

# Baryogenesis from R-parity breaking: the role of flavour

Giorgio Arcadi, Laura Covi, Fiona Kirk



## Starting point

We observe:

$$\frac{\Omega_{\Delta B}}{\Omega_{DM}} \sim \frac{1}{5} \quad \text{whereas} \quad \eta \equiv \frac{n_B}{n_\gamma} = \frac{n_b - n_{\bar{b}}}{n_\gamma} \approx 6 \times 10^{-10}$$

Why so similar?

Possible explanation:

Contemporary production of the BAU and of DM from the out-of-equilibrium decay of a mother particle

$\Rightarrow \Omega_{\Delta B} \sim \Omega_{DM}$  if both production mechanisms are suppressed by comparably small numbers.

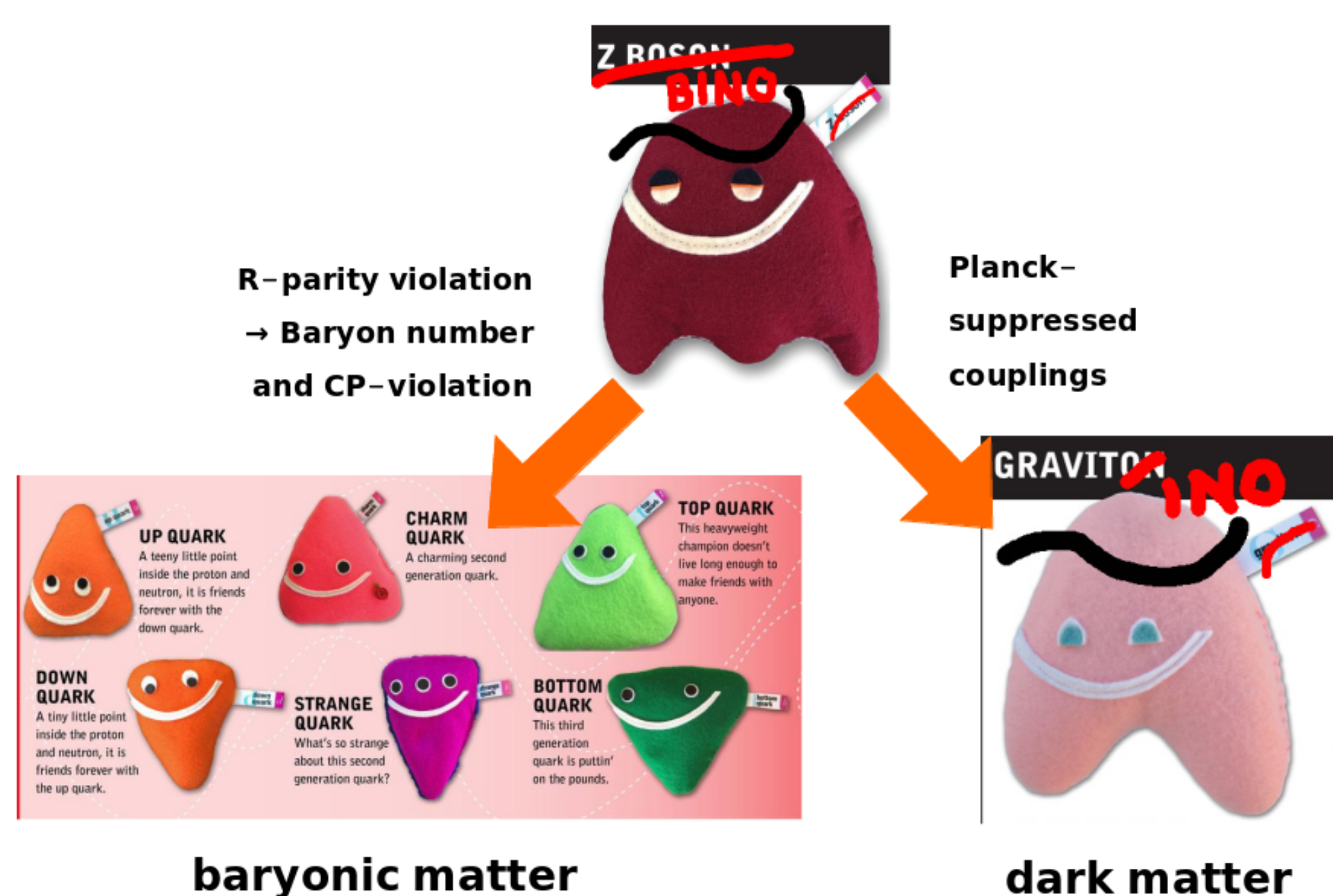
## Implementation

MSSM with the R-parity (and baryon number-) violating operator

$$\lambda'' U^c D^c D^c$$

$\Rightarrow$  avoid constraints from proton stability

$\Rightarrow$  R-parity violation: superpartners decay into SM-particles.



All three Sakharov conditions are satisfied:

**Baryon number violation** in the R-parity violating coupling  $\lambda'' U^c D^c D^c$

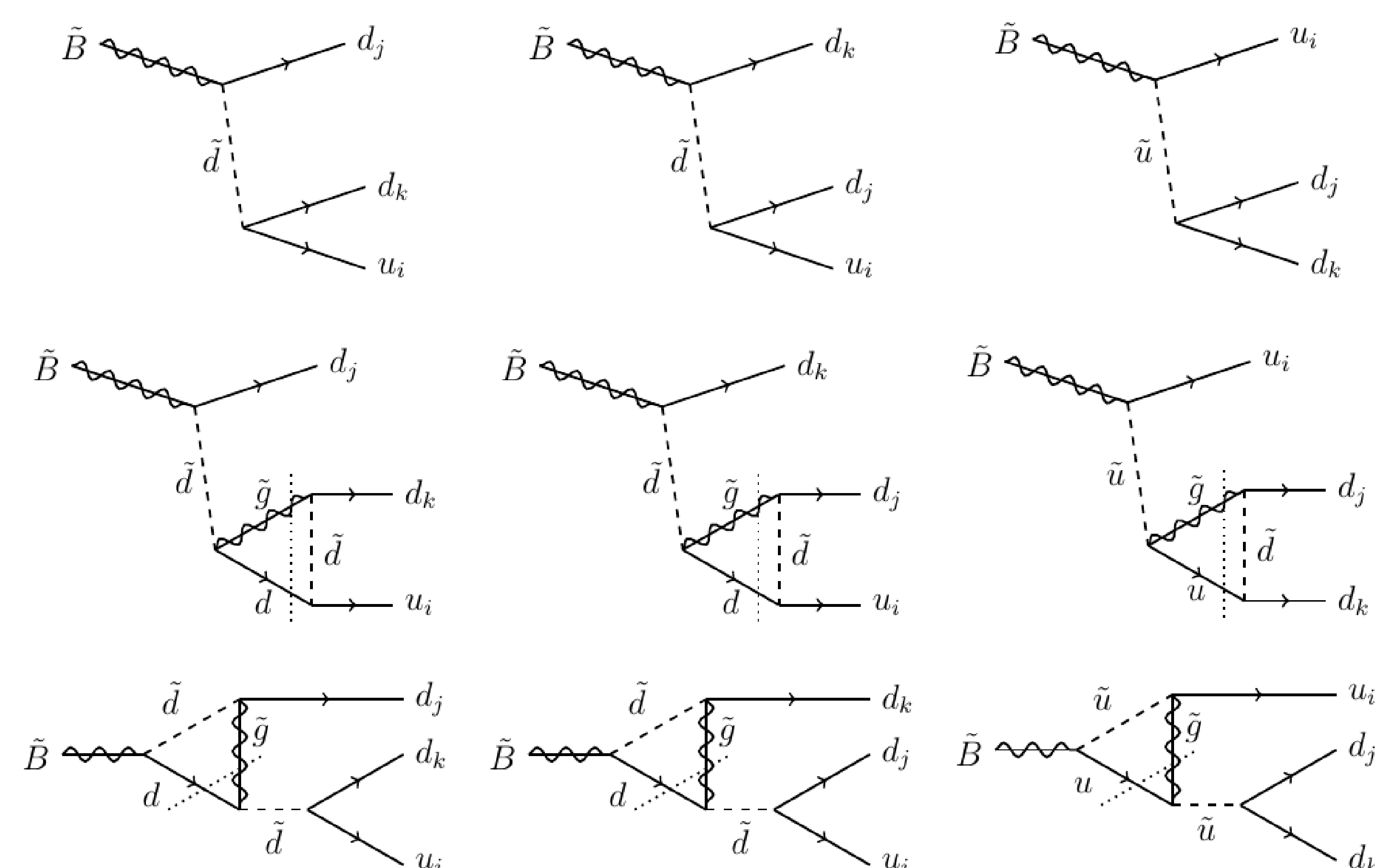
**CP-violation** from the interference between tree-level- and one-loop-contributions to the decay of the bino into quarks  
Nanopoulos-Weinberg equality at lowest order:

$$|\langle f|T|X \rangle|^2 - |\langle X|T|f \rangle|^2 = -2 \text{Im} [(\sum_h \langle f|T|h \rangle \langle h|T^\dagger|X \rangle) \langle f|T^\dagger|X \rangle]$$

**Departure from equilibrium** The bino  $\tilde{B}$  decays after freezing out.

## Generation of the BAU

The contributing diagrams are:



## Interactions

R-parity violating vertex:

$$\mathcal{L}_{udd} \ni 2\varepsilon^{\alpha\beta\gamma} \lambda''_{ijk} \left[ \Gamma_{RNi}^{u*} \tilde{u}_{N\alpha} \bar{d}_{j\beta}^c P_R d_{k\gamma} + \Gamma_{RNj}^{d*} \tilde{d}_{N\beta} \bar{u}_{i\alpha}^c P_R d_{k\gamma} \right] + h.c.$$

where  $\tilde{q}_N$ ,  $N = 1, \dots, 6$  are the squark mass eigenstates,  $\tilde{q}_{Li}$  and  $\tilde{q}_{Ri}$ ,  $i = 1, \dots, 3$  are the left- and right-handed squarks.  $\lambda''_{ijk}$  is antisymmetric in  $j, k$ .

Neutralino-sfermion-fermion vertex:

$$\mathcal{L}_{\tilde{\chi} f \tilde{f}} \ni \tilde{\chi} (\Gamma_{LNi}^f g_{\tilde{\chi}}^{LL} P_L + \Gamma_{RNj}^f g_{\tilde{\chi}}^{RR} P_R) f_i \tilde{f}_j^*$$

Here  $\tilde{\chi} = \tilde{B}, \tilde{G}$  and  $f, \tilde{f} = u, \tilde{u}, d, \tilde{d}$ .  $g_{\tilde{B}, \tilde{G}}^{LL, RR}$  are the usual coupling constants of the bino and the gluino:

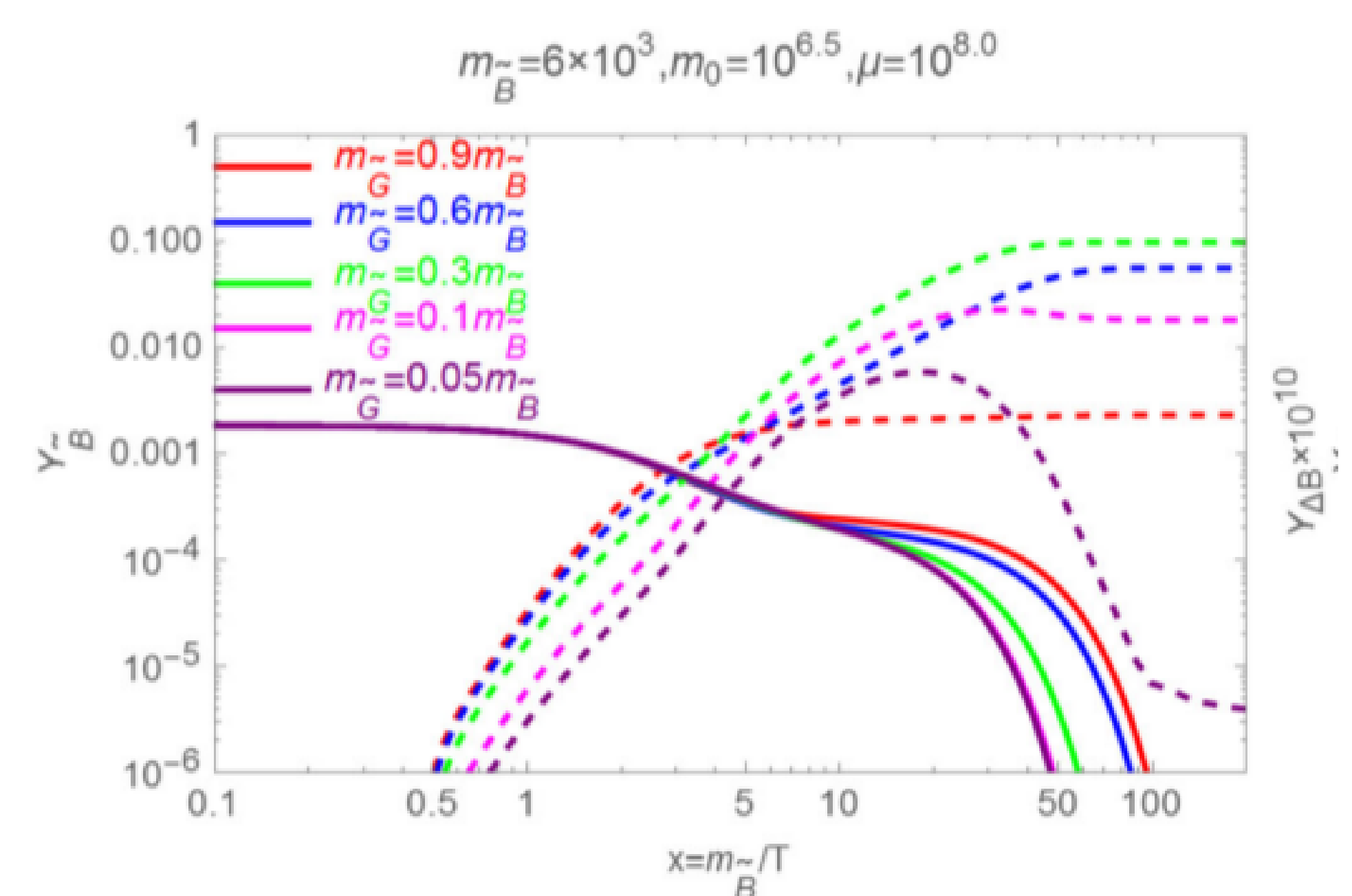
$$g_{\tilde{B}}^{LL} = -\sqrt{2}g_1(Q_f - T_3) \quad g_{\tilde{B}}^{RR} = -\sqrt{2}g_1 Q_f \quad g_{\tilde{G}}^{LL} = -\sqrt{2}g_3 \quad g_{\tilde{G}}^{RR} = \sqrt{2}g_3$$

## Previous works

- Y. Cui, JHEP **1312** (2013) 067 doi:10.1007/JHEP12(2013)067 [arXiv:1309.2952 [hep-ph]].
- G. Arcadi, L. Covi and M. Nardecchia, Phys. Rev. D **89** (2014) no.9, 095020 doi:10.1103/PhysRevD.89.095020 [arXiv:1312.5703 [hep-ph]].
- G. Arcadi, L. Covi and M. Nardecchia, Phys. Rev. D **92** (2015) no.11, 115006 doi:10.1103/PhysRevD.92.115006 [arXiv:1507.05584 [hep-ph]].

Arcadi et al. (2015): The correct DM and baryon relic densities can be obtained in this set-up with

- a bino mass,  $50 \text{ TeV} < m_{\tilde{B}} < 100 \text{ TeV}$
- a gluino NLSP mass,  $15 \text{ TeV} < m_{\tilde{G}} < 60 \text{ TeV}$  and
- a gravitino mass,  $100 \text{ GeV} < m_{\tilde{g}} < \text{a few TeV}$ .



Bino (solid lines) and baryon yields (dashed lines) for fixed  $m_{\tilde{B}}$ ,  $m_0$  (squark masses),  $\mu$  (the suppression factor of  $\tilde{B}\tilde{B} \rightarrow HH^*$  pair annihilation),  $\lambda = 0.3$  and for four values of  $m_{\tilde{G}}$  ranging from  $0.05 m_{\tilde{B}}$  to  $0.9 m_{\tilde{B}}$ . Plot taken from Arcadi et al. (2015)

These results were obtained for the case of

- degenerate right-handed down-type squarks
- flavour-diagonal mixing matrices  $\Gamma$
- zero quark masses.

## Outlook

Aim: Improve the study (enlarge the parameter space) by

- including additional diagrams
- allowing for non-degenerate squark masses
- considering the flavour structure (in the  $\Gamma$ s)

$\Rightarrow$  Flavour effects in the decay and wash-out?

$\Rightarrow$  Natural explanation of the ratio of baryon-to-DM densities?