Towards an Autonomous Trigger for the Detection of Air-Shower Radio Emission

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DPG Spring Meeting 2024

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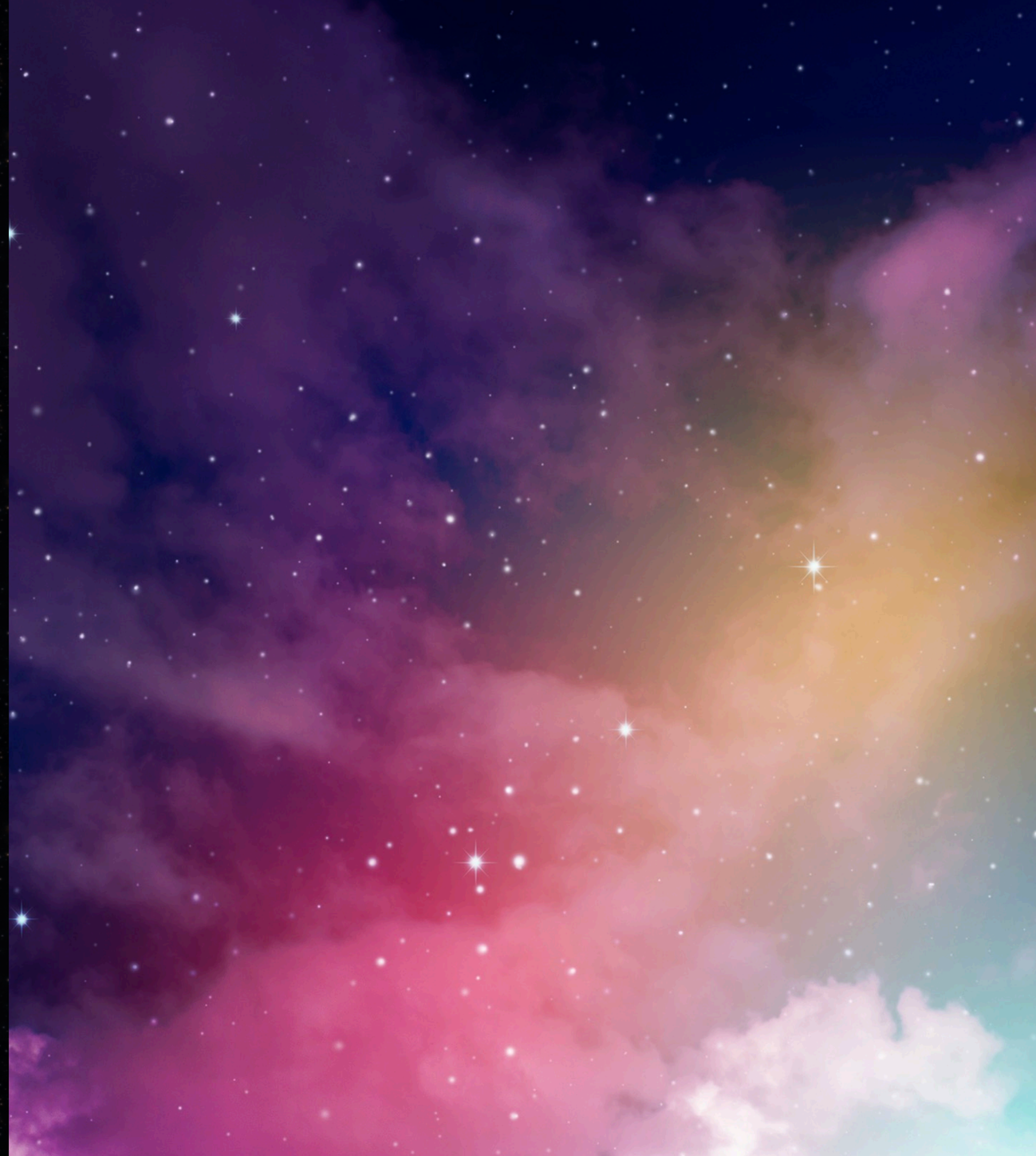
Physics Background: Cosmic Ray Air Showers



Extensive Air Showers

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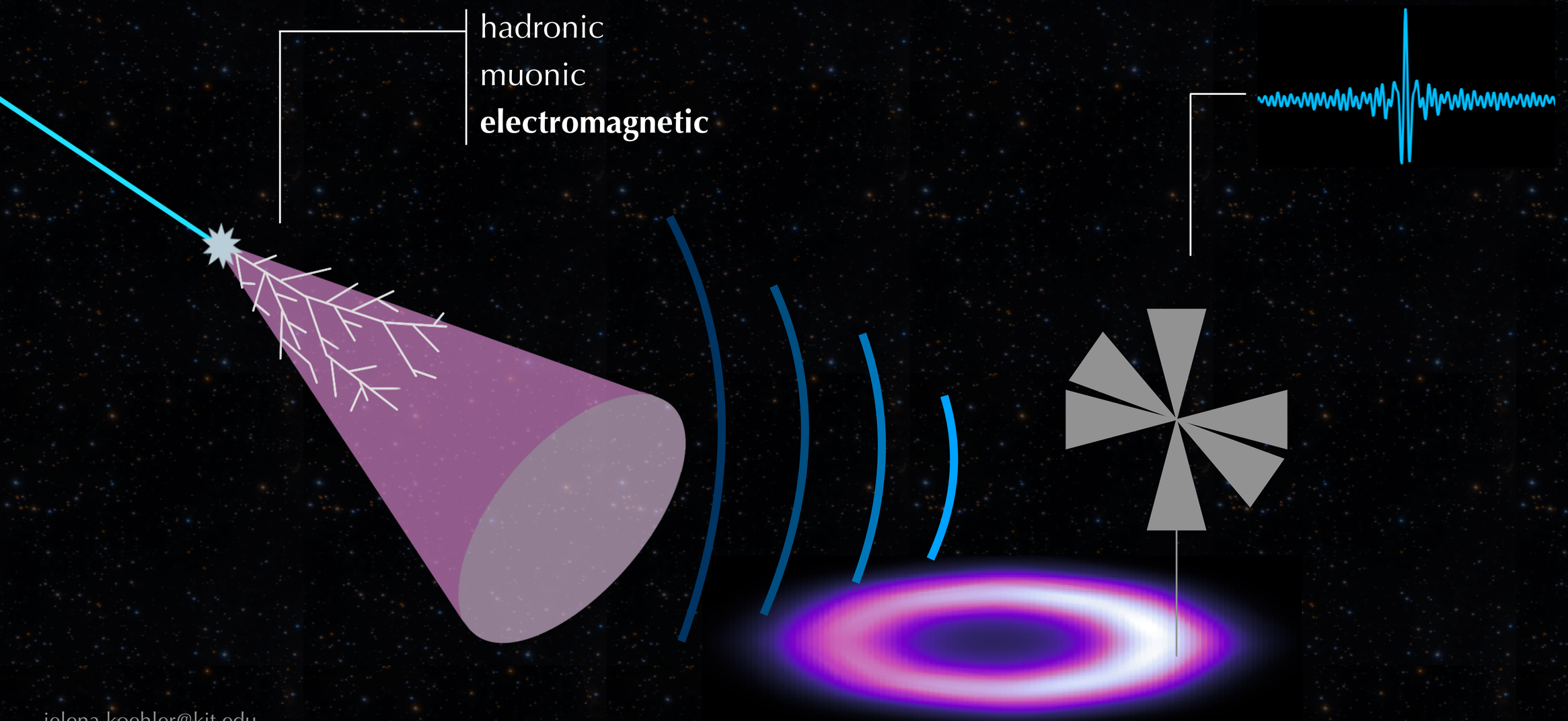
https://bigthink.com/starts-with-a-bang/speed-limit-below-speed-of-light/

Radio Emission of Air Showers

hadronic
muonic
electromagnetic



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GRAND
Giant Radio Array for Neutrino Detection

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Giant Radio Array for Neutrino Detection

- ★ CR and ν detection
- ★ pure radio experiment
- ★ combined detection area of 200 000 km^2
- ★ frequency band 50-200 MHz

GRAND
Site Status

GRAND@Auger

GRAND@Nançay

GP13 → GP80

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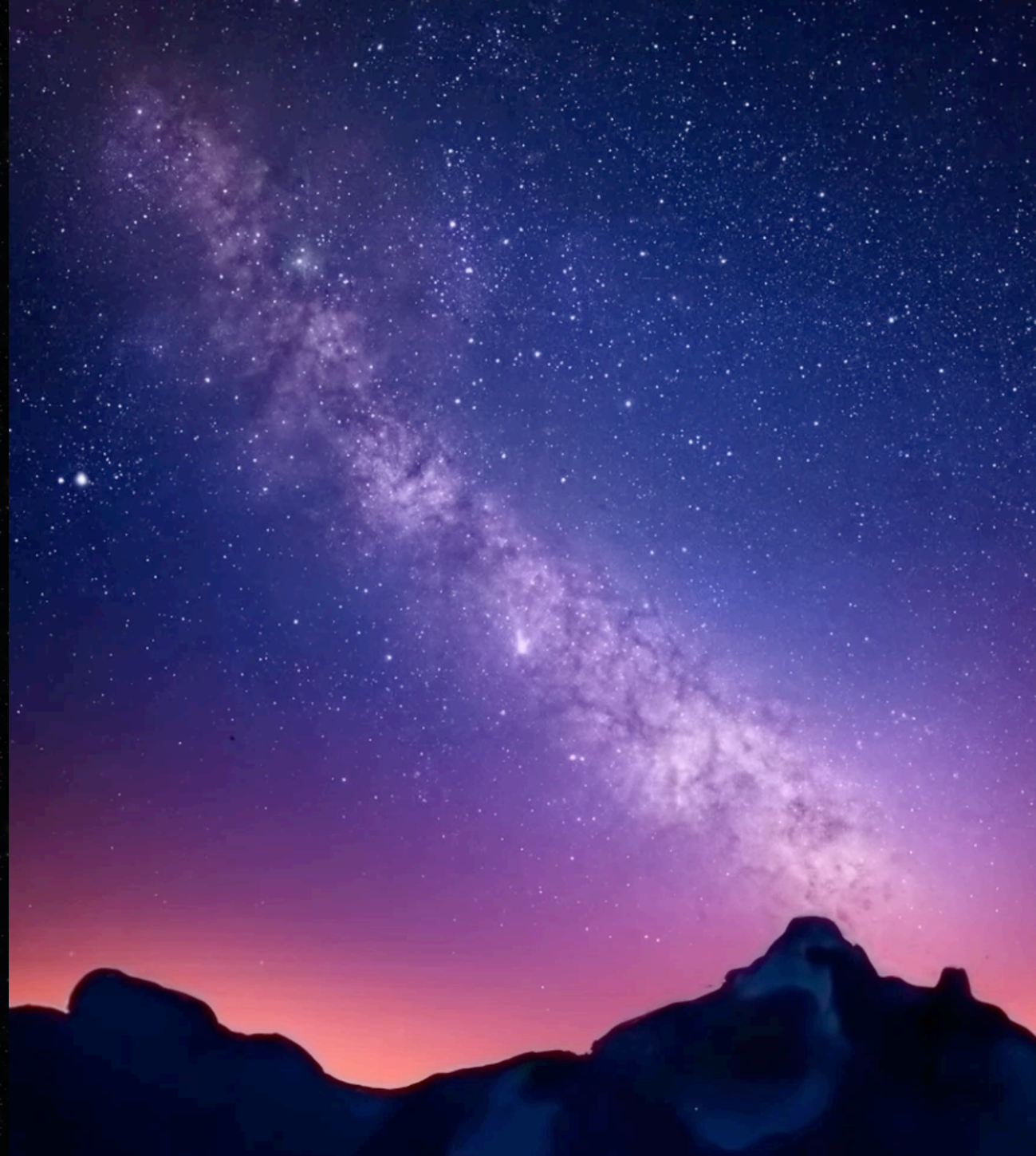
Developing an Autonomous Radio Trigger



Developing an Autonomous Radio Trigger

Multi-Level Trigger Systems

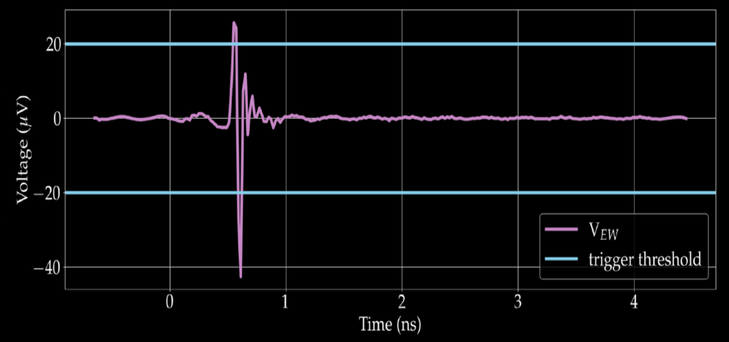
★ **step 0**: hardware trigger

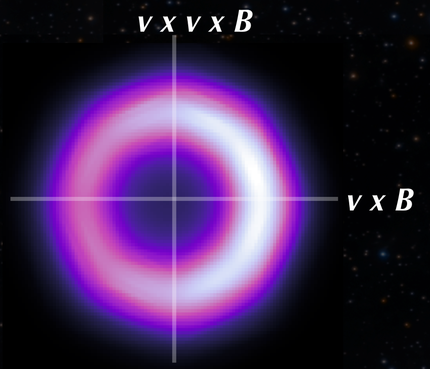
★ **step 1**: antenna level

- evaluate **individual antennas**

★ **step 2**: detector level

- evaluate **array of antennas**





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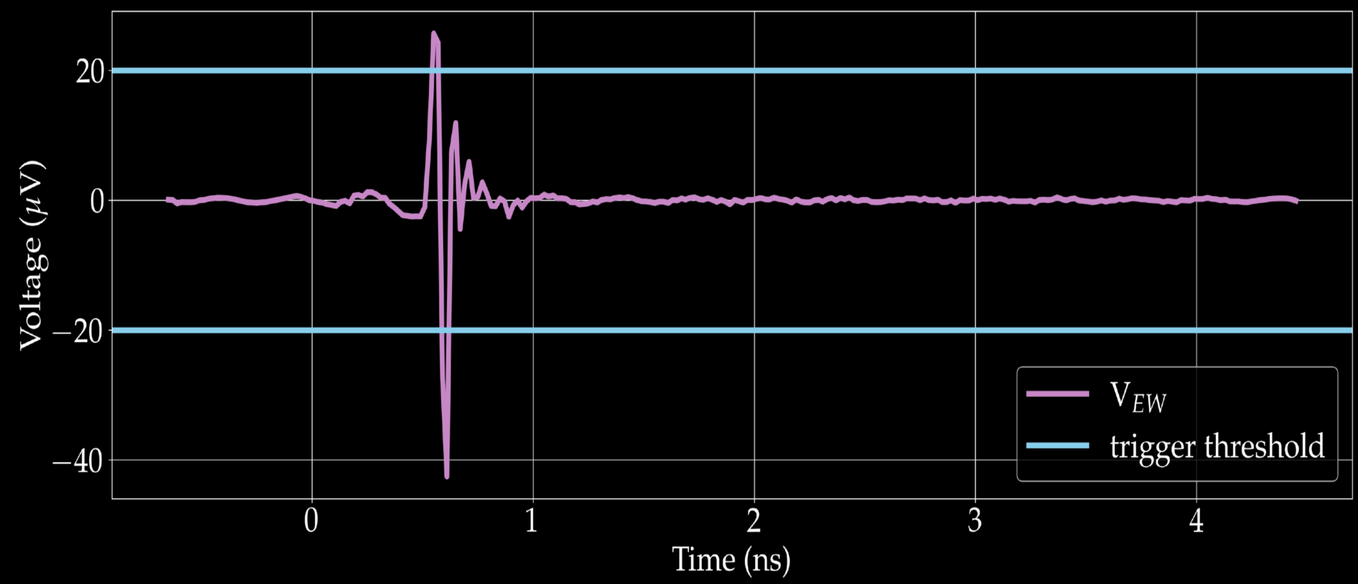
antenna level

Antenna Level Trigger

keep the classic methods

for each antenna



basic threshold trigger & template fitting

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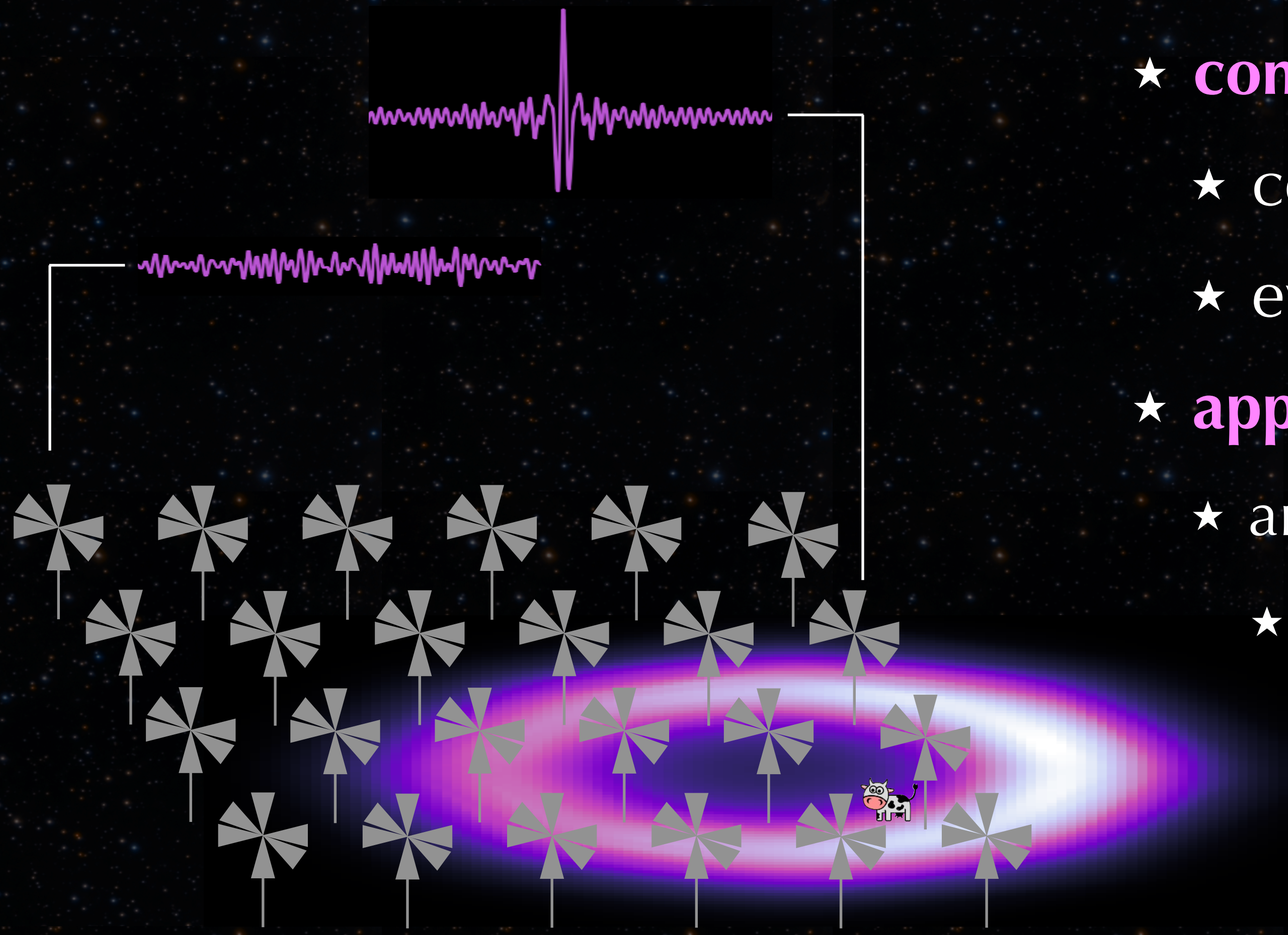
detector level

Event Level Trigger

add the new methods



- **combine antenna level info**
  - collect data from triggered antennas
  - evaluate distribution & signal arrival times
- **application**
  - analyze **radio emission patterns**
    - frequencies, energies, fluence, etc.

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Event Level Trigger Methods

Event Level Trigger Method

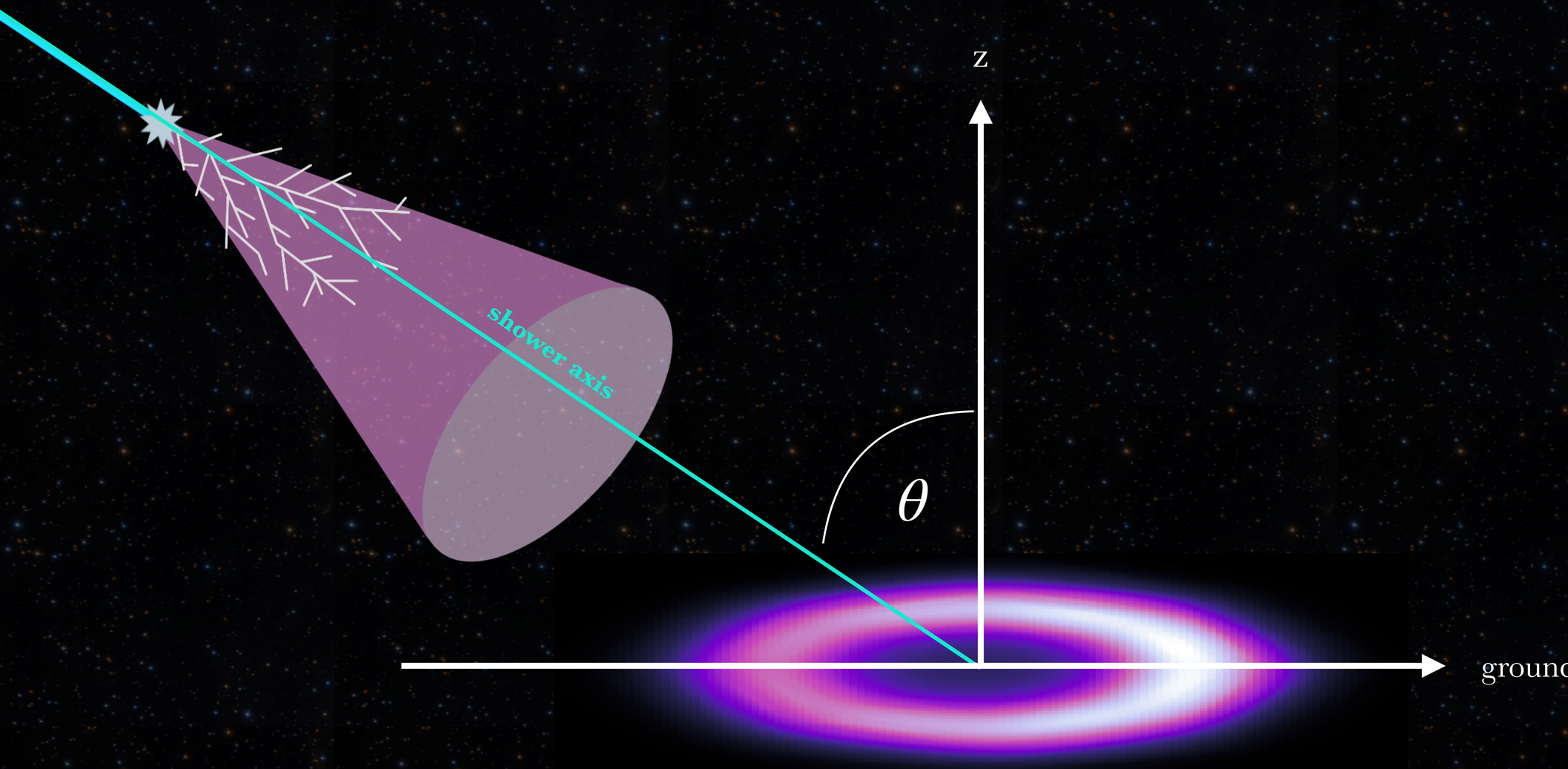
Do the reconstructed parameters match theoretical expectations?

- **trigger conditions:**
  - reconstructed **zenith angle** matches computed zenith angle
  - reconstructed **azimuth angle** matches computed azimuth angle
  - **signal arrival time distribution** matches theoretical expectations

- **real-life application:**
  - **template database** derived from simulations to approximate expected parameters in real time

Air Shower Parameters for Reconstruction

Zenith Angle



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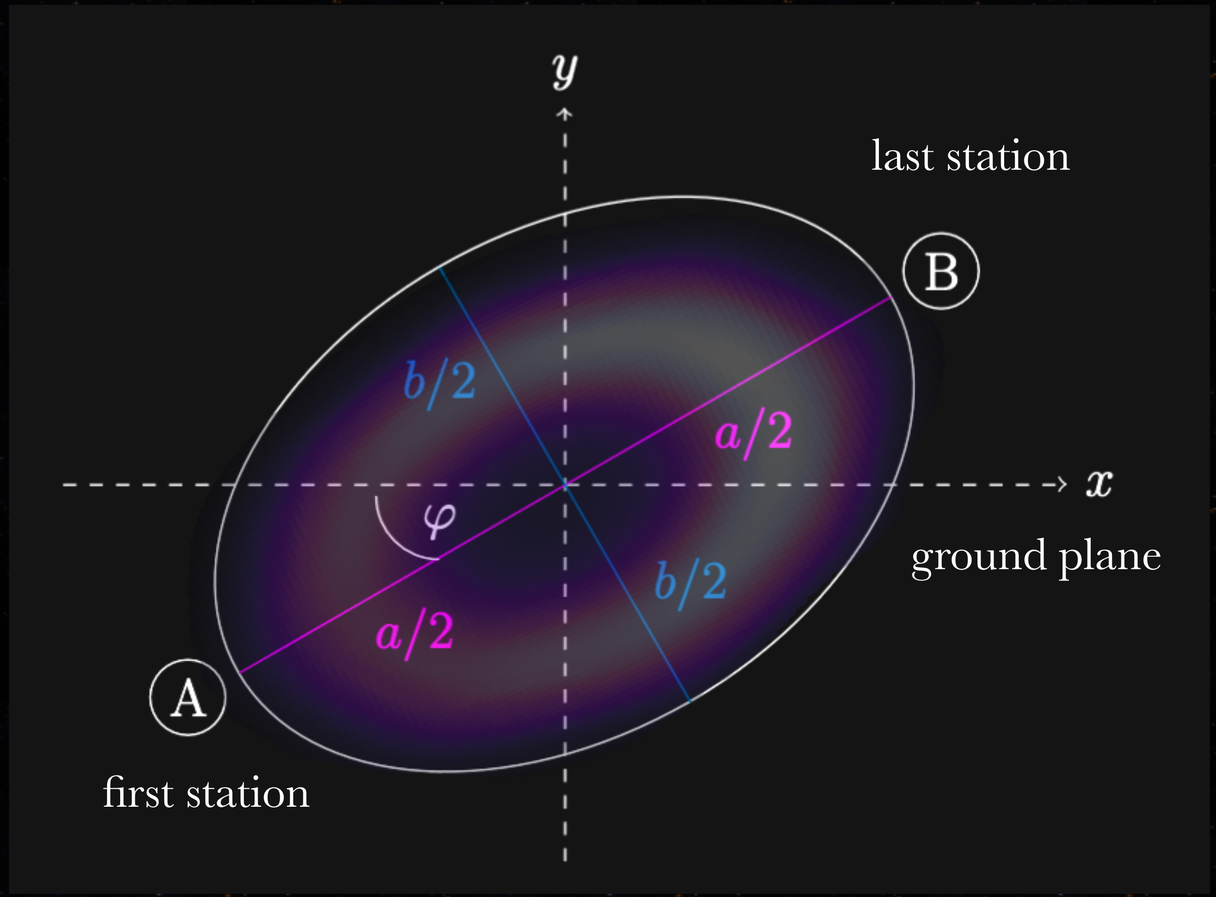
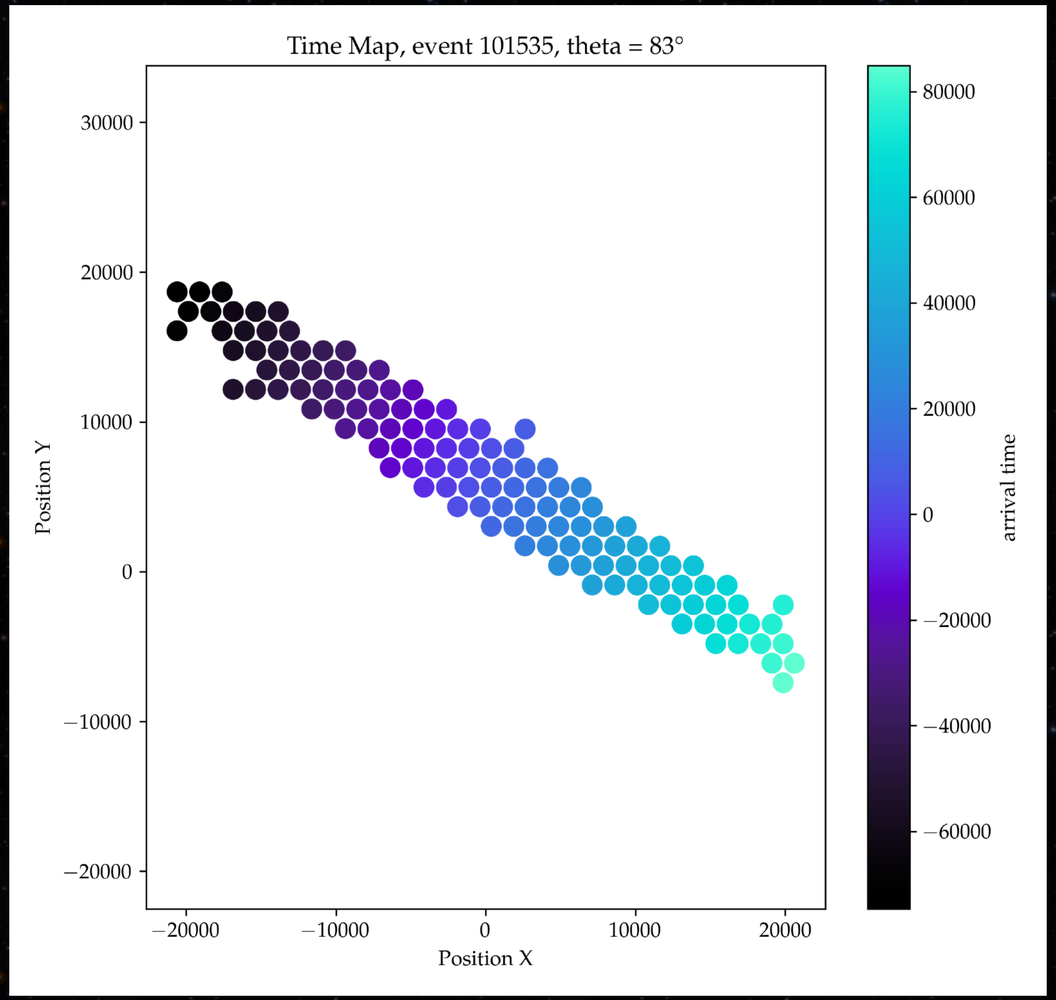
Air Shower Parameters for Reconstruction

Azimuth Angle



Air Shower Parameters for Reconstruction

Signal Arrival Times



Summary

- ★ autonomous radio triggers make cost-effective experiments at **large scales** possible
- ★ an **event-level trigger** enhances the classic methods with detailed understanding of radio emission

Outlook

- ★ **improve & validate** two-level trigger method in the next months using simulations & measurements

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Backup

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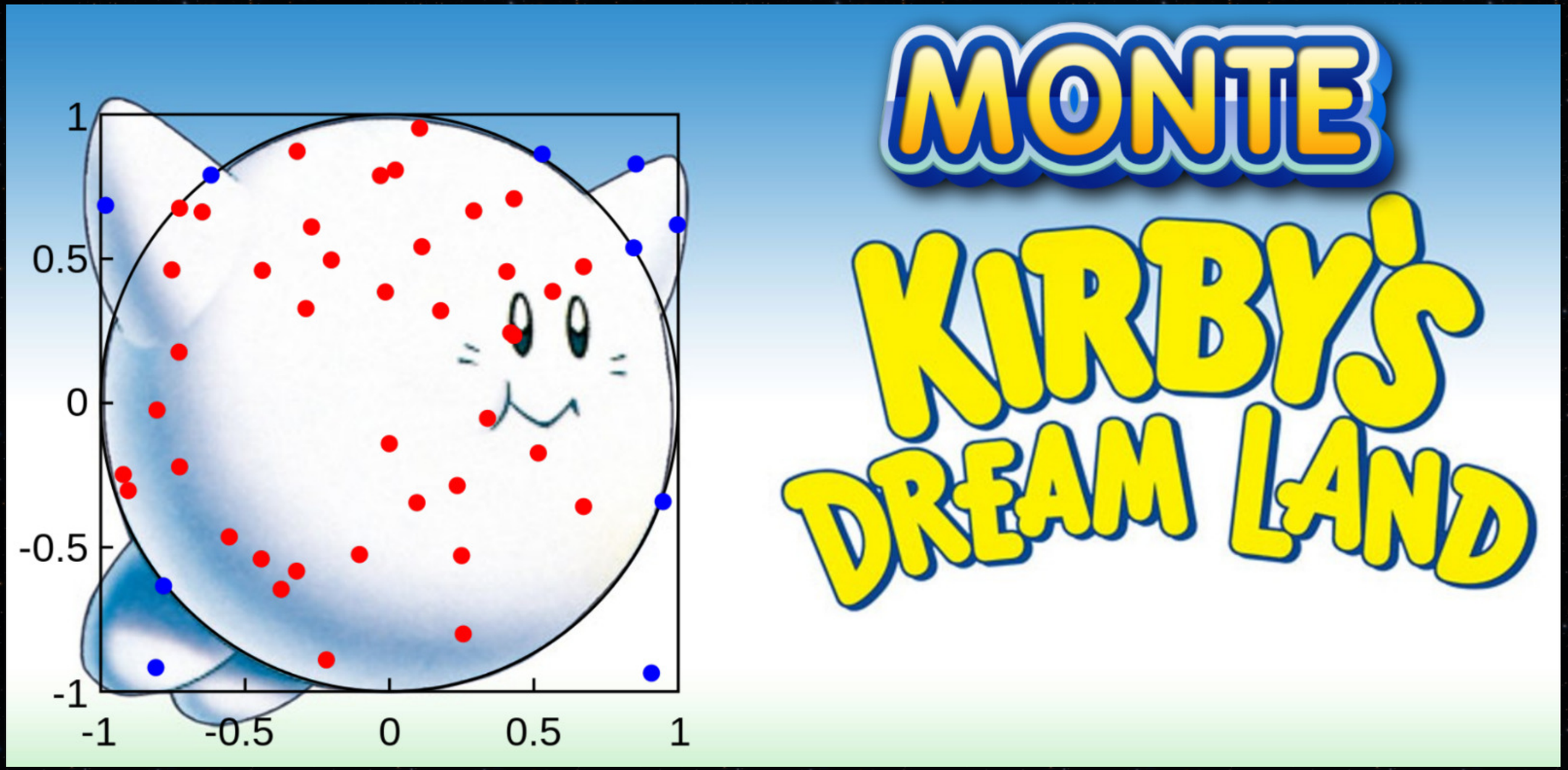
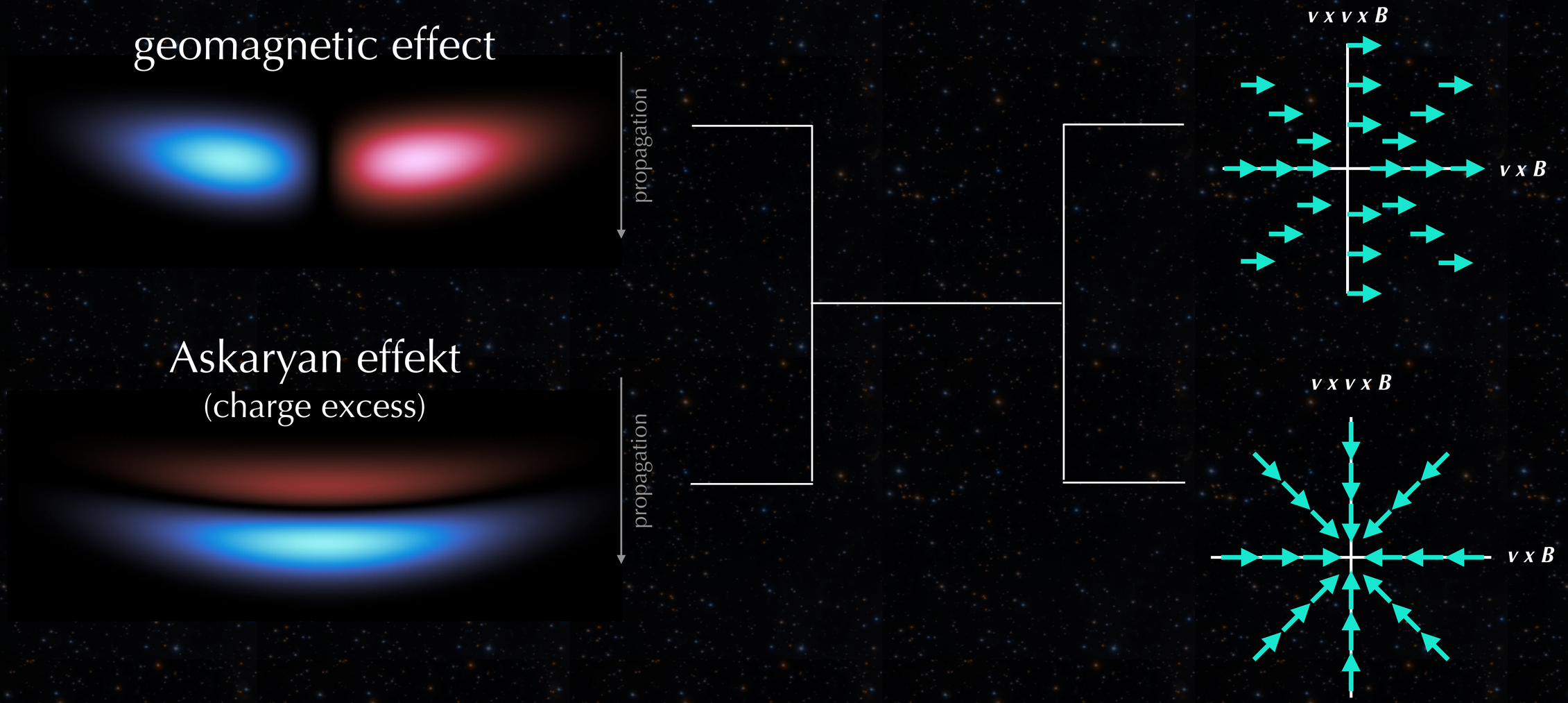

MONTE
KIRBY'S
DREAM LAND
1
0.5
0
-0.5
-1
-1
-0.5
0
0.5
1

Key Elements for New Trigger Methods

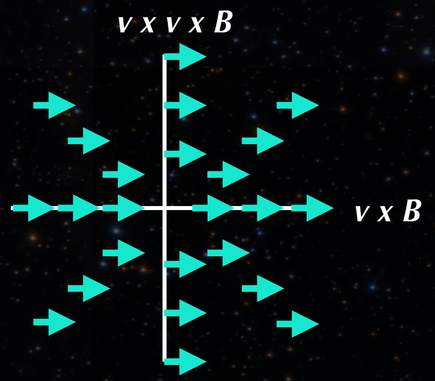
Understanding Radio Emission Patterns



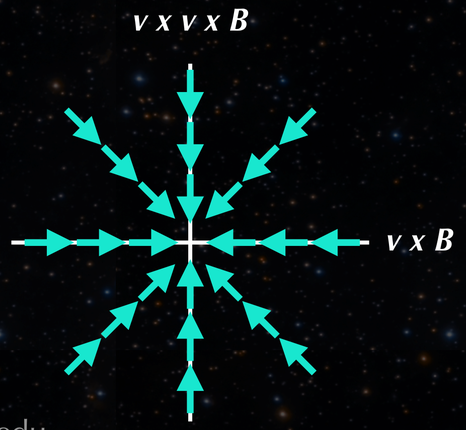
Key Elements for New Trigger Methods

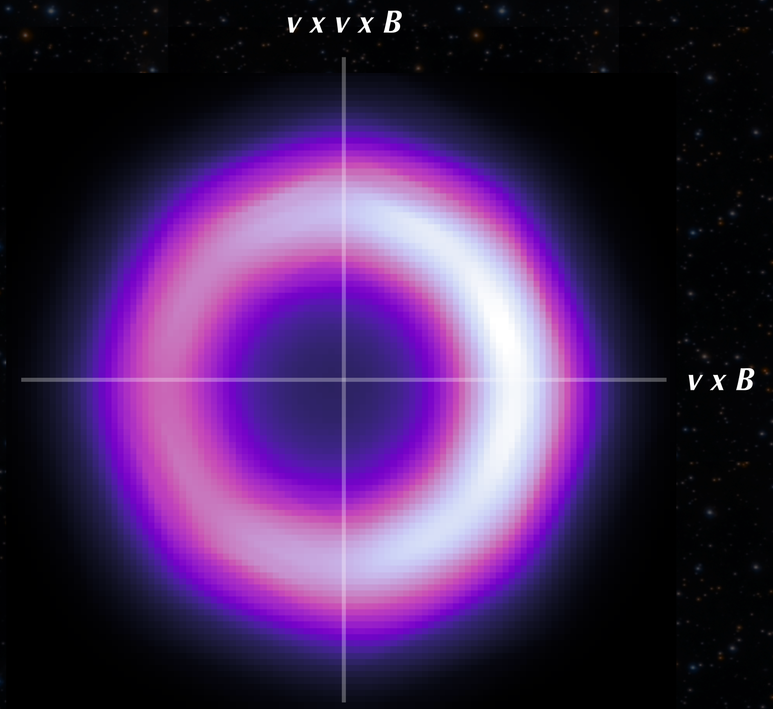
Understanding Radio Emission Patterns

geomagnetic effect



Askaryan effect
(charge excess)





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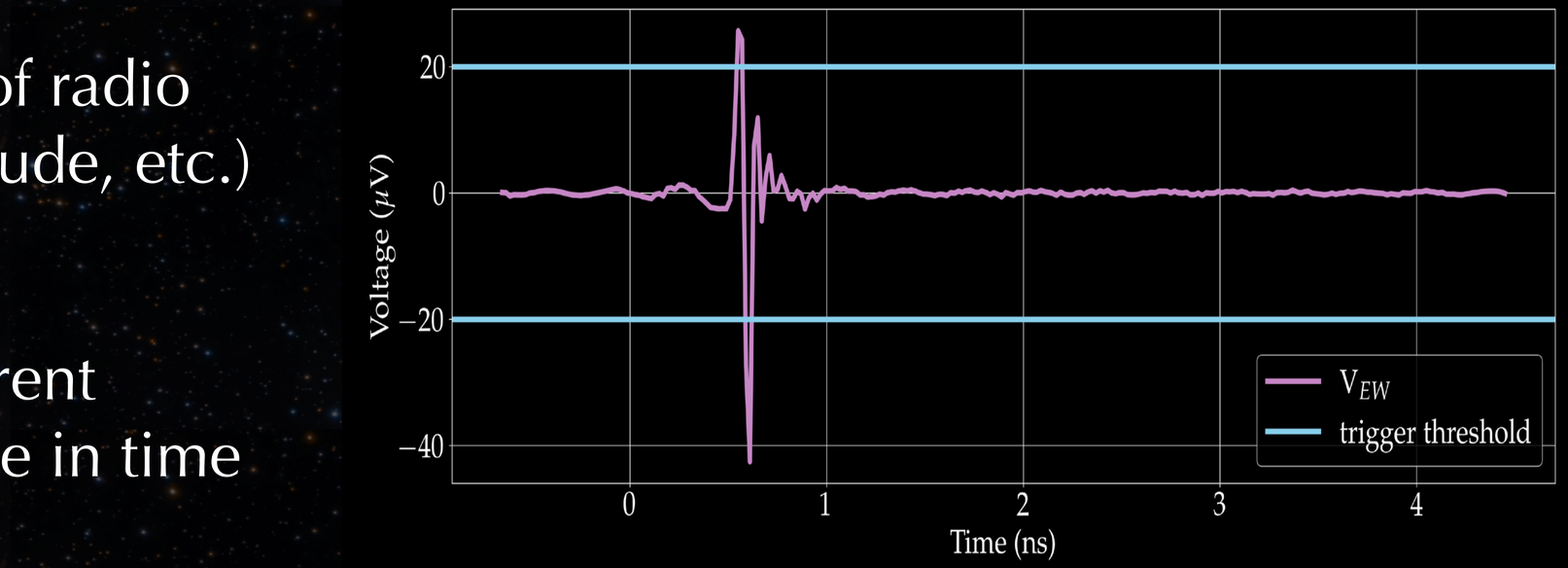
The Need for New Trigger Methods

Common Trigger Methods

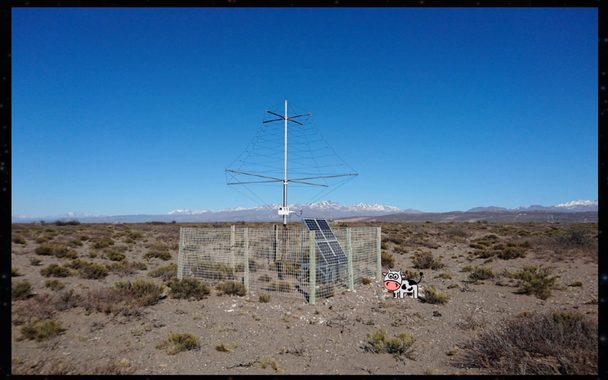
- **basic signal characteristics**
  - analyze basic shape of radio signals (width, amplitude, etc.)
- **temporal coincidence**
  - detect signals in different antennas that coincide in time
- **hybrid detection**
  - correlate radio signals with other detection methods



Limitations of Typical Radio Triggers

★ **strengths**

good performance for **very quiet regions**



effective in **small arrays** and in combination with **other detector types**

★ **challenges**

do not work for **noisier regions**

struggle to scale for **large arrays**

The Key to Larger Experiments

- ⋆ **advantages of radio**
  - ⋆ **cost-effective**: pure radio experiments are budget-friendly
  - ⋆ **scalability**: easily expandable for large arrays
- ⋆ **challenges of radio**
  - ⋆ **current trigger limitations**: existing triggers fall short for large experiments
- ⋆ **room for improvement**
  - ⋆ existing radio triggers do not apply the **understanding of radio** that we have

GRAND

Scale in Comparison to Auger's AERA

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GRAND

Scale in Comparison to LHC

LHC
27 km

GRAND
~450×450 km

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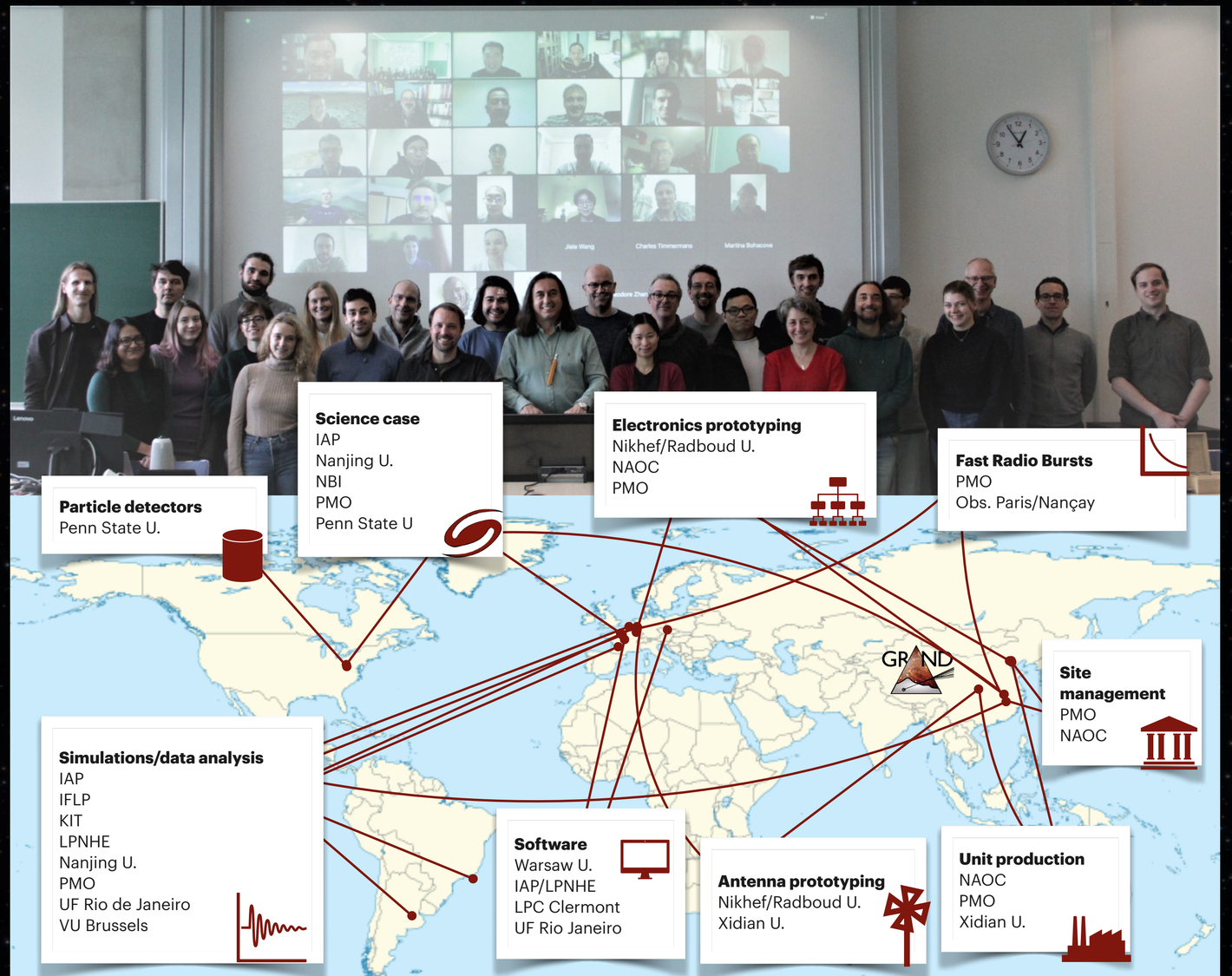
GRAND Today

93 collaborators from 12 countries



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