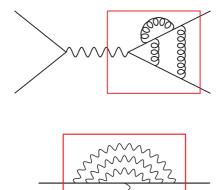
Young Scientists Meeting of the CRC TRR 257 | June 8 - 10, 2022

Fabian Lange

in collaboration with Matteo Fael, Kay Schönwald, Matthias Steinhauser | June 9, 2022

Motivation

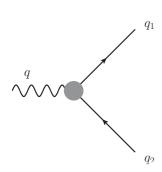




- Form factors are basic building blocks for many physical observables:
 - $t\bar{t}$ production at hadron and e^+e^- colliders
 - \blacksquare μe scattering
 - Higgs production and decay
 - ...
- Form factors exhibit an universal infrared behavior which is interesting to study

The process





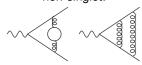
$$egin{align} X(q)
ightarrow Q(q_1) + ar{Q}(q_2) \ & \ q_1^2 = q_2^2 = m^2, \quad q^2 = s = \hat{s} \cdot m^2 \ & \ \end{array}$$

pseudo-scalar : $j^p = im\overline{\psi}\gamma_5\psi$, $\Gamma^p = imF^p(s)\gamma_5$

Status of massive non-singlet QCD corrections



non-singlet:





singlet:



(NNLO):

- fermionic contributions [Hoang, Teubner 1997]
- complete [Bernreuther, Bonciani, Gehrmann, Heinesch, Leineweber, Mastrolia, Remiddi 2004 2005]

(NNNLO):

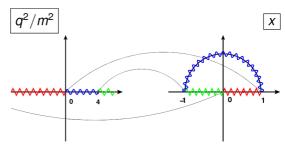
- large N_C [Henn, Smirnov, Smirnov, Steinhauser 2016; Lee, Smirnov, Smirnov, Steinhauser 2018; Ablinger, Blümlein, Marquard, Rana, Schneider 2 × 2018; Lee, Smirnov, Smirnov, Steinhauser 2018]
- n [Lee. Smirnov, Smirnov, Steinhauser 2018; Ablinger, Blümlein, Marguard, Rana, Schneider 2 × 2018]
- n_h (partially) [Blümlein, Marquard, Rana, Schneider 2019]

this talk: full (numerical) results for non-singlet contributions at NNNLO

Why numerical?



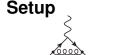
$$q^2 = s = -\frac{(1-x)^2}{x}$$



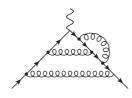
• Large- N_c and n_l contributions at NNNLO can be written as iterated integrals over letters

$$\frac{1}{x}$$
, $\frac{1}{1+x}$, $\frac{1}{1-x}$, $\frac{1}{1-x+x^2}$, $\frac{x}{1-x+x^2}$

- \blacksquare n_h terms already contain structures beyond iterated integrals (elliptic integrals)
- No ready-to-use tools available for analytic solution
- Instead: Full solution through analytic series expansions and numerical matching









- Generate diagrams with qgraf [Nogueira 1991]
- Map to predefined integral families with q2e/exp [Harlander, Seidensticker, Steinhauser 1998; Seidensticker 1999]
- FORM [Vermaseren 2000; Kuipers, Ueda, Vermaseren, Vollinga 2013; Ruijl, Ueda, Vermaseren 2017] for Lorentz, Dirac, and color algebra [van Ritbergen, Schellekens, Vermaseren 1998]
- Reduction to master integrals with Kira [Maierhöfer, Usovitsch, Uwer 2017; Klappert, FL, Maierhöfer, Usovitsch 2020] and Fermat [Lewis]
 - Construct good basis where denominators factorize in ϵ and \hat{s} with ImproveMasters.m [Smirnov, Smirnov 2020]
- Establish differential equations in ŝ with LiteRed [Lee 2012 + 2013]

non-singlet
271
34
302671
422

Algorithm to solve master integrals



$$\frac{\partial}{\partial \hat{\mathbf{s}}} M_n = A_{nm}(\epsilon, \hat{\mathbf{s}}) M_m$$

- Compute expansion around $\hat{s} = 0$ by:
 - Inserting an ansatz for the master integrals into the differential equation:

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\mathsf{max}}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- lacktriangle Compare coefficients in ϵ and \hat{s} to establish linear system of equations for $c_{ii}^{(n)}$
 - Solve system in terms of small number of boundary constants using Kira with FireFly [Klappert, FL 2019; Klappert, Klein, FL 2020]
- Compute boundary values for $\hat{s} = 0$ to fix remaining constants

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- Compute boundary values for $\hat{s} = 0$ to fix remaining constants
- Construct expansion around new point $\hat{s} = \hat{s}_0$ by modifying the ansatz and repeating the steps above
- Match both expansions numerically at a point where both expansions converge, e.g. $\hat{s}_0/2$
- Repeat





regular point (including static limit at
$$s=0$$
): $M_n(\epsilon, \hat{s}=\hat{s}_0)=\sum_{i=-3}^{\infty}\sum_{j=0}^{j_{\text{max}}}c_{ij}^{(n)}\,\epsilon^i\,(\hat{s}-\hat{s}_0)^j$



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$$s=\pm\infty \text{ (high-energy limit):} \qquad M_n(\epsilon, \hat{s}\to\pm\infty) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\text{min}}}^{j_{\text{max}}} \sum_{k=0}^{i+6} c_{ijk}^{(n)} \epsilon^i \, \hat{s}^{-j} \, \ln^k (\hat{s})$$



regular point (including static limit at
$$s=0$$
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$$s = 4m^2 \text{ (2-particle threshold):} \qquad M_n(\epsilon, \hat{\mathbf{s}} = 4) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\text{min}}}^{j_{\text{max}}} \sum_{k=0}^{i+3} c_{ijk}^{(n)} \, \epsilon^i \, \left[\sqrt{4-\hat{\mathbf{s}}} \right]^j \, \ln^k \left(\sqrt{4-\hat{\mathbf{s}}} \right)$$

$$s=4m^2$$
 (2-particle threshold):
$$M_n(\epsilon,\hat{\mathbf{s}}=4)=\sum_{i=-3}^{\infty}\sum_{j=-s_{\min}}^{\sum_{m=1}^{max}}\sum_{k=0}^{i-3}c_{ijk}^{(n)}\,\epsilon^i\,\left[\sqrt{4-\hat{\mathbf{s}}}\right]^j\,\ln^k\left(\sqrt{4-\hat{\mathbf{s}}}\right)^j$$



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$$S = 16m^2 \text{ (4-particle threshold):}$$

$$M_n(\epsilon, \hat{s} = 4) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\min}}^{j_{\max}} \sum_{k=0}^{i+3} c_{ijk}^{(n)} \epsilon^i \left[\sqrt{16 - \hat{s}} \right]^j \ln^k \left(\sqrt{16 - \hat{s}} \right)$$



Different ansätze for different points:

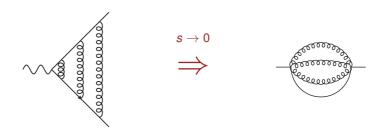
regular point (including static limit at
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): $M_n(\epsilon,\hat{s}=\hat{s}_0)=\sum_{i=-3}^{\infty}\sum_{j=0}^{j_{\text{max}}}c_{ij}^{(n)}\,\epsilon^i\,(\hat{s}-\hat{s}_0)^j$ $s=\pm\infty$ (high-energy limit): $M_n(\epsilon,\hat{s}\to\pm\infty)=\sum_{i=-3}^{\infty}\sum_{j=-s_{\text{min}}}^{j_{\text{max}}}\sum_{k=0}^{i+6}c_{ijk}^{(n)}\,\epsilon^i\,\hat{s}^{-j}\,\ln^k\,(\hat{s})$ $s=4m^2$ (2-particle threshold): $M_n(\epsilon,\hat{s}=4)=\sum_{i=-3}^{\infty}\sum_{j=-s_{\text{min}}}^{j_{\text{max}}}\sum_{k=0}^{i+3}c_{ijk}^{(n)}\,\epsilon^i\,\left[\sqrt{4-\hat{s}}\right]^j\,\ln^k\left(\sqrt{4-\hat{s}}\right)$ $s=16m^2$ (4-particle threshold): $M_n(\epsilon,\hat{s}=16)=\sum_{i=-3}^{\infty}\sum_{j=-s_{\text{min}}}^{j_{\text{max}}}\sum_{k=0}^{i+3}c_{ijk}^{(n)}\,\epsilon^i\,\left[\sqrt{16-\hat{s}}\right]^j\,\ln^k\left(\sqrt{16-\hat{s}}\right)$

• We construct expansions up to $j_{max} = 50$ around

$$\hat{s} = \{-\infty, -32, -28, -24, -16, -12, -8, -4, 0, 1, 2, \frac{5}{2}, 3, \frac{7}{2}, 4, \frac{9}{2}, 5, 6, 7, 8, 10, 12, 14, 15, 16, 17, 19, 22, 28, 40\}$$

Calculation of boundary conditions

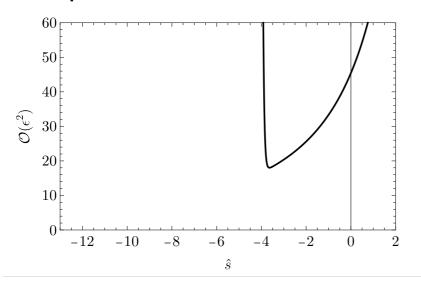


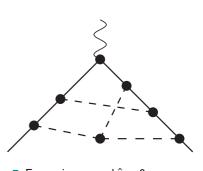


- For s = 0 the master integrals reduce to 3-loop on-shell propagators:
 - Well studied in the literature [Laporta, Remiddi 1996; Melnikov, van Ritbergen 1999; Lee, Smirnov 2010]
 - \blacksquare The reduction introduces high inverse powers in ϵ which requires some integrals up to weight 9
 - Using the dimensional-recurrence relations from [Lee, Smirnov 2010] we calculated the missing terms with SummerTime.m [Lee, Mingulov 2015] and PSLQ [Ferguson, Bailey, Arno 1999]

Example



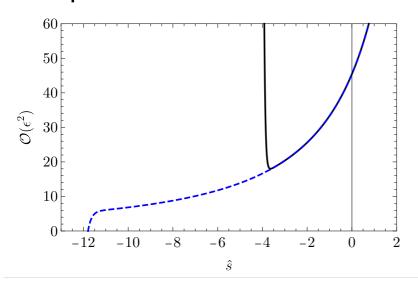


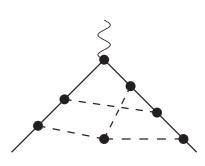


• Expansion around $\hat{s} = 0$

Example



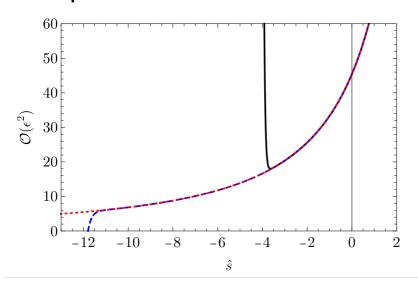


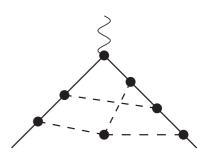


- Expansion around $\hat{s} = 0$
- Expansion around $\hat{s} = -4$, matched at $\hat{s} = -2$

Example







- Expansion around $\hat{s} = 0$
- Expansion around $\hat{s} = -4$, matched at $\hat{s} = -2$
- Expansion around $\hat{s} = -8$, matched at $\hat{s} = -6$





Other approaches based on differential equations and series expansions:

- SolveCoupledSystems.m [Blümlein, Schneider 2017]
- DESS.m [Lee, Smirnov, Smirnov 2017]
- DiffExp.m [Hidding 2020]
- SeaSvde.m [Armadillo, Bonciani, Devoto, Rana, Vicini 2022]

Our approach ...

- ... is tailored to problems with one real-valued kinematic variable
- ... does not require a special form for differential equations (except to be almost pole free on the diagonal)
- provides approximations over the whole kinematic range
- was successfully applied to physical quantities with 339 and 422 master integrals [Fael, FL, Schönwald,

Steinhauser 2021 + 20221

Renormalization and infrared structure



UV renormalization

- MS renormalization of α_s
- On-shell renormalization of mass Z_2^{OS} , wave function Z_2^{OS} , and (if needed) currents [Chetyrkin, Steinhauser 1999; Melnikov, van Ritbergen 20001

IR subtraction

- Structure of infrared poles given by cusp anomalous dimension Γ_{cusp} [Grozin, Henn, Korchemski, Marquard 2014]
- Define finite form factors $F = Z_{IR}F^{finite}$ with UV-renormalized form factor F and

$$Z_{\text{IR}} = 1 - \frac{\alpha_s}{\pi} \frac{1}{2\epsilon} \Gamma_{\text{cusp}}^{(1)} - \left(\frac{\alpha_s}{\pi}\right)^2 \left(\frac{\cdots}{\epsilon^2} + \frac{1}{4\epsilon} \Gamma_{\text{cusp}}^{(2)}\right) - \left(\frac{\alpha_s}{\pi}\right)^3 \left(\frac{\cdots}{\epsilon^3} + \frac{\cdots}{\epsilon^2} + \frac{1}{6\epsilon} \Gamma_{\text{cusp}}^{(3)}\right)$$

- $\Gamma_{\text{cusp}} = \Gamma_{\text{cusp}}(x)$ depends on kinematics
- Γ_{cusp} universal for all currents





$$\begin{split} F_1^{\text{v.f.}(3)}(\hat{\mathbf{s}} = 0) &= \Bigg\{ C_{\text{F}}^3 \bigg(-15a_4 - \frac{17\pi^2\zeta_3}{24} - \frac{18367\zeta_3}{1728} + \frac{25\zeta_5}{8} - \frac{5l_2^4}{8} - \frac{19}{40}\pi^2 l_2^2 + \frac{4957\pi^2 l_2}{720} + \frac{3037\pi^4}{25920} \\ &- \frac{24463\pi^2}{7776} + \frac{13135}{20736} \bigg) + C_{\text{A}}C_{\text{F}}^2 \bigg(\frac{19a_4}{2} - \frac{\pi^2\zeta_3}{9} + \frac{17725\zeta_3}{3456} - \frac{55\zeta_5}{32} + \frac{19l_2^4}{48} - \frac{97}{720}\pi^2 l_2^2 \\ &+ \frac{29\pi^2l_2}{240} - \frac{347\pi^4}{17280} - \frac{4829\pi^2}{10368} + \frac{707}{288} \bigg) + C_{\text{A}}^2C_{\text{F}} \bigg(-a_4 + \frac{7\pi^2\zeta_3}{96} + \frac{4045\zeta_3}{5184} - \frac{5\zeta_5}{64} - \frac{l_2^4}{24} \\ &+ \frac{67}{360}\pi^2 l_2^2 - \frac{5131\pi^2l_2}{2880} + \frac{67\pi^4}{8640} + \frac{172285\pi^2}{186624} - \frac{7876}{2187} \bigg) \Bigg\} \hat{\mathbf{s}} + \text{fermionic corrections} + \mathcal{O}(\hat{\mathbf{s}}^2) \end{split}$$

- $I_2 = In(2)$, $a_4 = Li_4(1/2)$ and $C_A = 3$, $C_F = 4/3$ for QCD
- Expansions for all currents are available up to $\mathcal{O}(\hat{s}^{67})$





$$\begin{split} F_{1}^{VI,(3)}\Big|_{s\to-\infty} &= 4.7318C_{\rm F}^{3} - 20.762C_{\rm F}^{2}C_{\rm A} + 8.3501C_{\rm F}C_{\rm A}^{2} + \left[3.4586C_{\rm F}^{3} - 4.0082C_{\rm F}^{2}C_{\rm A} - 6.3561C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S} \\ &+ \left[1.4025C_{\rm F}^{3} + 0.51078C_{\rm F}^{2}C_{\rm A} - 2.2488C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S}^{2} + \left[0.062184C_{\rm F}^{3} + 0.90267C_{\rm F}^{2}C_{\rm A} - 0.42778C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S}^{3} \\ &+ \left[-0.075860C_{\rm F}^{3} + 0.20814C_{\rm F}^{2}C_{\rm A} - 0.035011C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S}^{4} + \left[-0.023438C_{\rm F}^{3} + 0.019097C_{\rm F}^{2}C_{\rm A}\right]I_{\rm S}^{5} \\ &+ \left[-0.0026042C_{\rm F}^{3}\right]I_{\rm S}^{6} - \left\{-92.918C_{\rm F}^{3} + 123.65C_{\rm F}^{2}C_{\rm A} - 47.821C_{\rm F}C_{\rm A}^{2} + \left[-10.381C_{\rm F}^{3} + 2.3223C_{\rm F}^{2}C_{\rm A}\right]I_{\rm S}^{2} + 17.305C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S} + \left[4.9856C_{\rm F}^{3} - 19.097C_{\rm F}^{2}C_{\rm A} + 8.0183C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S}^{2} + \left[3.0499C_{\rm F}^{3} - 6.8519C_{\rm F}^{2}C_{\rm A} + 1.9149C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S}^{3} \\ &+ \left[0.67172C_{\rm F}^{3} - 0.91213C_{\rm F}^{2}C_{\rm A} + 0.24069C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S}^{4} + \left[0.13229C_{\rm F}^{3} - 0.051389C_{\rm F}^{2}C_{\rm A} + 0.0043403C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S}^{5} \\ &+ \left[0.0041667C_{\rm F}^{3} - 0.0010417C_{\rm F}^{2}C_{\rm A} - 0.00052083C_{\rm F}C_{\rm A}^{2}\right]I_{\rm S}^{6} + \mathcal{O}\left(\frac{m^{4}}{\rm S^{2}}\right) + \text{fermionic contributions} \end{aligned}$$

Dedicated calculation of leading logarithms [Liu, Penin, Zerf 2017]:

$$F_1^{\text{v,f},(3)} = -\frac{C_{\text{F}}^3}{384} I_s^6 - \frac{m^2}{s} \left(\frac{C_{\text{F}}^3}{240} - \frac{C_{\text{F}}^2 C_{\text{A}}}{960} - \frac{C_{\text{F}} C_{\text{A}}^2}{1920} \right) I_s^6 + \dots, \quad \text{with } I_s = \ln \left(\frac{m^2}{-s} \right)$$

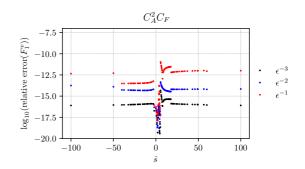
We reproduce these terms with high precision

Results - pole cancellation



- We use the pole cancellation to estimate the precision
- To estimate the number of significant digits we use

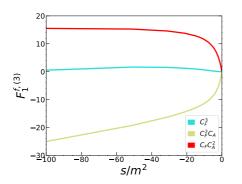
$$\log_{10} \left(\left| \frac{\text{expansion} - \text{analytic CT}}{\text{analytic CT}} \right| \right)$$

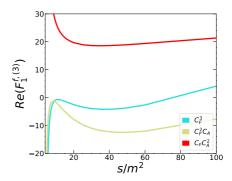


- ⇒ We estimate at least 8 correct digits for the finite terms
- Most regions much more precise

Results – some plots











Close to threshold we can construct cross-sections and decay rates like

$$\sigma(e^{+}e^{-} \to Q\bar{Q}) = \sigma_{0}\beta \underbrace{\left(\left|F_{1}^{v} + F_{2}^{v}\right|^{2} + \frac{\left|(1 - \beta^{2})F_{1}^{v} + F_{2}^{v}\right|^{2}}{2(1 - \beta^{2})}\right)}_{=3/2\Delta^{v}}$$

with the quark velocity $\beta = \sqrt{1 - 4m^2/s}$

- Real radiation suppressed by β^3
- ⇒ Direct phenomenological relevance
- We find (with $I_{2\beta} = \ln(2\beta)$)

$$\begin{split} \Delta^{\text{v},(3)} &= \textit{C}_{\text{F}}^{3} \Big[-\frac{32.470}{\beta^{2}} + \frac{1}{\beta} \big(14.998 - 32.470 \textit{I}_{2\beta} \big) \Big] + \textit{C}_{\text{A}}^{2} \textit{C}_{\text{F}} \frac{1}{\beta} \big[16.586 \textit{I}_{2\beta}^{2} - 22.572 \textit{I}_{2\beta} + 42.936 \big] \\ &+ \textit{C}_{\text{A}} \textit{C}_{\text{F}}^{2} \Big[\frac{1}{\beta^{2}} \big(-29.764 \textit{I}_{2\beta} - 7.7703 \big) + \frac{1}{\beta} \big(-12.516 \textit{I}_{2\beta} - 11.435 \big) \Big] \\ &+ \mathcal{O}(\beta^{0}) + \text{fermionic contributions} \end{split}$$

Agrees with dedicated calculation [Kiyo, Maier, Maierhöfer, Marguard 2009]

Conclusions and outlook



Conclusions

- Calculated non-singlet contributions to massive guark form factors at NNNLO in QCD
 - Vector current partially published in [Fael, FL, Schönwald, Steinhauser 2022]
 - Other currents follow soon
- Applied a semianalytic method by constructing series expansions and matching numerically
- Reproduce known results from the literature, e.g.
 - large- N_c limit, n_l and partial n_h contributions
 - static, high-energy, and threshold expansions
- Estimate precision to at least 8 significant digits over the whole real axis
- Extracted matching coefficients between QCD and NRQCD ⇒ talk by Manuel Egner

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Outlook

- Contributions of singlet diagrams
- Singlet contributions to NRQCD matching coefficients
- The method can be applied to other one-scale problems







E.g. extension of G_{66} (given up to and including $\mathcal{O}(\epsilon^3)$ in [Lee, Smirnov 2010]):

$$=\cdots+\epsilon^4\left(-4704s_6-9120s_{7a}-9120s_{7b}-547s_{8a}+9120s_6\ln(2)+28\ln^4(2)+\frac{112\ln^5(2)}{3}-\frac{808}{45}\ln^6(2)\right)$$

$$-\frac{347}{9}\ln^8(2)+672\text{Li}_4\left(\frac{1}{2}\right)-\frac{5552}{3}\ln^4(2)\text{Li}_4\left(\frac{1}{2}\right)-22208\text{Li}_4\left(\frac{1}{2}\right)^2-4480\text{Li}_5\left(\frac{1}{2}\right)-12928\text{Li}_6\left(\frac{1}{2}\right)+\ldots\right)$$

$$+\epsilon^5\left(14400s_6-\frac{377568s_{7a}}{7}-\frac{93984s_{7b}}{7}-2735s_{8a}+7572912s_{9a}-3804464s_{9b}-\frac{5092568s_{9c}}{3}-136256s_{9d}\right)$$

$$+681280s_{9e}+272512s_{9f}+\frac{377568}{7}s_6\ln(2)-\frac{32465121}{20}s_{8a}\ln(2)-10185136s_{8b}\ln(2)+136256s_{7b}\ln^2(2)+\ldots\right)$$

$$+\mathcal{O}(\epsilon^6)$$

Moebius Transformations



- The radius of convergence is at most the distance to the closest singularity.
- We can extend the radius of convergence by changing to a new expansion variable.
- If we want to expand around the point x_k with the closest singularities at x_{k-1} and x_{k+1} , we can use:

$$y_k = \frac{(x - x_k)(x_{k+1} - x_{k-1})}{(x - x_{k+1})(x_{k-1} - x_k) + (x - x_{k-1})(x_{k+1} - x_k)}$$

■ The variable change maps $\{x_{k-1}, x_k, x_{k+1}\} \to \{-1, 0, 1\}$.