

# Thinning

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building on work with Jannik Augscheller

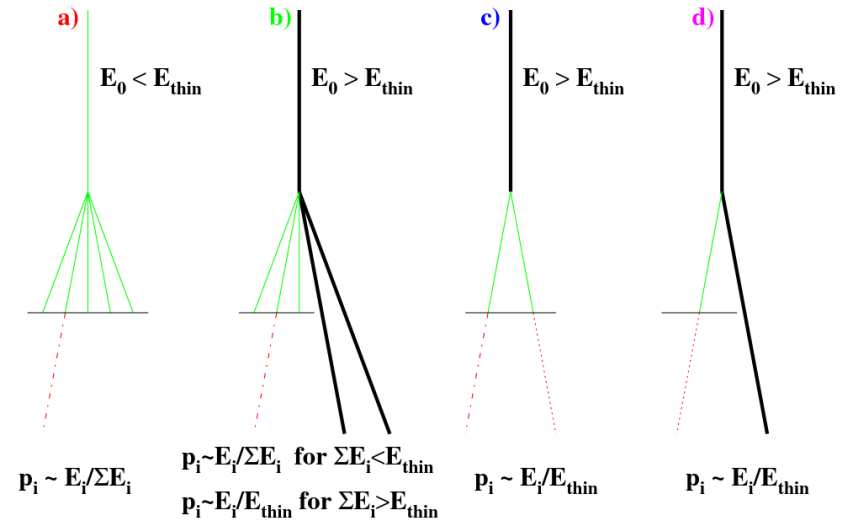
# Motivation

- runtime  $\sim$  primary energy / energy cut
- track fewer particles
- assign stat. weights, avoid bias
- downside: *artificial fluctuations* (vs. Shower-to-shower fluctuations)
- belongs to *Russian roulette* methods

# Hillas thinning

- original description ICRC 1981
- threshold energy  $E_{\text{th}} = \varepsilon E_0$  ( $\varepsilon \sim 10^{-5}, 10^{-6}, \dots$ )
- in vertices below  $E_{\text{th}}$ , select **one** out of all secondaries with
 
$$p_i = E_i / \sum_k E_k$$

$$w_i = w_0 / p_i$$
- (weighted) energy conserved exactly
- improved further by Hillas (1997), Heck & Knapp (1998)



# Artificial fluctuations

- related to weight distribution (narrower = better)

$$\langle N \rangle = \langle N_t \rangle \langle w \rangle$$

$$\sigma^2(N) = \langle N_t \rangle \sigma^2(w) + \langle w \rangle^2 \sigma^2(N_t)$$

- width can be controlled with **weight limitation**
- "optimum weight"  $w_{\max} = E_{\text{th}} / 1 \text{ GeV}$
- cumbersome with Hillas thinning  
(C7 implementation obscure...)

# Statistical thinning

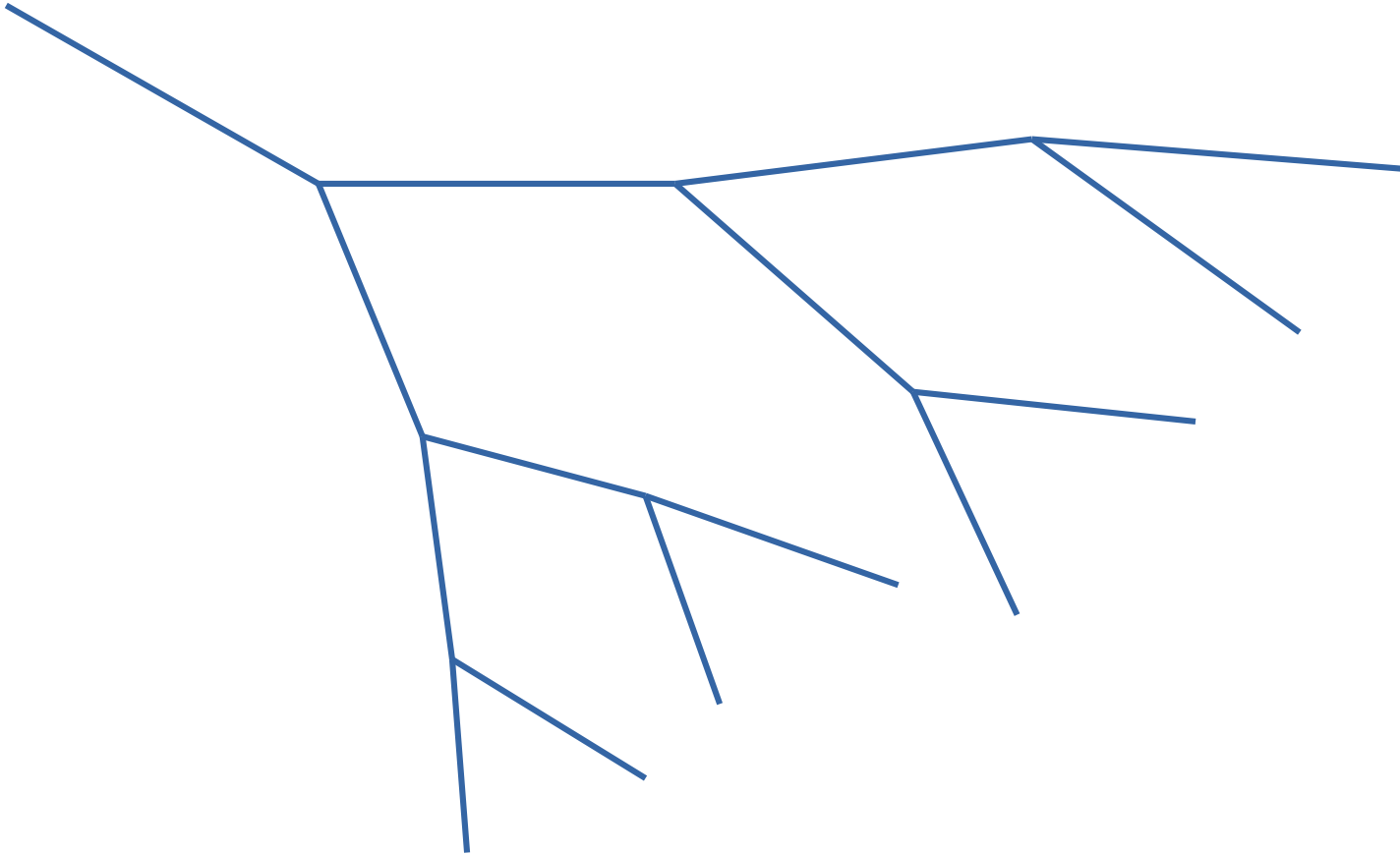
- keep/discard-test for each secondary independently
- no. of kept particles not fixed
- energy conserved *on average*
- bigger fluctuations (according to Hillas)
- acceptance factor arbitrary

# C8 implementation

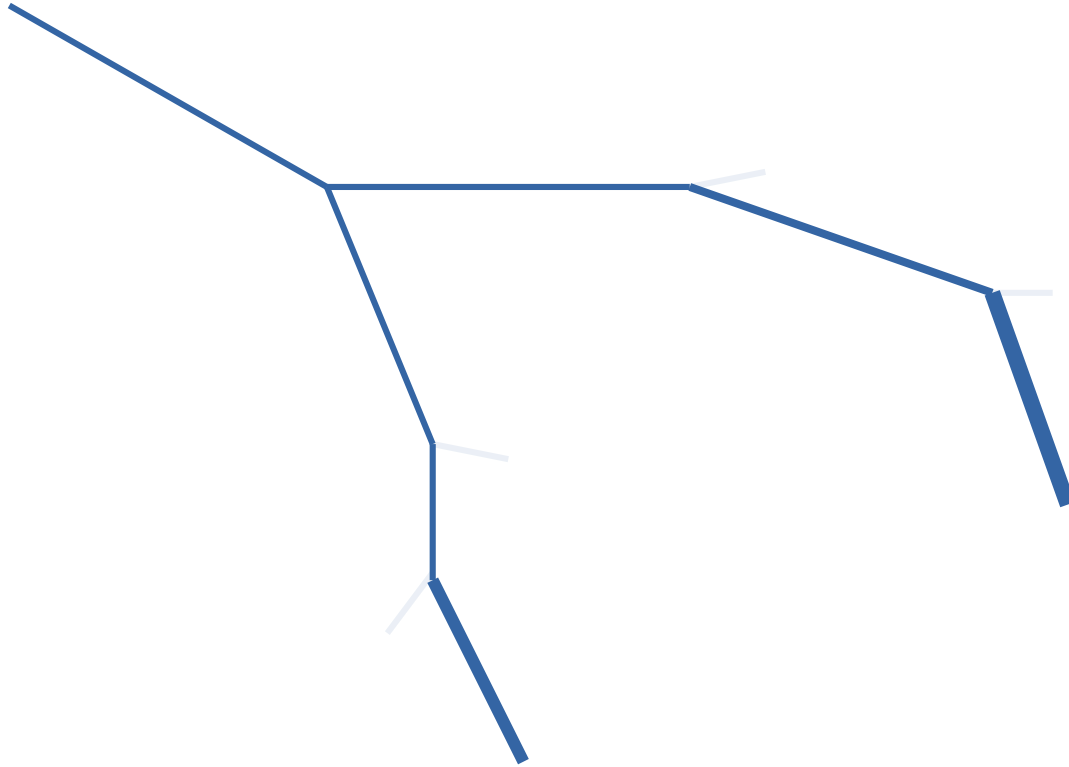
- only EM vertices ( $1 \rightarrow 2$  splitting)
- Hillas thinning while no secondary can reach  $w_{\max}$
- otherwise stat. thinning for both secondaries with

$$p_i = \max( p_i / \sum_k E_k, w_0 / w_{\max} )$$

# Unthinned

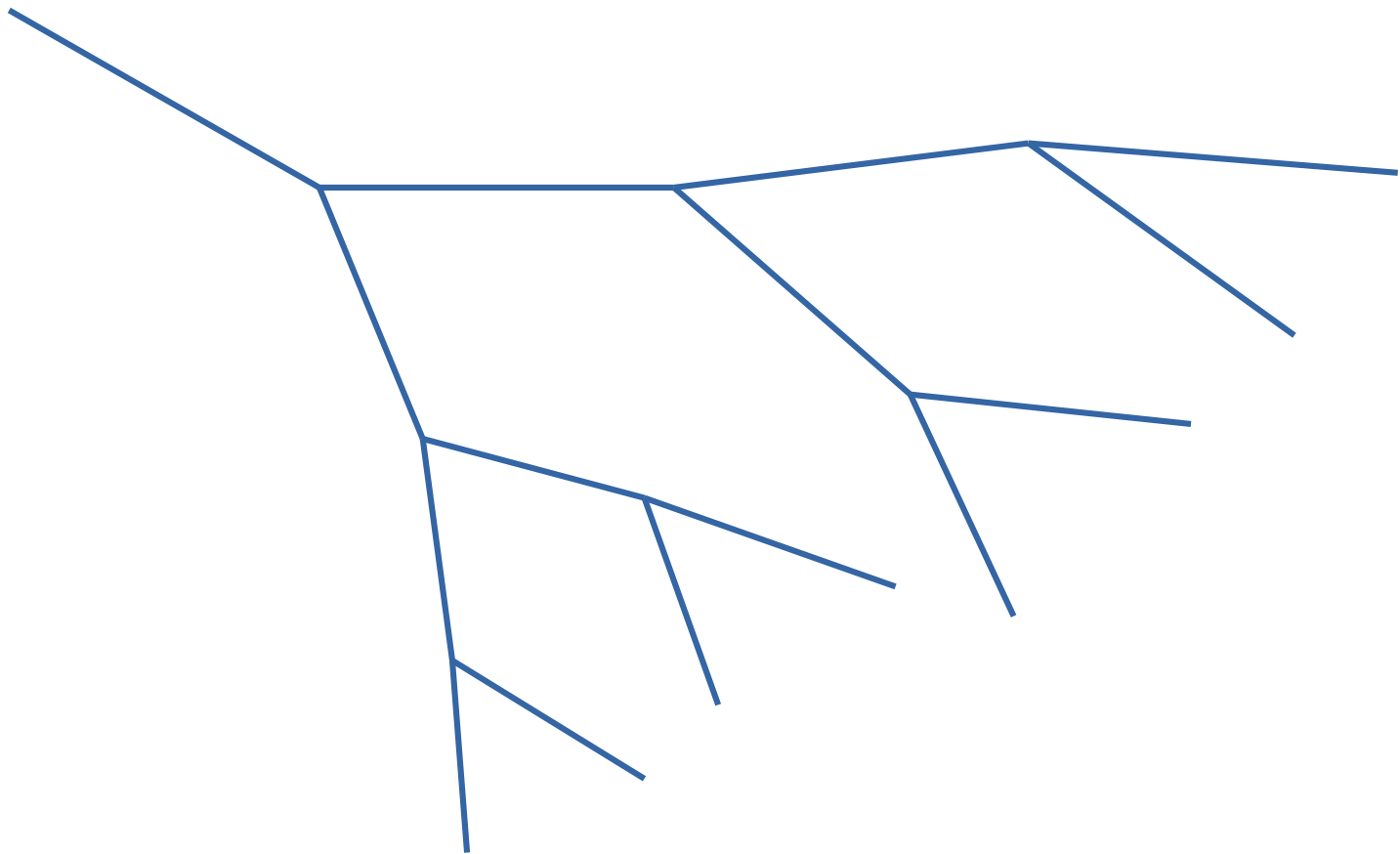


# "real" thinning



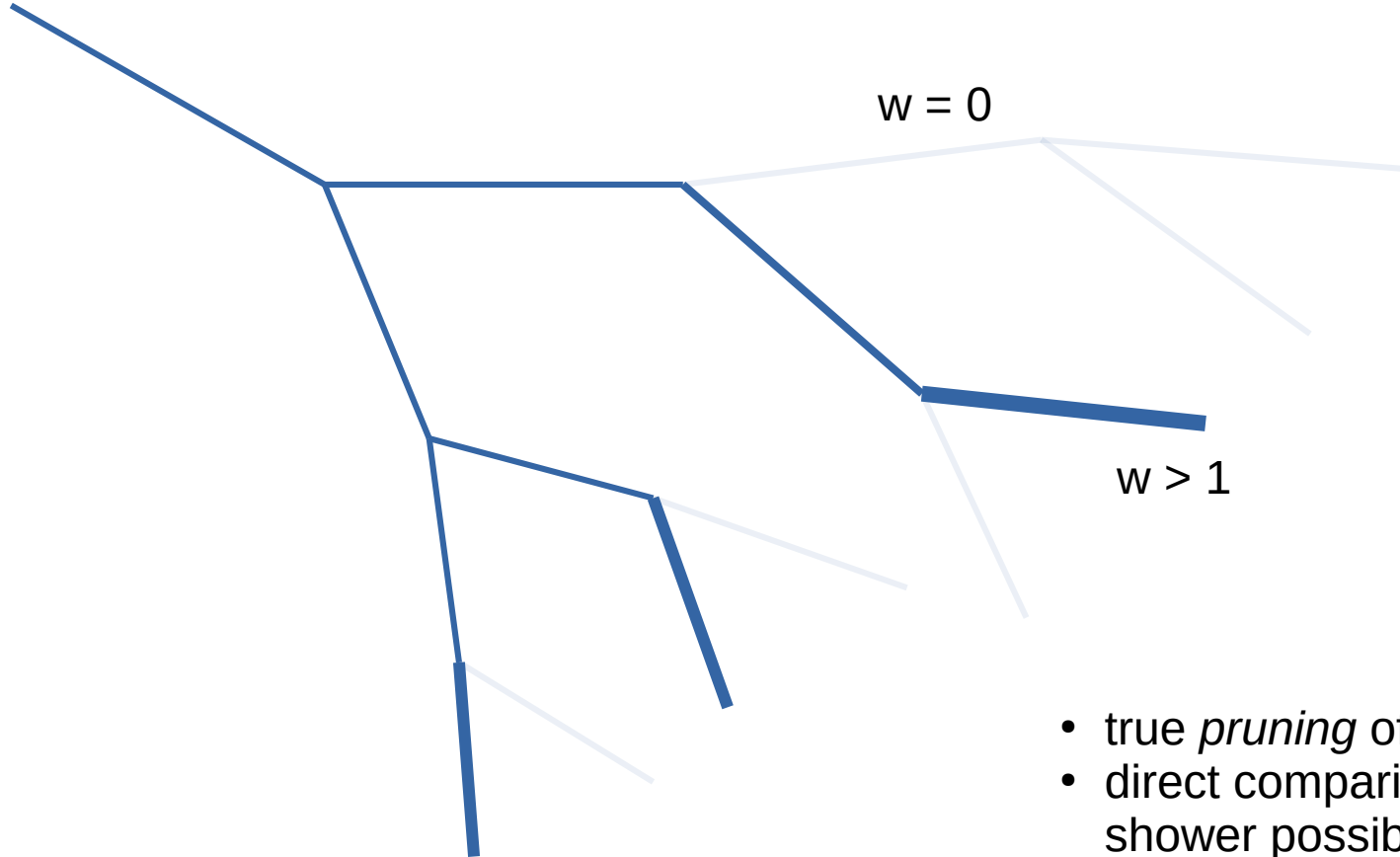
does not correspond to unthinned shower!





# "fake" thinning

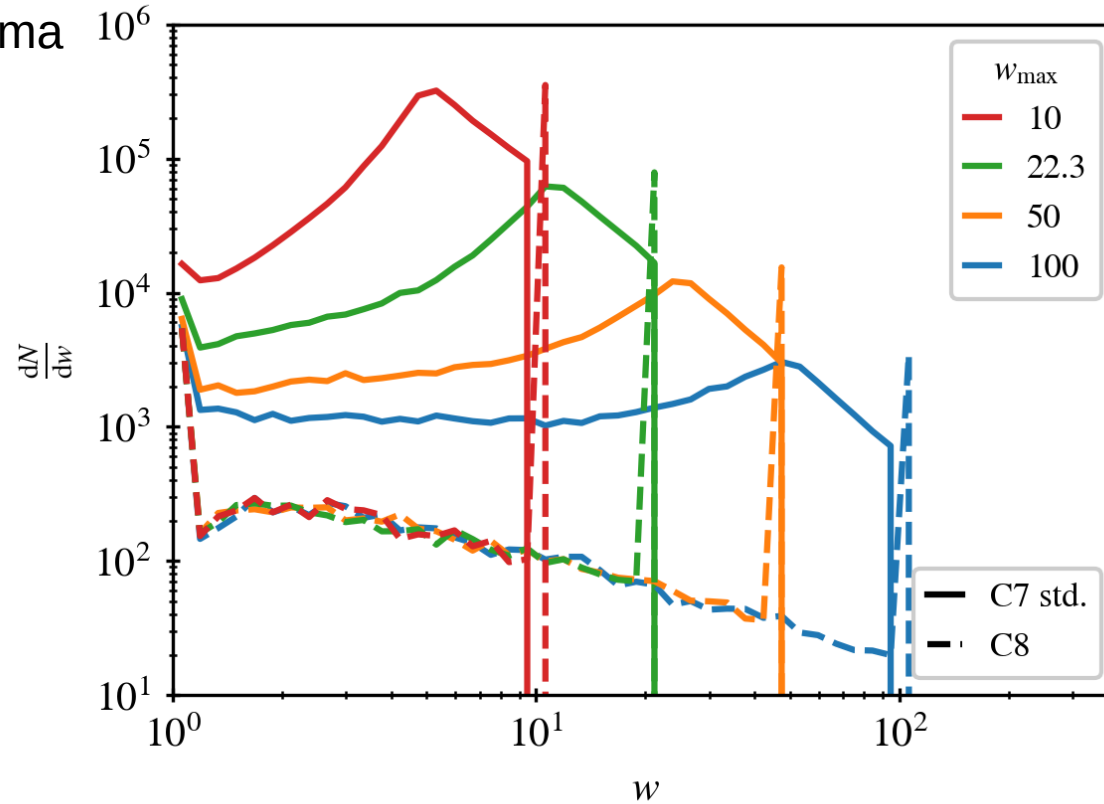
bin/corsika --multithin



- true *pruning* of shower
- direct comparison with unthinned shower possible

# Weight distribution

vertical  $10^{16}$  eV gamma  
 $\varepsilon = 10^{-5}$



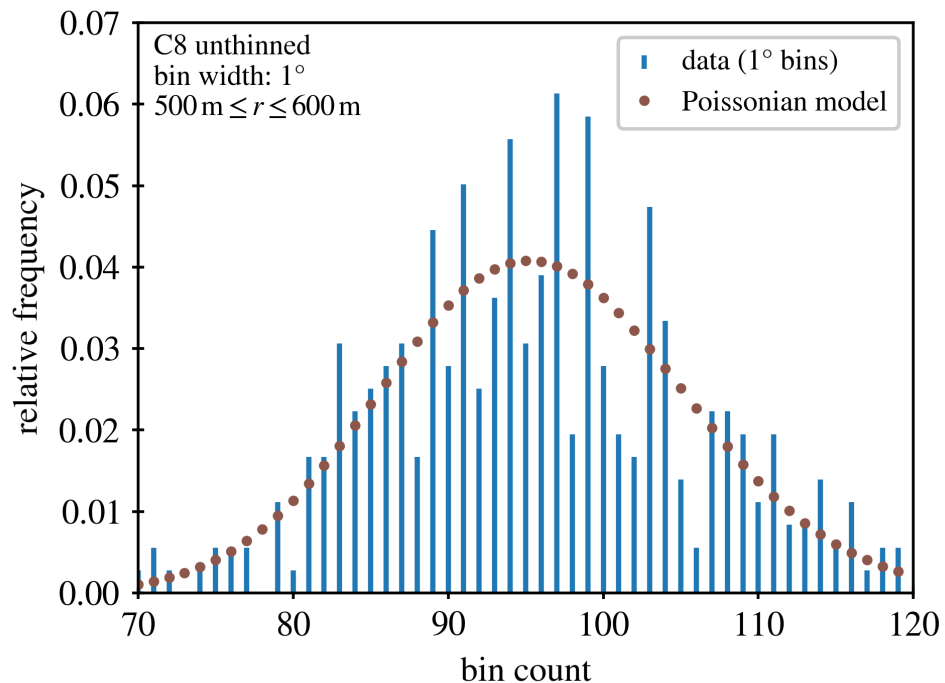
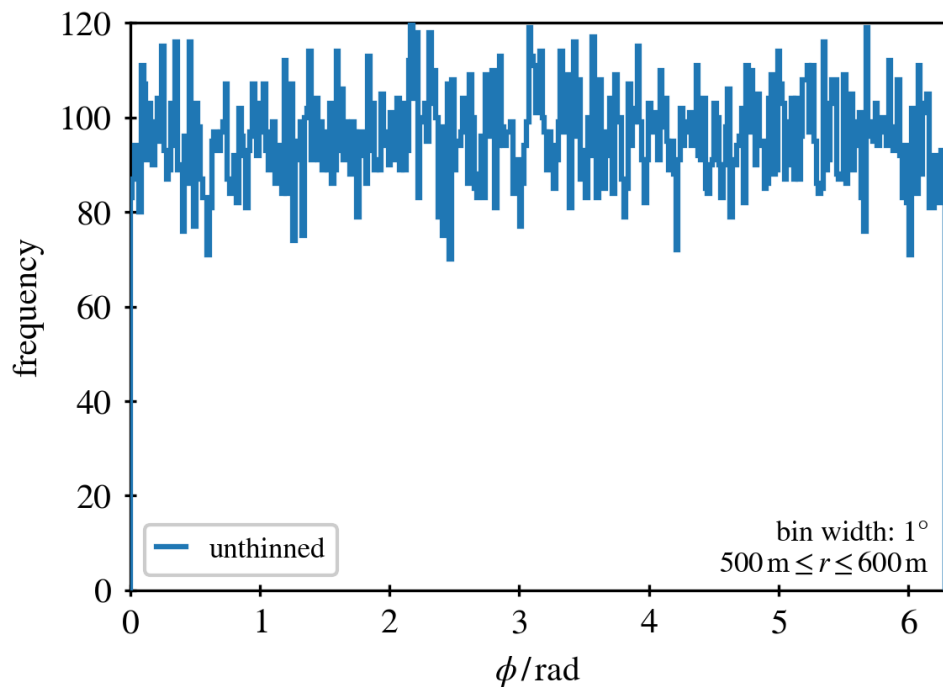
# Backport to C7

- MR 11, THIN and MULTITHIN
- new coconut option?

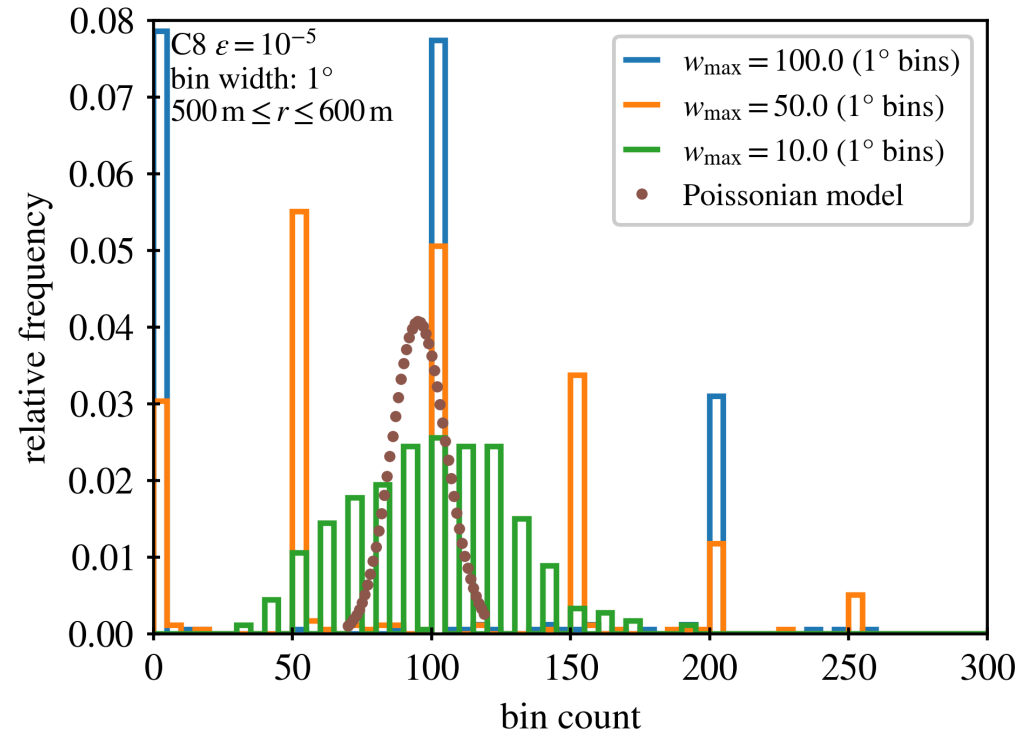
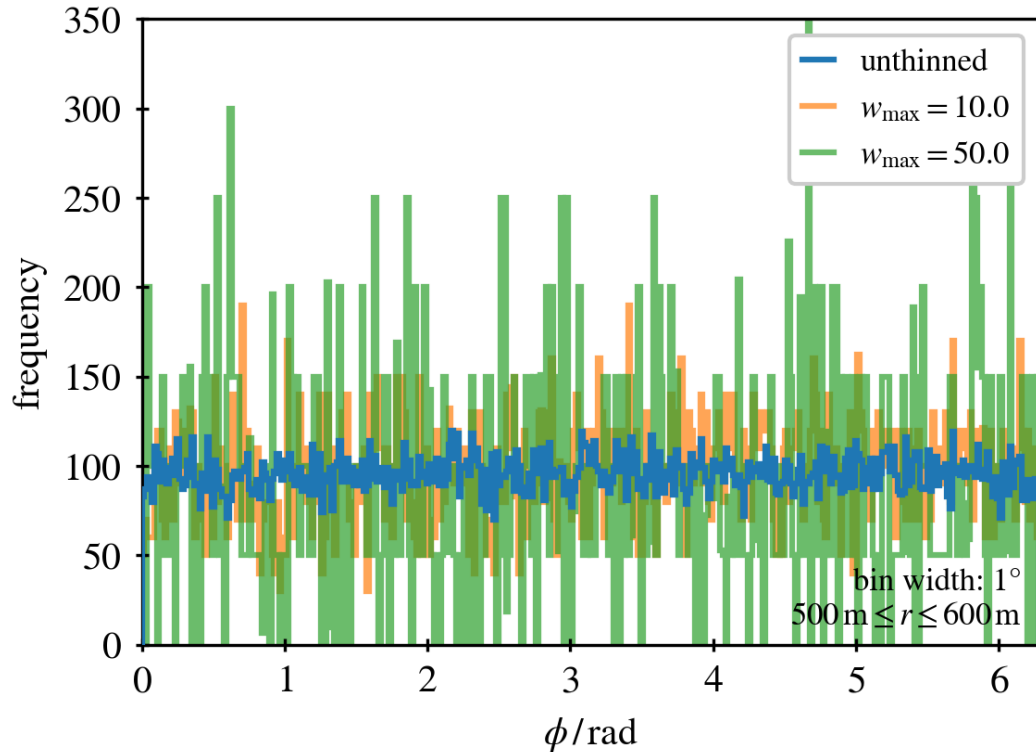
```
49942 2 FORMAT( ' THIN : EK,E1,IQ1,E2,IQ2=',IP,2E10.3,IX,12,E10.3,IX,12)
49943
49944 + ! weight not updated yet -> equal to parent weight
49945 +     wparnt = wt(np)
49946 +     fmax = wmaxem / wparnt
49947 +     f1 = ek / ekhigh
49948 +     f2 = ek / eklow
49949 +
49950 +     if (ek <= ethinn .and. (f1 >= fmax .or. f2 >= fmax)) then
49951 +     ! do statistical thinning
49952 +         f1 = min(ek / ekhigh, fmax)
49953 +         f2 = min(ek / eklow, fmax)
49954 +         call rmmard(rnd, 2, 2)
49955 +
49956 +     ! consider higher-energy secondary (at NP-1)
49957 +     ! keep with probability p = 1/f1
49958 +     if (rnd(1) * f1 <= 1.) then
49959 +     ! keep, update weight
49960 +         wt(np-1) = wparnt * f1
49961 +     else
49962 +     ! discard, move next particle here
49963 +     ! note: time, weight, etc. of both secondaries are equal, no need to overwrite
49964 +         E(NP-1) = E(NP)
49965 +         IQ(NP-1) = IQ(NP)
49966 +         U(NP-1) = U(NP)
49967 +         V(NP-1) = V(NP)
49968 +         W(NP-1) = W(NP)
49969 +         NP = NP-1
49970 +     endif
49971 +
49972 +     ! consider lower-energy secondary (at NP)
49973 +     ! keep with probability p = 1/f2
49974 +     if (rnd(2) * f2 <= 1.) then
49975 +     ! keep, update weight
49976 +         wt(np) = wparnt * f2
49977 +     else
49978 +     ! discard
49979 +         NP = NP-1
49980 +     endif
49981 +
49982 +     return
49983 +     endif
49984 +
49985 CALL RMMARD( RD,1,2 )
49986
```

# (Artificial) fluctuations

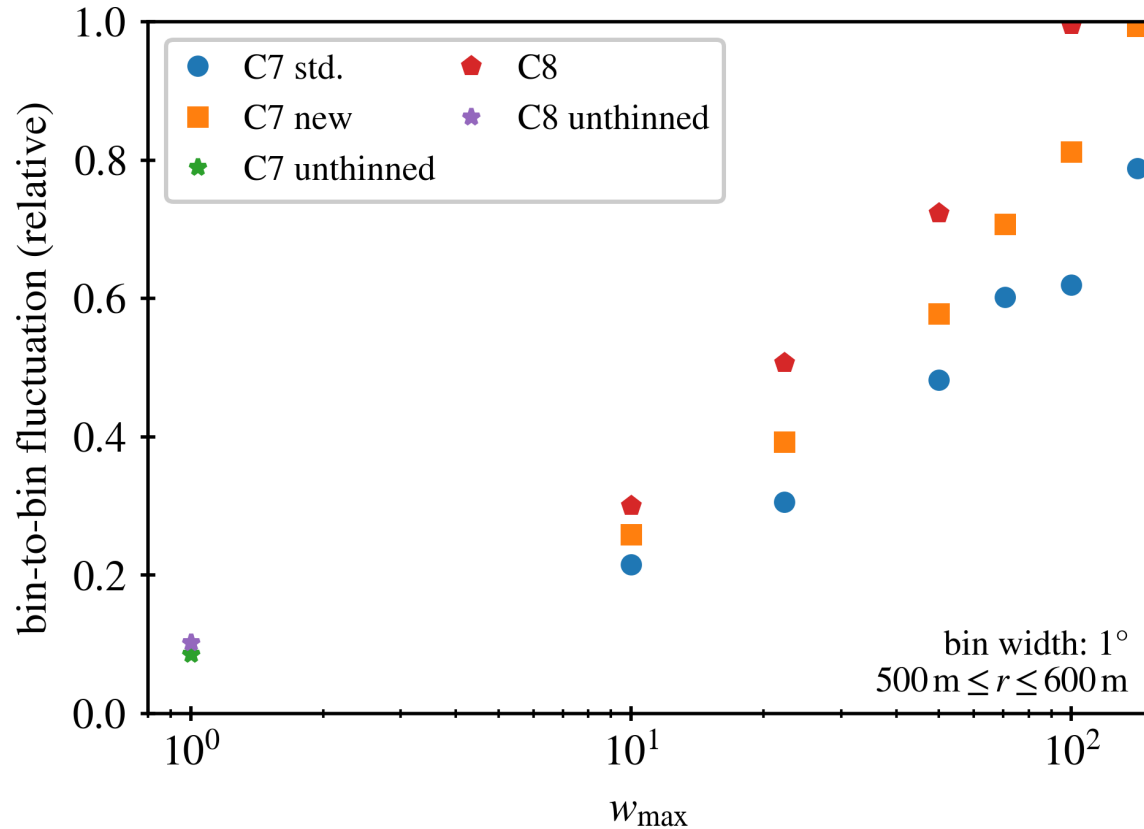
- here: azimuthal (bin-to-bin) fluctuations



# With thinning



# Performance comparison



impact on runtime  
not checked yet!

# Conclusions

- Thinning works
- unbiased, check with `--multithin`
- narrow weight distribution possible
- Performance? New optimum weight prescription?
- Paper?