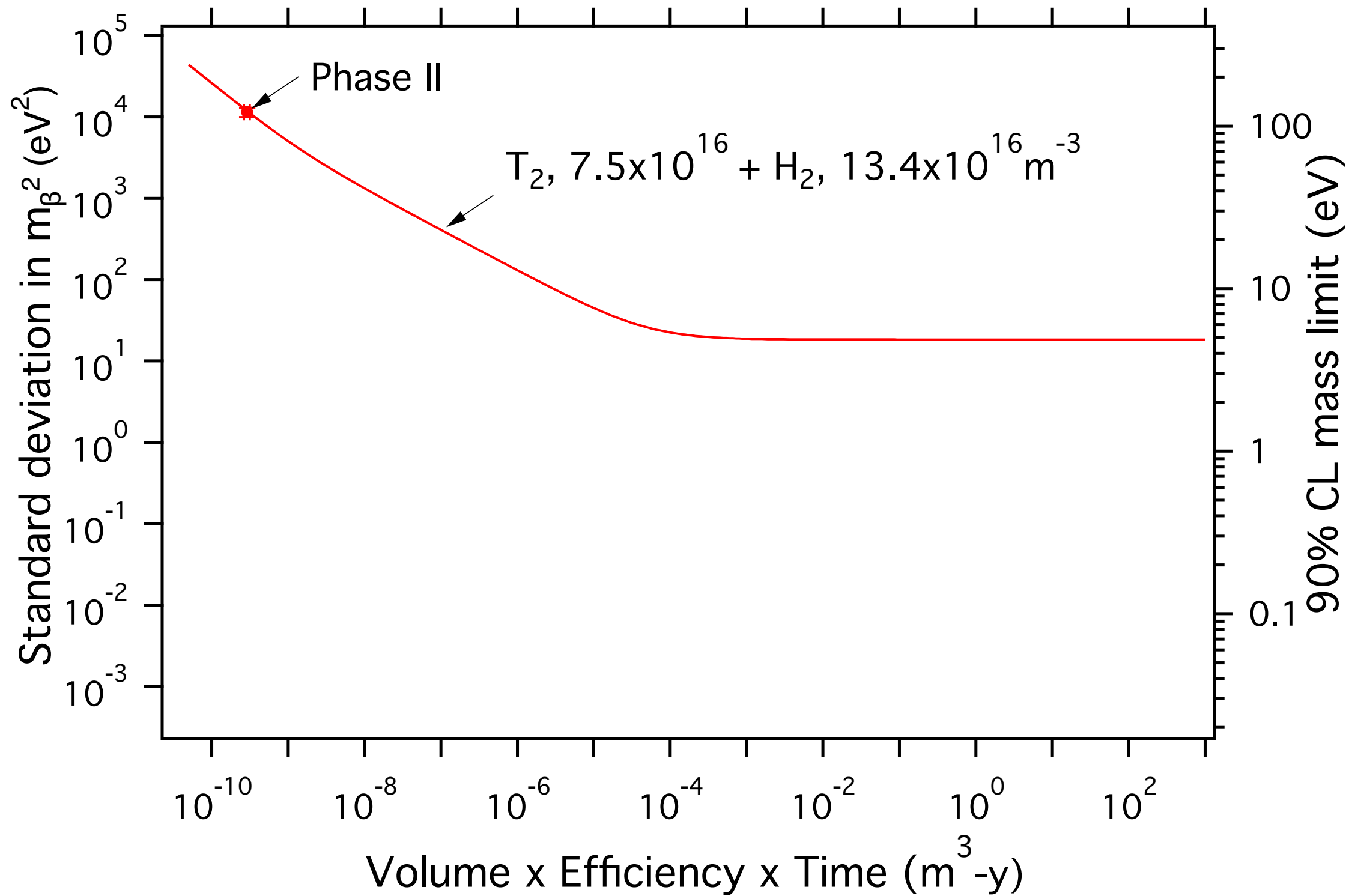


# Towards an atomic tritium source for Project 8

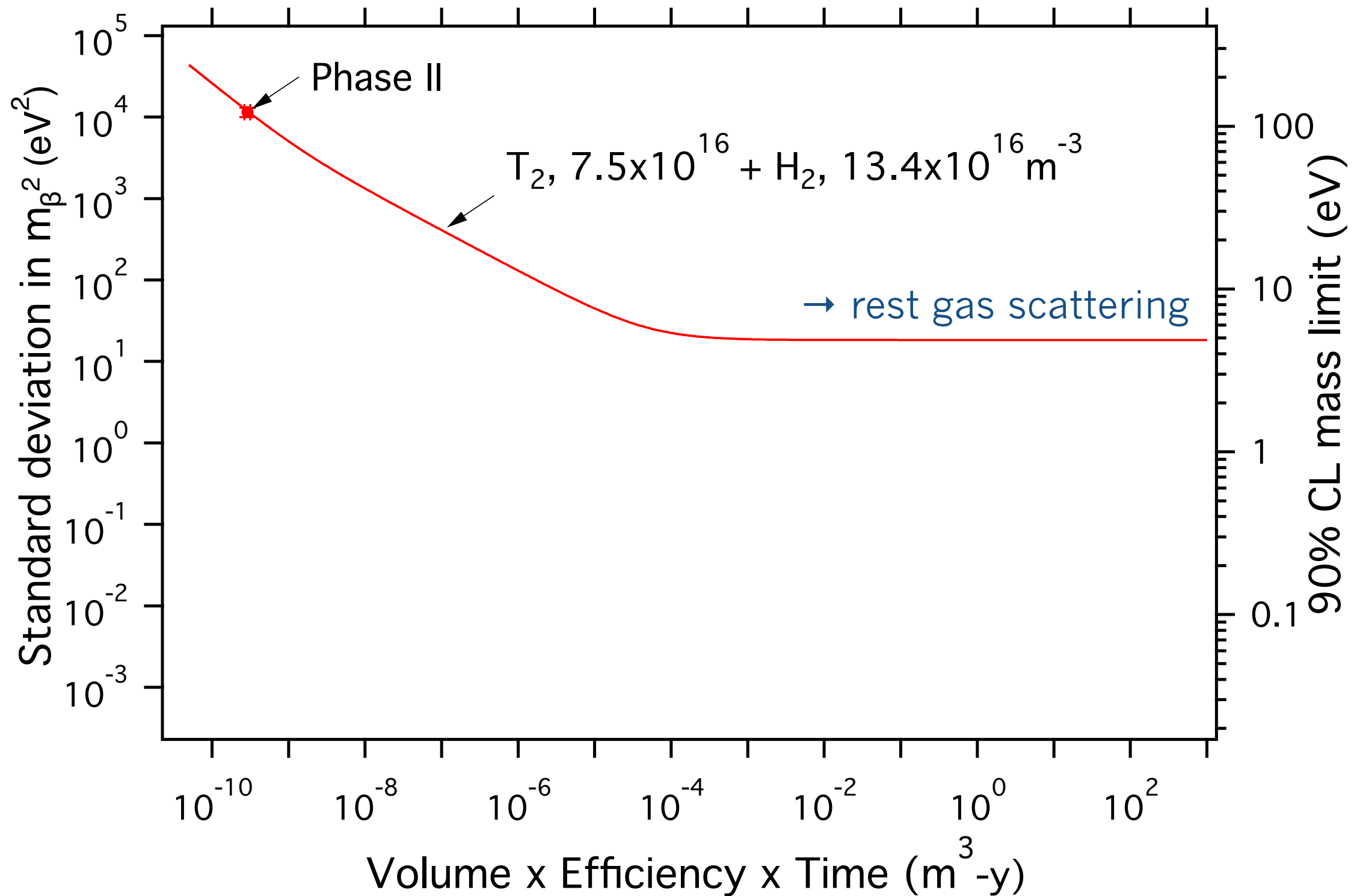
Sebastian Böser

30 years TLK Symposium | Karlsruhe | May 24<sup>nd</sup> 2023

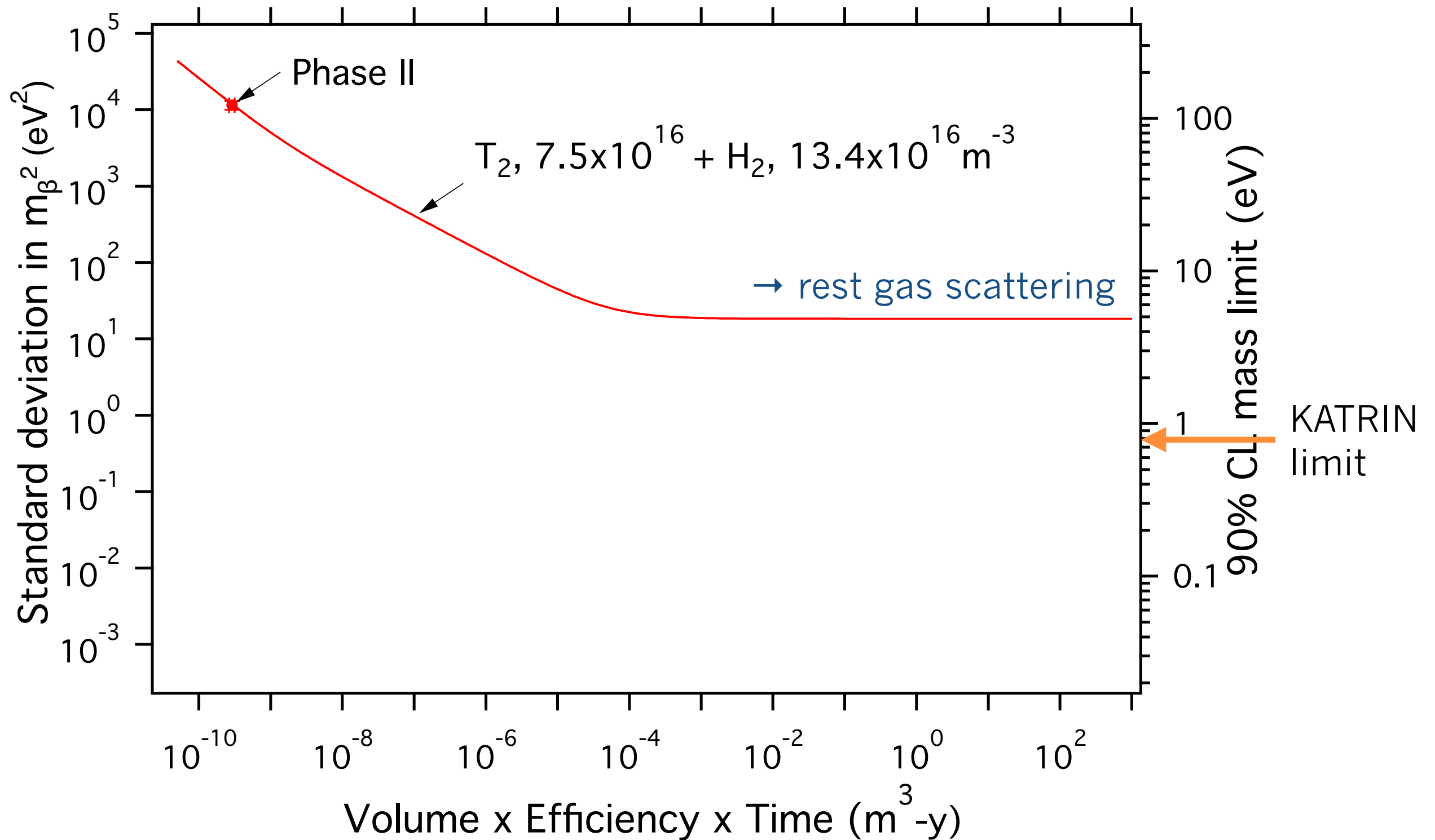
# Potential for neutrino mass?



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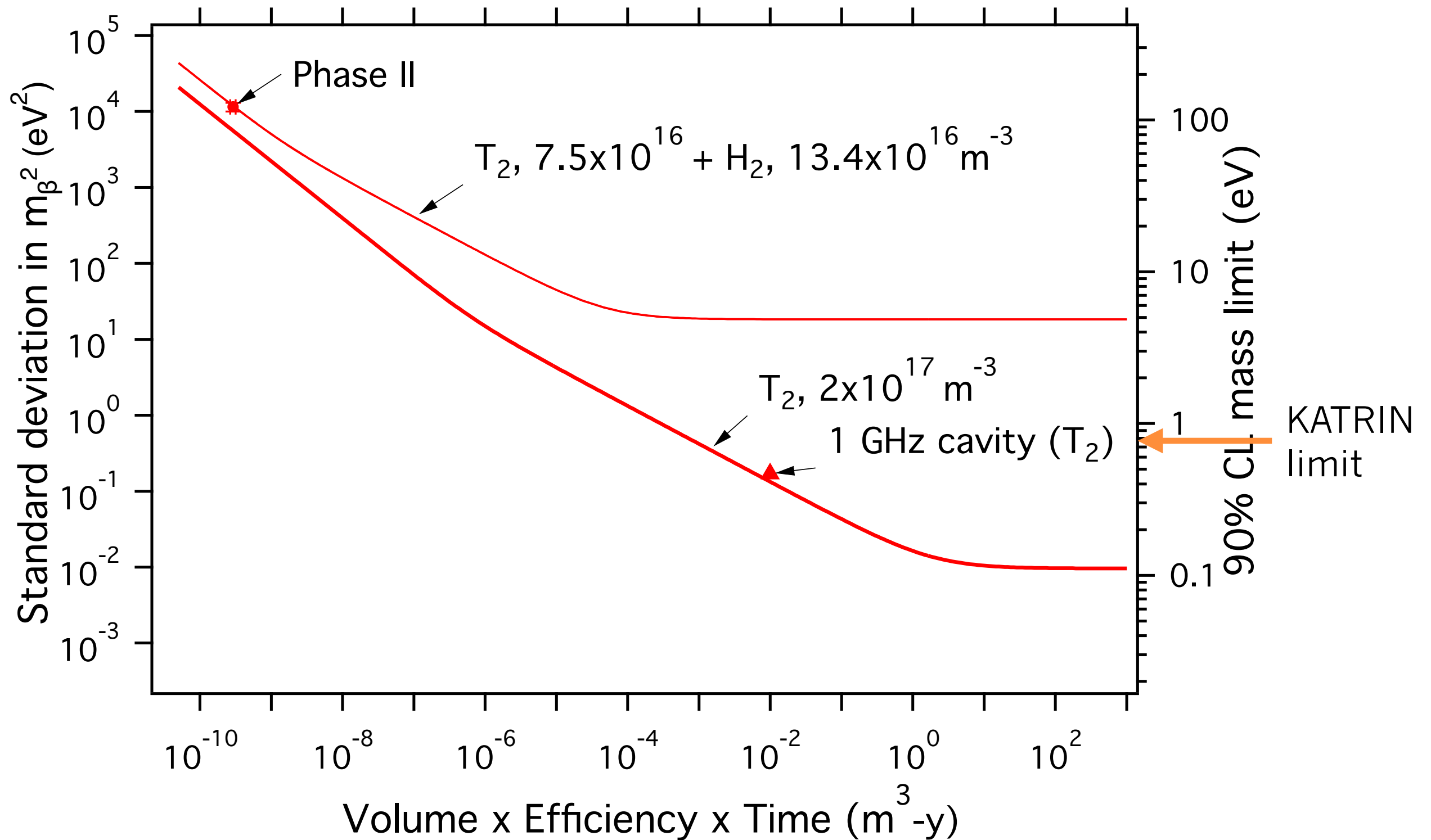


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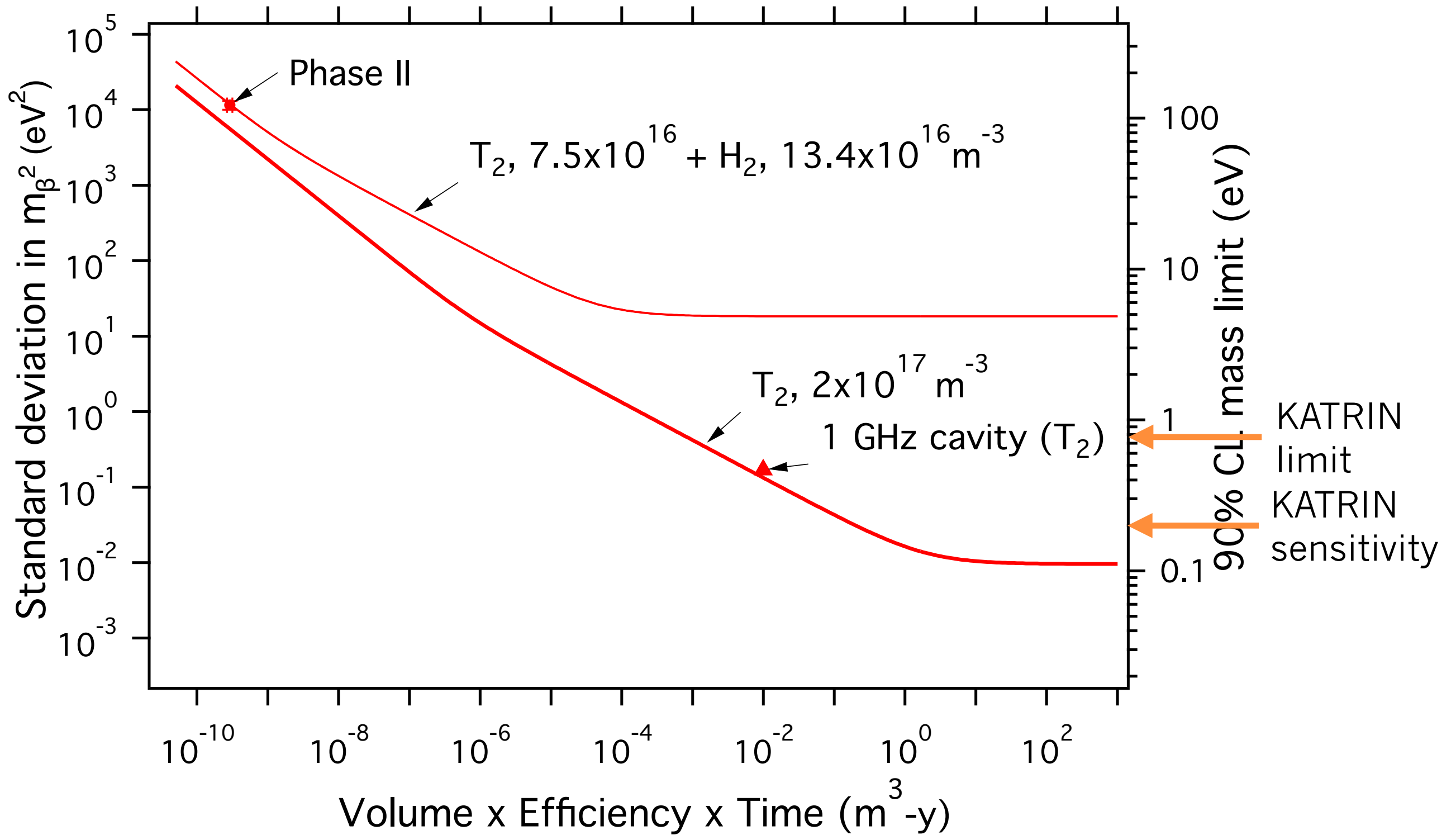




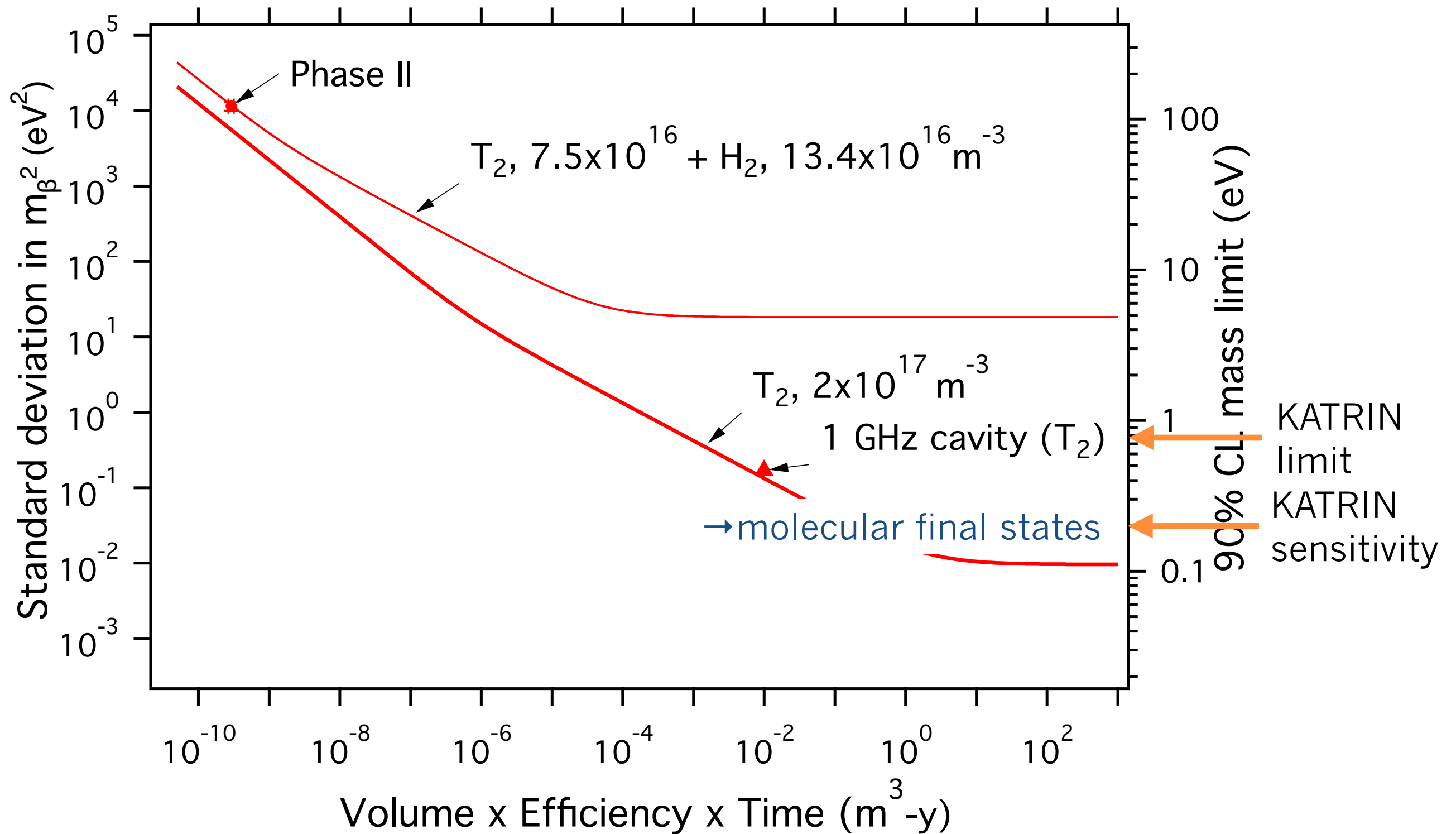
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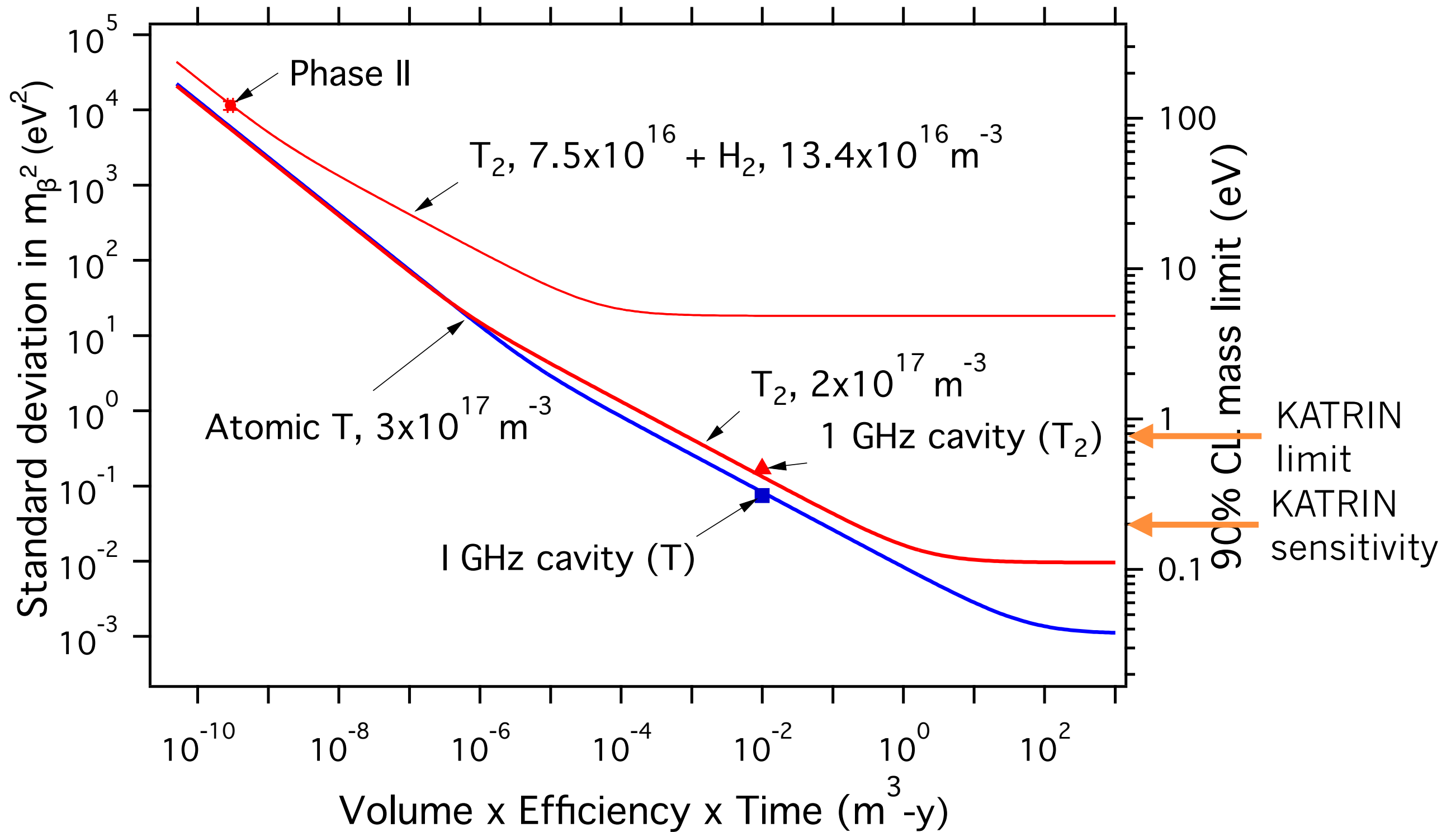
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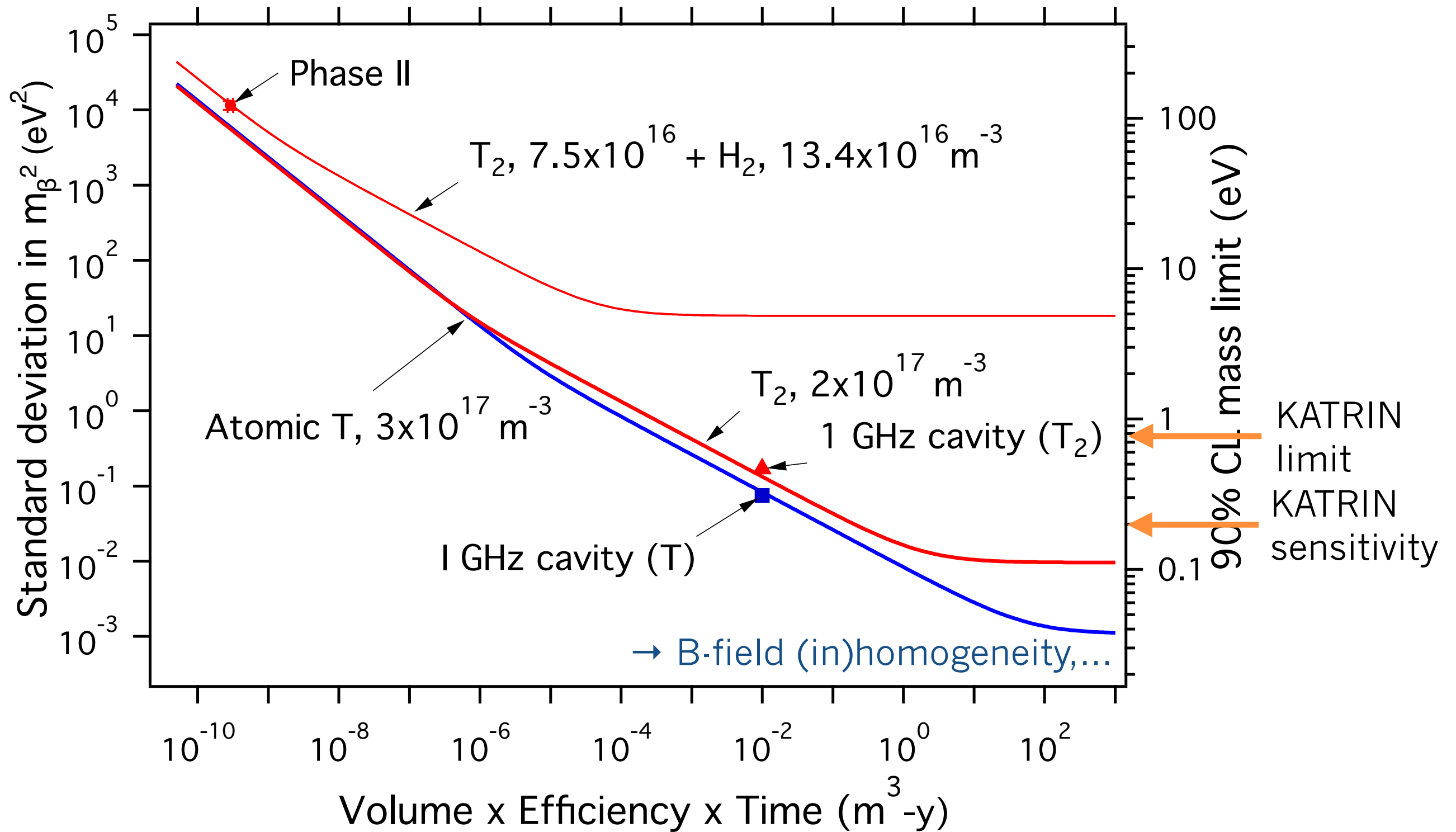
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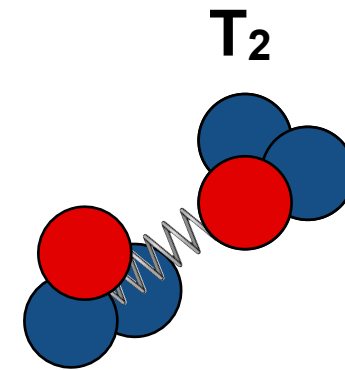
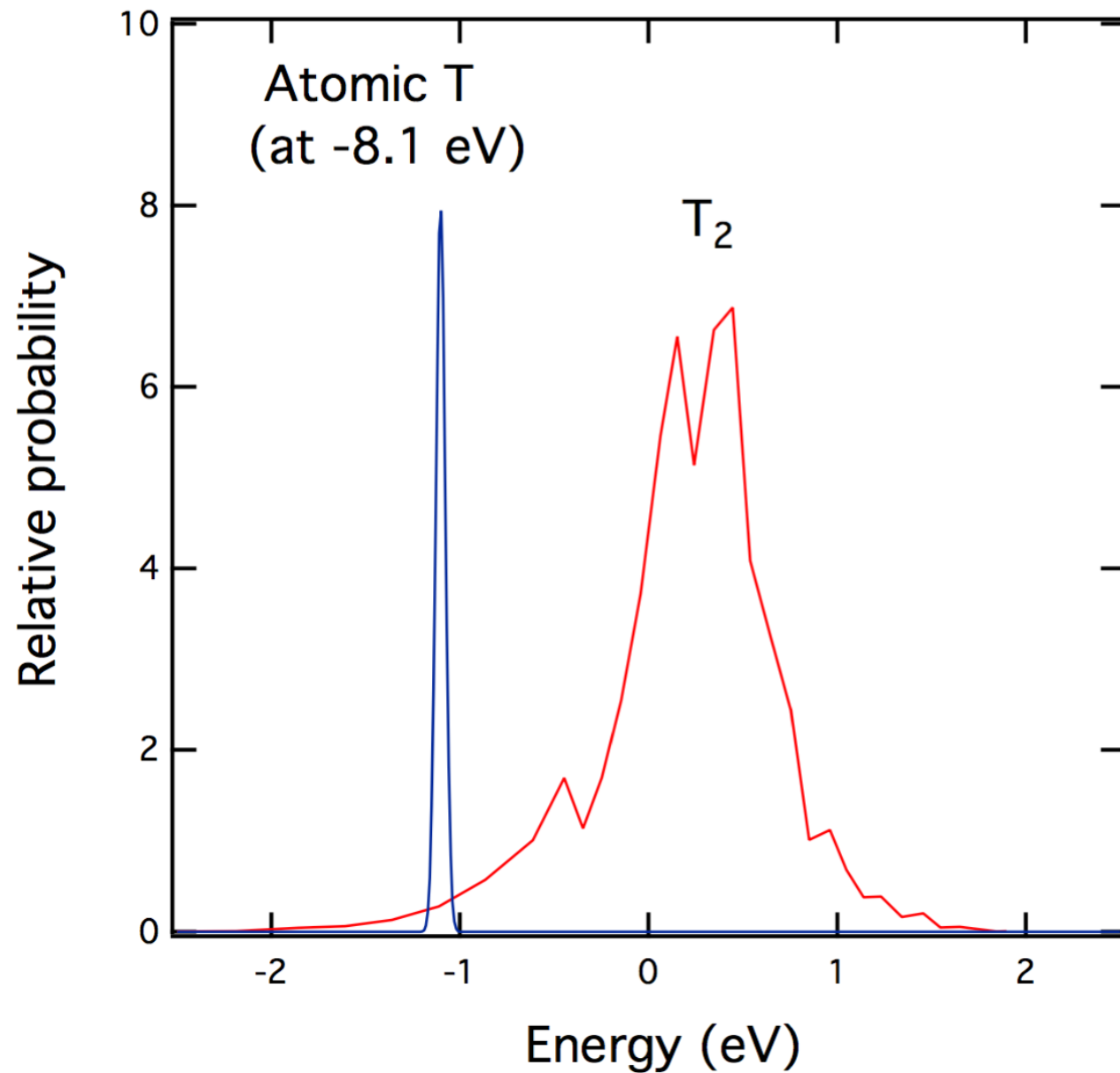
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# Potential for neutrino mass?



# Molecular tritium limitations

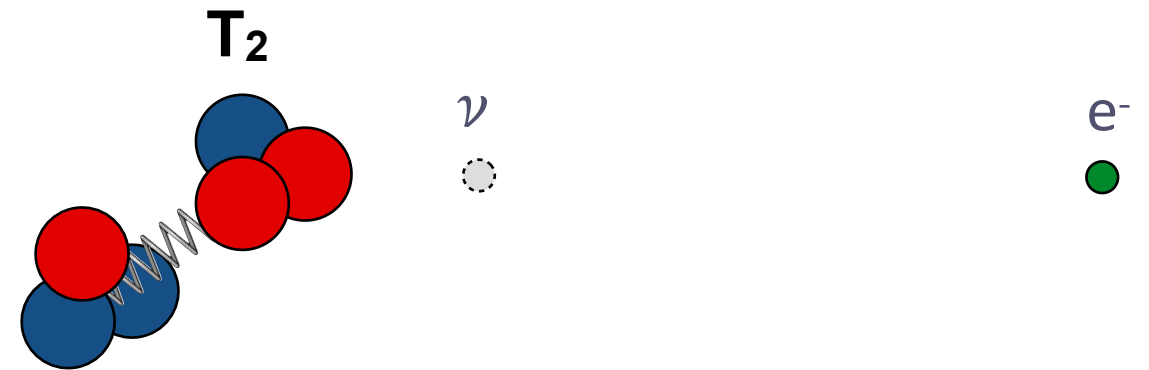
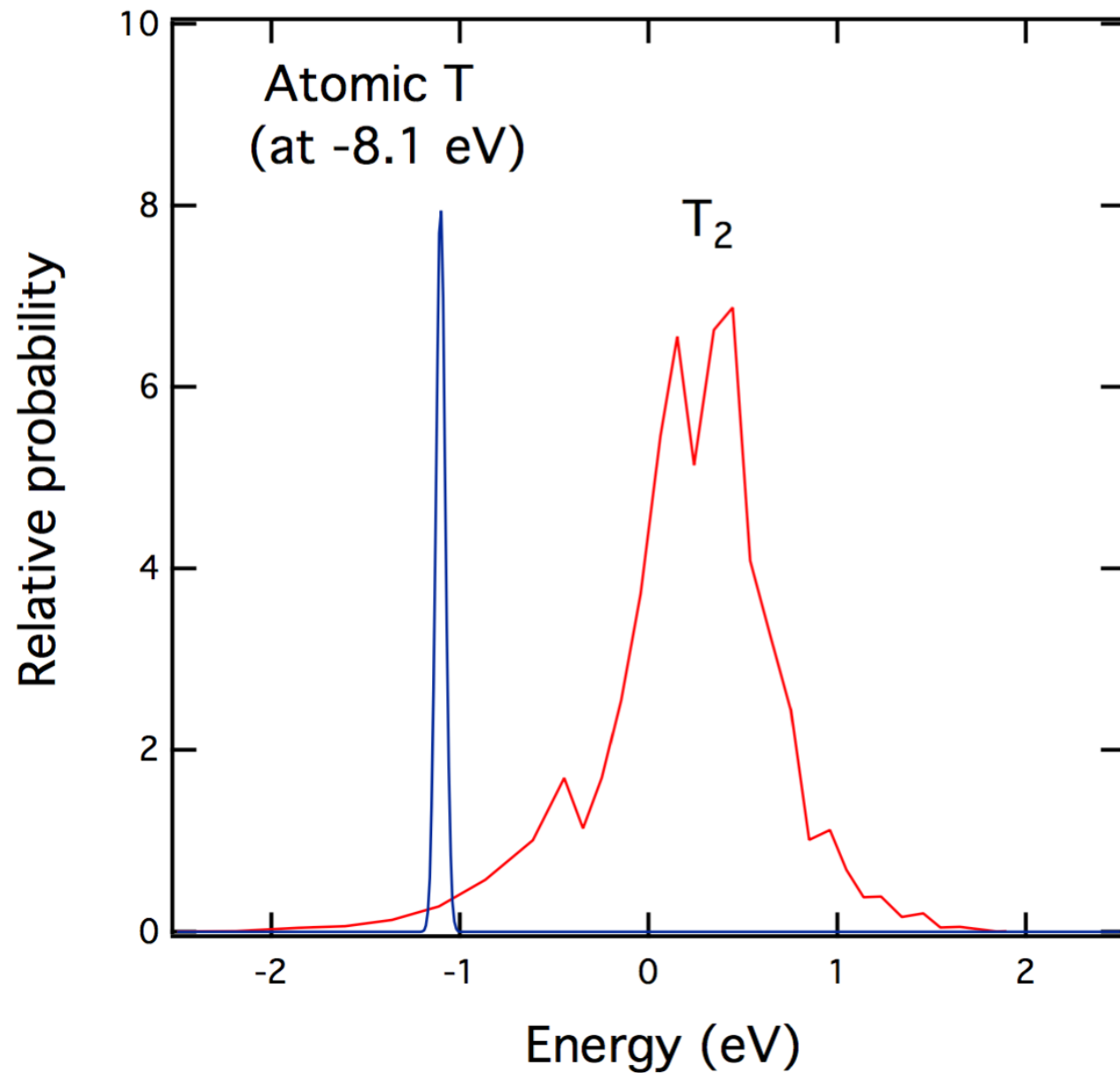


Molecular excitations  
in <sup>3</sup>HeT daughter molecule

- blur tritium endpoint
- ▶ fundamental limit to measurement of  $\nu$ -mass

Need atomic tritium  
for **ultimate** experiment!

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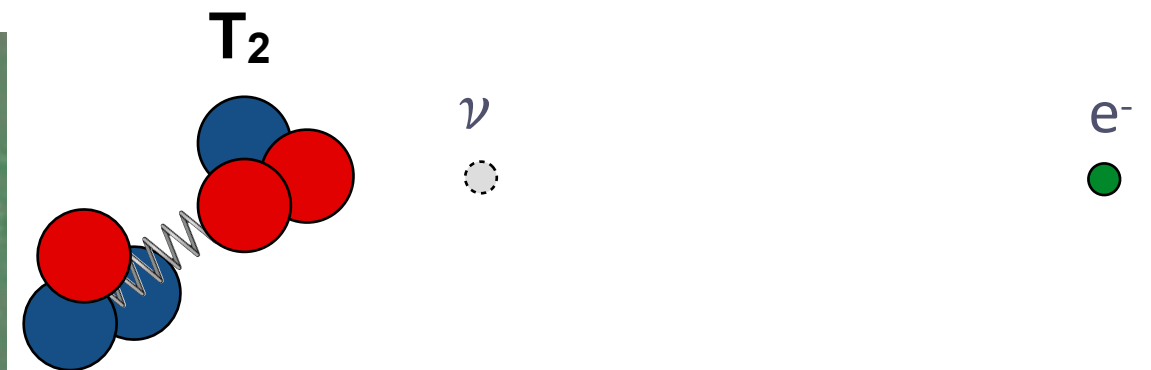
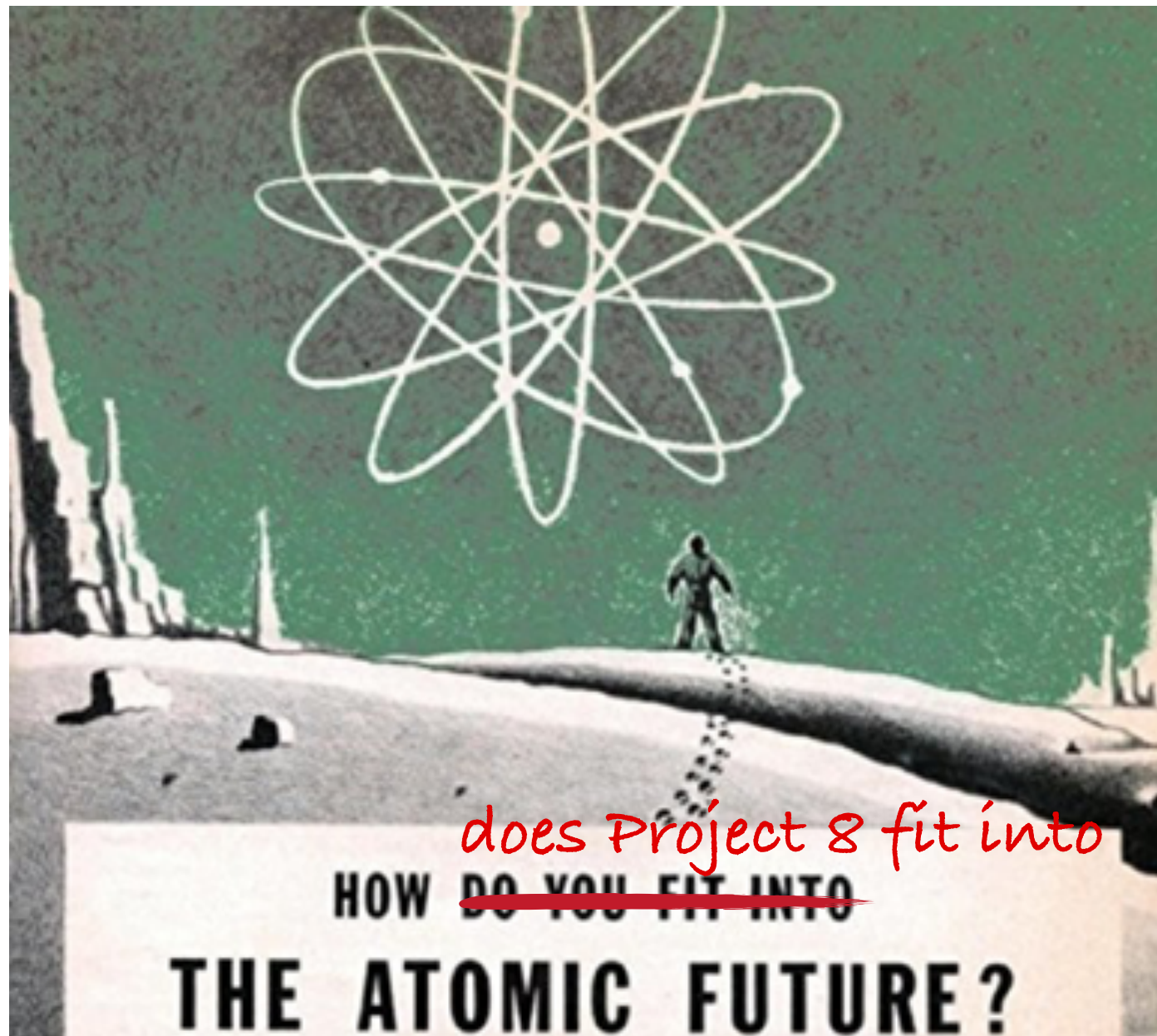


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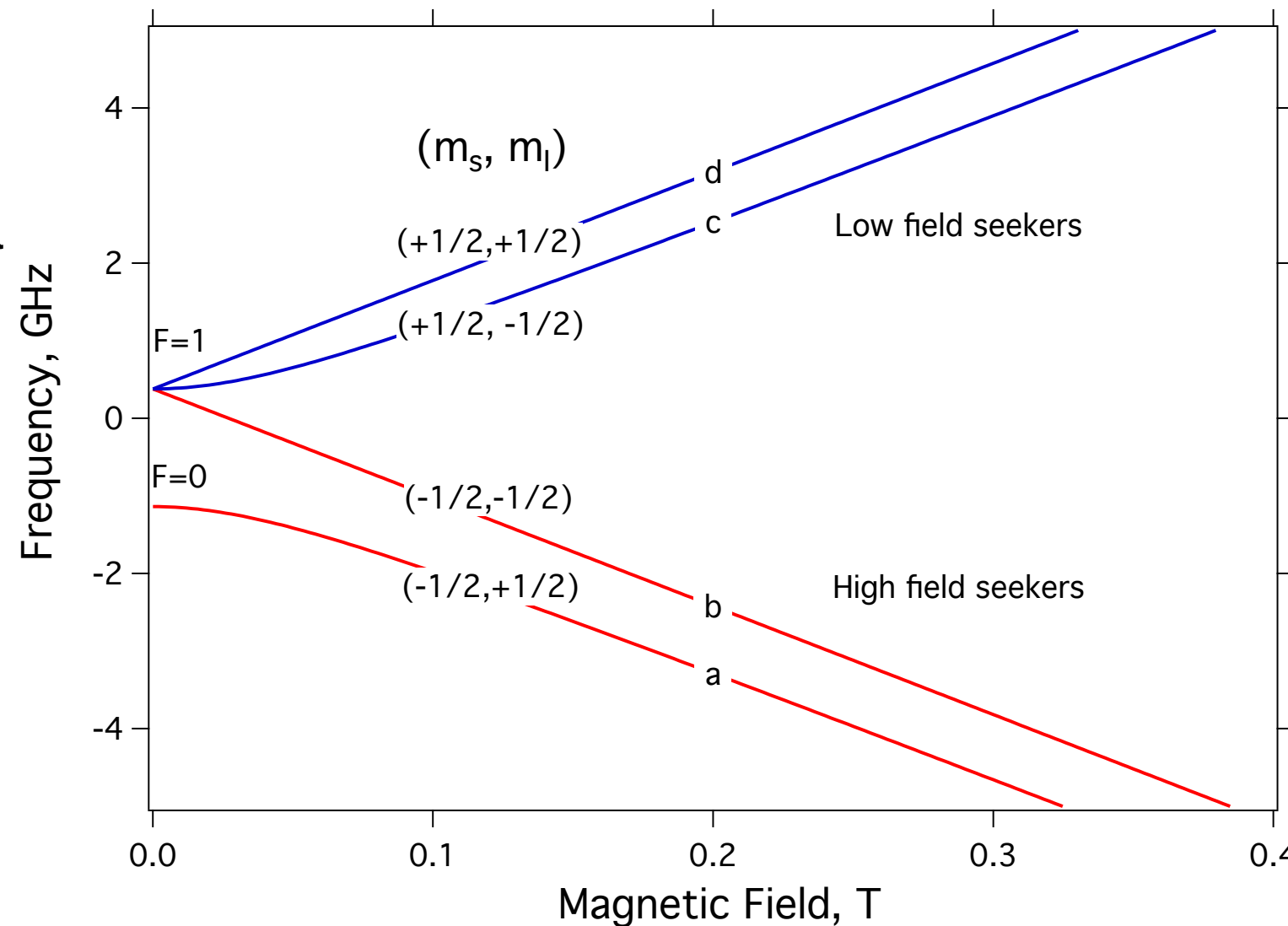
# Magnetic storage of neutral atoms

## Storage of atomic T

- recombination catalyzed by walls  
→ difficult!
- H, D and T have unpaired e<sup>-</sup>
  - ▶ non-zero magnetic moment  $\mu$
  - ▶ tend to (anti-)align with B-field if change is adiabatic

## Potential energy

- $\Delta E = -\vec{\mu} \cdot \vec{B}$   
→ half of spin states seek field minimum



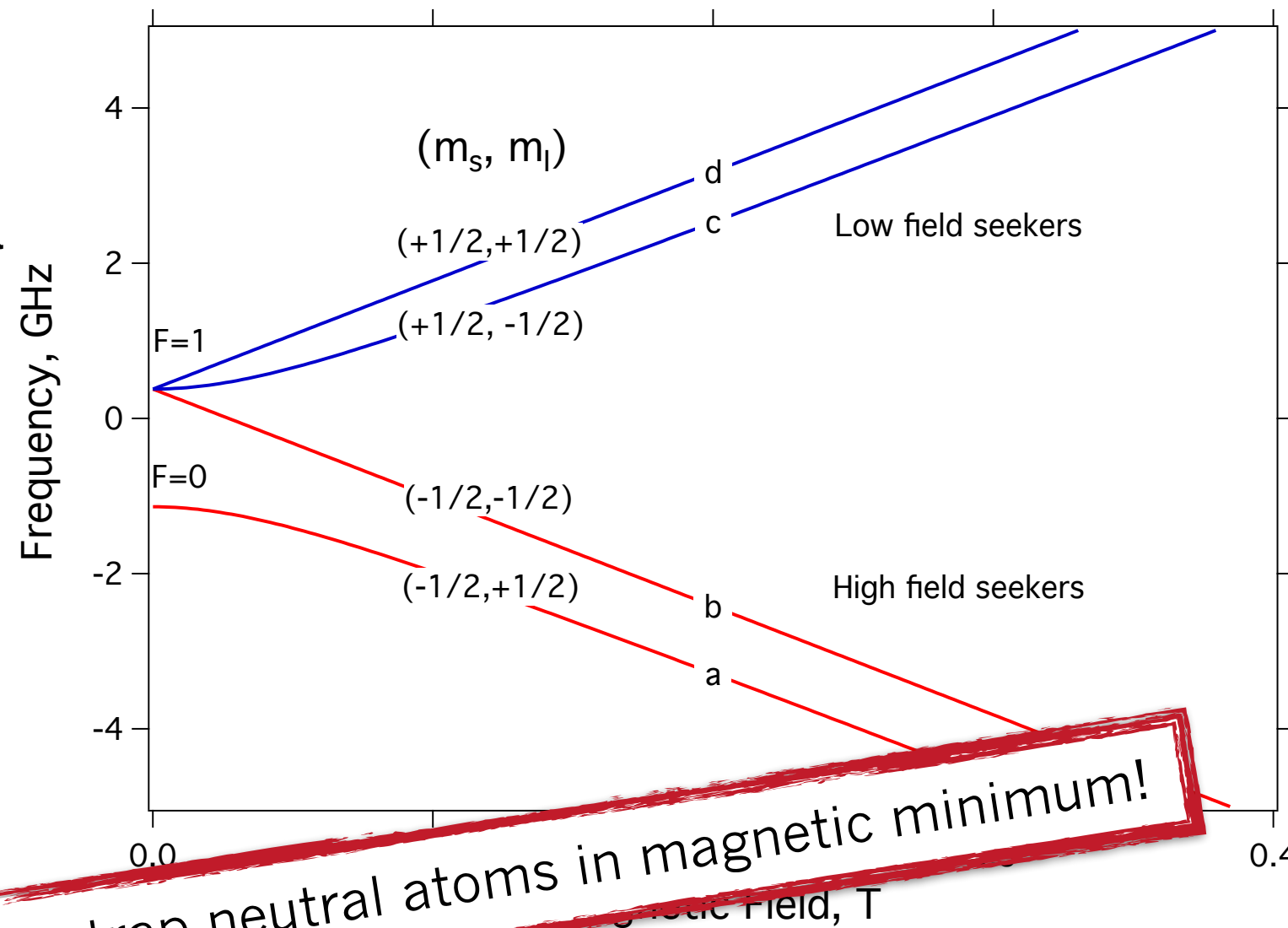
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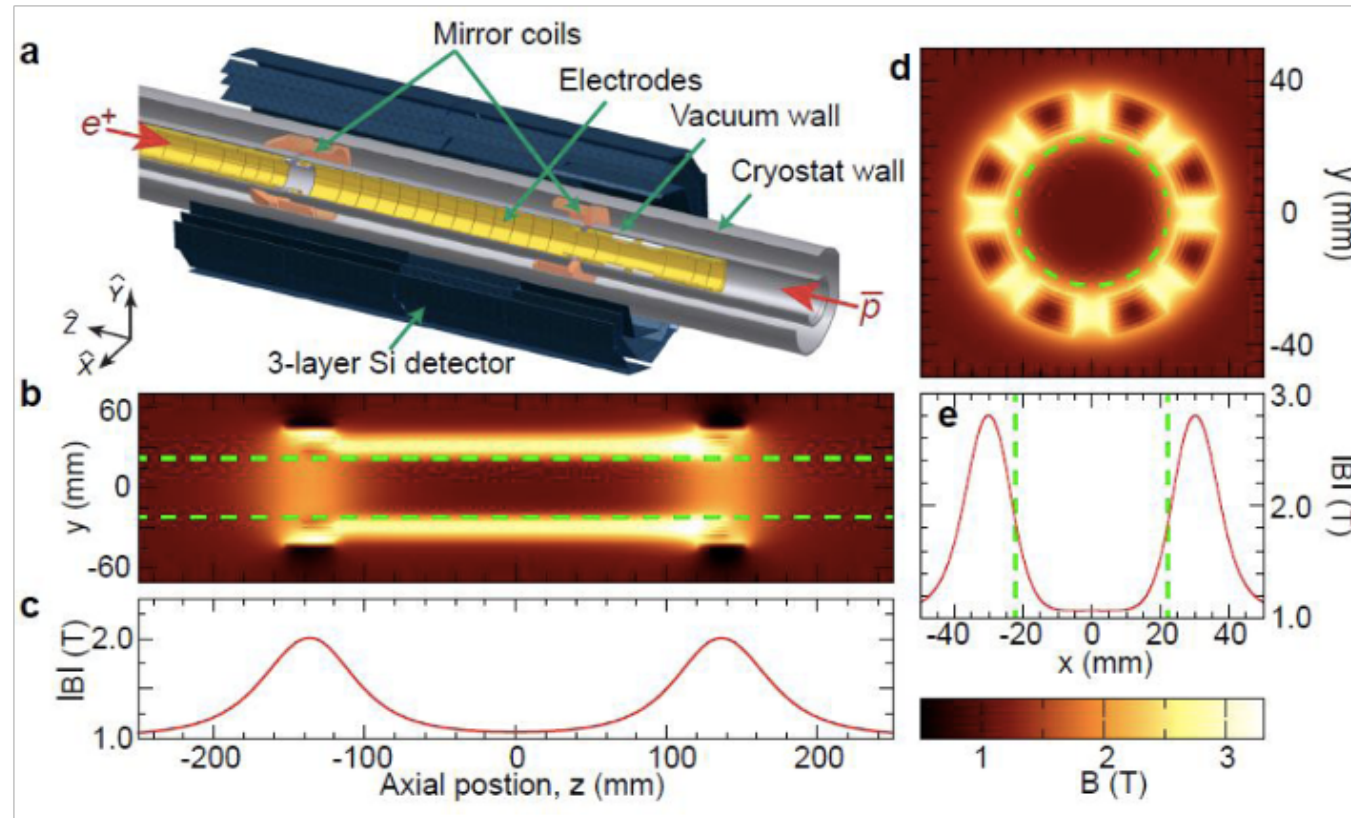
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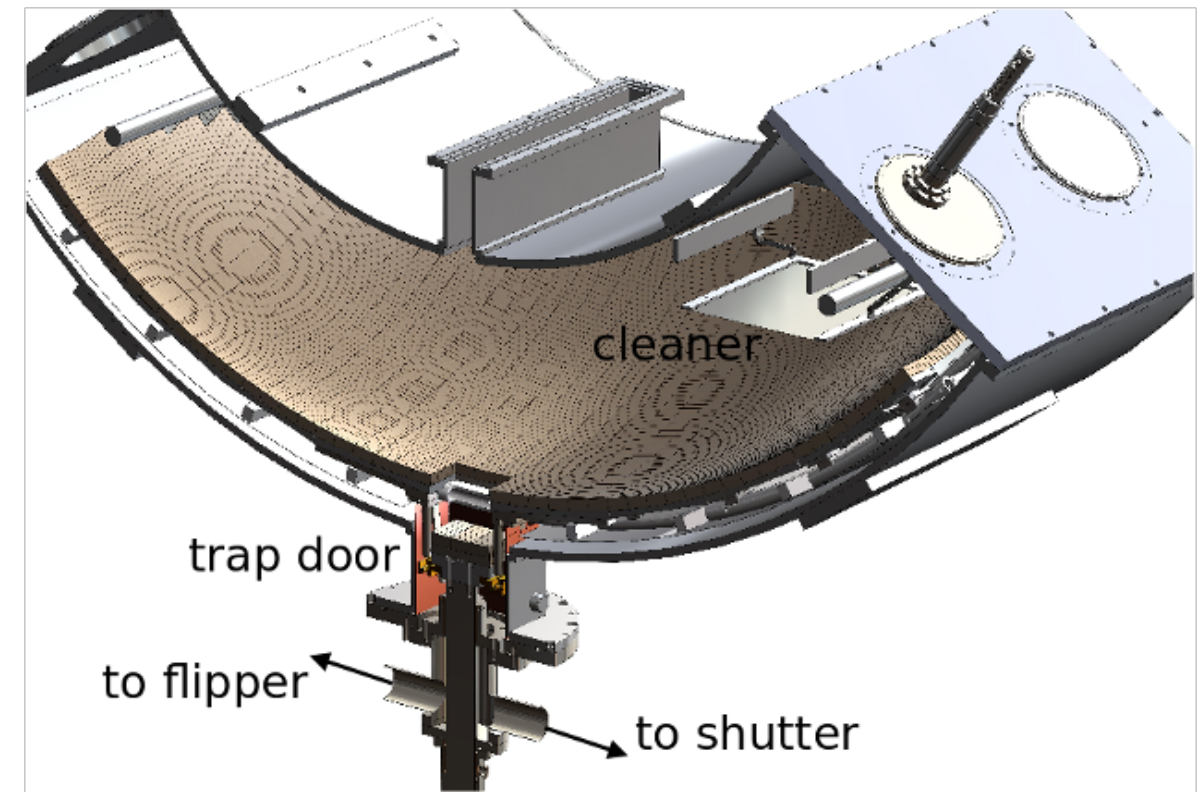
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# Neutral particle traps



ALPHA Collaboration: Nature Phys 7:558, 2011;  
arXiv 1104.4982

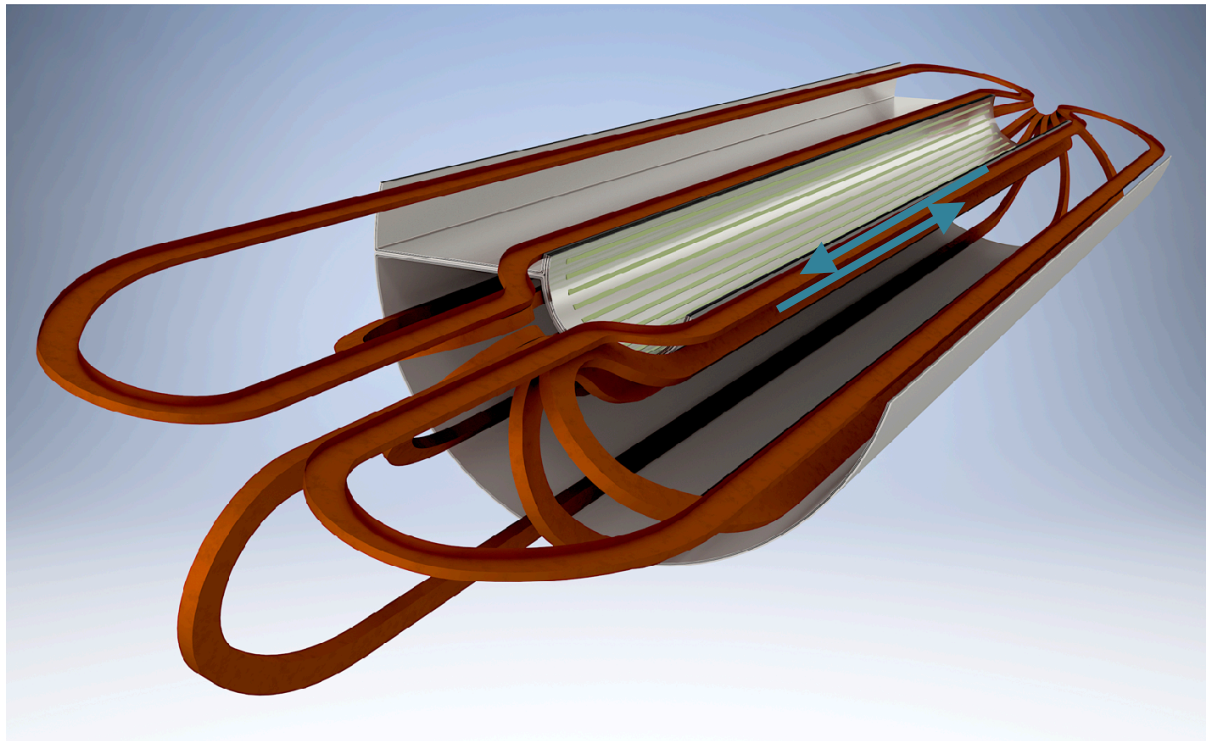


UCNtau Collaboration: Phys Rev C89, 052501, 2014;  
arXiv 1310.5759v3

## General design

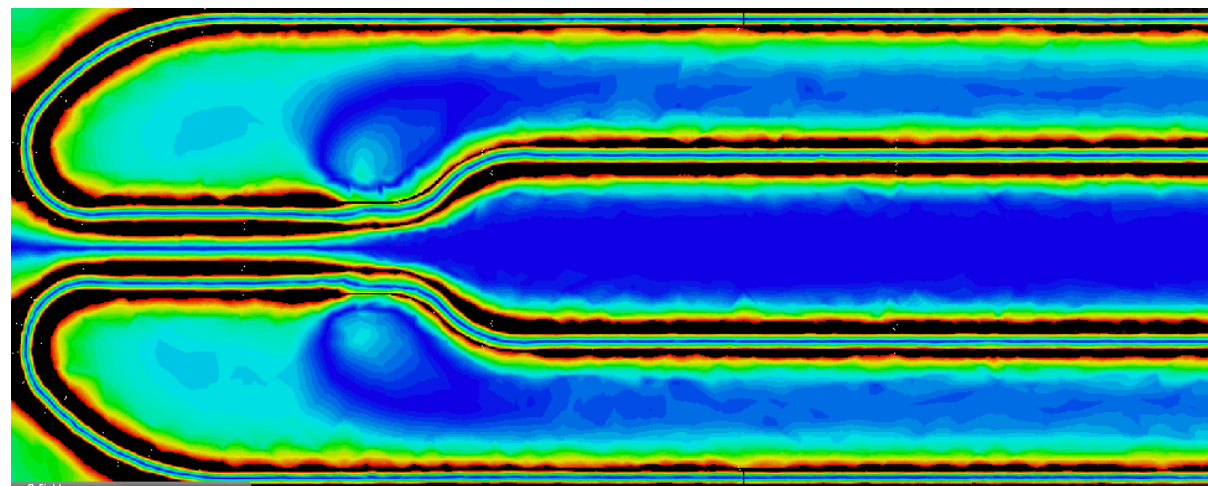
- high magnetic field at walls
- low magnetic fields in the center
  - ▶ near-field to far-field transition with opposing fields

# Atom trapping



## Studying Ioffe-Pritchard trap

- plausible field step
  - ▶  $\Delta B = 2 \text{ T}$
- limit thermal loss fraction
  - ▶  $\epsilon_{\text{loss}} = 10^{-10}$
- maximum allowed temperature
  - ▶  $T_{\text{max}} = \mathbf{30 \text{ mK}}$

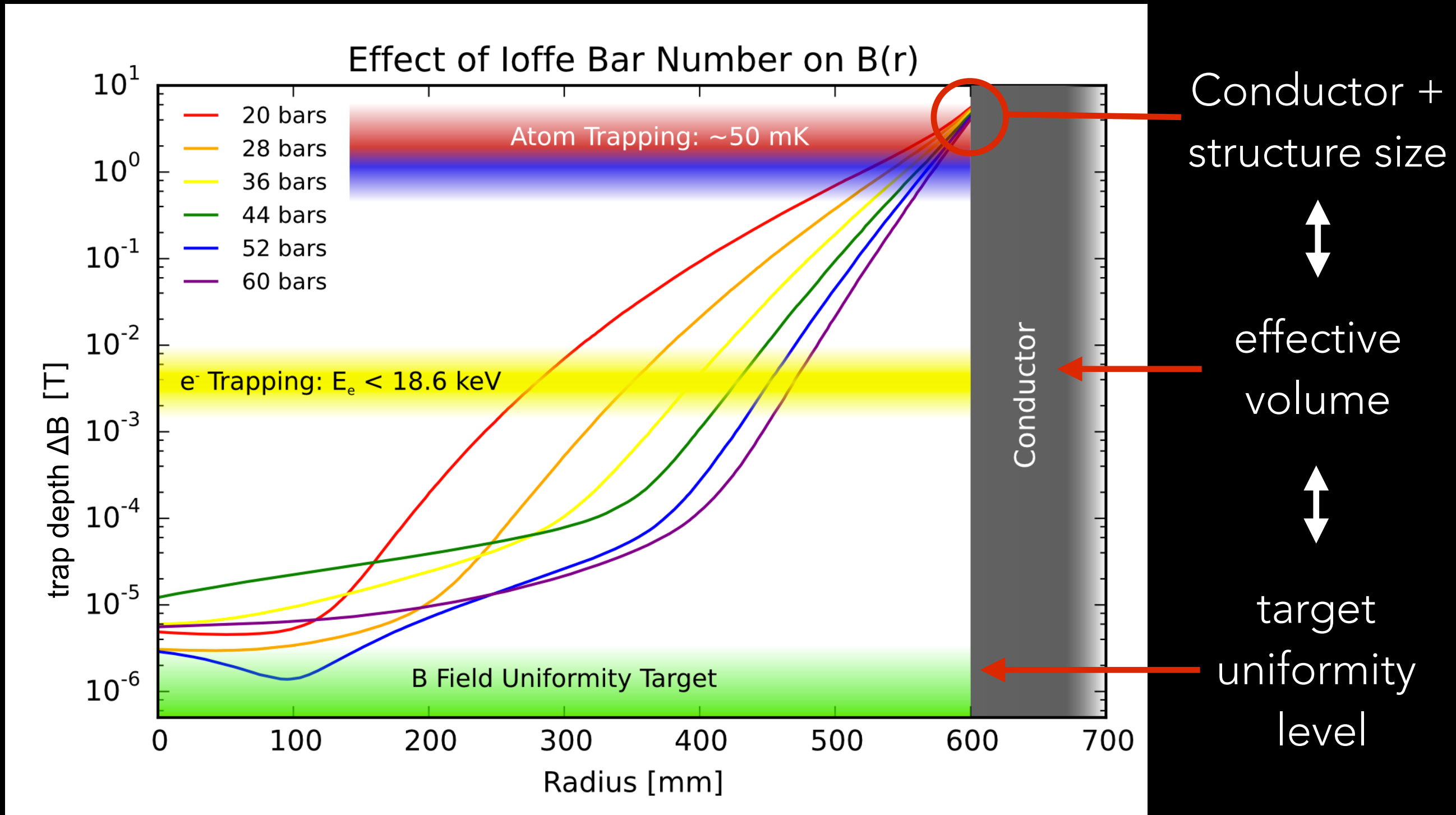


## Challenges

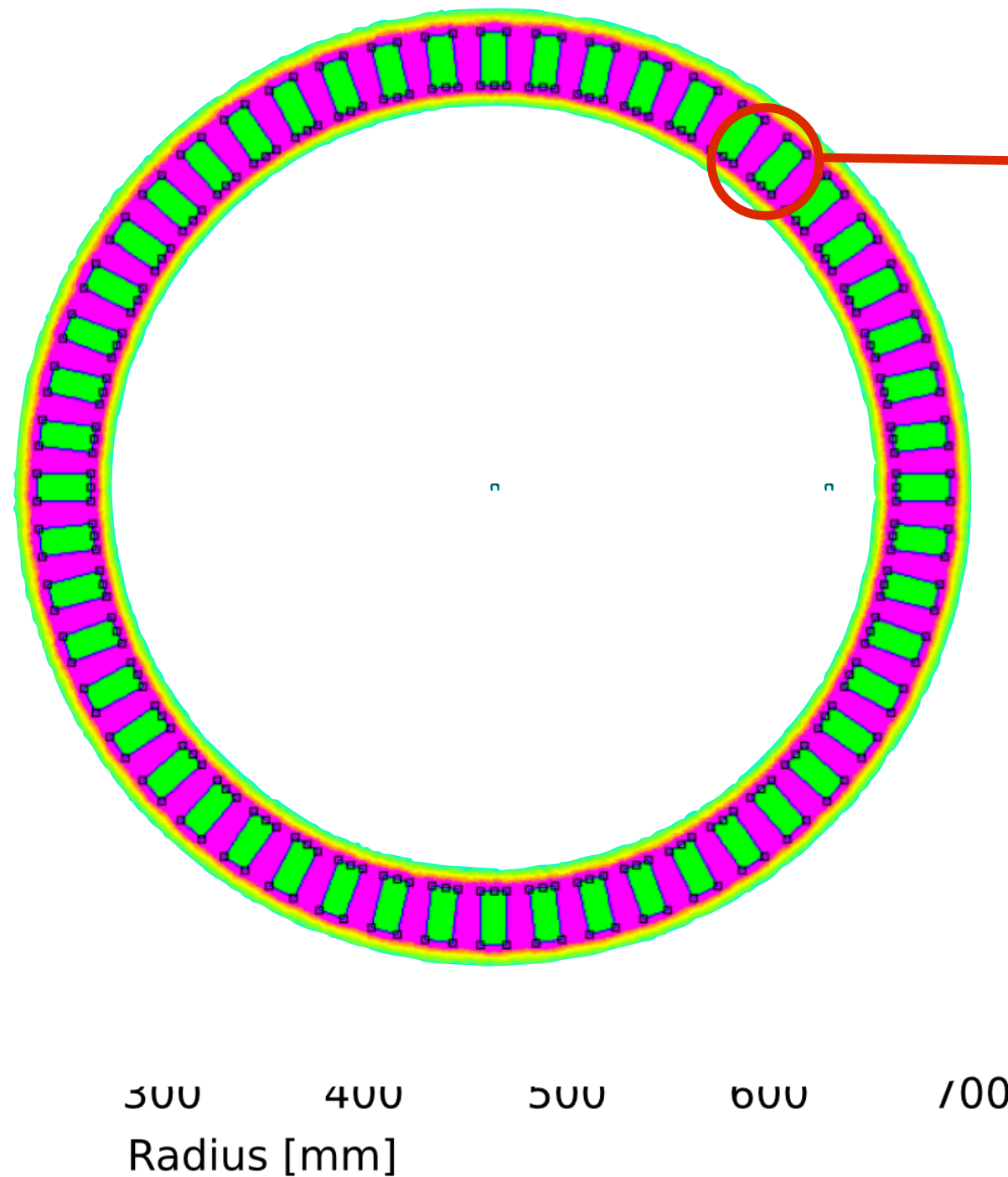
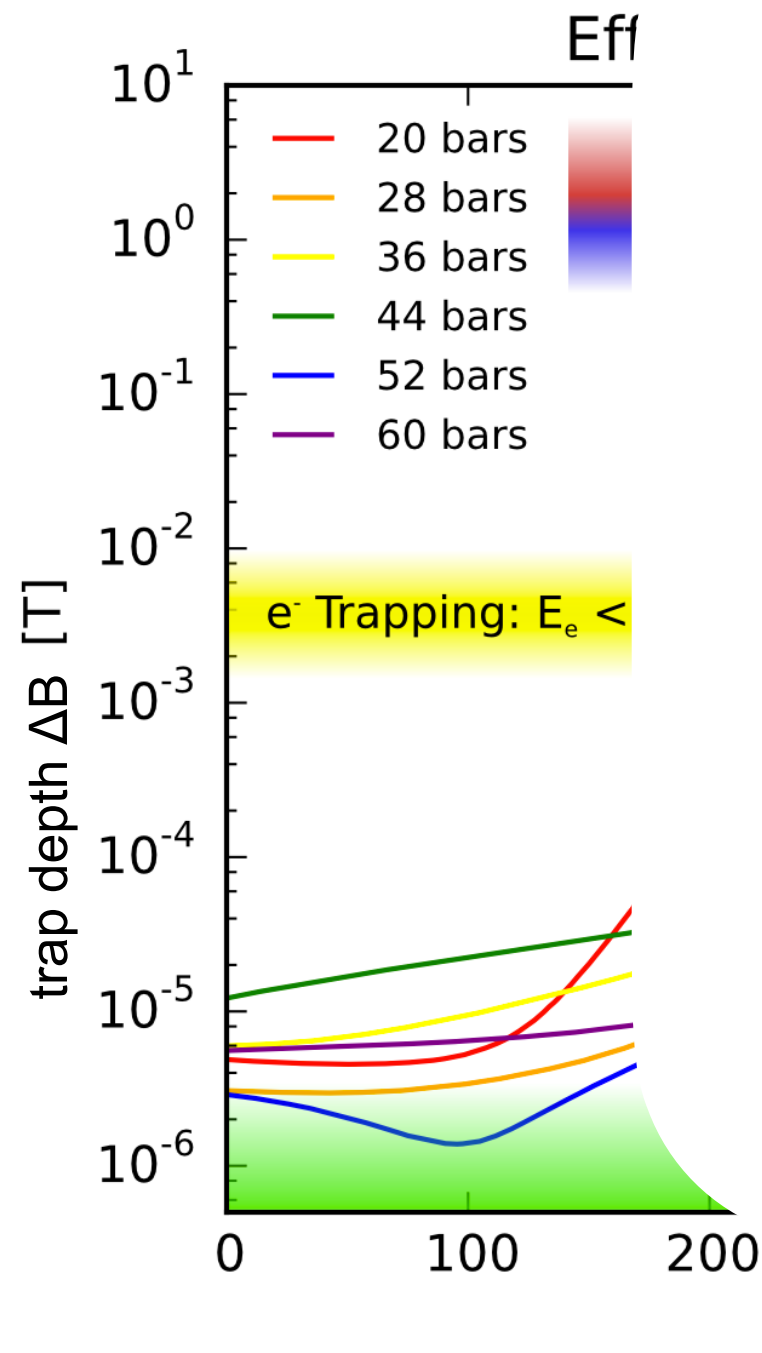
- cooling to sub-Kelvin level
- keep high  $T/T_2$  purity
  - ▶ molecular  $T_2$  not trapped!
- field uniformity in central region



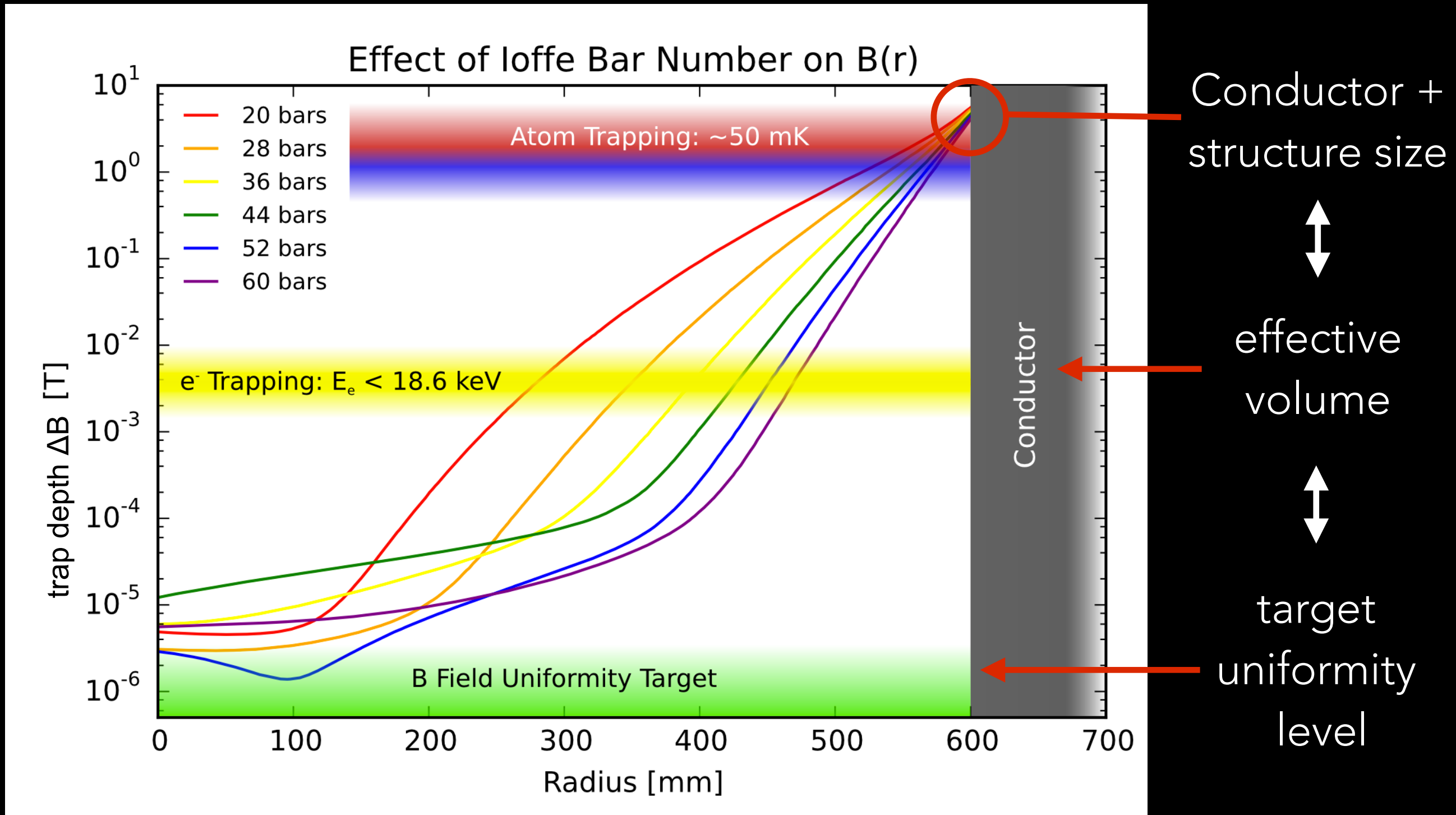
# Atomic trapping: trap challenges



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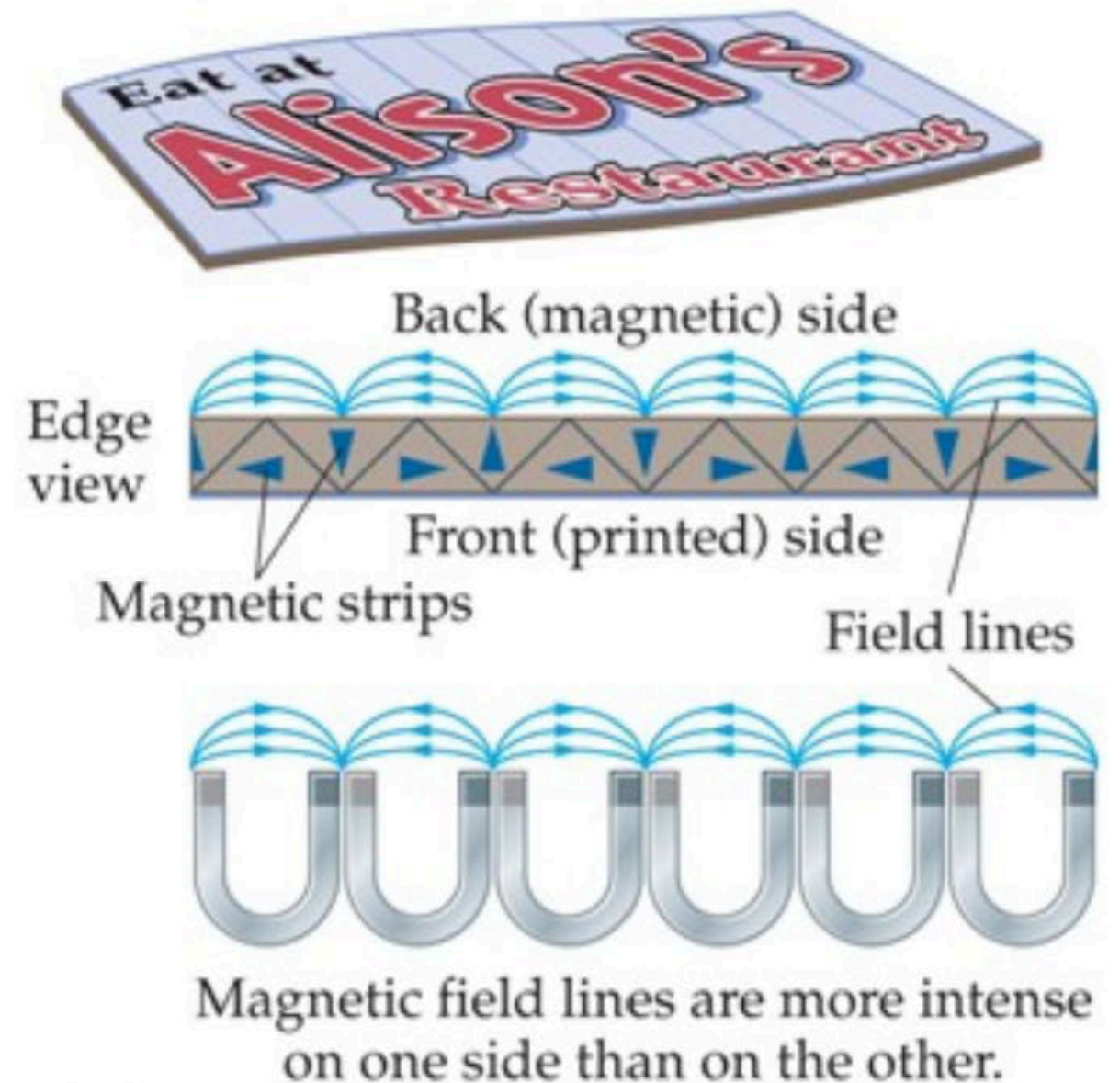
# Atomic trapping: trap challenges



# Halbach arrays

Permanent magnet configuration

- alternate orientations
  - ▶ strong near field
  - ▶ weak far field
- circular flux configuration
  - ▶ one “*magnetic*” side
  - ▶ one “*non-magnetic*” side

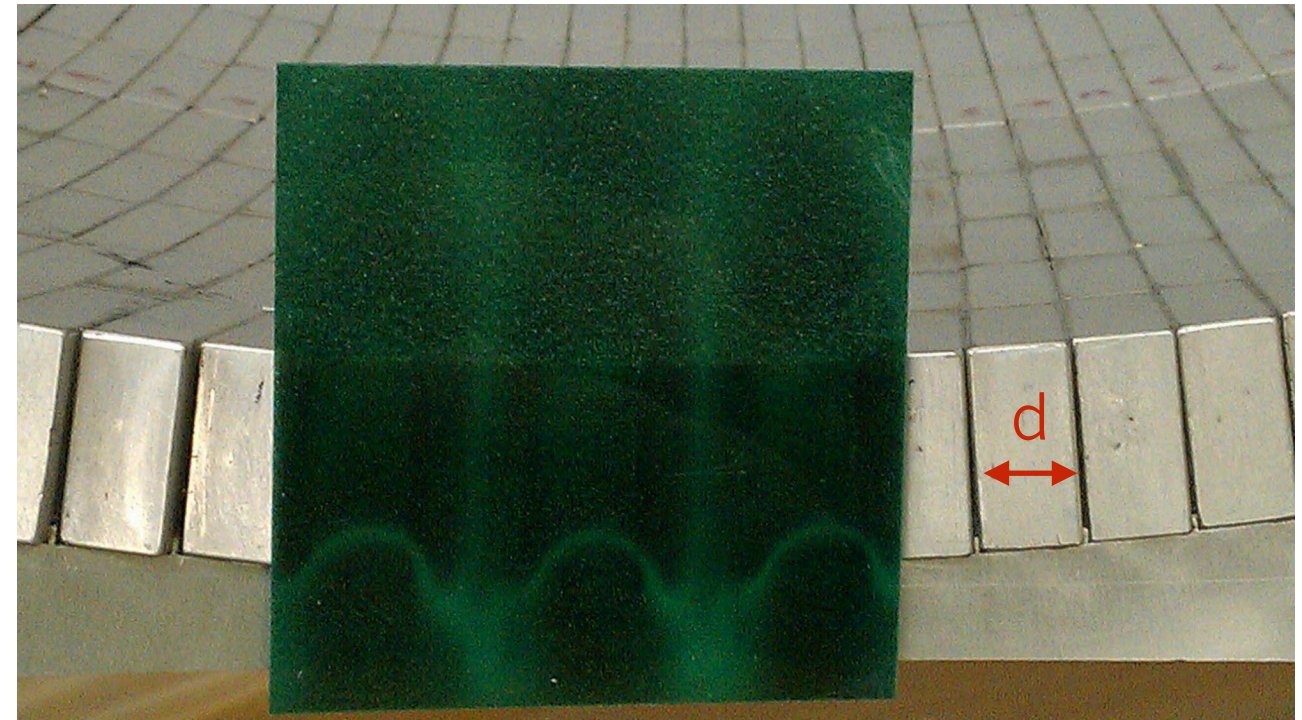




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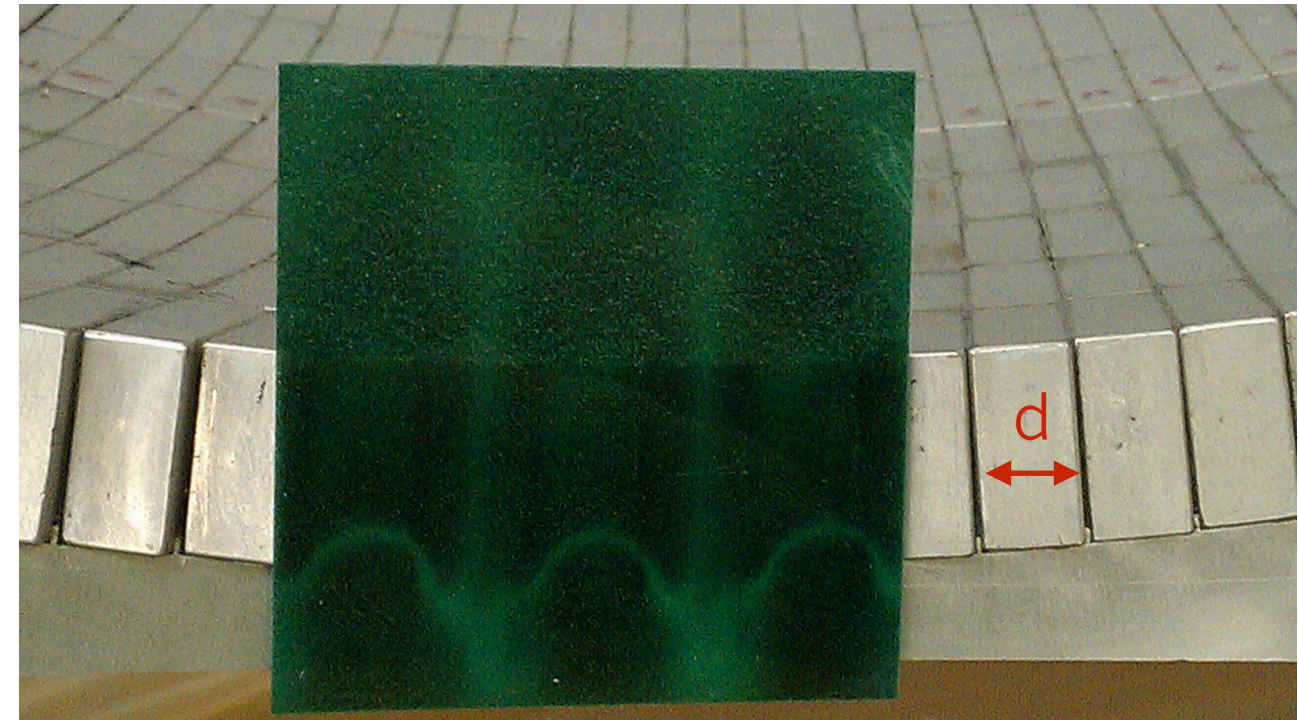
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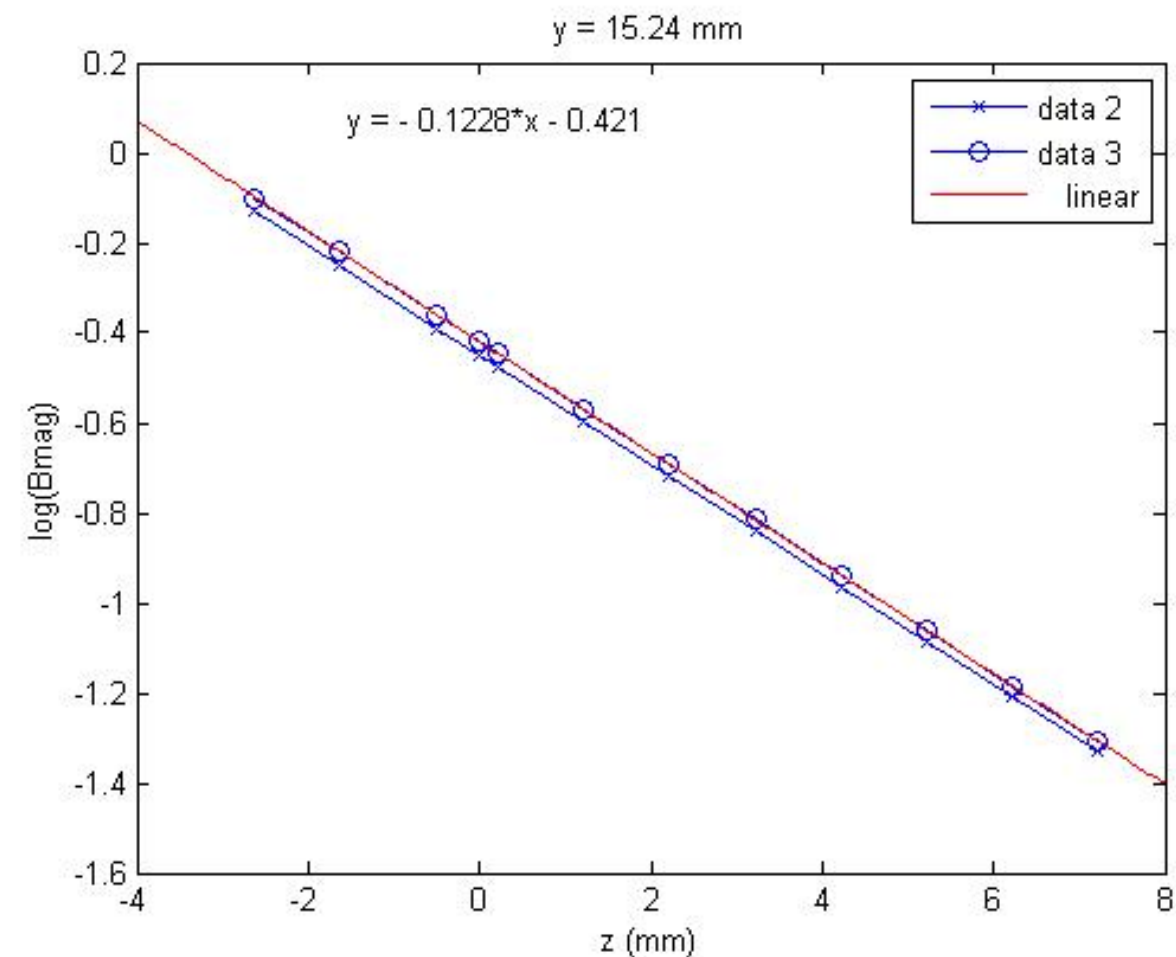
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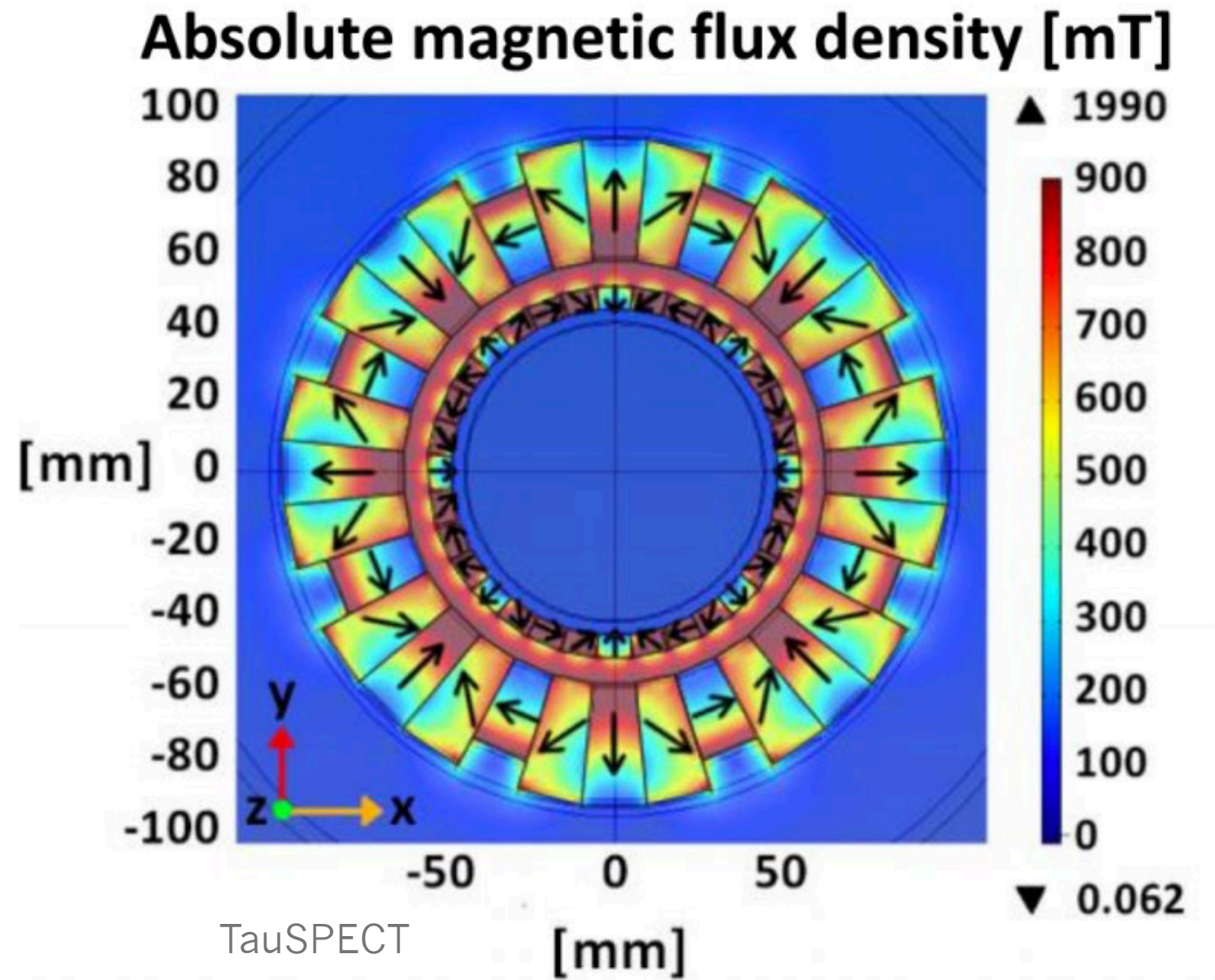
## Flat Halbach array

- field falls exponentially with characteristic length  $l = \frac{2d}{\pi}$ 
  - ▶ weak far field





# Halbach arrays for Project 8



# Halbach arrays for Project 8

TauSPECT

# Halbach arrays for Project 8

Field configuration

- cylindrical configuration

$$\vec{B}_{\text{Halbach}} = \vec{B}_{\rho} + \vec{B}_{\varphi}$$

- ▶ no cancellation with solenoid electron field

$$\vec{B}_{\text{solenoid}} = \vec{B}_z$$

TauSPECT

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## Choice of permanent magnets

- rare earth give highest field strength
- phase transition in NdFeB
  - ▶ not suitable below 140K
- SmCo, PrFeB,...

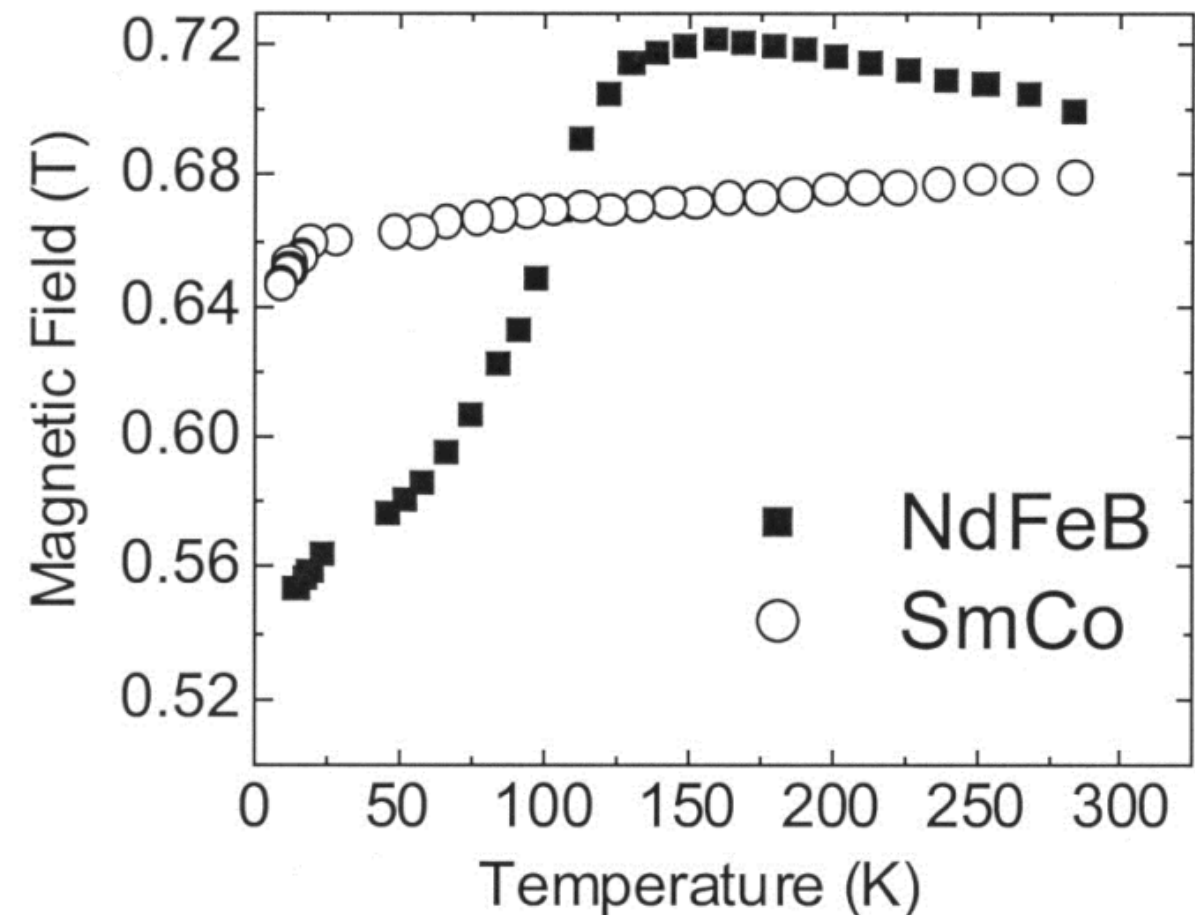


Figure 4. Magnetic Field as a Function of Temperature for NdFeB and SmCo. [6]

# Magneto-gravitational trap

Magnetic trapping

- $E_m = \mu_B B = 58 \mu\text{eV/T}$

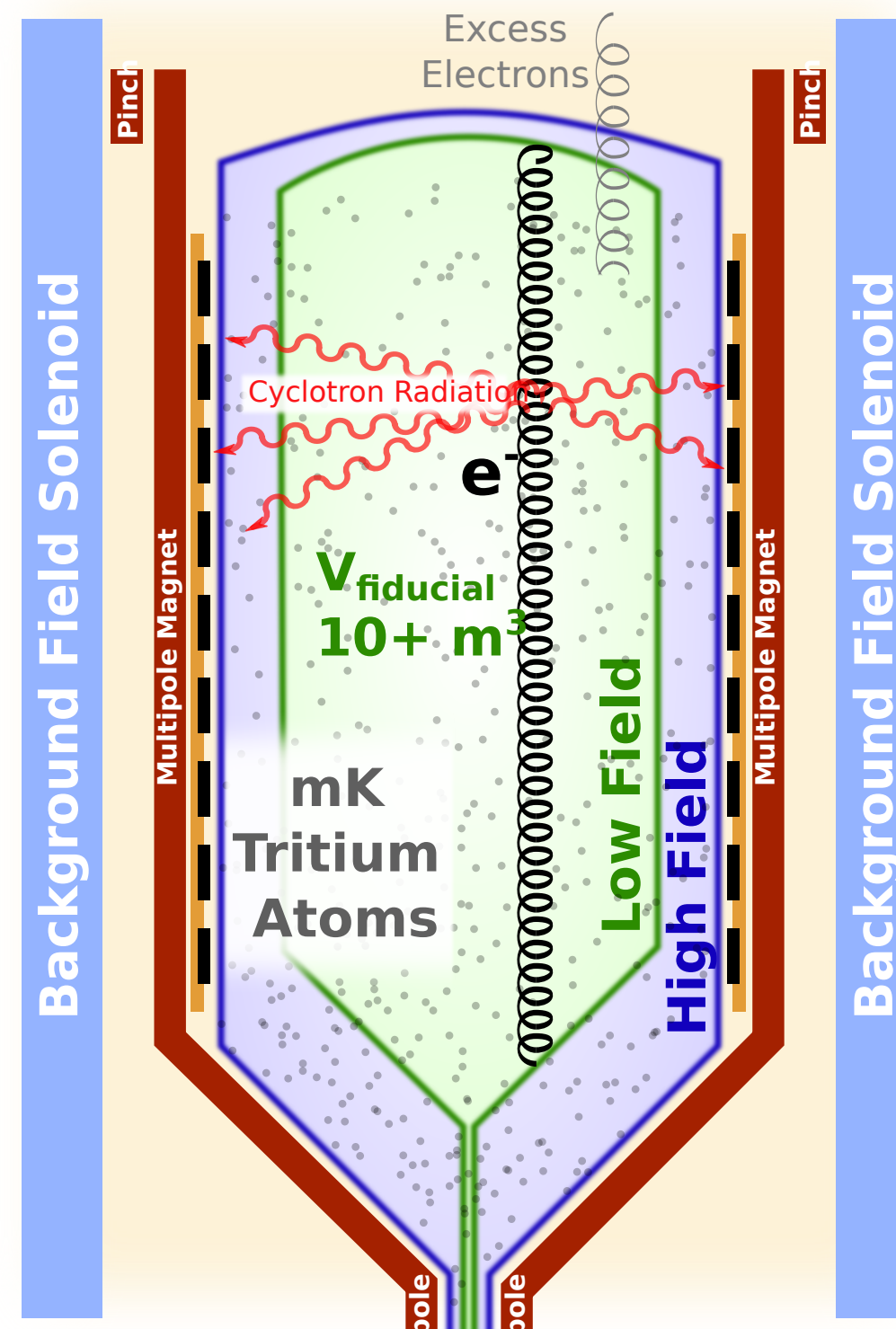
Gravitational trapping

- $E_g = mgh = 0.3 \mu\text{eV/m}$

Energy of cold atomic beam

- $E_k = k_B T = 64 \mu\text{eV/K}$

→ For 1mK cold beam, it takes 0.21 meters of gravity and 11 gauss of B-field to trap.



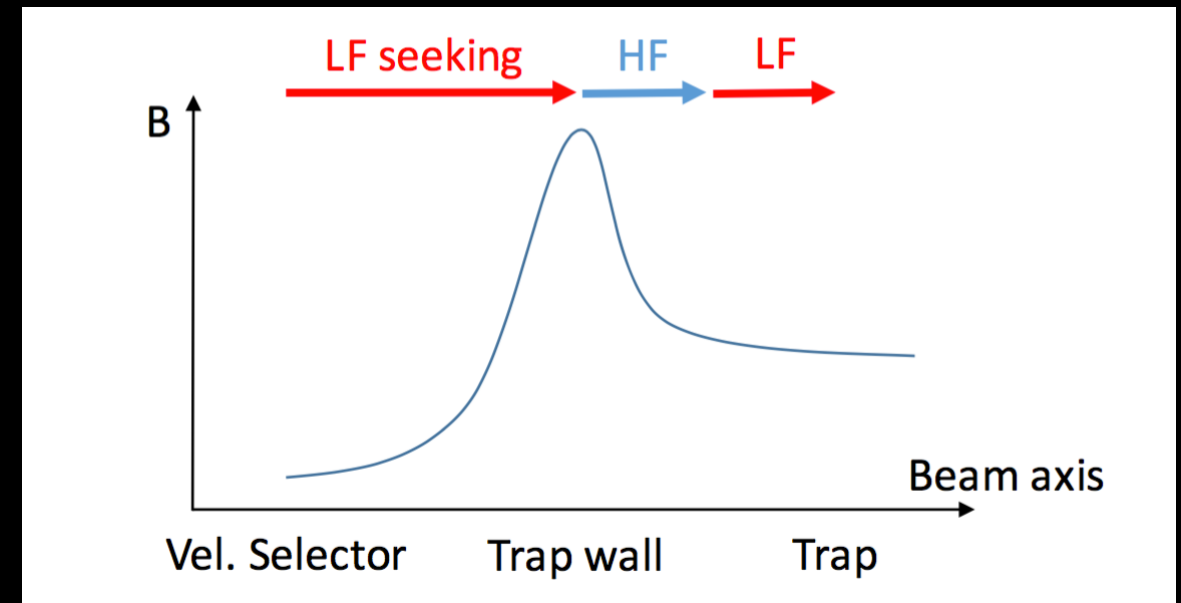


# How to fill the trap?

# How to fill the trap?

## Spin-flip loading ?

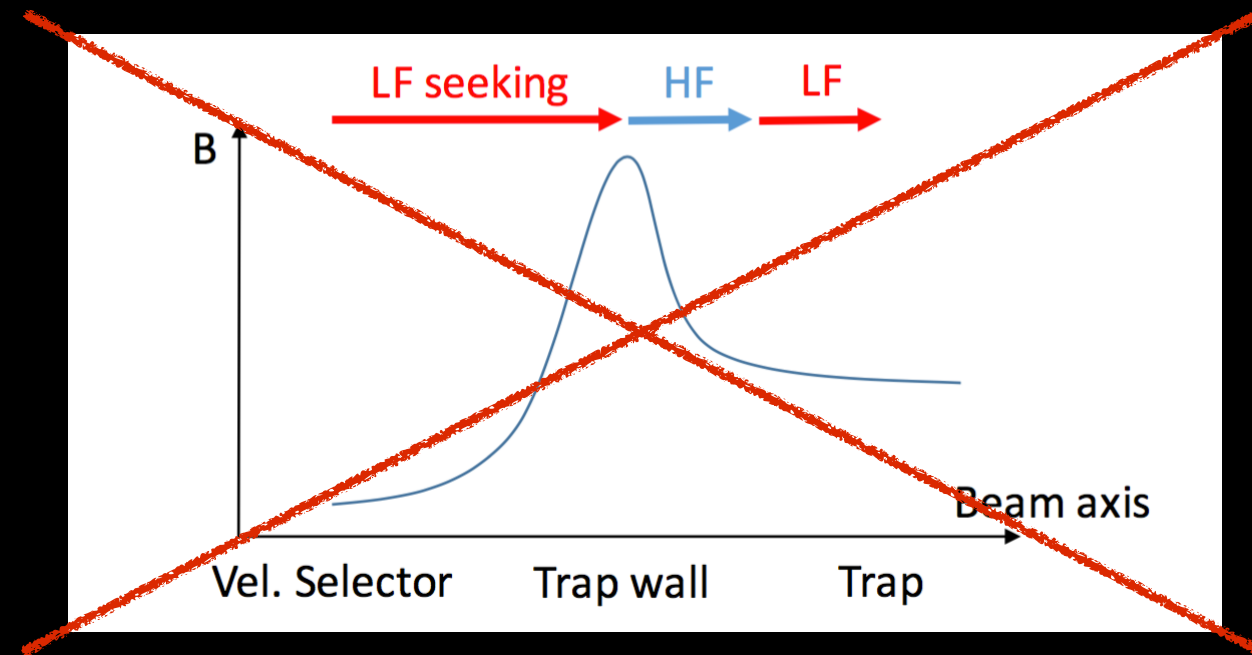
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- But: stimulated emission
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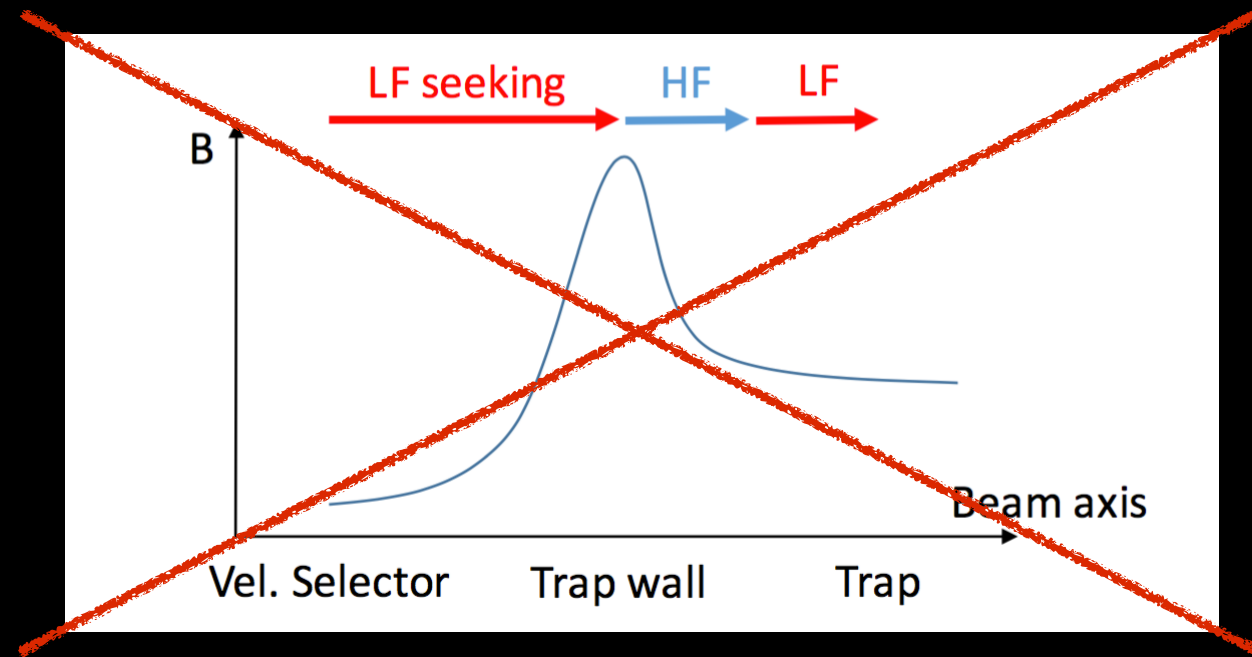
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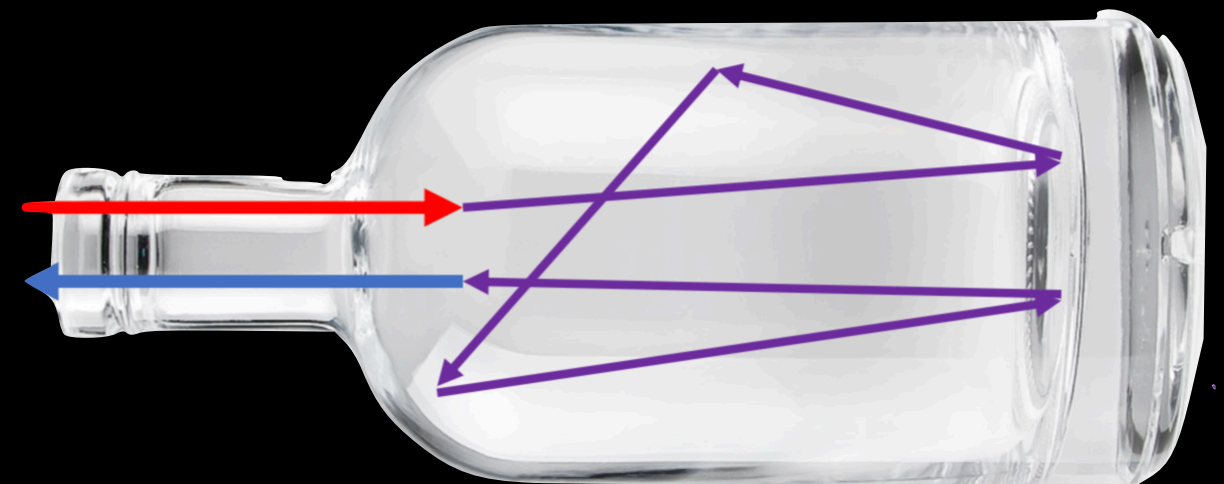
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## Cornucopia\* loading

- Blow cold atoms into trap
  - accept loss through entrance hole
- required input flux for 1cm hole @ 50mK
  - ▶  $5 \cdot 10^{12}$  atoms/sec

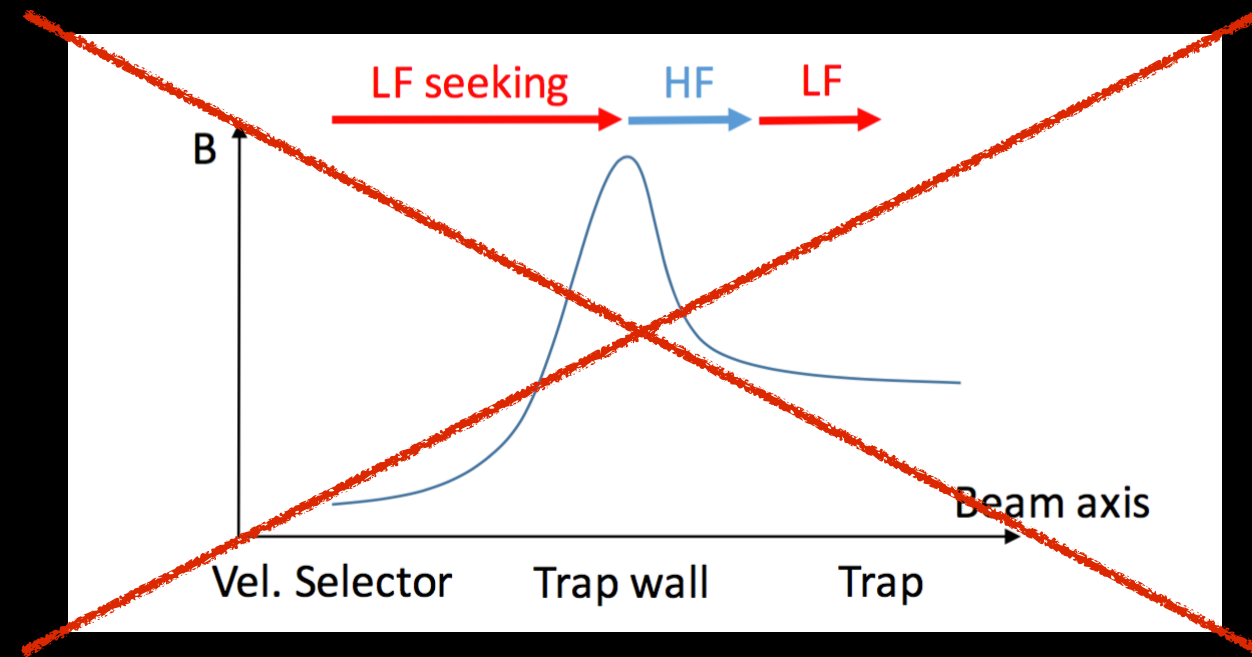


\* horn of plenty

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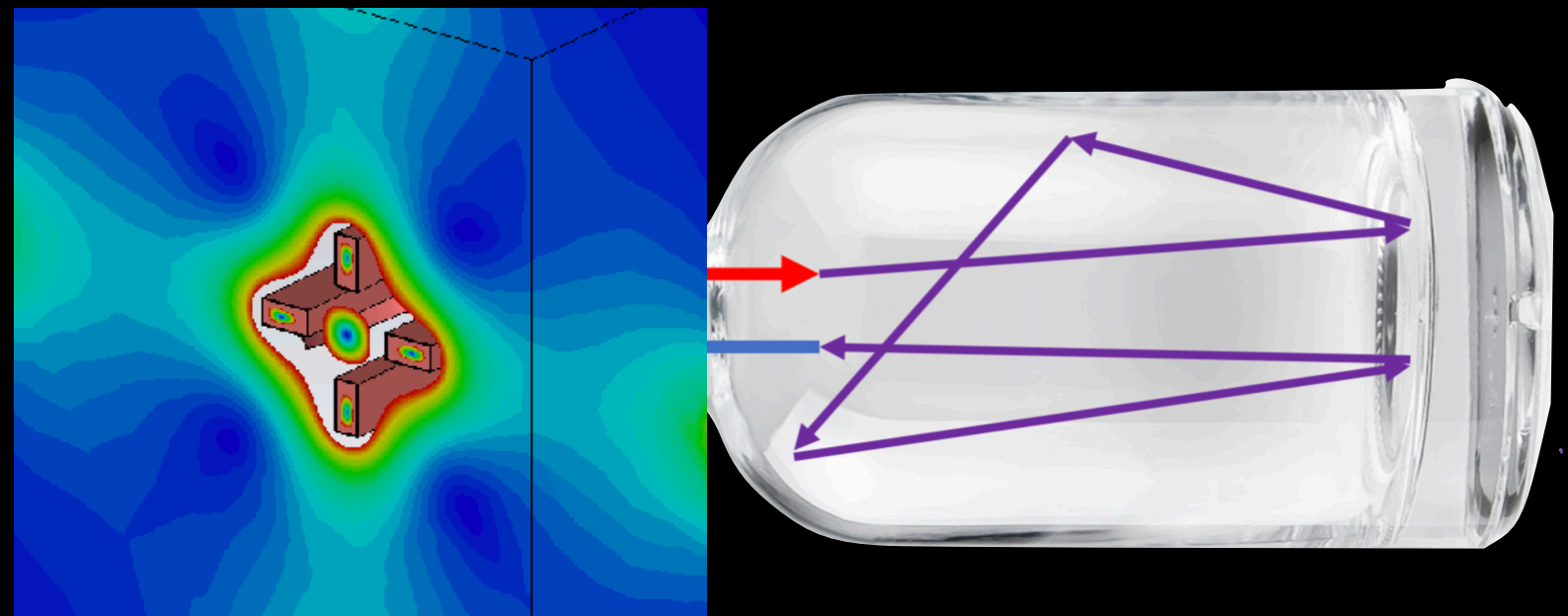
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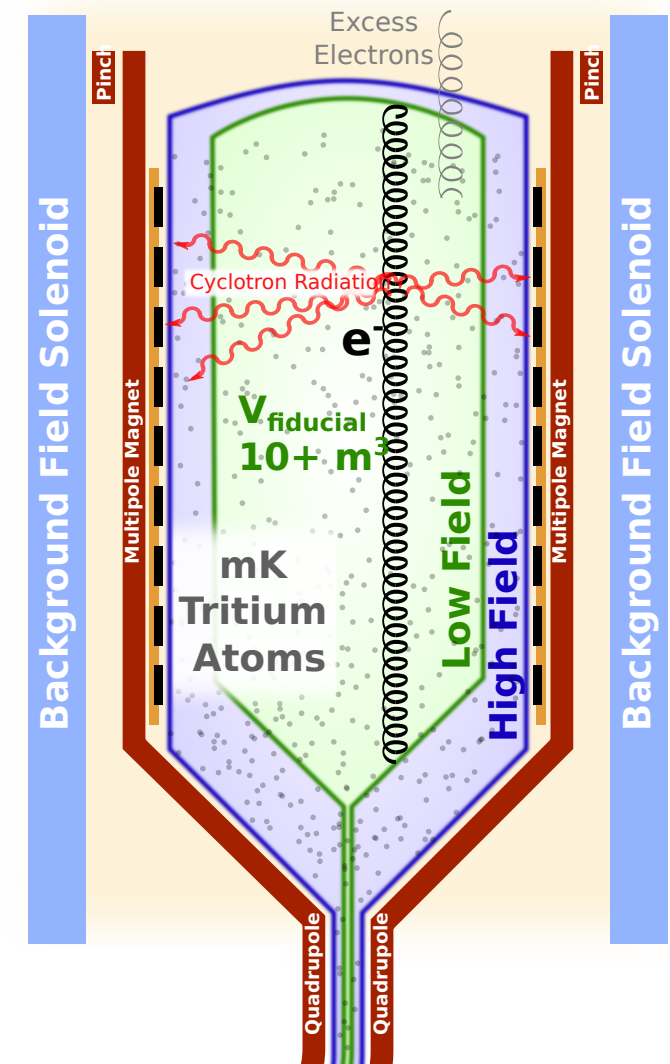
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# Project 8: Designs concepts

Atomic T source

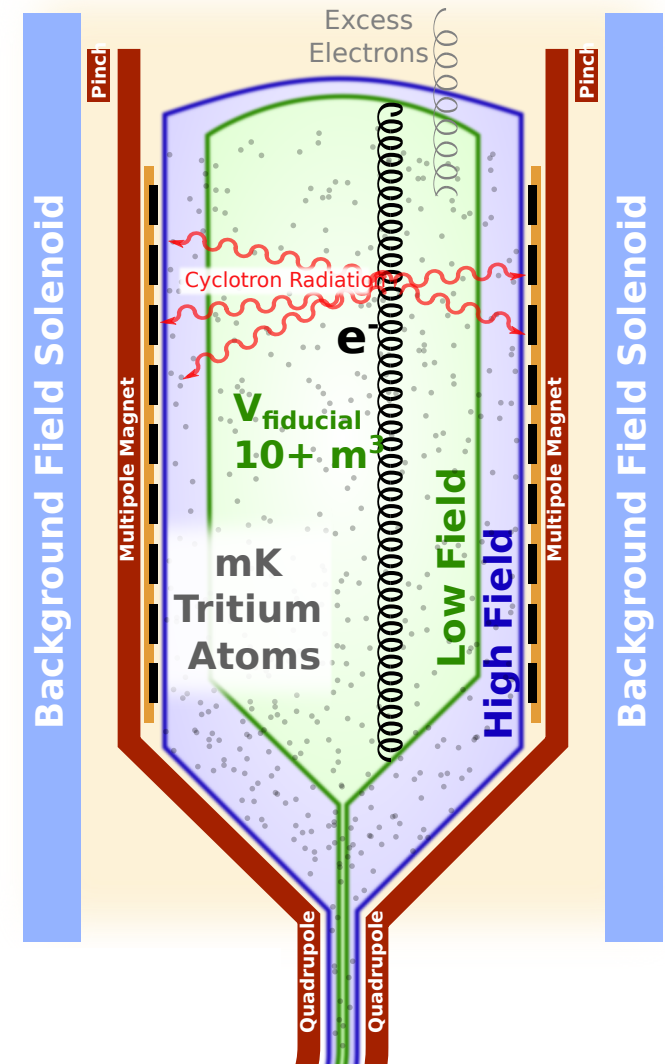
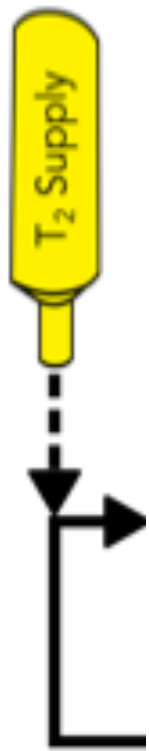
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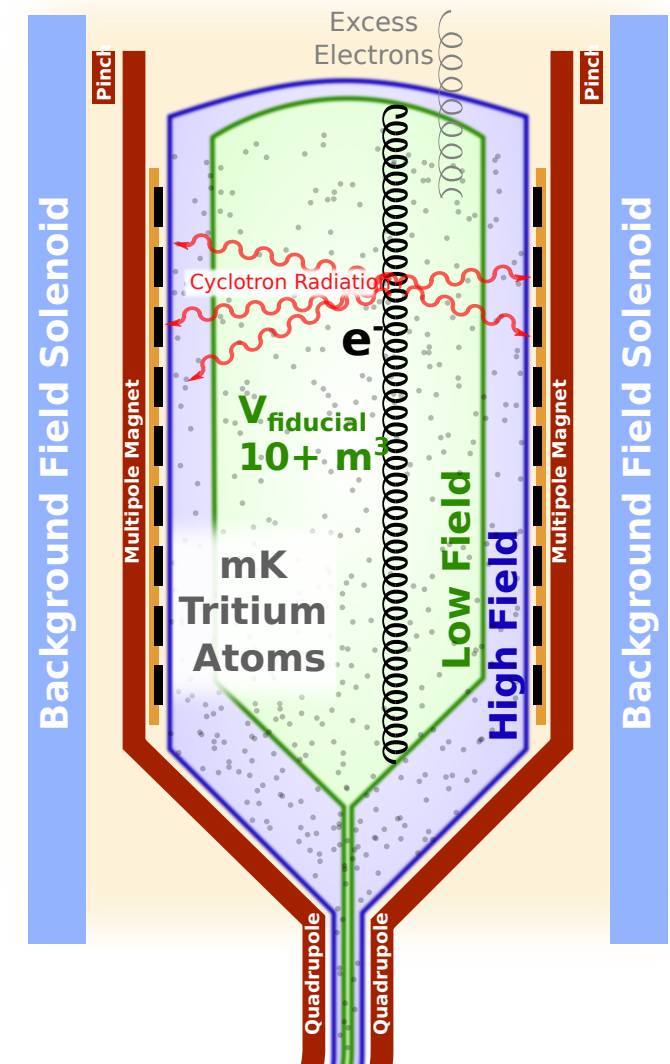
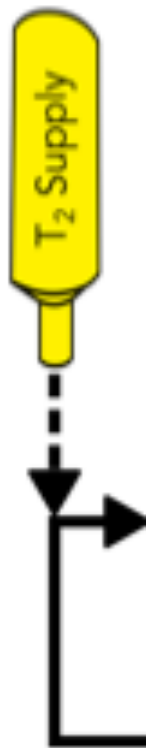




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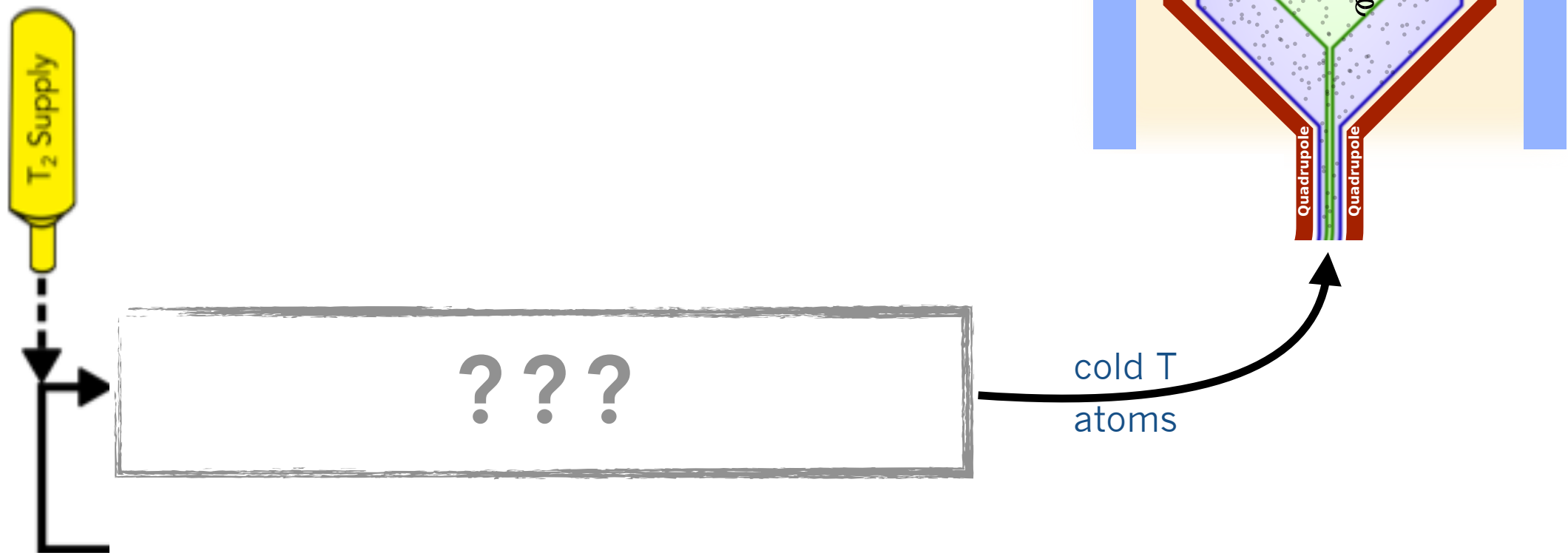
- Injection ✓
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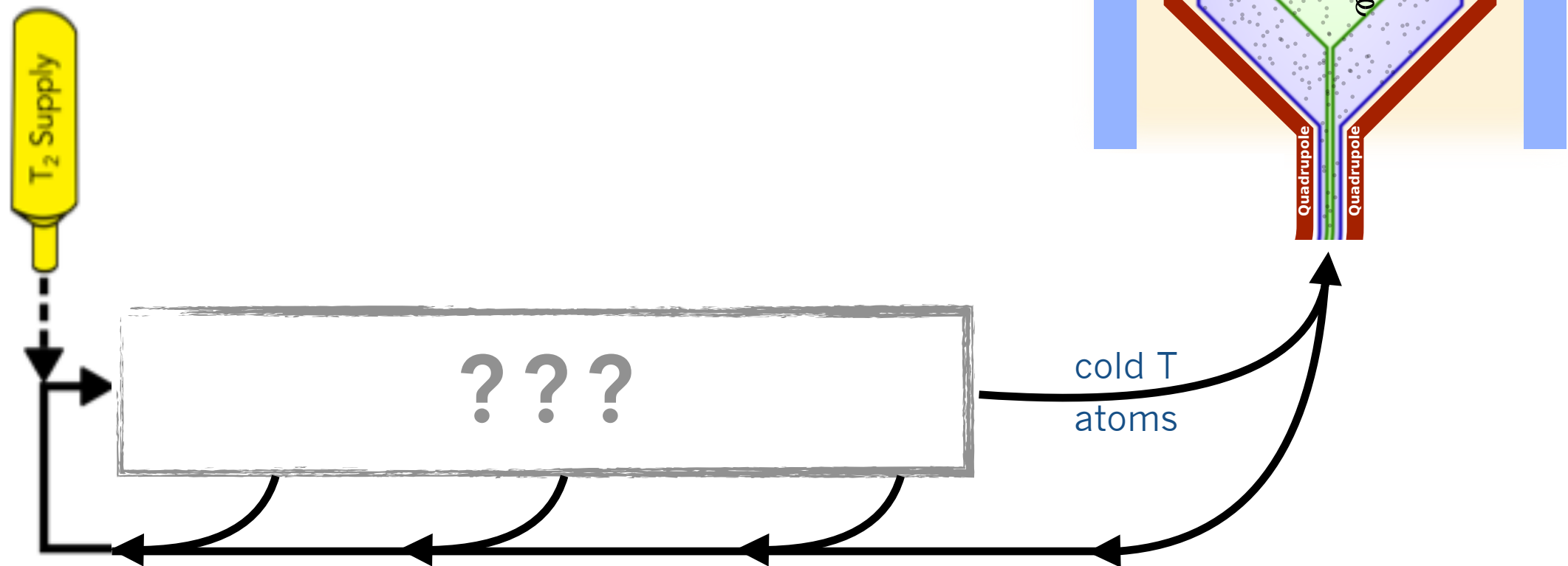
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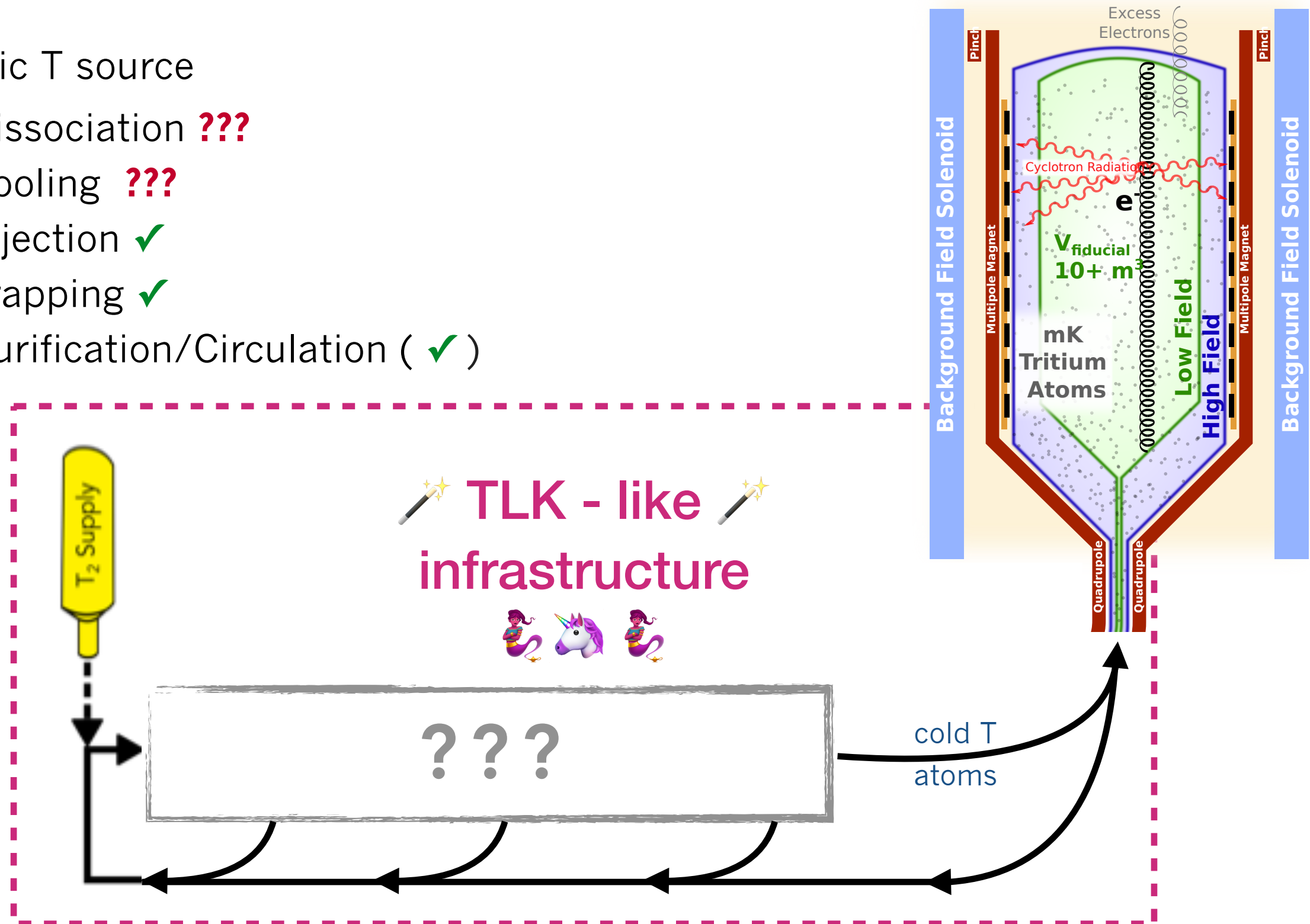
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# Project 8: Designs concepts

## Atomic T source

- Dissociation ???
- Cooling ???
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- Trapping ✓
- Purification/Circulation ( ✓ )



# T<sub>2</sub> dissociation schemes

Microwave dissociation @ 151MHz

- well tested for hydrogen
- chemical reaction with glass
  - ▶ not feasible with tritium!

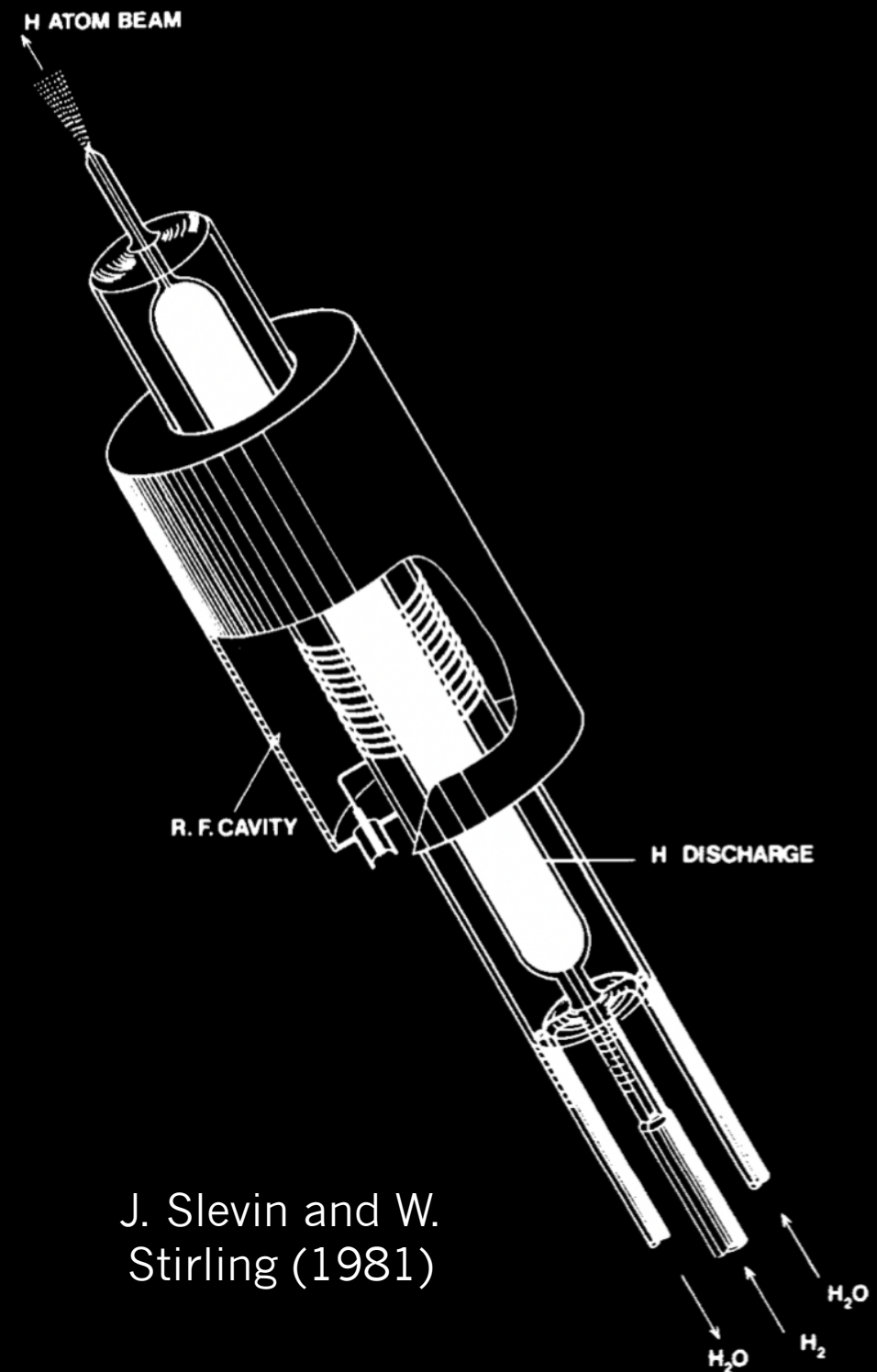
Laser dissociation

- dissociation energy 4.52eV
  - ▶ wavelength < 274nm
- required laser power ~ kW!

Coulomb explosion

- difficult to re-neutralize

...

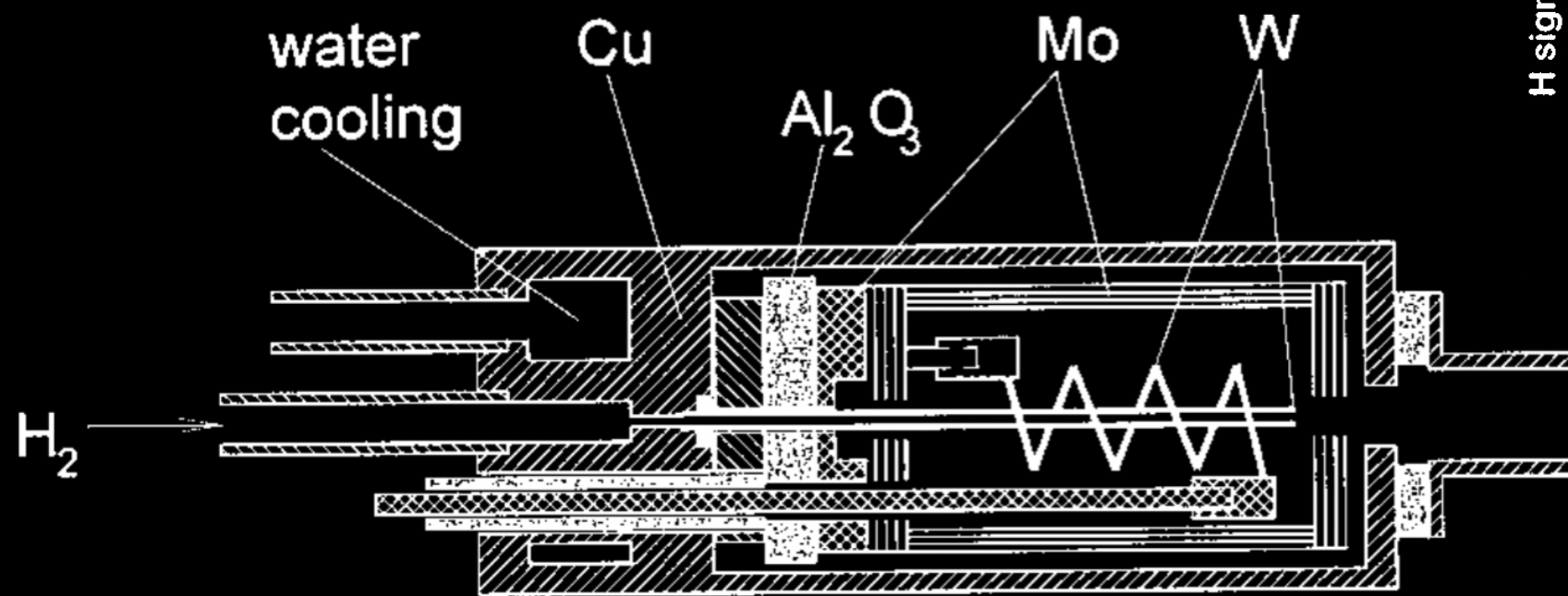


J. Slevin and W. Stirling (1981)

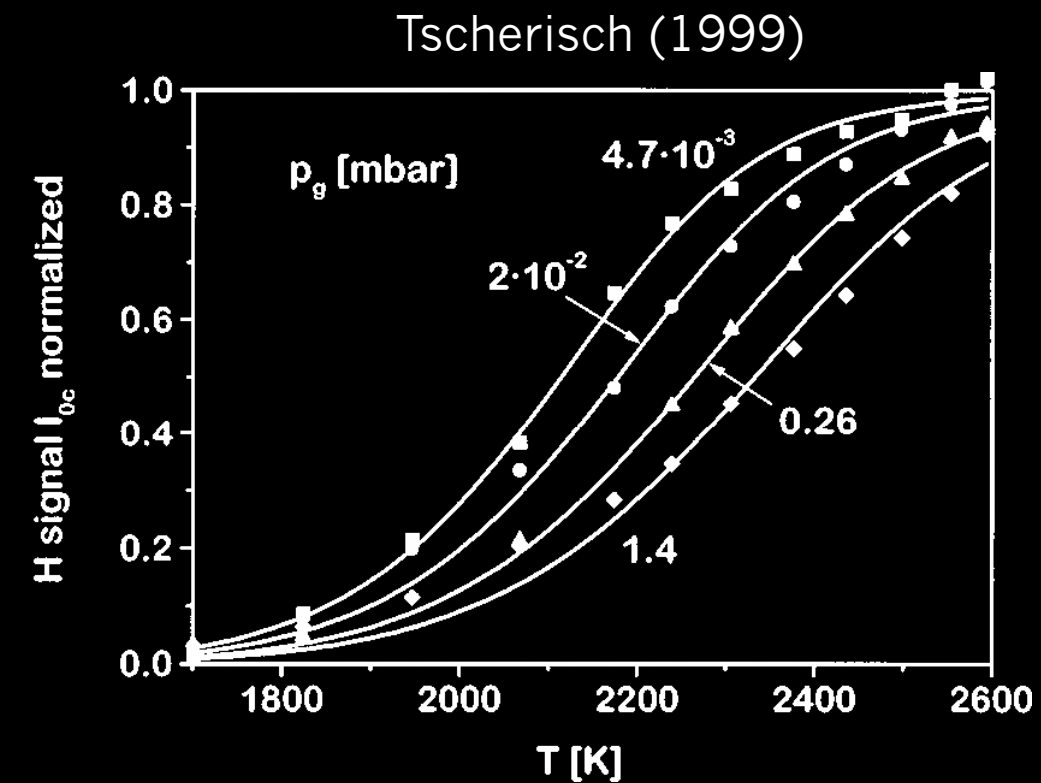
# Thermal Dissociator

Hot tungsten tube heated to 2500K

- radiatively or
- by electron bombardment
  - ▶ commercial devices available



K.G. Tschersich and V. von Bonin (1998)



HABS hydrogen cracker



# Cooling atomic hydrogen

## Hydrogen recombination

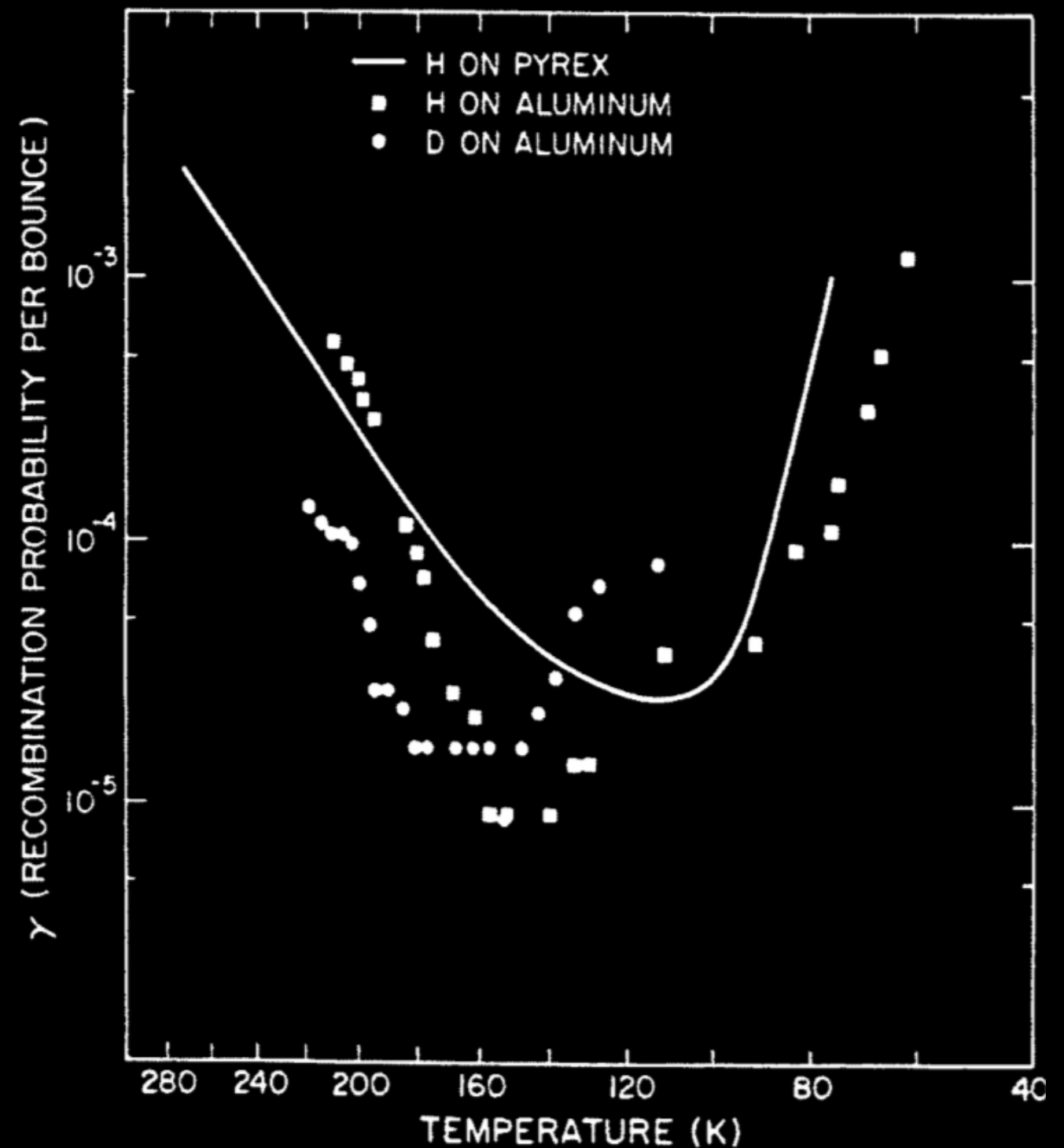
- three-body gas interactions
  - ▶ small rate
- wall interactions
  - long *sticking* time
  - ▶ dominates recombination

## Probability depends on

- temperature
- material
- hydrogen isotope

## Superfluid He containment

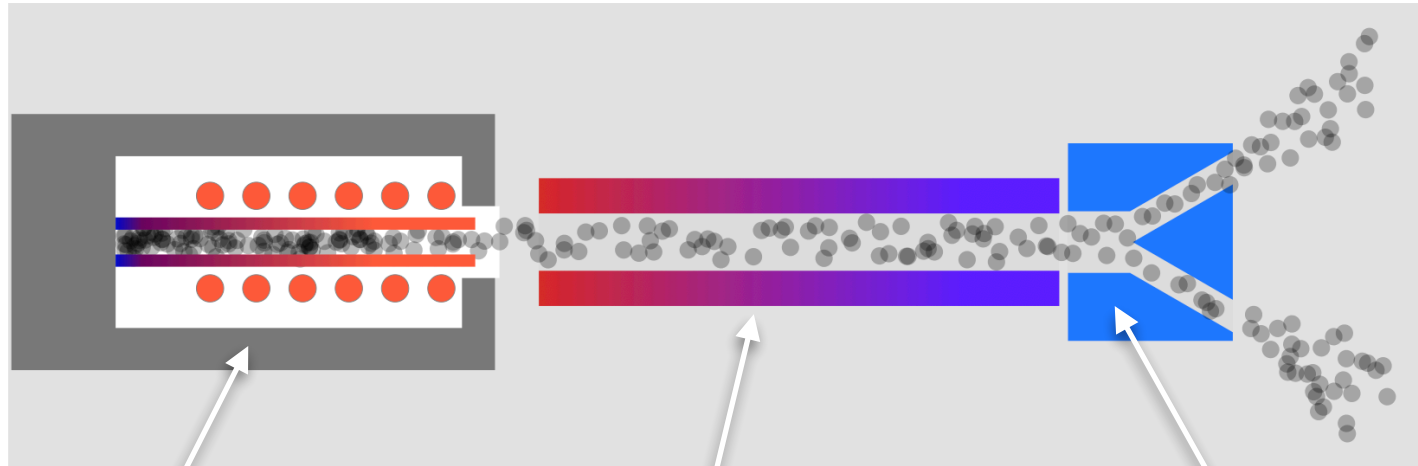
- ▶ does not work for tritium!



D. A. Knapp et al., AIP conference proceedings (1984)  
Wood and H. Wise, J. Chem. Phys. 66, 1049, (1962)

# Cooling tritium atoms

Cracker 2500 K  
Accommodator 160 K  
Nozzle ~10 K



Cracker

→ purity vs. flow

Accommodator  
(liquid nitrogen)

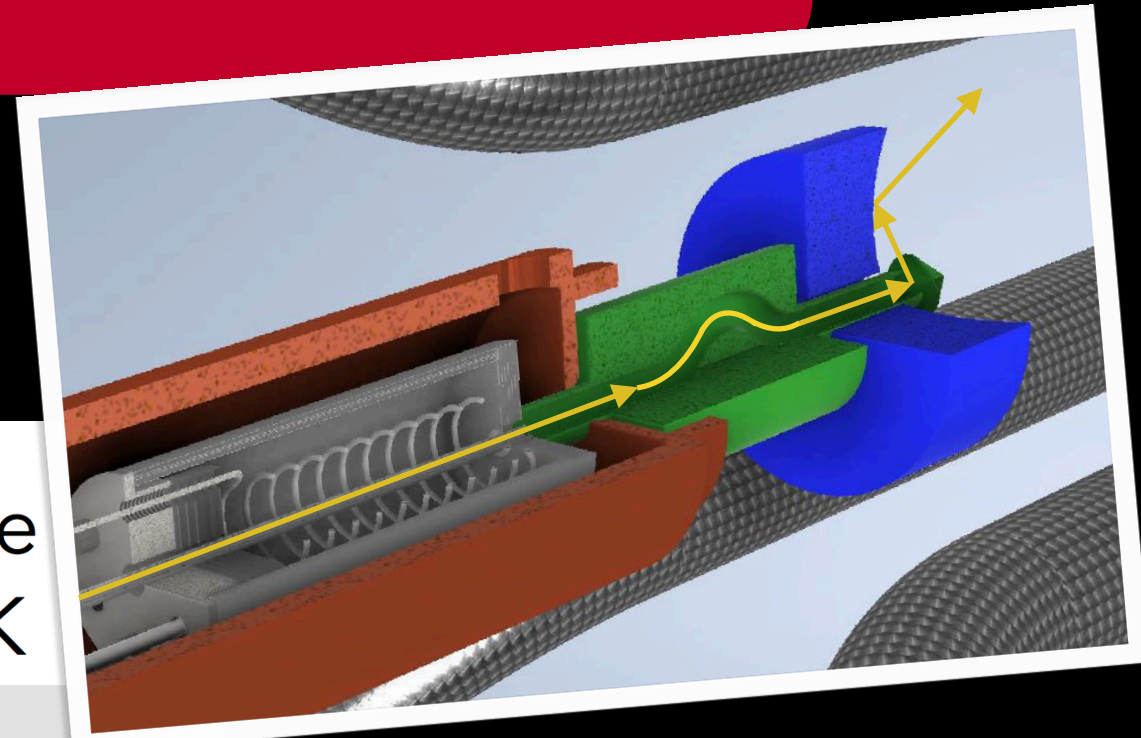
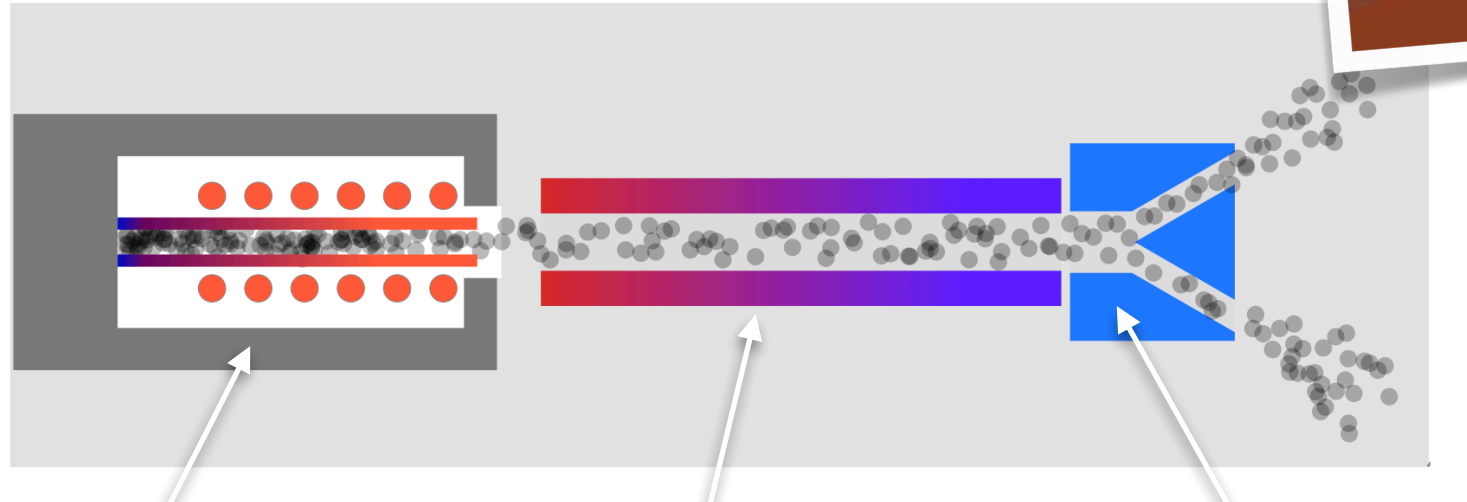
Final nozzle

- design for few bounces
- freeze-out 30K  
→ periodic purging



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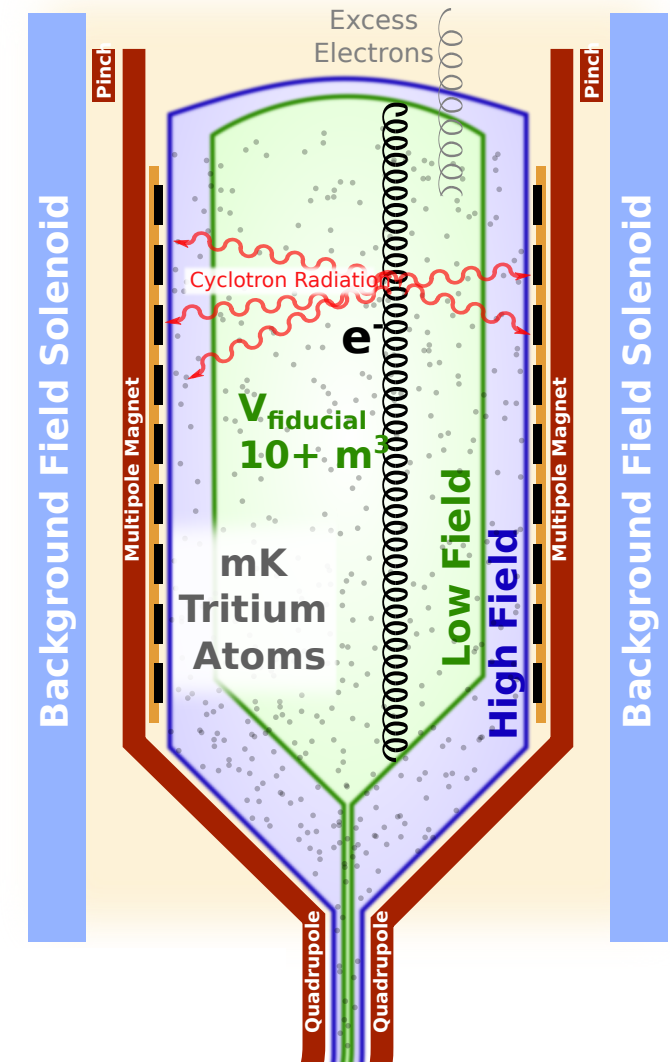
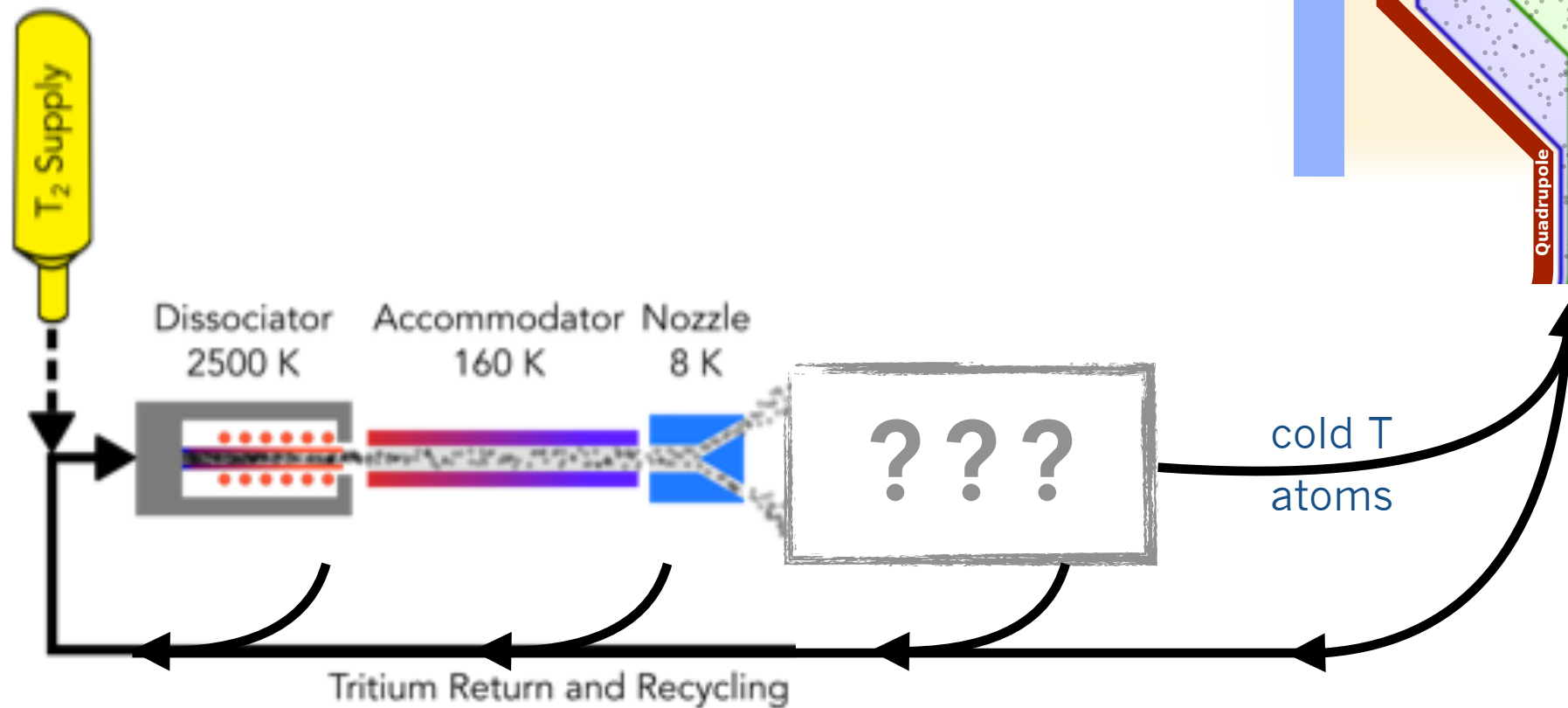
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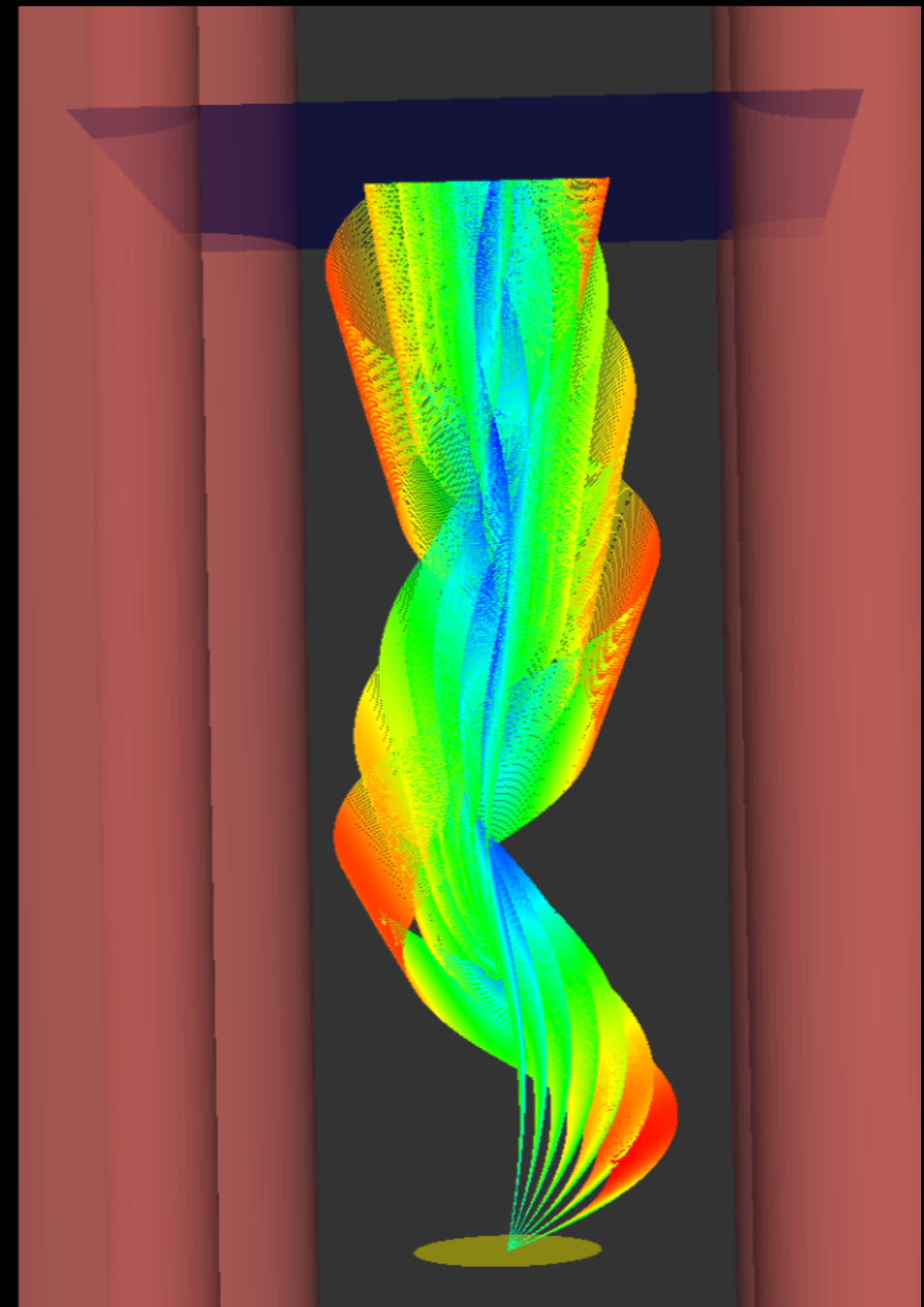
- Dissociation ✓
- Cooling ???
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# Velocity and State Selectors

Only atomic T guided magnetically

- (bend) quadrupole with skimmers



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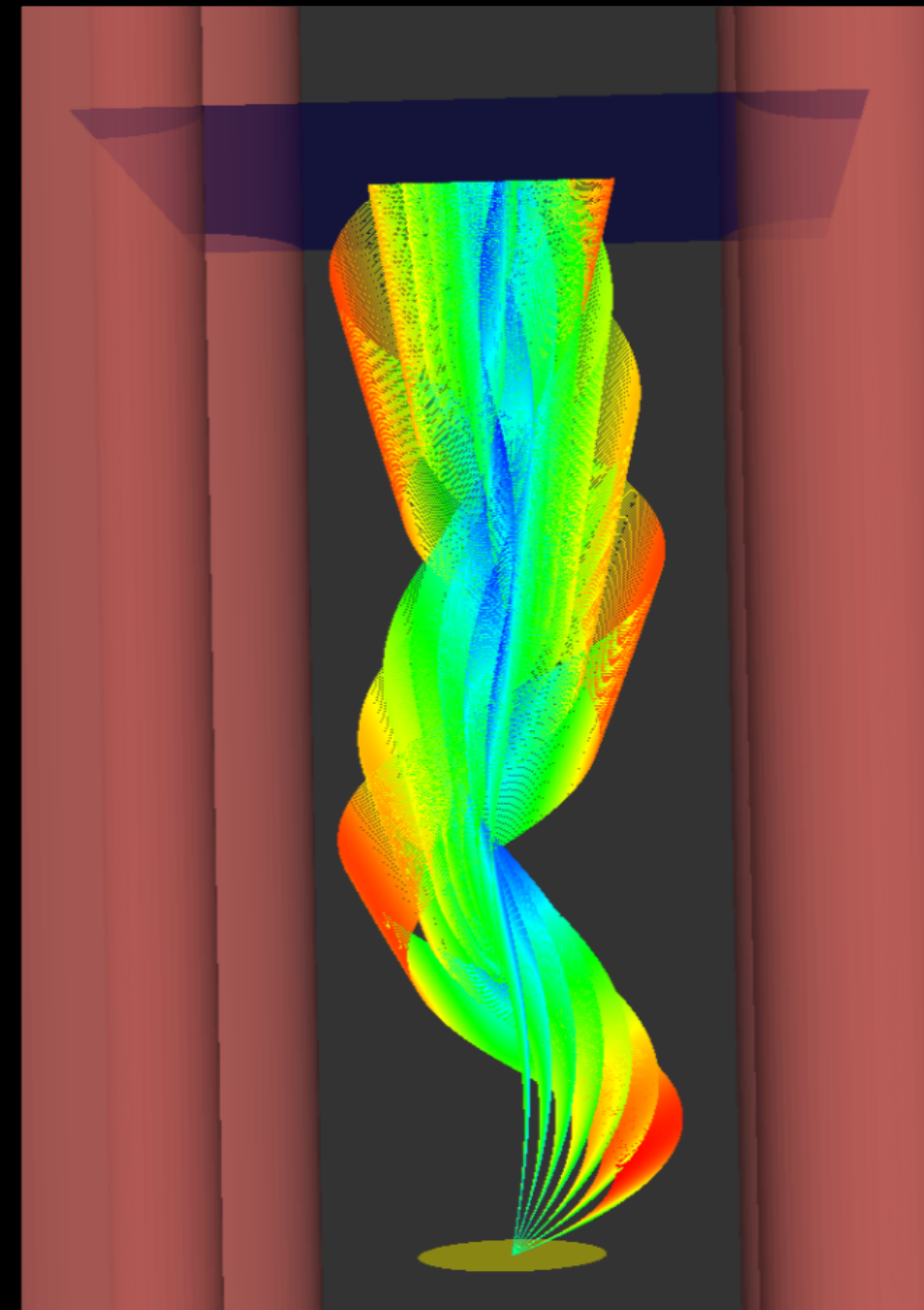
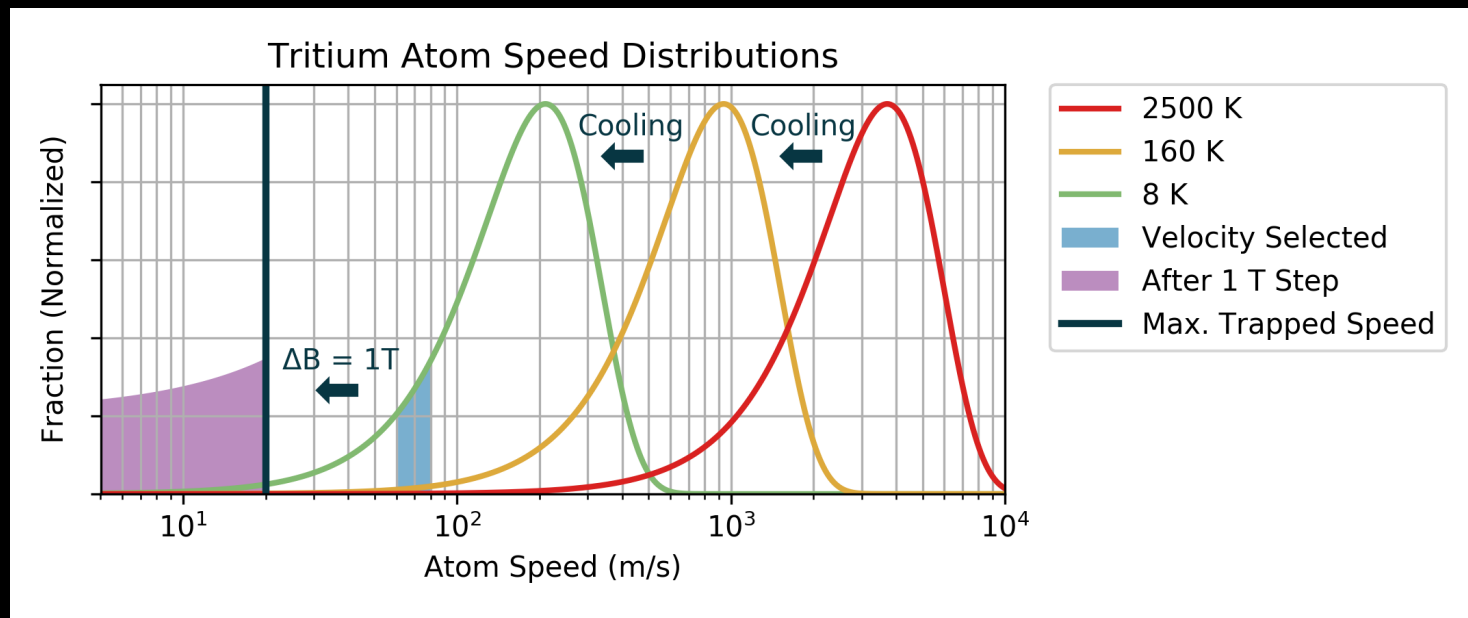
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Tune acceptance for

- $T_{\text{out}} = \mathcal{O}(50\text{mK})$
- $T_2$  contamination  $< 10^{-5}$ 
  - ▶ efficiency  $\varepsilon_{\text{cold}} \sim 25\%-100\%$





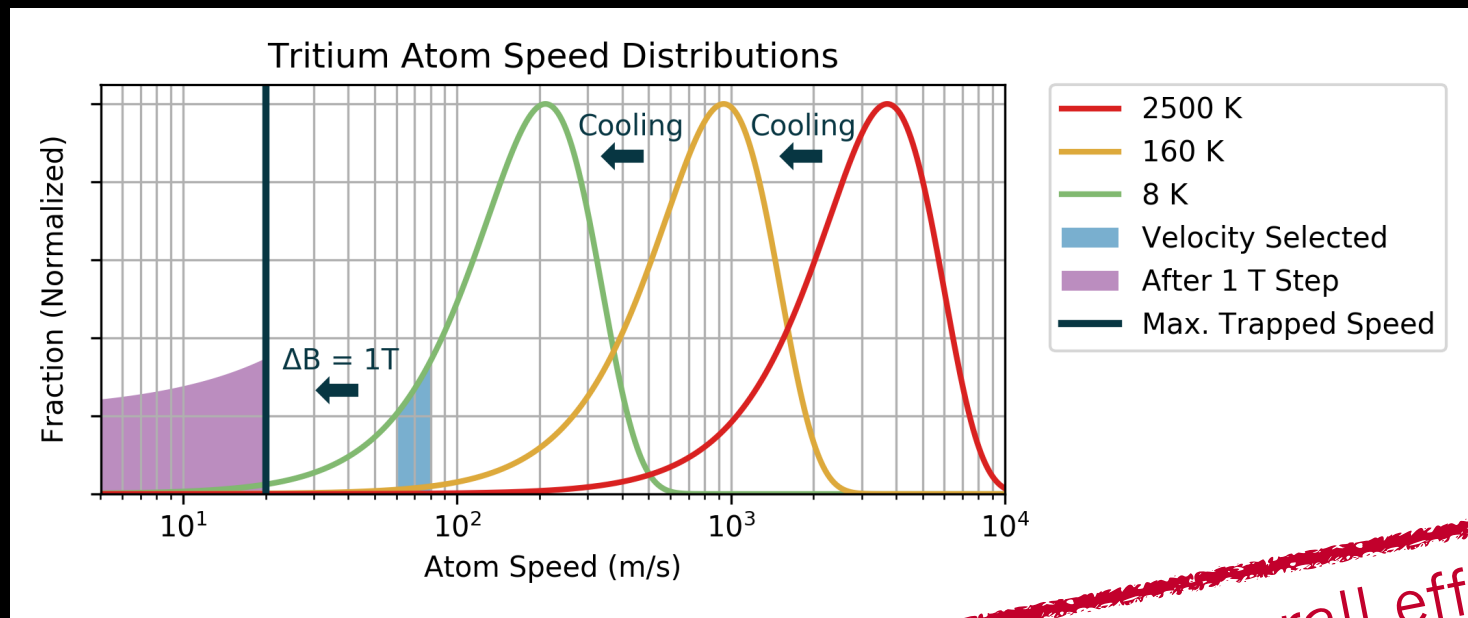
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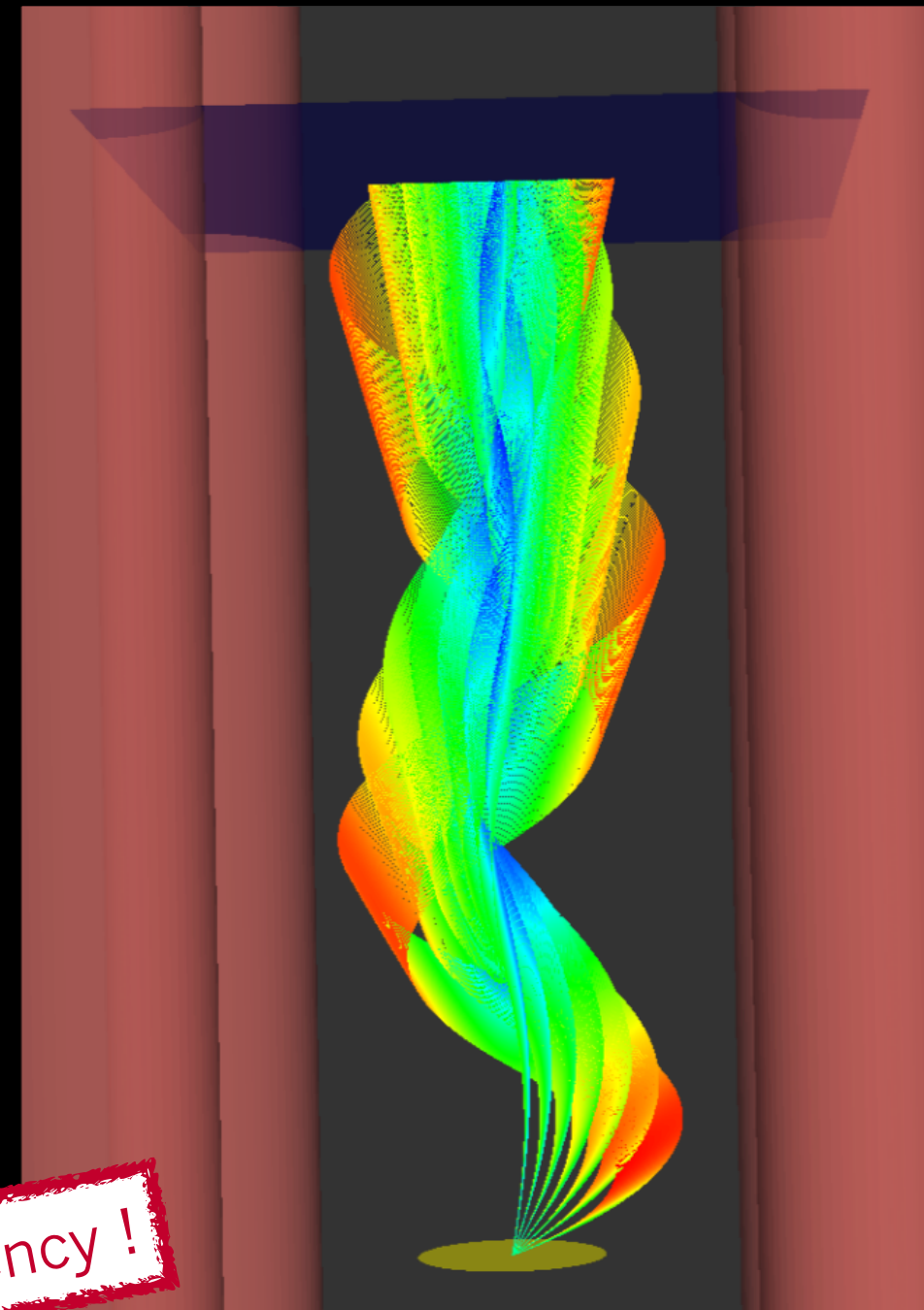
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→ but poor overall efficiency !



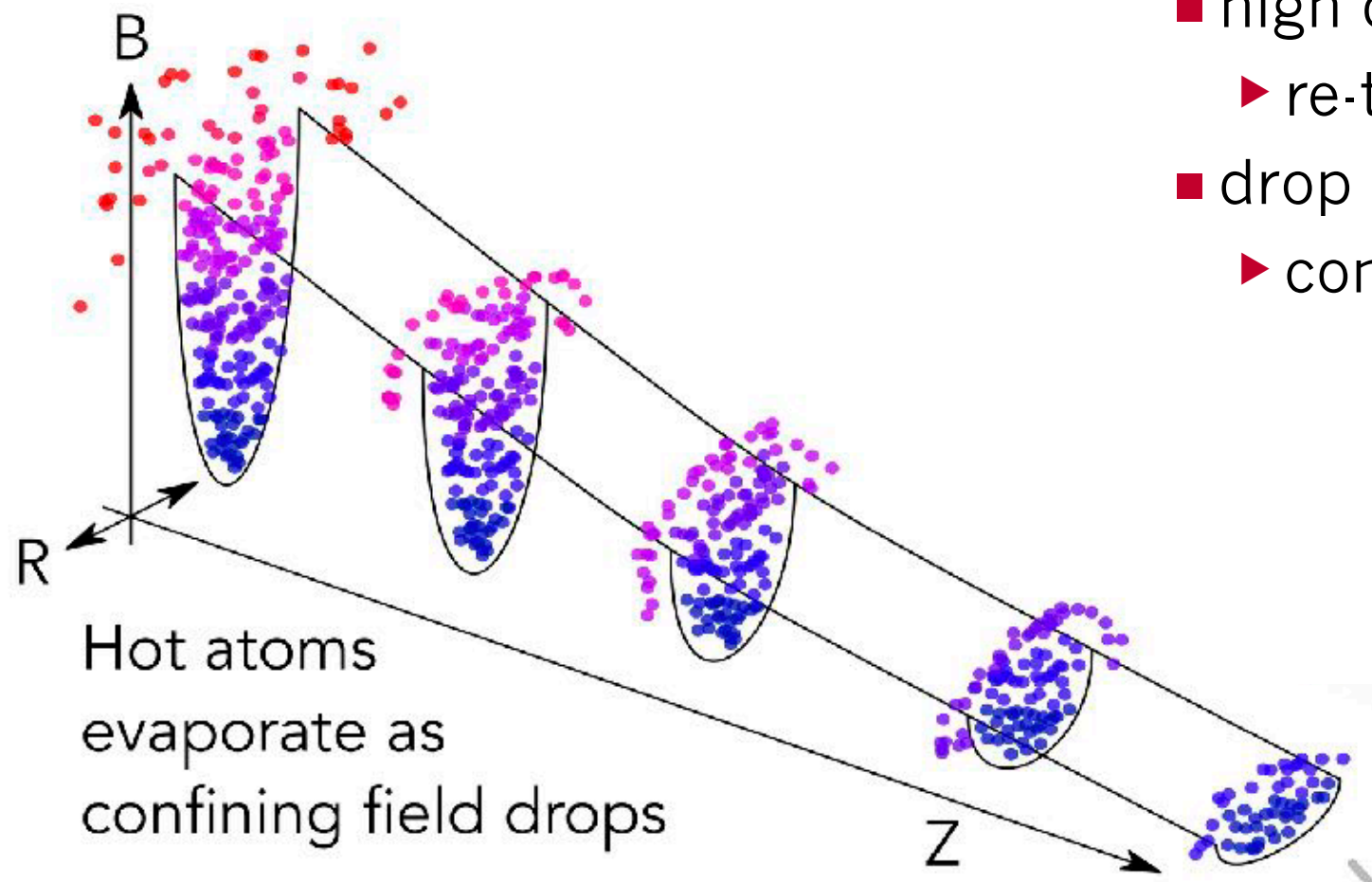
# Evaporative cooling

## Evaporative cooling of an atomic beam

E. Mandonnet<sup>1</sup>, A. Minguzzi<sup>2</sup>, R. Dum<sup>1</sup>, I. Carusotto<sup>2</sup>, Y. Castin<sup>1</sup>, and J. Dalibard<sup>1,a</sup>

<sup>1</sup> Laboratoire Kastler Brossel, 24 rue Lhomond, 75005 Paris, France

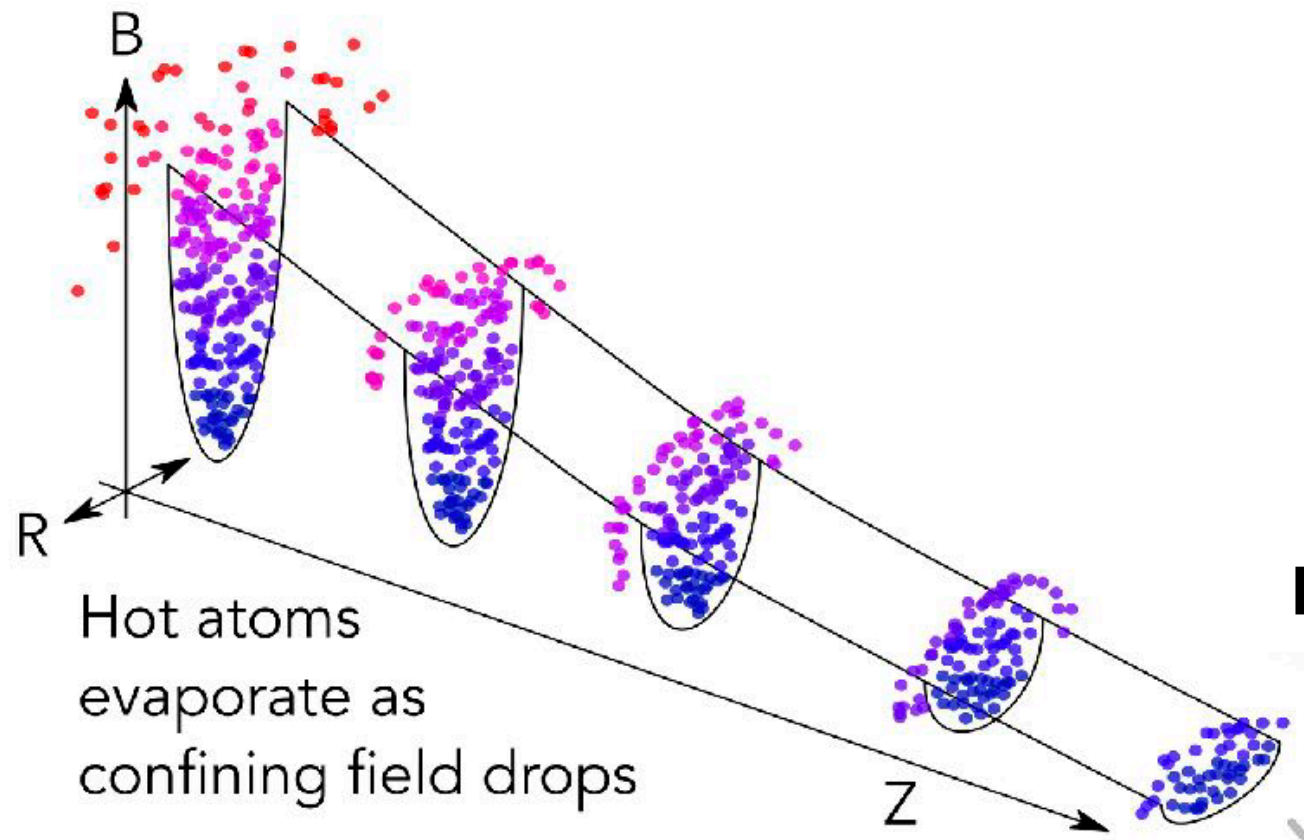
<sup>2</sup> INFN, Scuola Normale Superiore, Piazza dei Cavalieri 7, 56126 Pisa, Italy



## Basic idea

- magnetic wall
  - ▶ loose *hot* atoms (high  $p_{\perp}$ )
- high density
  - ▶ re-thermalization
- drop magnetic field along beam
  - ▶ continuous cooling

# Evaporative cooling: the math



→ assumes equilibrium !

## Evaporation efficiency $\eta_{ev}$

$$\mu B(z) = \eta_{ev} k_B T(z)$$

- large  $\eta_{ev}$ 
  - ▶ slow cooling, high efficiency
- small  $\eta_{ev}$ 
  - ▶ fast cooling, low efficiency

## Number density

$$n \propto T^{1/\gamma} \text{ with } \gamma = \frac{2}{3} \left( \eta_{ev} + \frac{1}{2} - C \right)$$

- optimal value ( $C = 2$ ,  $\eta_{ev} = 4.5$ )
  - ▶  $\gamma = 2 \rightarrow n \propto T^{\frac{1}{2}}$



# Evaporative cooling: the hard math

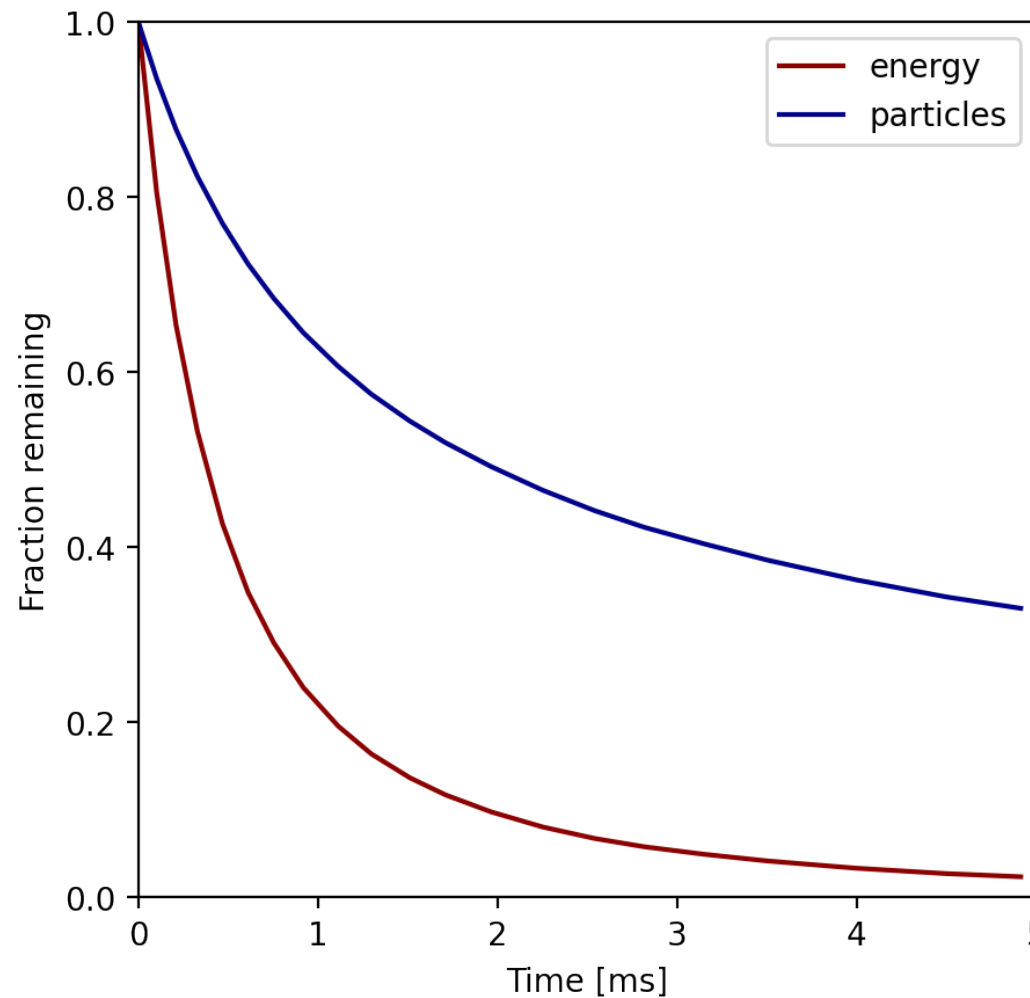
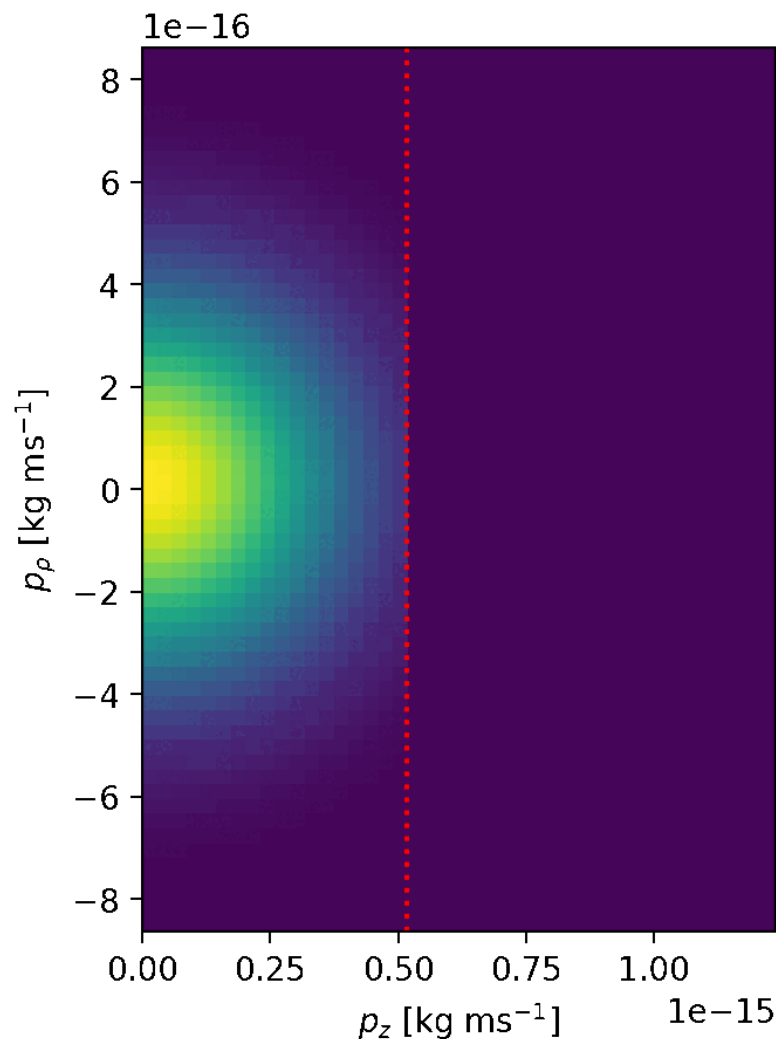
Boltzman transport equation

$$\left( \frac{p}{m} \nabla_r - \nabla_r U \cdot \nabla_p + \frac{\partial}{\partial t} \right) f(r, p) = \frac{1}{\pi m^2} \int d^3 p_2 \int d^3 p_3 \int d^3 p_4 \frac{d\sigma}{d\Omega}(p_1, p_2, p_3, p_4) (f(r, p_3) f(r, p_4) - f(r, p) f(r, p_2))$$

**kinetic terms**

**collision integral**

► solve numerically



- Losing energy faster than particles means cooling
- Radial cooling also cools longitudinal motion since collisions redistribute momentum

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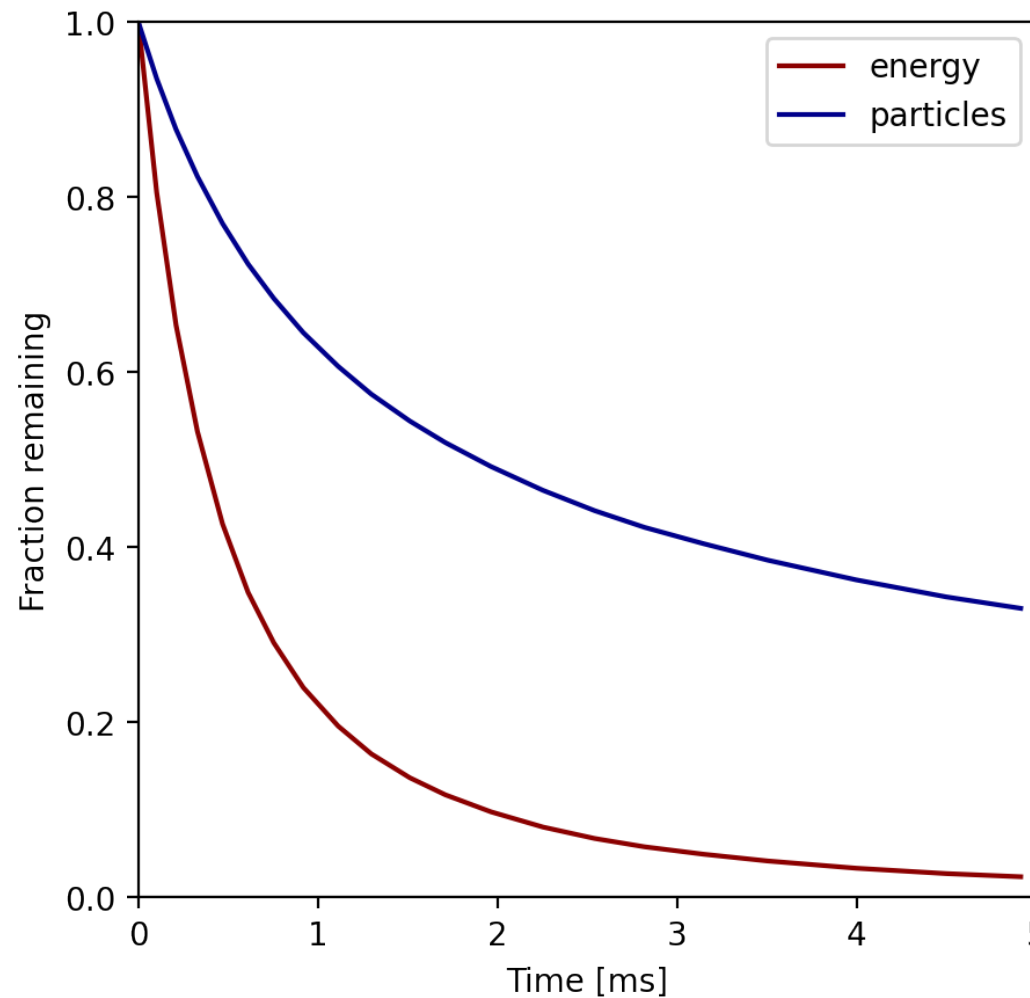
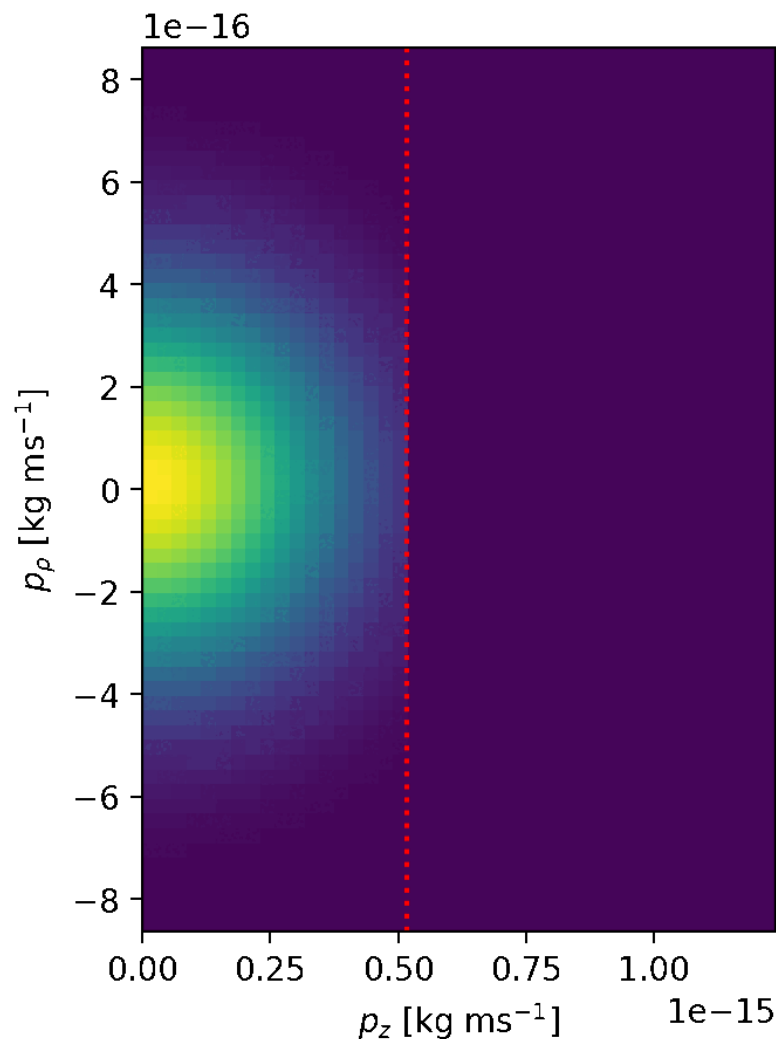
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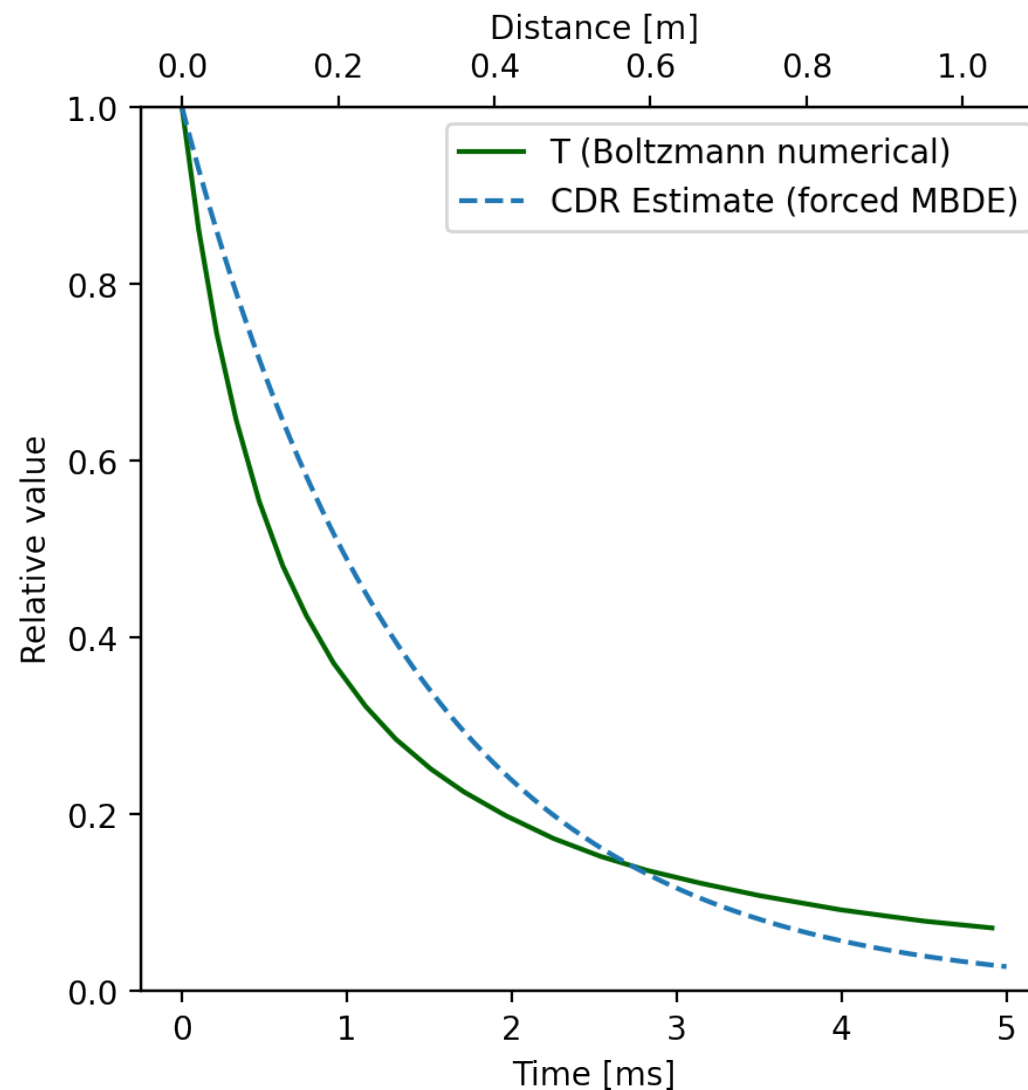
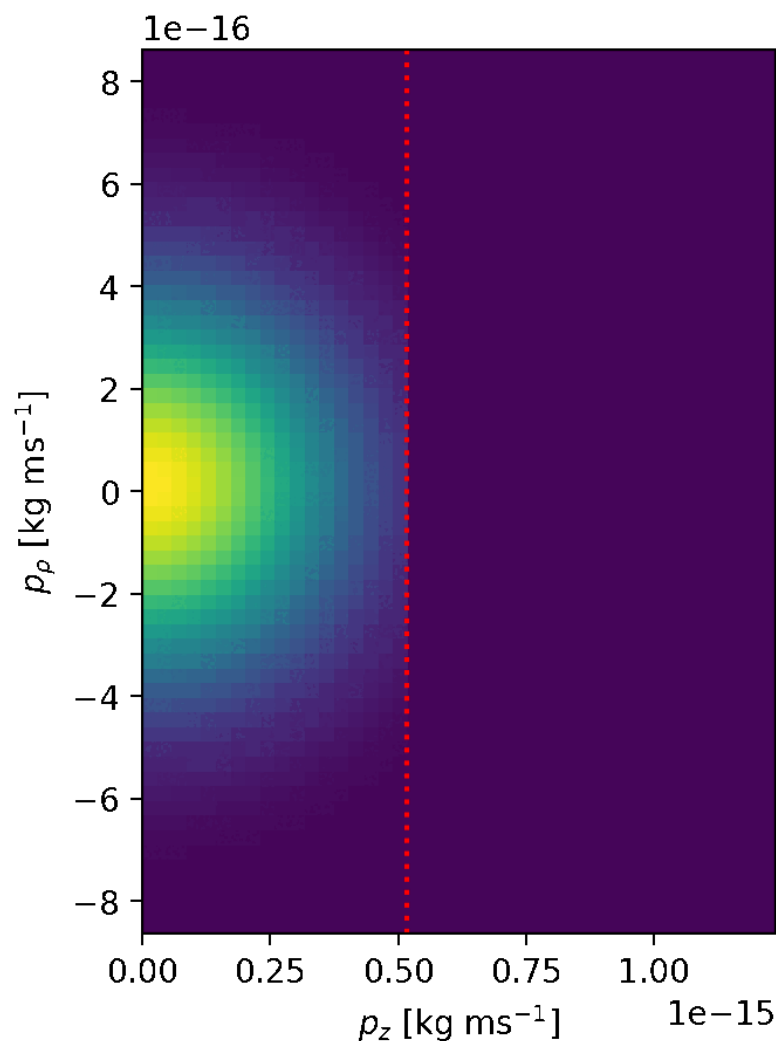
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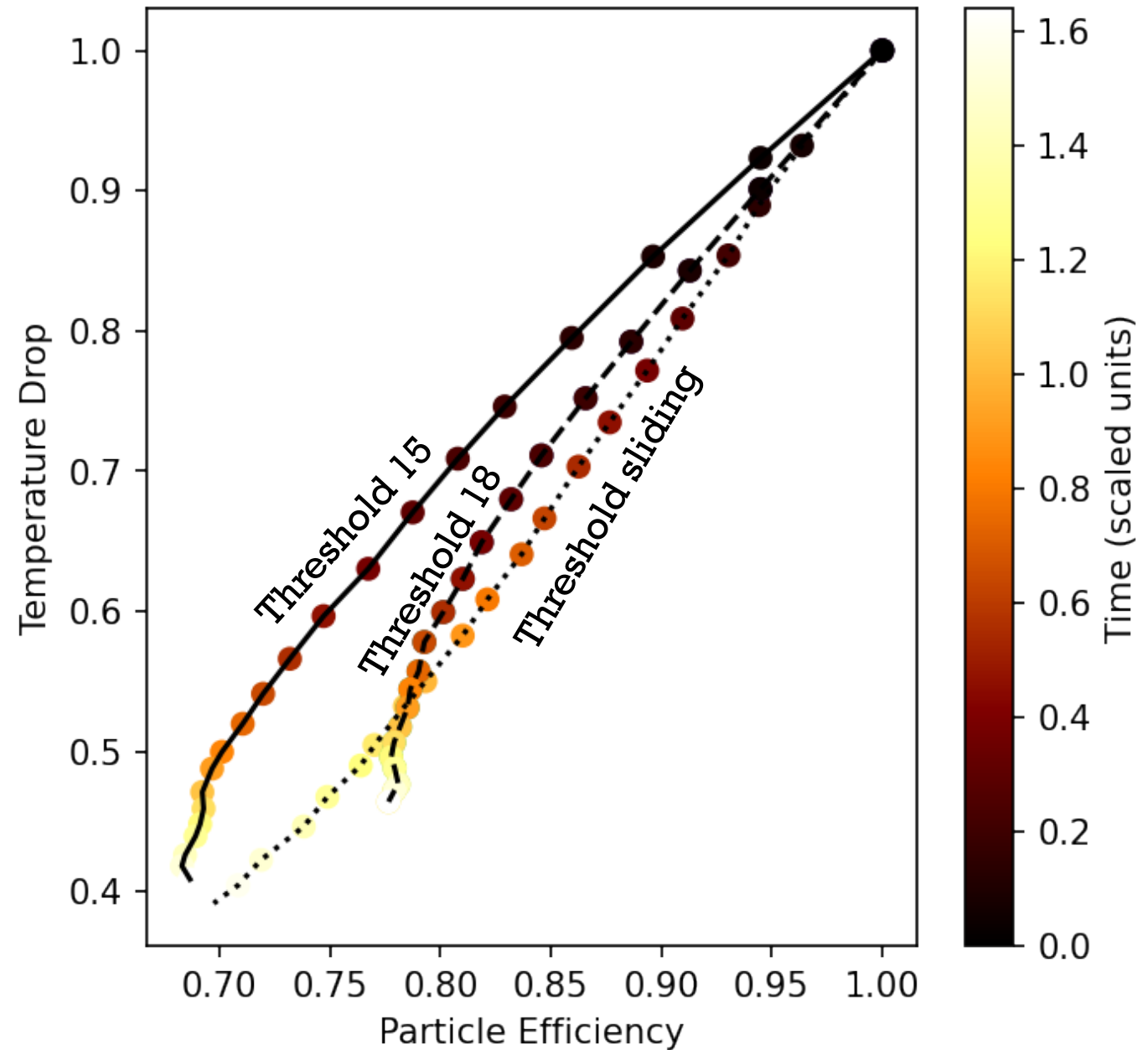
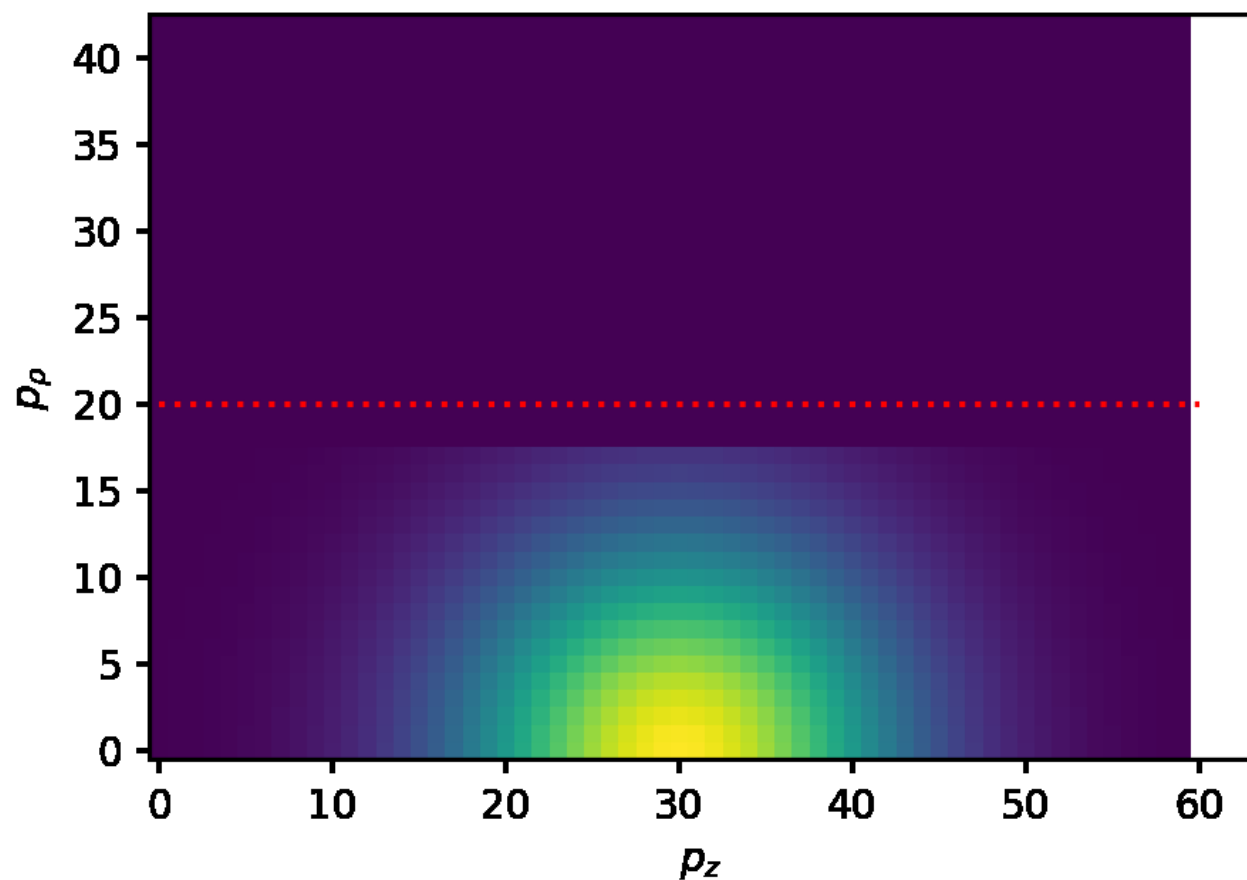
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# Evaporative cooling: sliding threshold

## Original idea

- decrease wall height with time
- ▶ faster cooling (initially)

→ large phase space left to explore

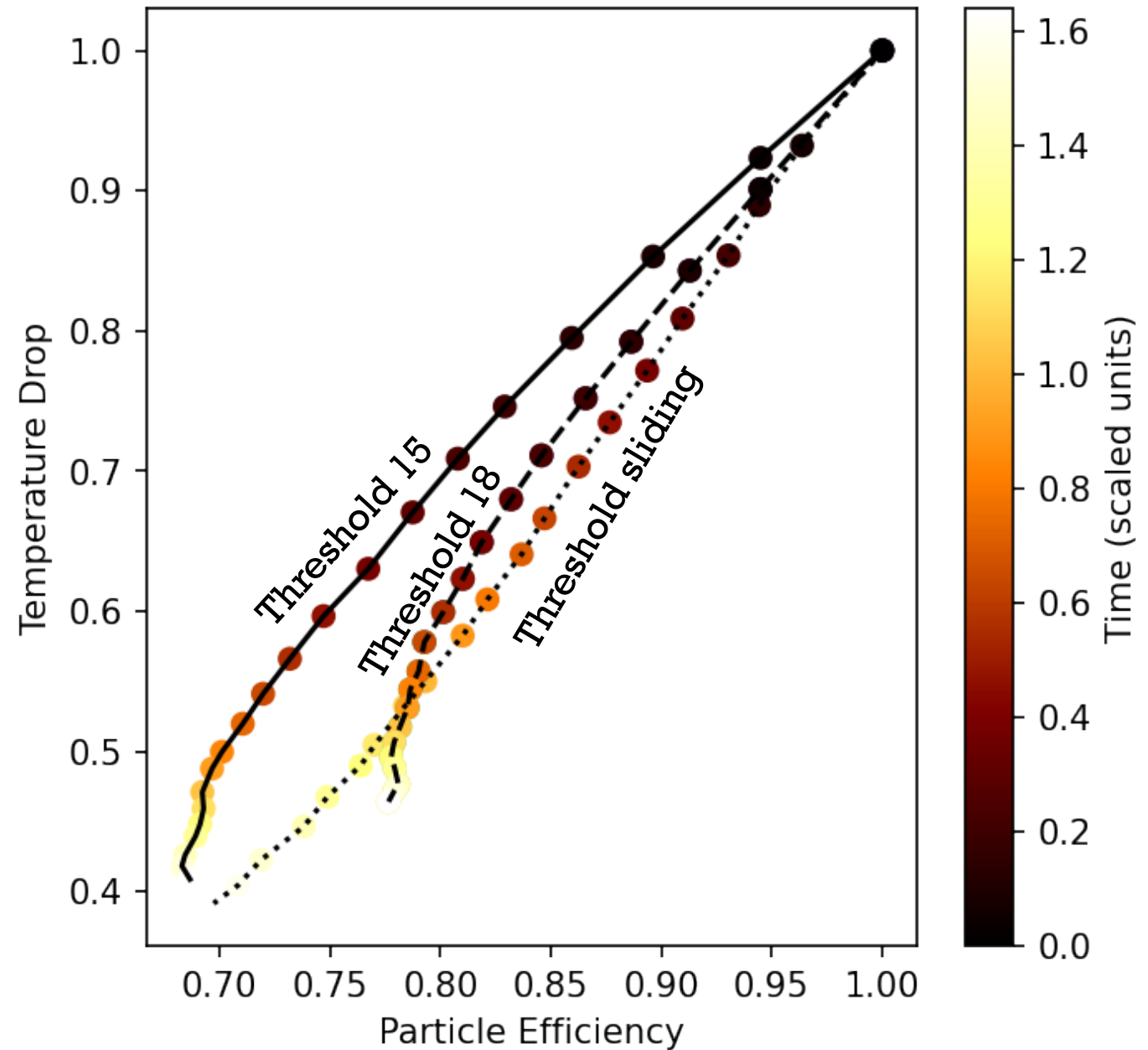
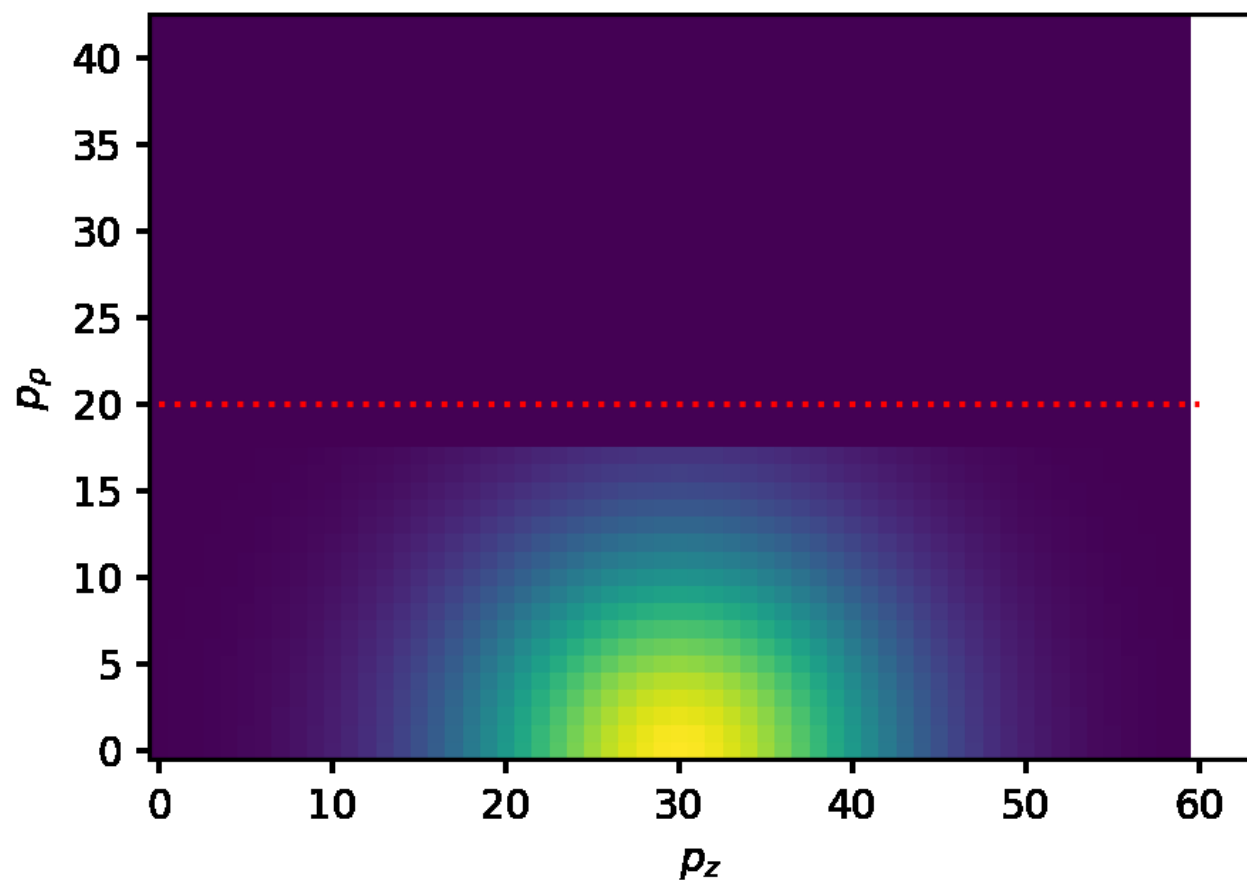


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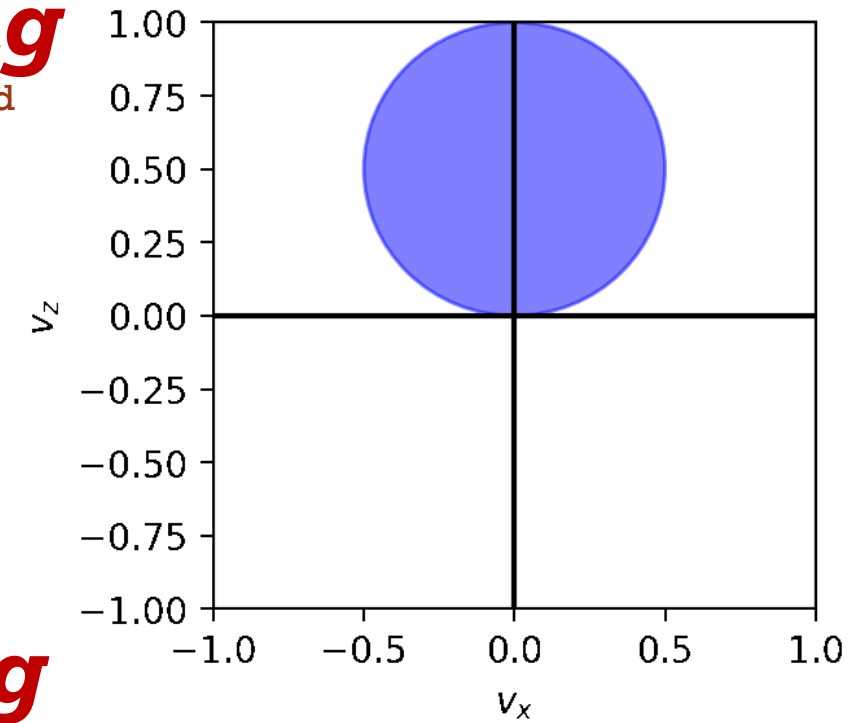
# Cooling vs slowing

Initial atomic *beam*

- net forward momentum
  - ▶ need **cooling** and **slowing**

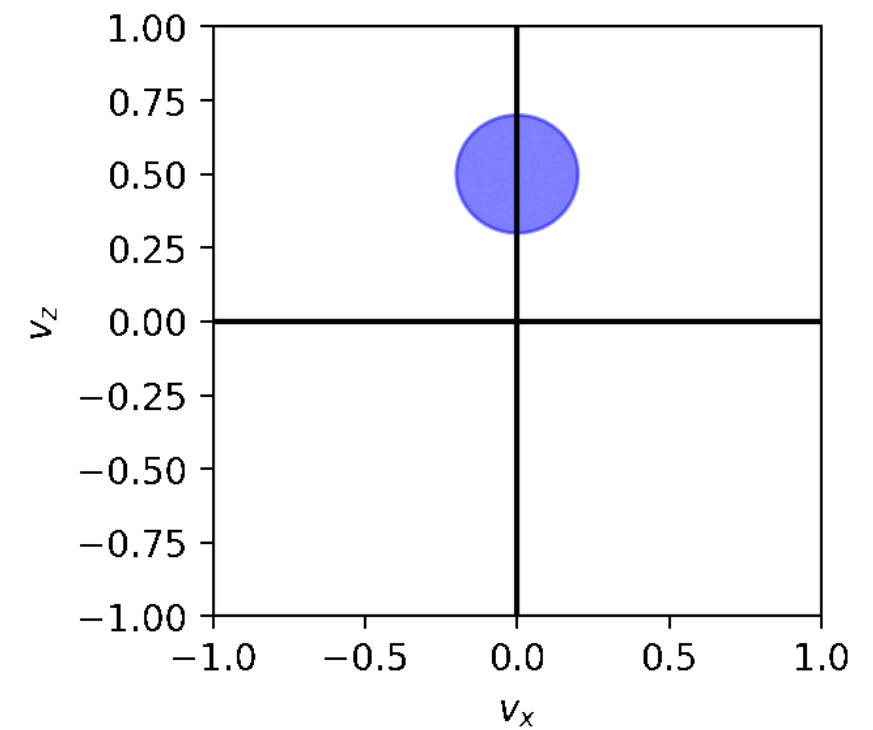
**Cooling**

Mean(p) conserved



**Slowing**

Mean(p) reduced





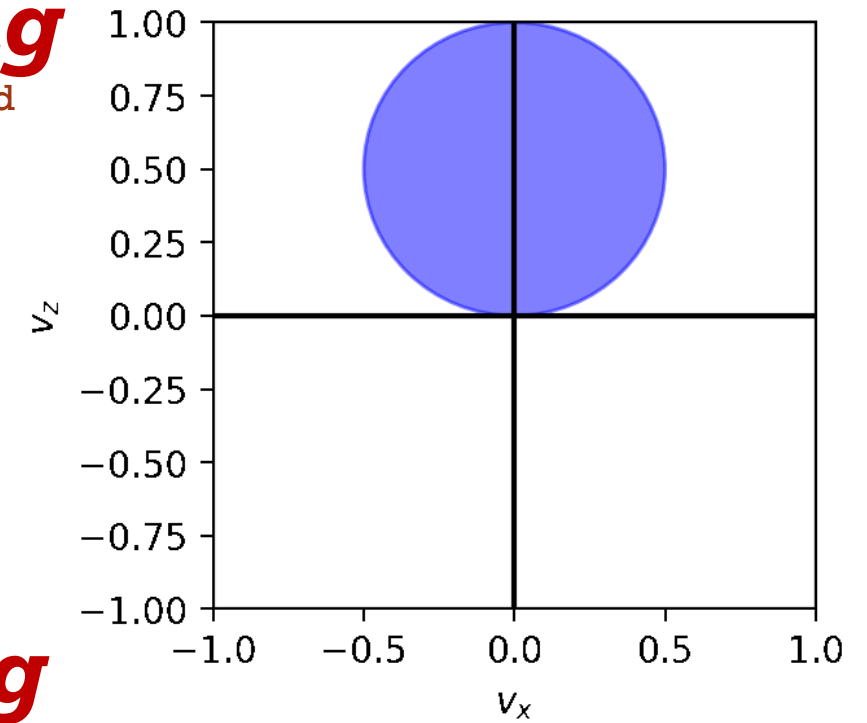
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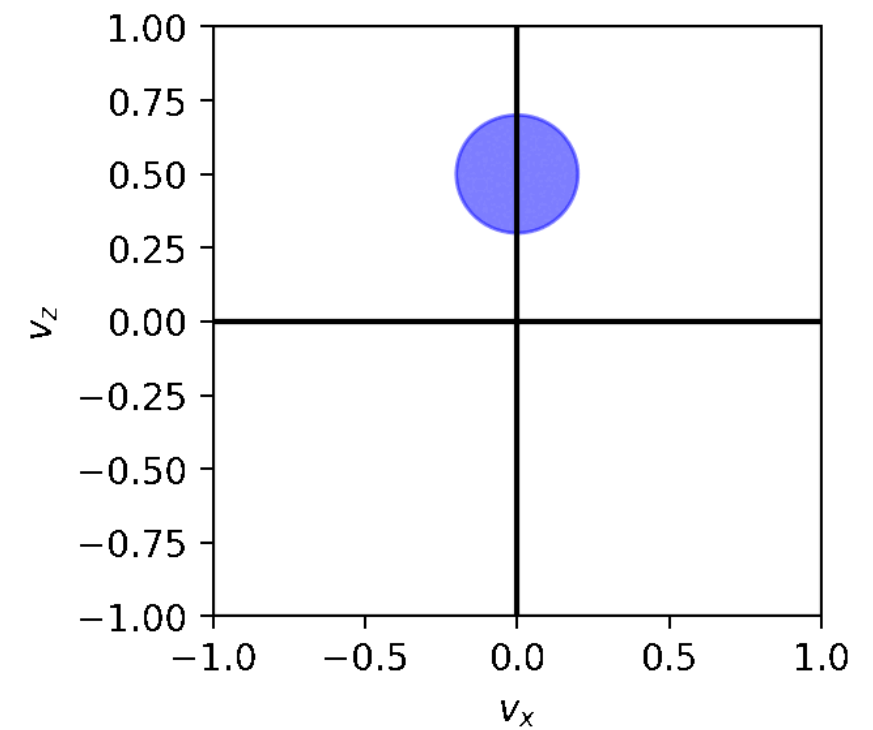
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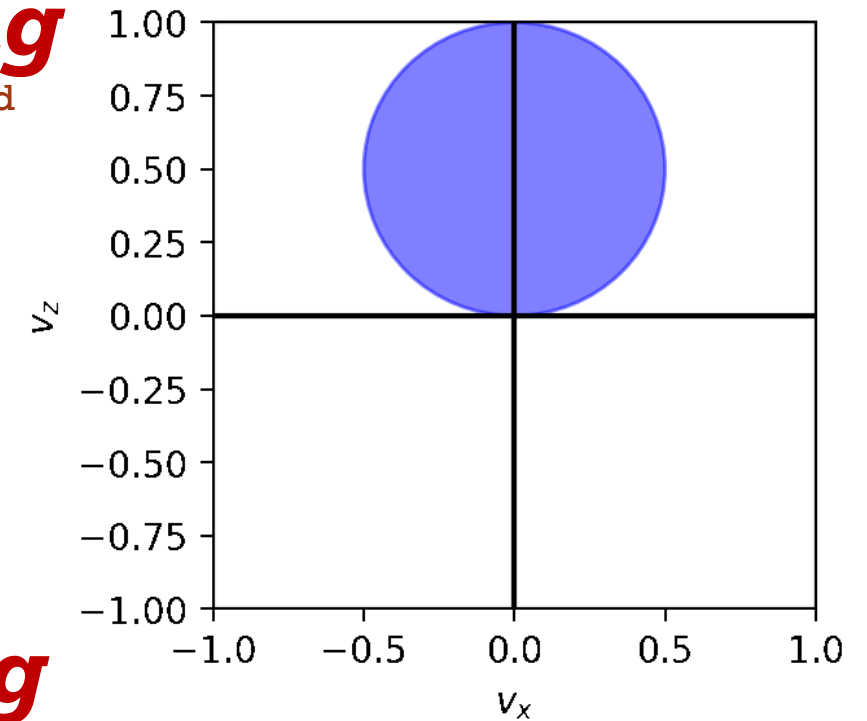
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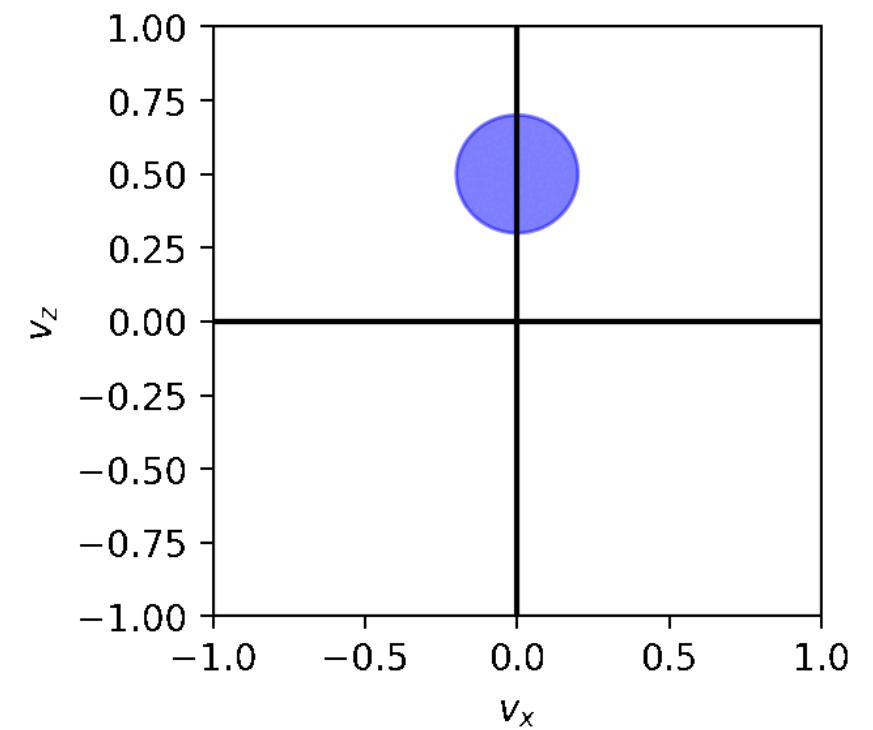
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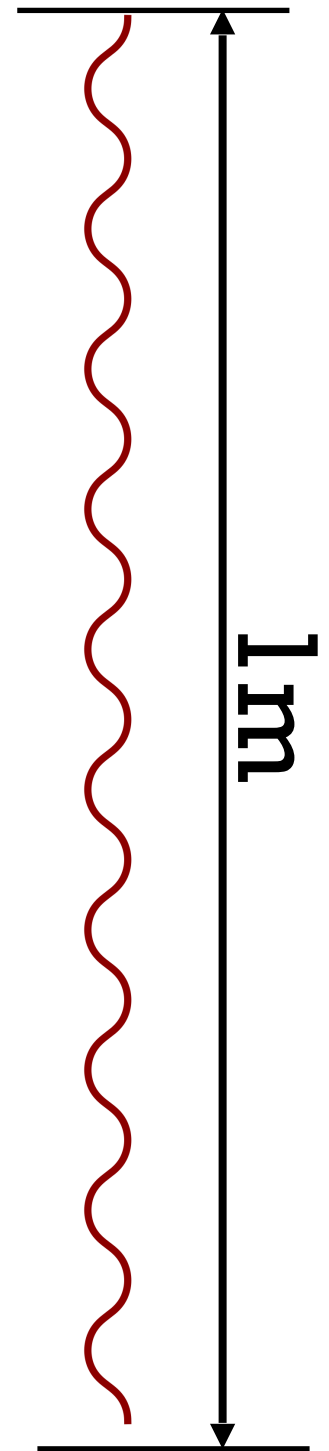
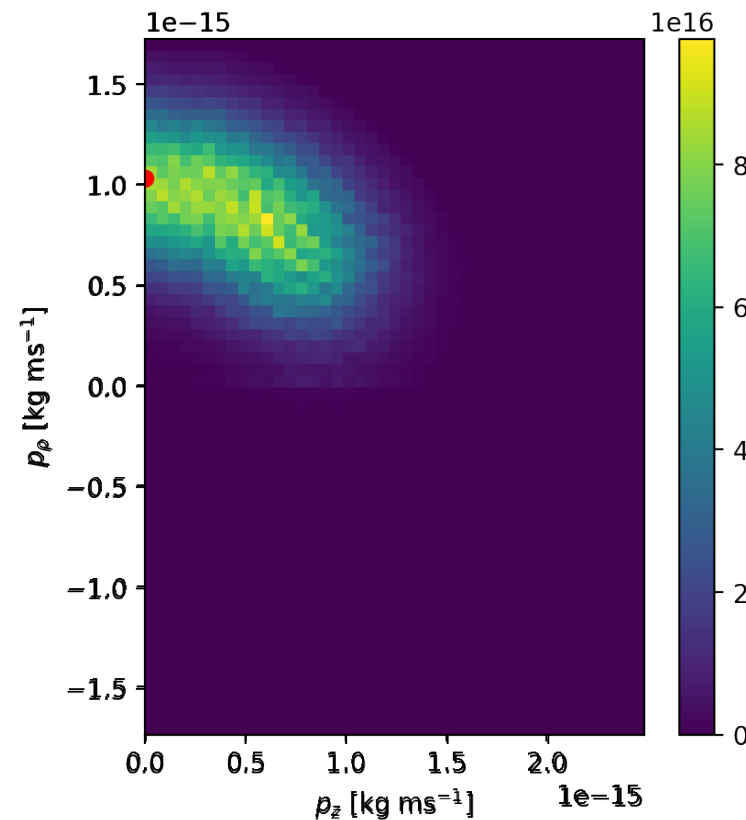
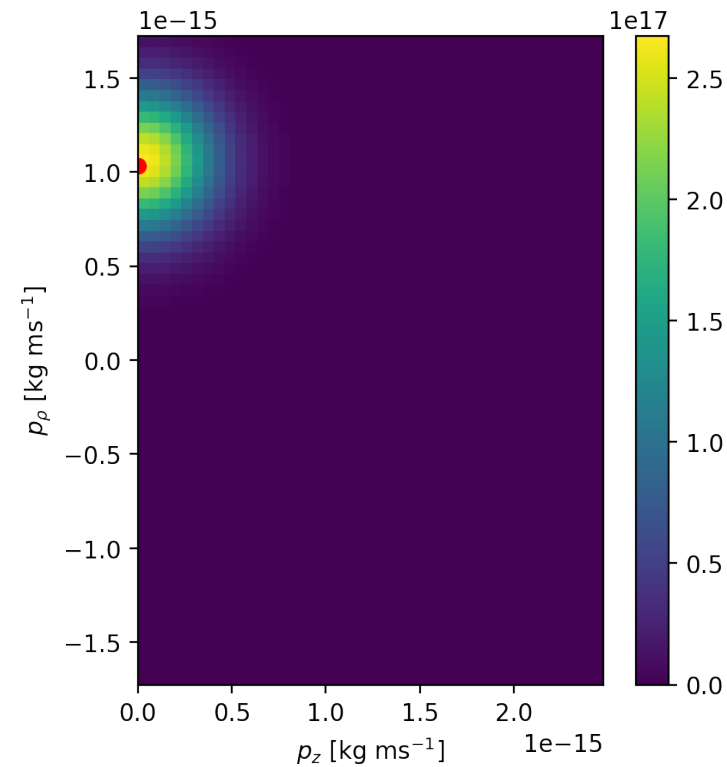
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The *wiggler*

- several wiggles within mean free path
  - ▶ transfer longitudinal to perpendicular momentum
- re-thermalization
  - ▶ slows down beam



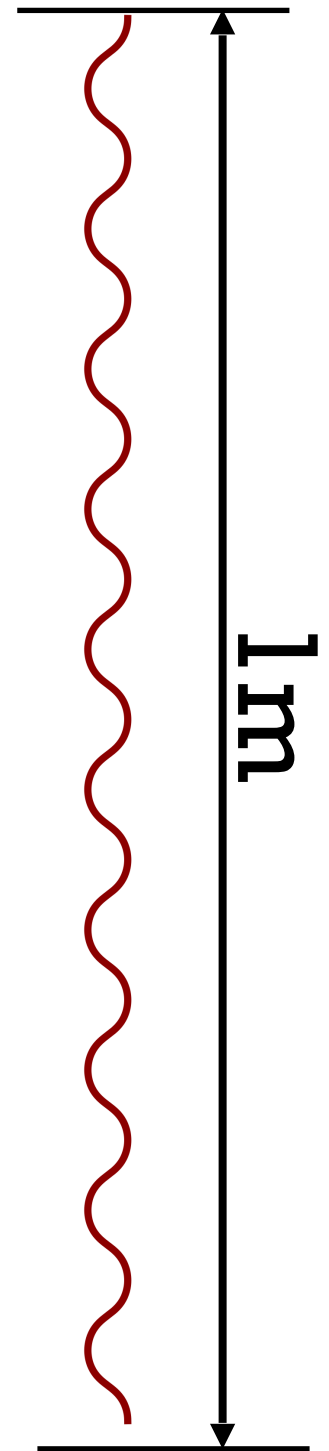
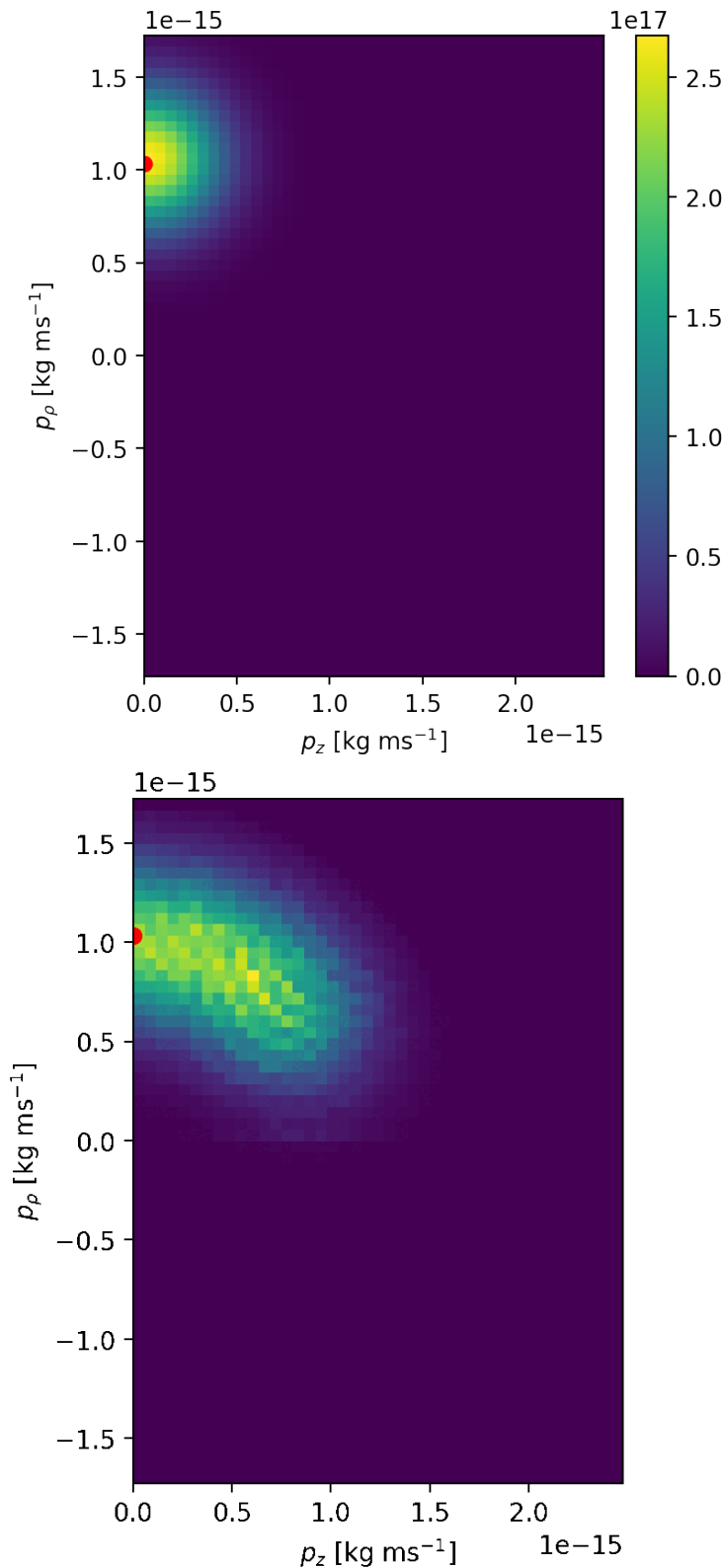
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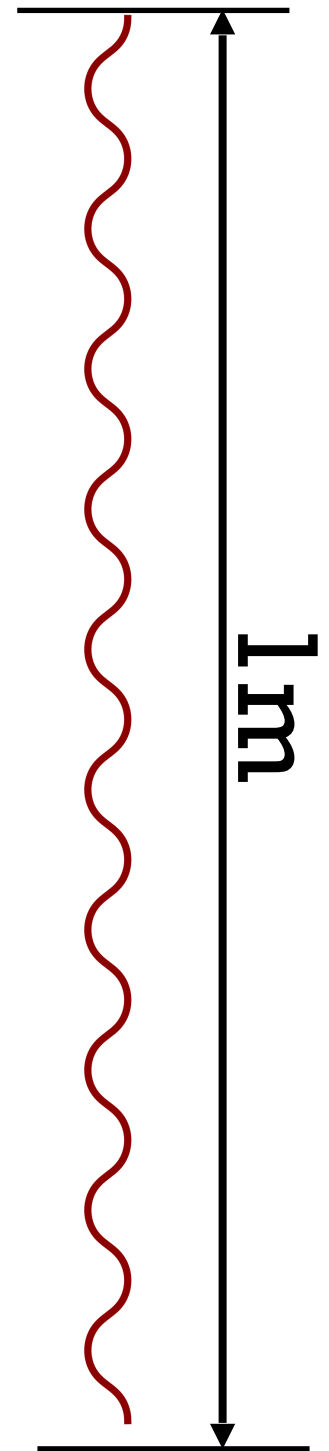
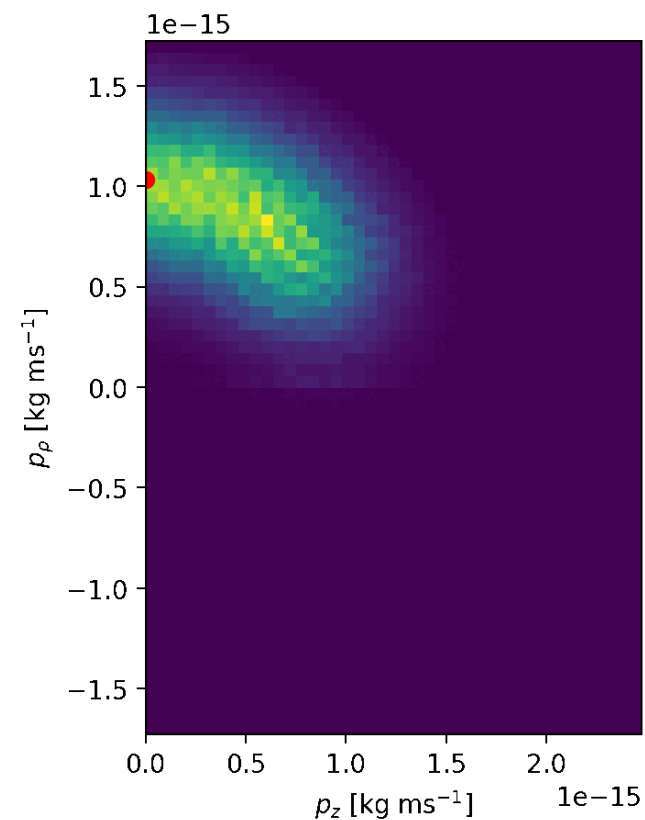
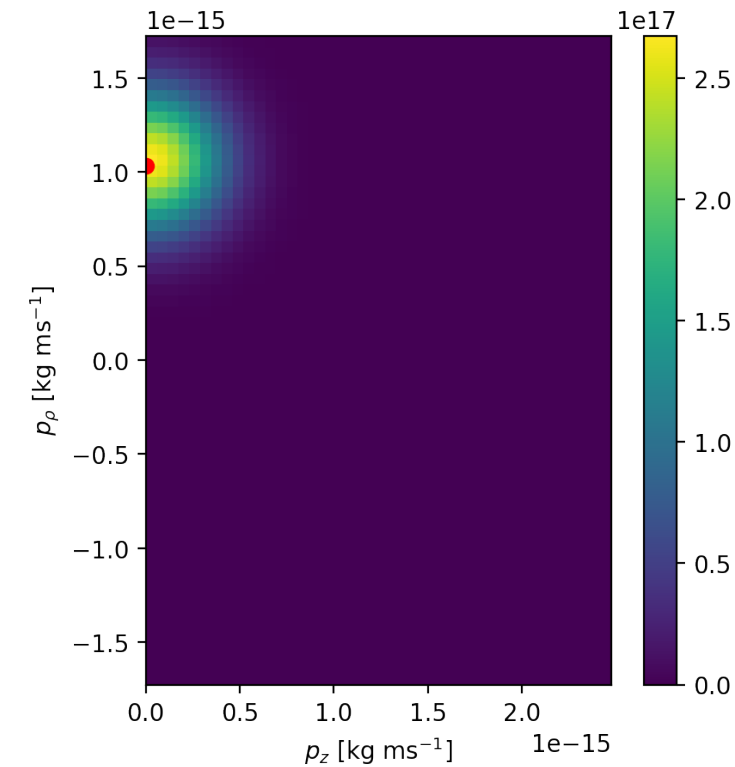
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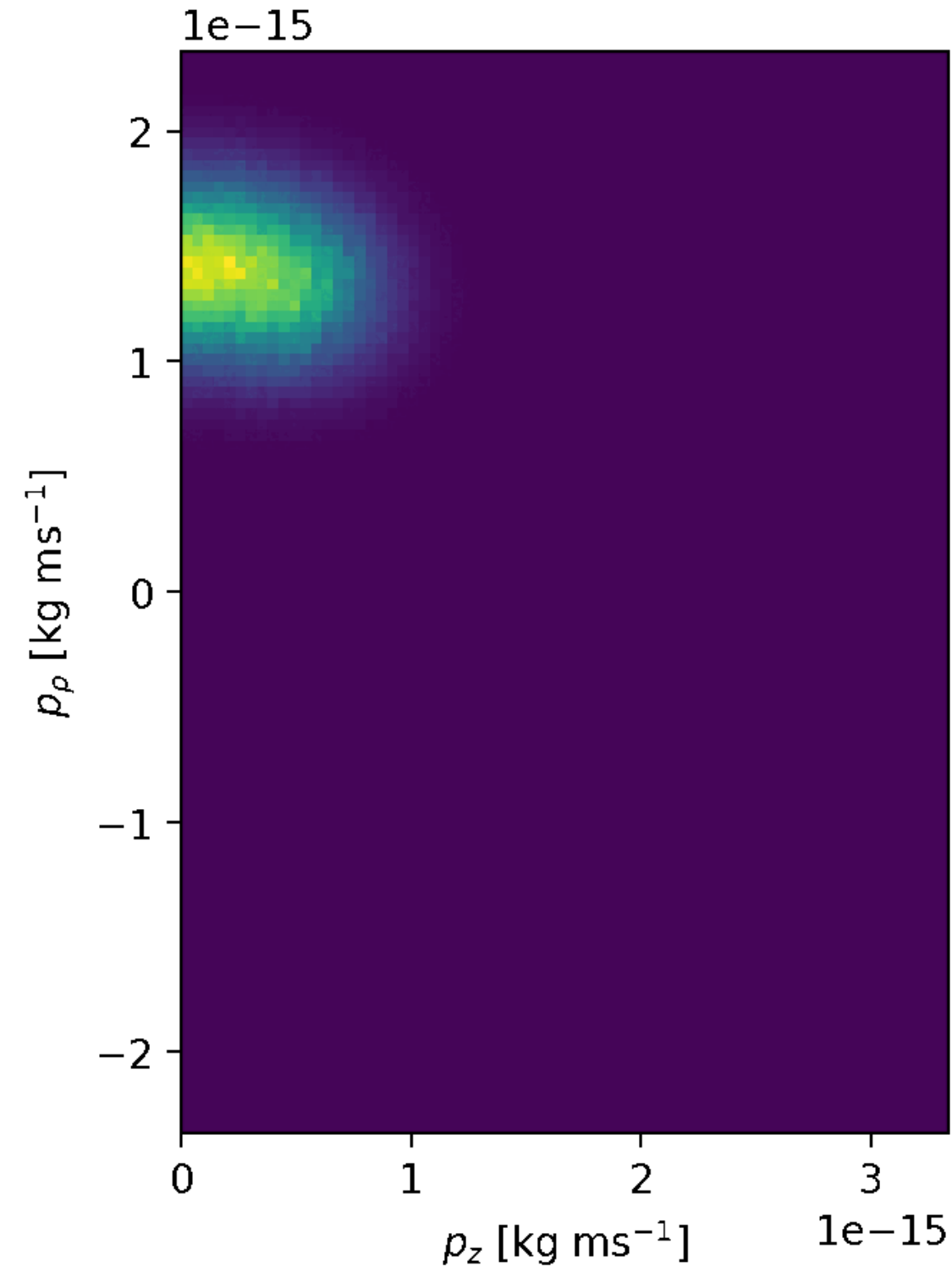
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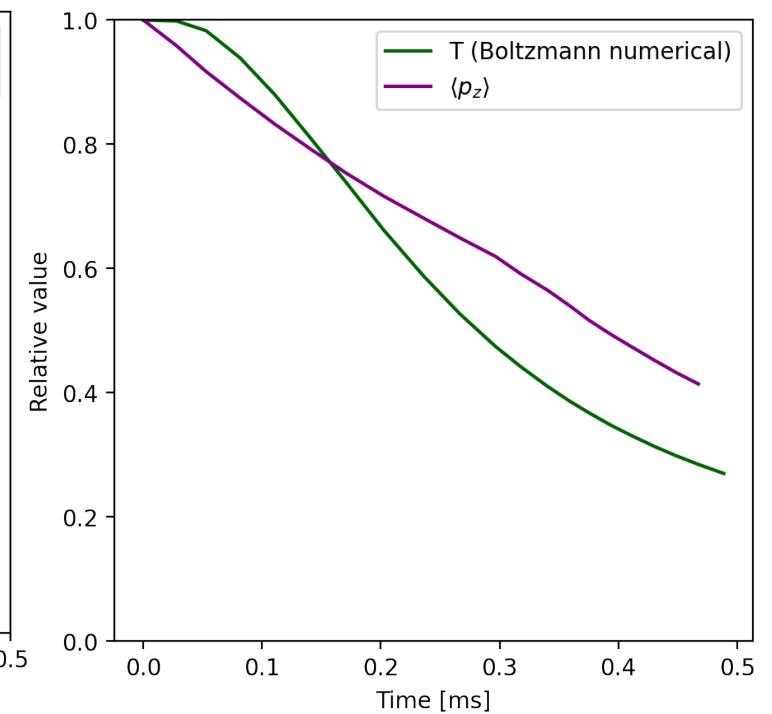
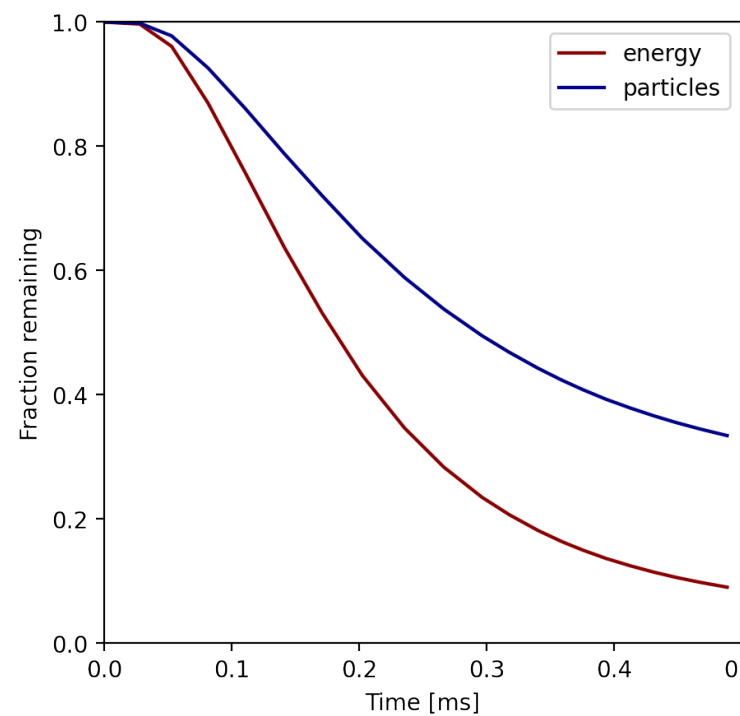


# Evaporative wiggle-cooling (or cool wiggeling?)



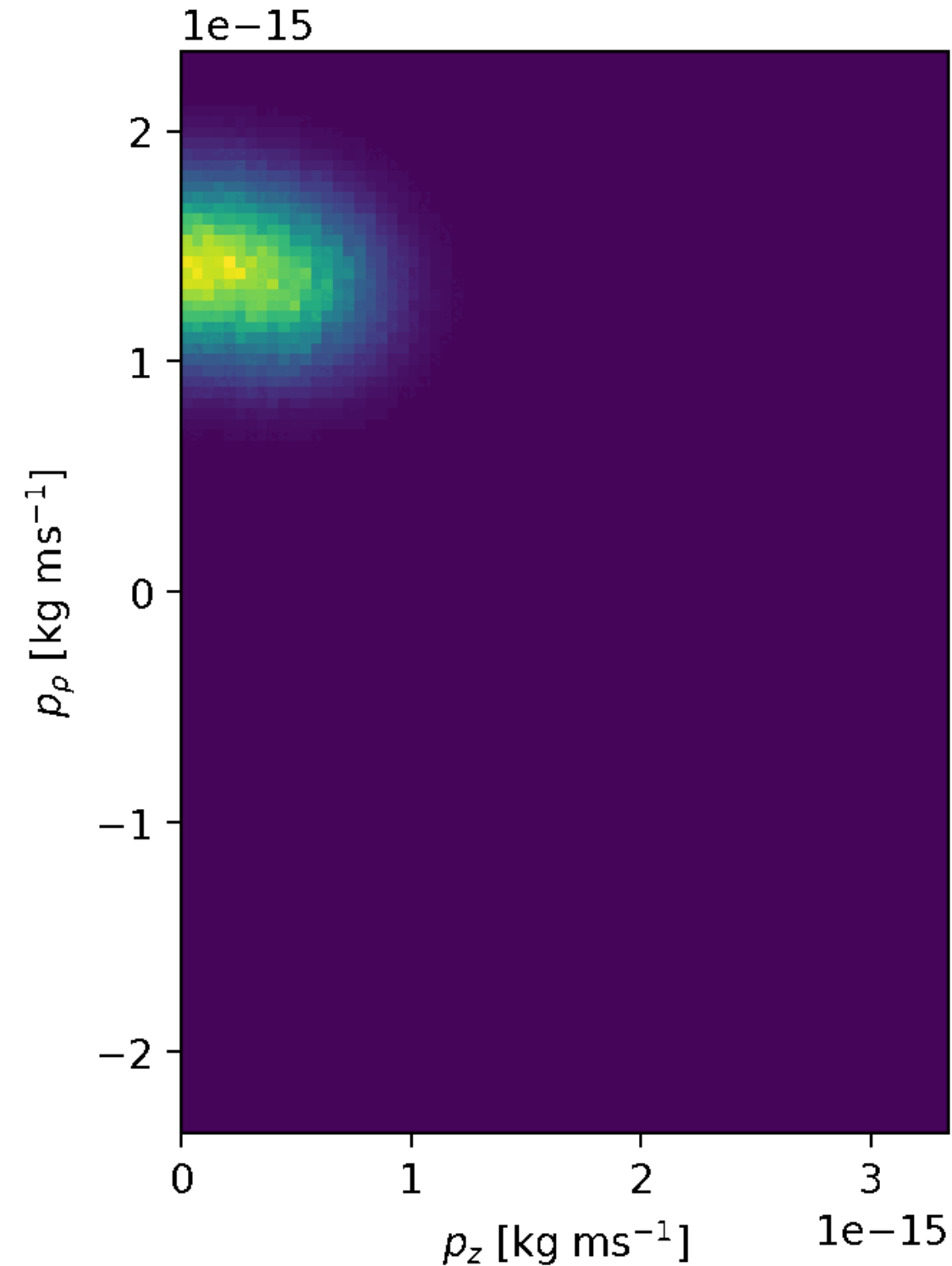
Simultaneous

- evaporative cooling
  - wiggeling
  - re-thermalization
- ▶ distribution gets **colder** and **slower**



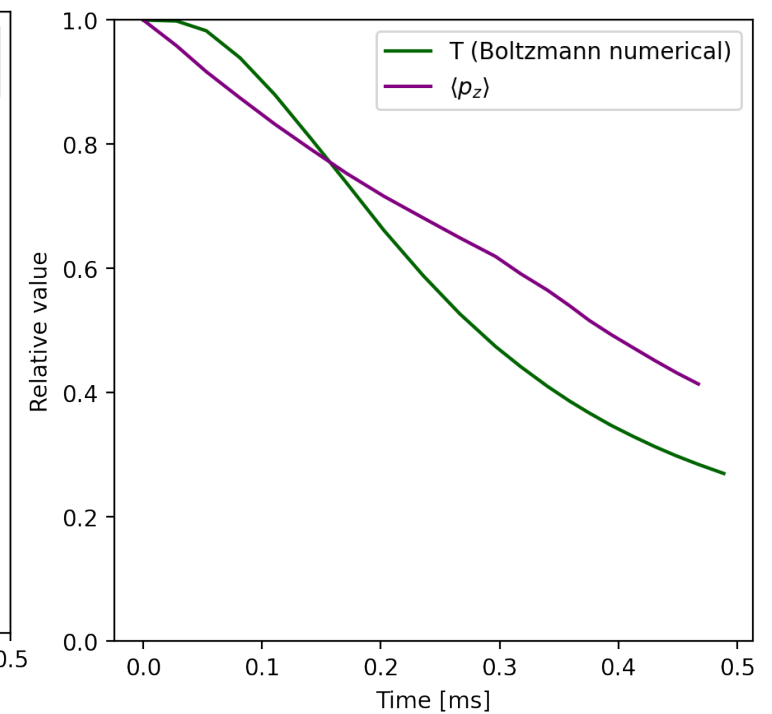
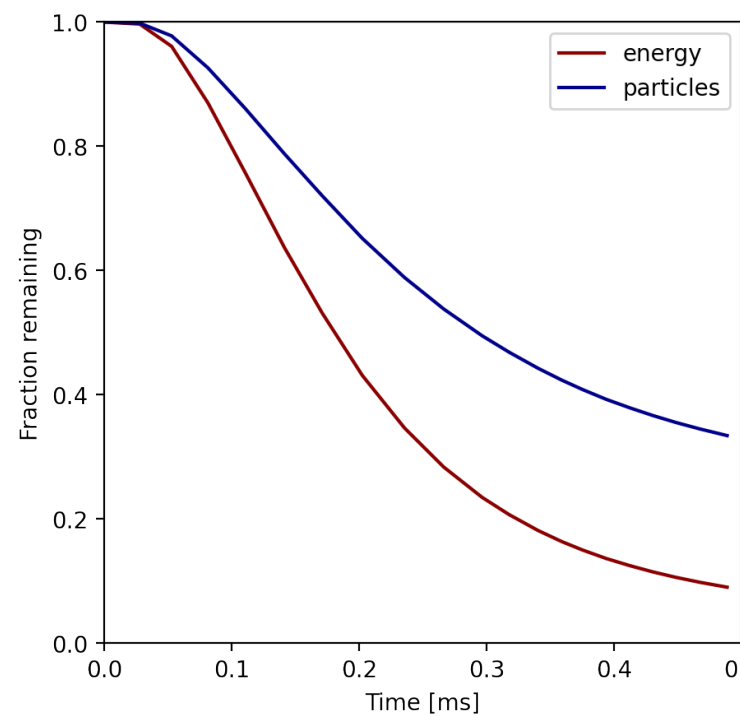


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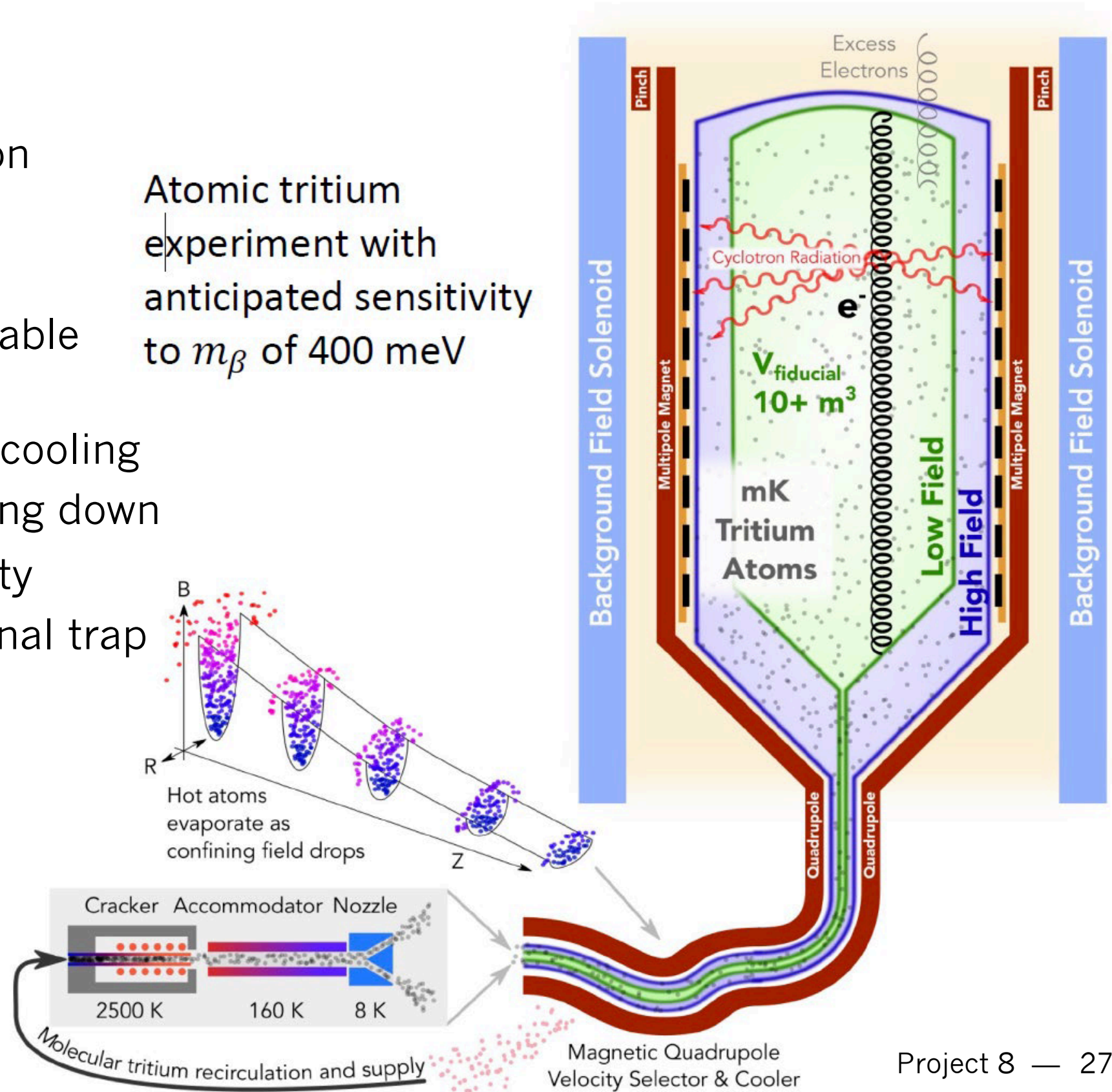
# Atomic tritium experiment

PROJECT 8

## Conceptual design

- thermal dissociation
  - ▶  $T_2 \rightarrow T$
- accommodation
  - ▶  $\emptyset(K)$  with acceptable recombination
- evaporative wobble cooling
  - ▶  $\emptyset(mK)$  and slowing down
  - ▶ high atomic purity
- Magneto-gravitational trap
  - ▶ Halbach array

Atomic tritium experiment with anticipated sensitivity to  $m_\beta$  of 400 meV





# Atomic beam demonstration: next steps

## Dissociation and accommodation

- surface physics critical
  - ▶ establish experimentally

## Evaporative wobble cooling

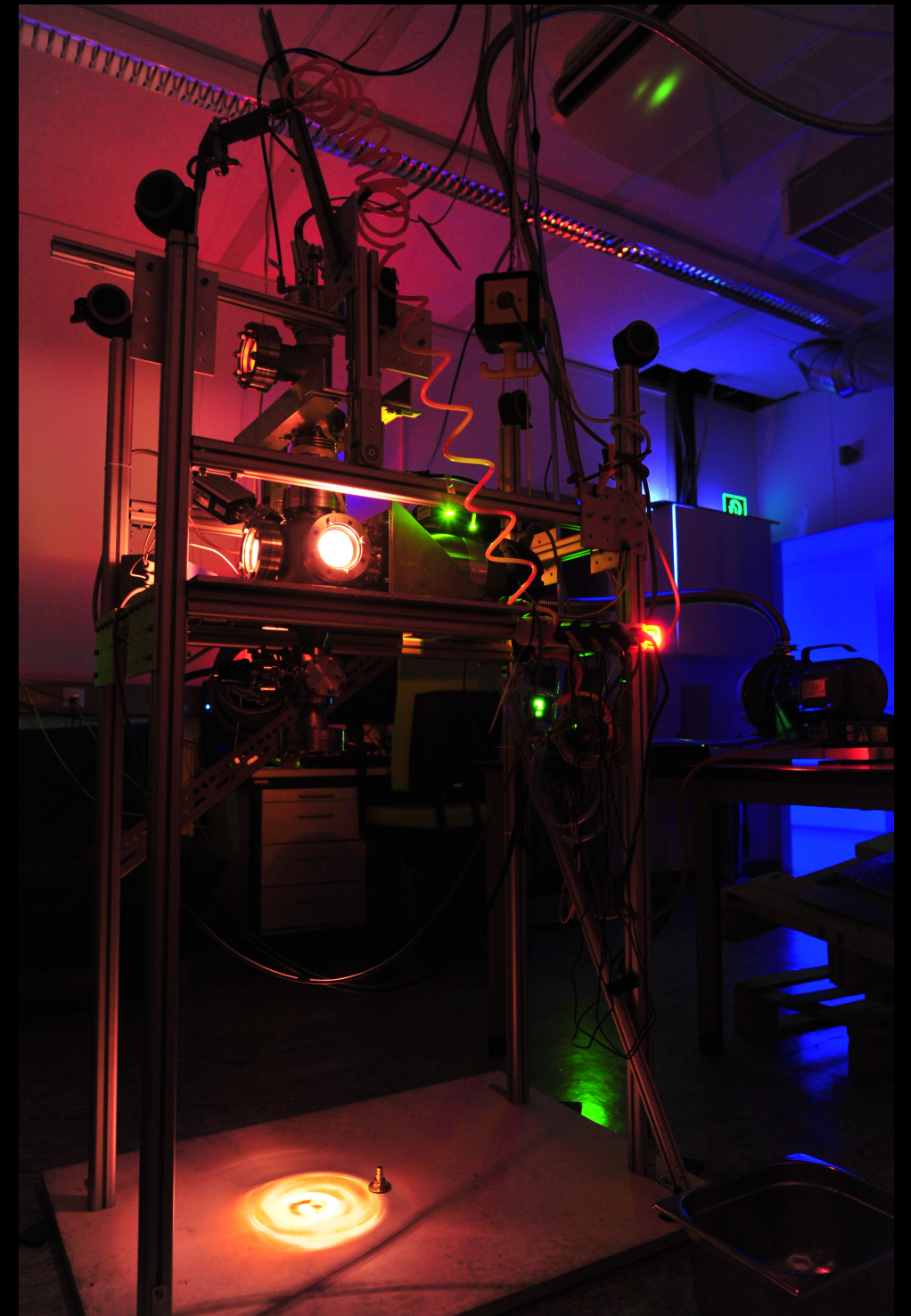
- magneto-thermodynamics
  - ▶ enter design phase

## Trapping

- fix temperature and density
  - ▶ enter design phase

## Overall approach

- start with  $H_2$  and  $D_2$ 
  - ▶ see talk by A. Lindman
- vet with  $T_2$ 
  - ▶ see the future





Thank you!







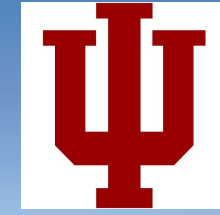
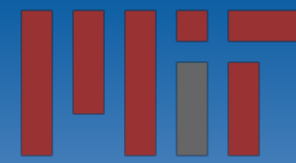
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# Project 8 collaboration