



# Search for eV sterile neutrino with the KATRIN experiment

#### ISAPP 24, Bad Liebenzell

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## Non-standard or Sterile Neutrino

Sterile neutrino = SM neutral singlet fermion

 Existence could be revealed through effects of mass and mixing with active neutrinos (neutrino oscillations, β- decay, 0νββ-decay)



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- Singlet fermions naturally appear in the dark sector
- Members of dark sector could mix with active neutrinos via neutrino portal coupling
- Sterile neutrinos can live at any mass scale: GeV...,

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Lecture by J.Kopp



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- Experimental hints for eV scale :
  - Appearance LSND ( $3\sigma$ ) and MiniBooNE ( $4.8\sigma$ ) excess observations Explained by ( $\nu_{\mu} \rightarrow \nu_{s} \rightarrow \nu_{e}$ )
  - Disappearance SAGE and GALLEX: Gallium anomaly (2.9 $\sigma$  deficit) Explained by  $\nu_e \rightarrow \nu_s$
  - The Gallium anomaly reaffirmed by BEST experiment

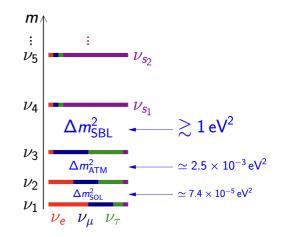
Lecture by T.Lasserre



#### Interpretation



- SBL anomalies could be explained by an additional neutrino flavor (v<sub>s</sub>)
- There must be at least one additional mass squared difference,  $3\nu + 1$  framework  $\Delta m_{SBI}^2 \approx (1-2) \text{ eV}^2$
- Allowed by solar, atmospheric and long baseline experiments, achieved with  $|U_{e4}|^2 \ll 1$



# $R_{\beta}(E, m_{u}^{2}, m_{4}^{2}, |U_{e4}|^{2})$

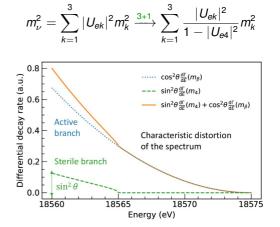
Active branch  

$$= \cos^2 \theta \cdot R_{\beta}(E, m_{\nu}^2) + \sin^2 \theta \cdot R_{\beta}(E, m_4^2)$$
Sterile branch

 $= (1 - |U_{e4}|^2) \cdot R_{\beta}(E, m_{\nu}^2) + |U_{e4}|^2 \cdot R_{\beta}(E, m_4^2)$ 

Sterile neutrino branch leads to a kink at  $E_0 - m_4$ 

Differential decay rate:

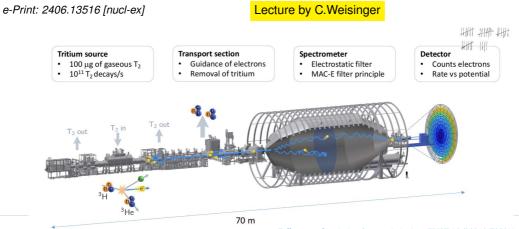




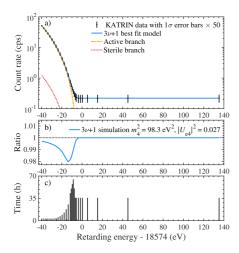


## **KATRIN Experiment**

- Kinematics-based neutrino mass experiment (expected sensitivity is better than 0.3 eV (90% CL) with 1000 days of measurement time)
- Current result:  $m_\beta$  < 0.45 eV (90%) CL, from 259 days of data only



## Sterile Signal in $\beta$ -decay Spectrum



Measured integral spectrum N<sub>exp</sub>(qU) is fitted to the model N<sub>model</sub>(qU, Θ):

$$N_{ ext{model}}(qU, \Theta) = A \cdot \int R_{eta}(E, \Theta) \cdot f(E, qU) + Bg$$

- 6 model parameters:
  - A Signal amplitude
  - E<sub>0</sub> effective endpoint energy
  - m<sup>2</sup> effective mass of electron anti-neutrino
  - Bg Background rate
  - m<sub>4</sub><sup>2</sup> sterile neutrino mass
  - $|U_{e4}|^2$  sterile neutrino mixing



## **Dataset and Bias-free Analysis**



#### Data Combination Approach:

- Summed  $\beta$ -spectrum counts at identical set points.
- Combined spectra from different detector pixels.

#### Modeling and Likelihood:

- Maximum-likelihood fit with Gaussian/Poisson likelihoods.
- Minimize χ<sup>2</sup> = -2 log(L) with respect to nuisance parameters (Θ = (m<sup>2</sup><sub>ν</sub>, E<sub>0</sub>, A, B, ..)).

#### Systematic Uncertainties:

- Pull terms in likelihood for experimental parameters.
- Calibration measurements for external estimates.

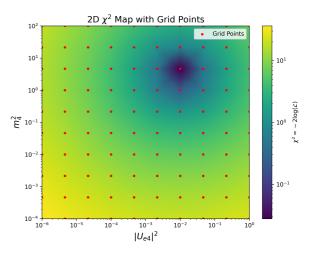
#### Unblinding Procedure

- Code validation on Monte Carlo twins.
- Tritium spectrum, response model, systematics treatment and budget (pull term approach) same as active neutrino mass analysis.
- Two Independent Analysis Teams with Independent Codes:
  - KaFit (exact model evaluation).
  - Netrium (uses neural nets for swift model interpolation).



#### **Analysis Method**

- Extend β- spectrum model from 3 active to 3+1 (active + sterile) framework
- Grid Scan:  $50 \times 50 [\log(|U_{e4}|^2), \log(m_4^2)]$  plane
- Contours are drawn at  $\Delta \chi^2 = \chi^2 \chi^2_{BF}$  = 5.99 (95% CL, 2 dof)
- Energy range: [*E*<sub>0</sub> − 40, *E*<sub>0</sub> + 135] eV
- $\blacksquare$  Sensitive to  $\mathit{m}_{4}^{2} \leq$  1600 eV^{2} and  $|\mathit{U}_{e4}|^{2} \leq$  0.5
- Two complementing analyses
  - Case-I Fixed neutrino mass:  $m_{\nu}^2$  = 0 ( $m_{1,2,3} \ll m_4$ )
  - Case-II Free neutrino mass: m<sup>2</sup><sub>ν</sub> as nuisance parameter





#### **Results from First Two Science Runs**

- 5.24 $\times$ 10<sup>6</sup> electrons for 40 eV below E<sub>0</sub>, 1265 hours of data  $10^{3}$ Best fit:  $-m_4^2 = 59.9 \text{ eV}^2$ ,  $|U_{e4}|^2 = 0.011$ ,  $10^{2}$  $m_4^2 (eV^2)$  $m_{v}^2 = 0.0 \text{ eV}^2$  $-\Delta \chi^2_{null} = 0.66$ Active neutrino mass set free  $m_{\nu}^2$  free  $m_{\rm u}^2 = 0 \, {\rm eV}^2$ Best fit: ..... KNM1 KNM1  $- m_4^2 = 87.4 \text{ eV}^2, |U_{e4}|^2 = 0.019,$ ---KNM2 ---KNM2  $m_{\nu}^2 = 0.57 \text{ eV}^2$ -KNM1+2 --- KNM1+2  $-\Delta\chi^2_{null}=1.69$  $10^{-2}$
- Signal-to-background ratio of up to 235

Figure: Phys. Rev. D 105 (7 2022)

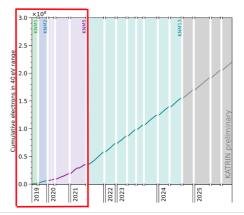
 $|U_{e^4}|^2$ 

 $10^{-1}$ 



#### **Data Collection Status**

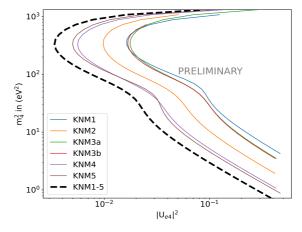
- KNM1 to KNM5: 20 % of expected KATRIN data
- Significant experimental development: Shifted Analyzing Plane (SAP) background reduction method Lokhov et al., EPJ C 82 (2022) 3, 258





## **Sensitivity Projection From Five Science Runs**

- **Case-I**:  $m_{\nu}^2$  = 0 eV<sup>2</sup>
- 40 eV fit range,  $|U_{e4}|^2 \in [0, 0.5]$
- Stat. only + all systematics 95% CL
- Gain in overall sensitivity with increased statistics
   S. Mohanty, PoS EPS- HEP2023 (2024)



Campaign	KNM1	KNM2	KNM3a	KNM3b	KNM4	KNM5	KNM1-5
No. of signal electrons ( $\times 10^6$ )	2.0	4.3	1.1	1.4	10.2	16.8	35.8



#### **Impact of Systematics**

Calculating 68% CL uncertainty on 
$$|U_{e4}|^2$$
:  $\sigma_{syst} = \sqrt{\sigma_{Stat+Syst}^2 - \sigma_{Stat}^2}$ 

- Statistically dominated uncertainties
- Largest systematic contribution: Penning Bg (low m<sup>2</sup><sub>4</sub>), Column Density (high m<sup>2</sup><sub>4</sub>)





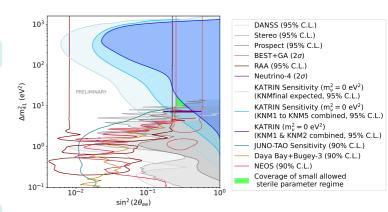
## **Comparison to Other Experimental Results**

Translation of parameters:

 $\sin^2(2\theta) = 4|U_{e4}|^2(1-|U_{e4}|^2)$ 

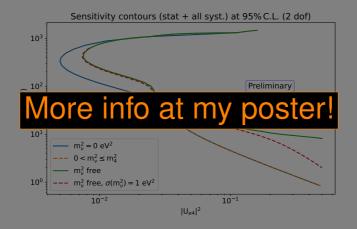
- Large △m<sup>2</sup><sub>41</sub> solutions of RAA and BEST+GA anomalies excluded
- Current KATRIN data extends exclusion bounds from SBL oscillation experiments for  $\Delta m_{41}^2 \ge 10 \text{ eV}^2$
- Probing large parameter space for light sterile neutrino anomalies

Analysis has been applied to the real data and results to be released soon!





#### Impact of active neutrino on sterile neutrino search



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- New physics beyond the SM can include sterile neutrinos at all mass scales
- KATRIN uniquely addresses SBL anomalies via spectral shape analysis
- Results from first two science runs (KNM1 + KNM2):
  - No significant sterile-neutrino signal observed
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Outlook:

- Analysis on data for first five science runs ongoing
- Stay tuned for upcoming release!









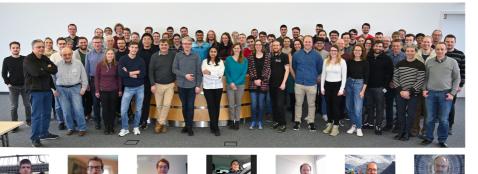












# Thank You

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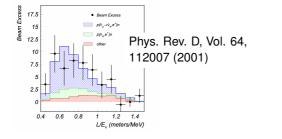
# Backups

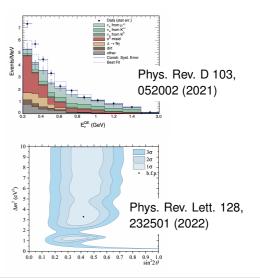
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#### **Experimental hints**

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# Testing applicability of Wilks' Theorem

- Generate O(10<sup>3</sup>) twins with statistical fluctuations for particular choice of MC truth
- Perform fitting for sterile parameter values on a grid and for MC truth for each sample (m<sup>2</sup><sub>ν</sub> = 0)
- Evaluate  $\Delta \chi^2 = \chi^2_{\rm MC \ truth} \chi^2_{\rm best \ fit}$  for each sample
- Compare distribution of  $\Delta \chi^2$  values to  $\chi^2$ -distribution with 2 dof

