



Monitoring the non-thermal Universe 2018





ASTRO WÜRZBURG

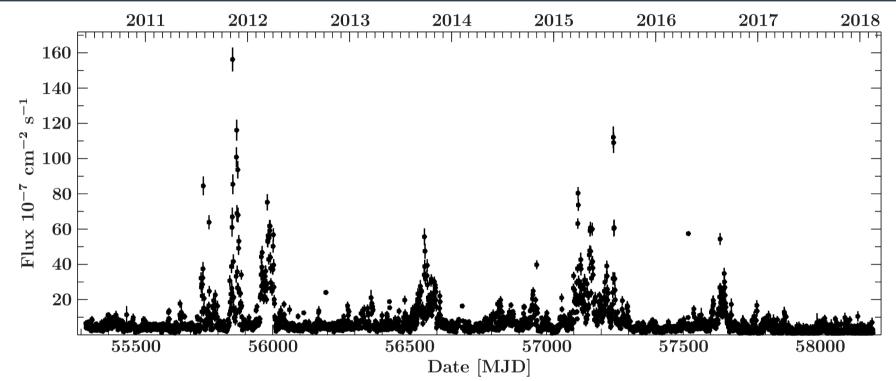
# Time-resolved SEDs of the variable blazar PKS 1510-089

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## The blazar PKS 1510-089



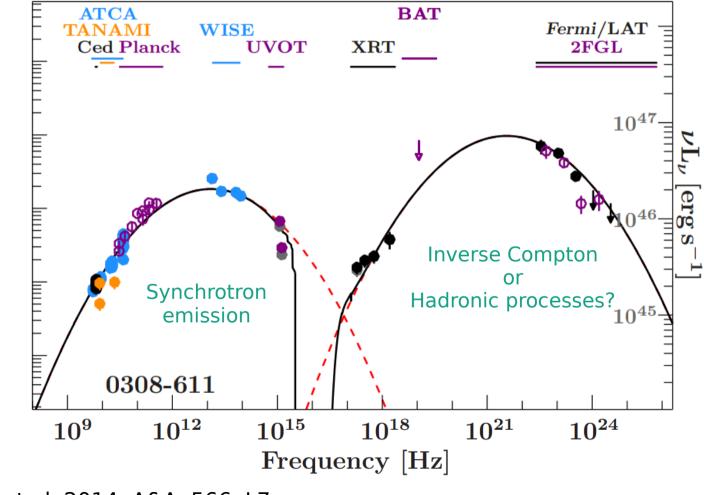
- FSRQ at *z* = 0.36 (Tanner et al. 1996)
- Second FSRQ being detected in VHE regime (Abramowski et al. 2013)



 One-zone scenario unlikely (e.g. Nalewajko et al. 2012, Brown 2013)

## **Spectral Energy Distribution (SED)**

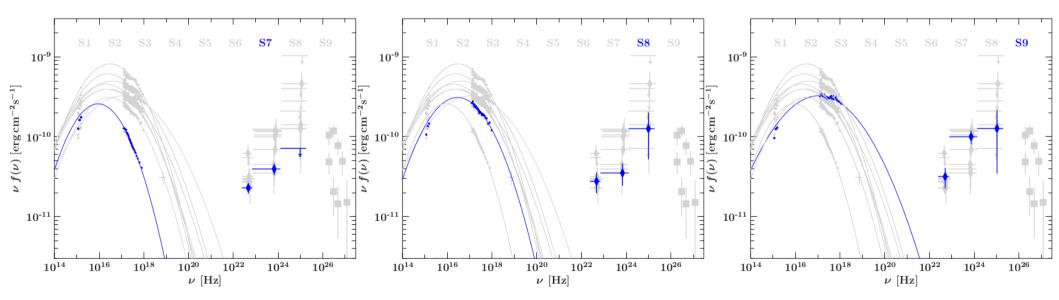
#### Broadband emission from Radio to Gamma-rays



Credit: Krauss et al. 2014, A&A, 566, L7

### **Changes in SEDs**

- Retrieving insights in the connection between flares and emission processes
- Example: Outburst of Mrk 421

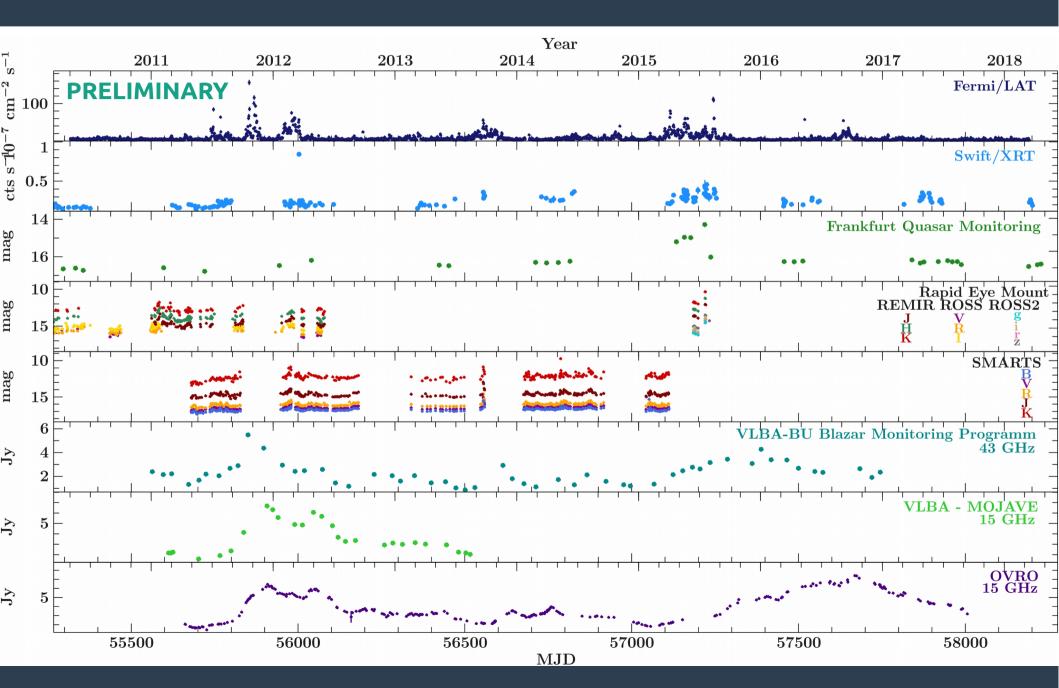


Credit: Kreikenbohm et al., in prep.

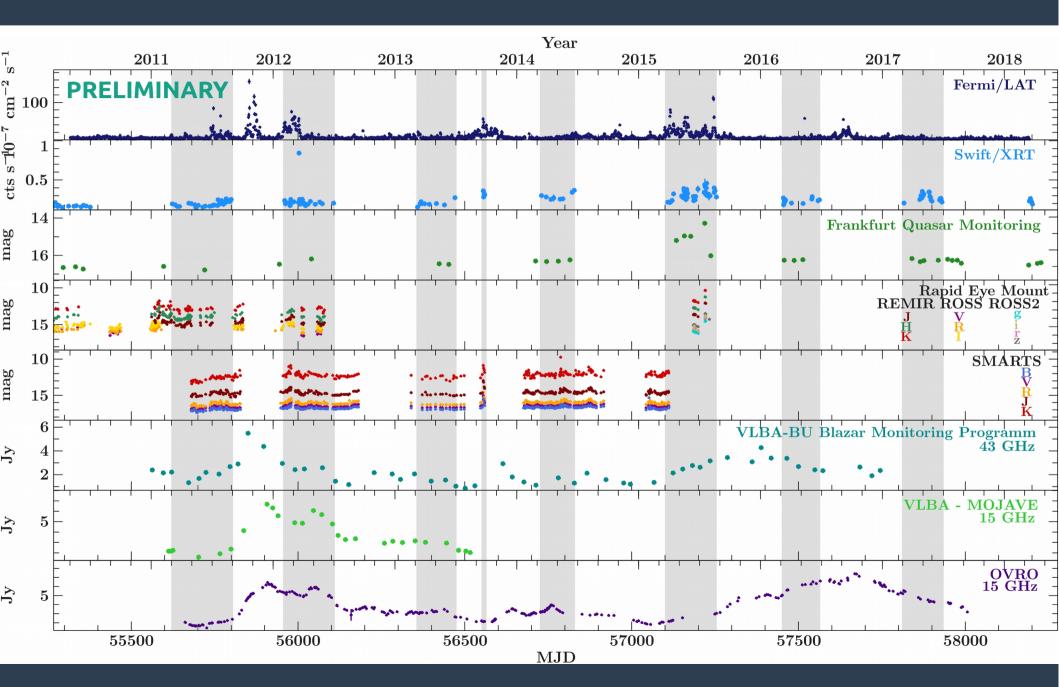
### All SED data from PKS 1510-089



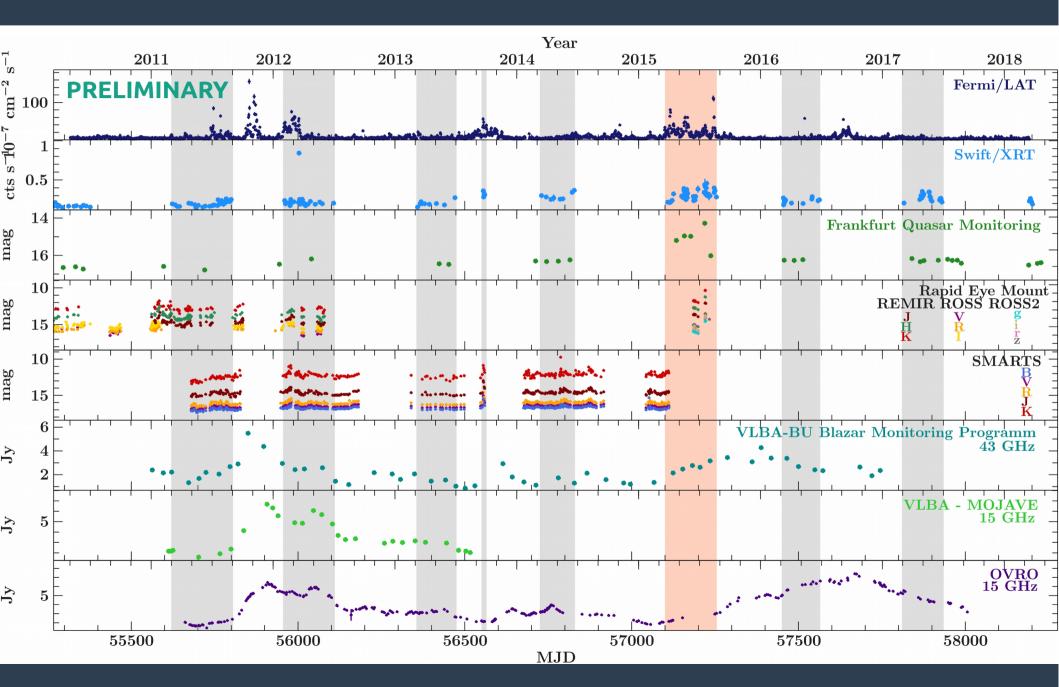
#### Long-term monitoring of PKS 1510-089



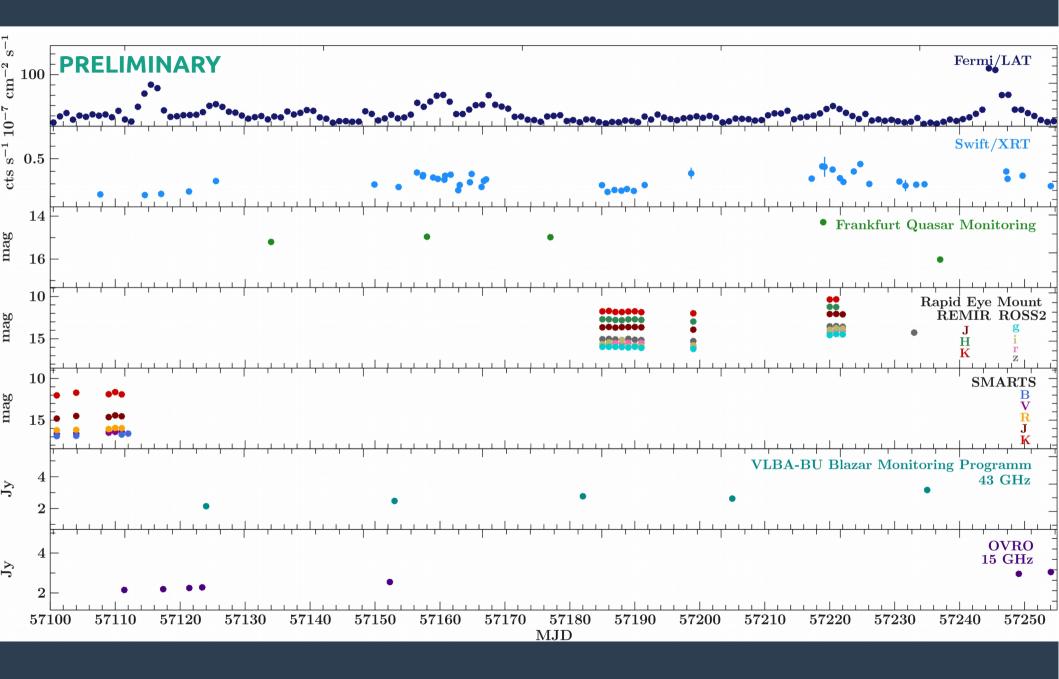
#### Long-term monitoring of PKS 1510-089



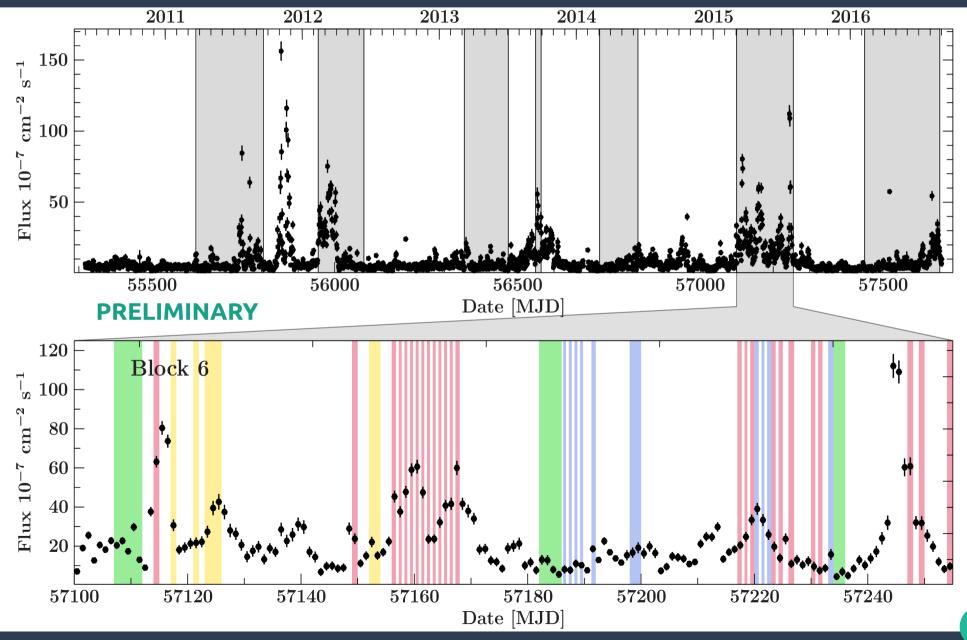
#### Long-term monitoring of PKS 1510-089



### Zoom in on selected time range

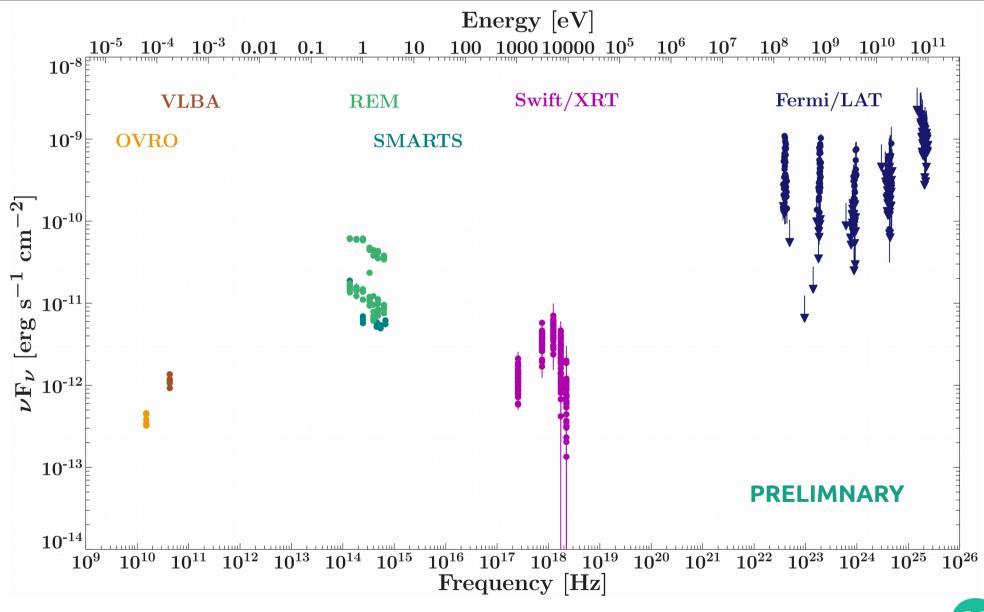


#### **Simultaneous observations**



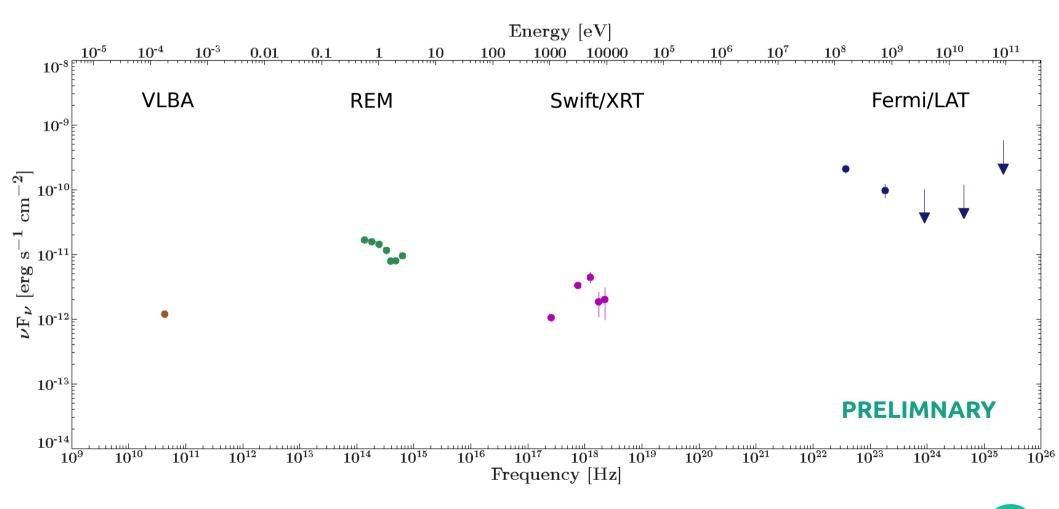
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#### Average SED for selected time range



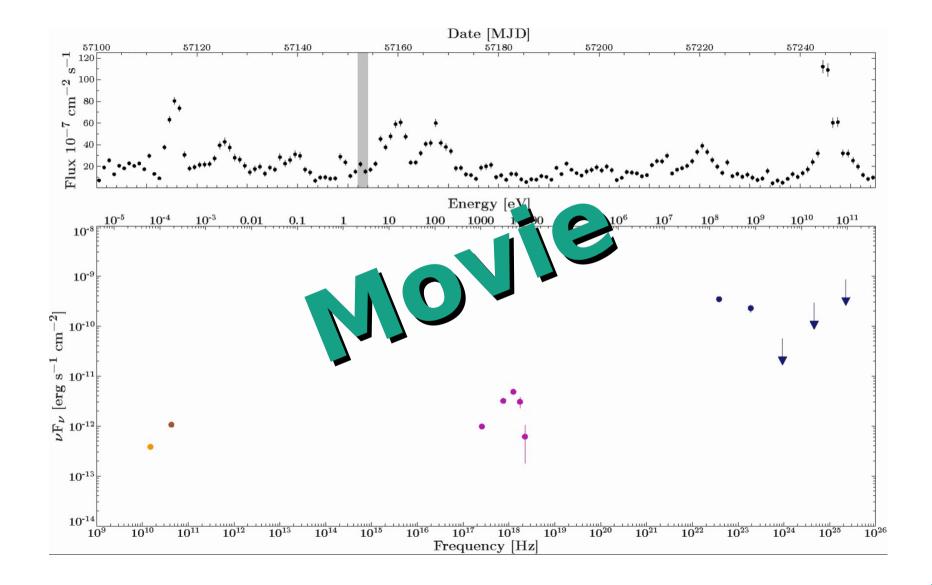
### SED of PKS 1510-089

#### Example SED from chosen time range



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### **Time-resolved SEDs**



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### Summary

- Selected time range shows changing behaviour in SED of PKS 1510-089
- Outlook:
  - Include more data (e.g. Swift/UVOT, H.E.S.S.?, MAGIC?)
  - Analysing and modeling the time-resolved SEDs for all selected time ranges
  - Testing for signs of hadronic emission processes
- Full broad range monitoring essential for understanding variability of blazars

