# Quasi-Dirac HNLs in the Left-Right Symmetric Model

## BLV 2024



Oleksii Mikulenko

Leiden University, the Netherlands

Quasi-Dirac HNLs in the LRSM

# Beyond the Standard Model



#### **Standard Model of Elementary Particles**

### Still missing:



Dark matter



Baryon asymmetry



#### Neutrino masses

# Heavy Neutral Leptons (HNL)

 Minimal solution – right-handed neutrinos with Majorana mass:

$$\mathcal{L}_N \supset -F_{\alpha I} \bar{L}_{\alpha} \tilde{H} N_I + \frac{M_I}{2} \bar{N}_I^c N_I + \text{h.c.},$$

• Mixing angles after spontaneous symmetry breaking

$$heta_{lpha I} = rac{F_{lpha I} v}{\sqrt{2} M_I}, \qquad U_{lpha I} = | heta_{lpha I}| \ll 1$$







- T

-

・ロン ・日 ・ ・ ヨン・

∃ ► Ξ|= <00</p>

• Two HNLs  $N_2$ ,  $N_3$  to produce two active neutrino masses via the seesaw mechanism. The mixing angles are at least

$$U^2 \gtrsim U_{
m seesaw}^2 \sim rac{m_
u}{m_N}$$

. • 1	2011	T	50

 EL OQO

• Two HNLs  $N_2$ ,  $N_3$  to produce two active neutrino masses via the seesaw mechanism. The mixing angles are at least

$$U^2 \gtrsim U_{
m seesaw}^2 \sim rac{m_
u}{m_N}$$

• The same two HNLs generate baryon asymmetry via leptogenesis. Bonus: HNLs can be as light as GeV if mass degenerate

= - 000

• Two HNLs  $N_2$ ,  $N_3$  to produce two active neutrino masses via the seesaw mechanism. The mixing angles are at least

$$U^2 \gtrsim U_{
m seesaw}^2 \sim rac{m_
u}{m_N}$$

- The same two HNLs generate baryon asymmetry via leptogenesis. Bonus: HNLs can be as light as GeV if mass degenerate
- One HNL  $N_1$  keV dark matter. Negligible contribution to neutrino masses from stability considerations

## Experimental constraints



[2204.08039]

Quasi-Dirac HNLs in the LRSM

BLV 2024

< □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > <

# Experimental constraints



Future experiments have the potential to observe millions of events

 *precision physics*

(日) (문) (문) (문) (문)

# Experimental constraints



Future experiments have the potential to observe millions of events

 $\implies$  precision physics

• (If we can have  $U^2 \gg U_{\text{seesaw}}^2$ )

## Quasi-Dirac limit

- Approximate lepton symmetry: cancellation of seesaw contributions
- The mixing angles for two degenerate HNLs converge to the same value

$$\Delta m_N \ll m_N, \qquad U_{\alpha 2}^2 = U_{\alpha 3}^2 \left[ 1 + O\left(\frac{U_{\text{seesaw}}^2}{U^2}, \frac{\Delta m_N}{m_N}\right) \right] \equiv U_{\alpha}^2$$



If a signal has been observed, what can we tell about the underlying physics of what has been found?

 ELE NOR

If a signal has been observed, what can we tell about the underlying physics of what has been found?

Have we solved any problem, or just added more?

# The SHiP experiment



Requirement	Value
Track momentum	> 1.0 GeV/c
Track pair distance of closest approach	< 1 cm
Track pair vertex position in decay volume	> 5 cm from inner wall
Impact parameter w.r.t. target (fully reconstructed)	< 10 cm
Impact parameter w.r.t. target (partially reconstructed)	< 250 cm

Background source	Expected events
Neutrino DIS	< 0.1 (fully) / $< 0.3$ (partially)
Muon DIS (factorisation)	$< 6 \times 10^{-4}$
Muon combinatorial	$1.2  imes 10^{-2}$

#### [2112.01487]

< □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > <



	Physics model	Final state
	HNL, SUSY neutralino	$\ell^{\pm}\pi^{\mp}, \ \ell^{\pm}K^{\mp}, \ \ell^{\pm}\rho^{\mp}(\rho^{\mp} \rightarrow \pi^{\mp}\pi^{0})$
	DP, DS, ALP (fermion coupling), SUSY sgoldstino	l+l-
DS	DP, DS, ALP (gluon coupling), SUSY sgoldstino	$\pi^{+}\pi^{-}, K^{+}K^{-}$
	HNL, SUSY neutralino, axino	$\ell^+\ell^-\nu$
	ALP (photon coupling), SUSY sgoldstino	77
	SUSY sgoldstino	$\pi^{0}\pi^{0}$
	LDM	Electron, proton, hadronic shower
D	$v_{\tau}, \overline{v}_{\tau}$ measurements	T <sup>±</sup>
	Neutrino-induced charm production $(v_e, v_{\mu}, v_{\tau})$	$D_{\delta}^{\pm}, D^{\pm}, D^{0}, \overline{D^{0}}, \Lambda_{c}^{+}, \overline{\Lambda_{c}}^{-}$

#### $m_N = 1.5 \, { m GeV}$



	decay mode	mixing	$\Gamma_{\alpha} \times 10^{13},  \text{GeV}$
0)	$N \rightarrow 3\nu$	$U_{e,\mu,\tau}^2$	1.7
1)	$N \rightarrow \nu ee$	$(U_e^2, U_{\mu,\tau}^2)$	(1.0, 0.2)
2)	$N \rightarrow \nu e \mu$	$U_{e,\mu}^{2}$	1.7
3)	$N \rightarrow \nu \mu \mu$	$(U^2_{\mu}, U^2_{e,\tau})$	(1.0, 0.2)
4)	$N \to \nu h^0 (\text{NC})$	$U_{e,\mu,\tau}^2$	2.5
5)	$N \to eh^+ (CC)$	$U_e^2$	5.0
6)	$N \to \mu h^+$ (CC)	$U_{\mu}^2$	5.0

 $Br_i \sim x_e \Gamma_e + x_\mu \Gamma_\mu + x_\tau \Gamma_\tau$ 

< ロ > < 同 > < 回 > < 回

三日 のへの

## Probe two-HNL seesaw



BLV 2024

A¶ ▶

10/29



< □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □



- T

# Measuring PMNS Majorana phase



$\sim$			
		m	20
0.0	-Noll		

**A →** 

13/29

# With such a tool, what can we say about **nonminimal interactions** of HNLs?

# With such a tool, what can we say about **nonminimal interactions** of HNLs?

For example:

- Dipole portal
- Axion-HNL coupling

• ...?

[1803.03262]

고 노

[1911.12394] [2212.11290]

# Add right-handed interactions: Left-Right Symmetric Model

$$\begin{split} \mathcal{S}\mathcal{U}_{\mathcal{C}}(3) imes \mathcal{S}\mathcal{U}_{\mathcal{L}}(2) imes \mathcal{S}\mathcal{U}_{\mathcal{R}}(2) imes \mathcal{U}_{\mathcal{B}-\mathcal{L}}(1) & 
ightarrow & \mathcal{S}\mathcal{U}_{\mathcal{C}}(3) imes \mathcal{S}\mathcal{U}_{\mathcal{L}}(2) imes \mathcal{U}_{\mathcal{Y}}(1) \ & 
ightarrow \mathcal{S}\mathcal{S}\mathcal{U}_{\mathcal{C}}(3) imes \mathcal{U}_{\mathsf{EM}}(1) \end{split}$$

• Effective (below EW) left

$$\begin{split} \mathcal{L} \supset \quad \theta_{\alpha I}^{L} \frac{\mathcal{G}_{F}}{\sqrt{2}} \bar{l}_{\alpha} \mathcal{N}_{I}^{c} \times [\bar{\nu}_{\beta} l_{\beta} + V_{ij}^{\mathsf{CKM}} \bar{u}_{i,L} d_{j,L}] \\ \qquad \qquad + \theta_{\alpha I}^{L} \frac{\mathcal{G}_{F}}{\sqrt{2}} \bar{\nu}_{\alpha} \mathcal{N}_{I}^{c} J_{Z} \end{split}$$



Quasi-Dirac HNLs in the LRSM

# Add right-handed interactions: Left-Right Symmetric Model

$$\begin{split} SU_{C}(3) imes SU_{L}(2) imes SU_{R}(2) imes U_{B-L}(1) &
ightarrow \ &
ightarrow SU_{C}(3) imes SU_{L}(2) imes U_{Y}(1) \ &
ightarrow SU_{C}(3) imes U_{EM}(1) \end{split}$$

• Effective (below EW) left

$$\begin{split} \mathcal{L} \supset \quad \theta^{L}_{\alpha I} \frac{G_{F}}{\sqrt{2}} \bar{l}_{\alpha} N^{c}_{I} \times [\bar{\nu}_{\beta} l_{\beta} + V^{\mathsf{CKM}}_{ij} \bar{u}_{i,L} d_{j,L}] \\ \quad + \theta^{L}_{\alpha I} \frac{G_{F}}{\sqrt{2}} \bar{\nu}_{\alpha} N^{c}_{I} J_{Z} \end{split}$$

• ... and right-handed interactions

$$+\theta_{\alpha I}^{R}\frac{G_{F}}{\sqrt{2}}\bar{l}_{\alpha}N_{I}\times\left[\tilde{V}_{J\beta}^{R}\bar{N}_{J}l_{\beta}+V_{ij}^{R,\mathsf{CKM}}u_{i,R}d_{i,R}\right]$$





Two sets of couplings  $|\theta| \ll 1$ 

(LH): 
$$\theta_{\alpha I}^L$$
, (RH):  $\theta_{\alpha I}^R \sim \frac{m_W^2}{m_{W_R}^2}$ 

				-
-	1.1.1.1		 ын	

Quasi-Dirac HNLs in the LRSM

BLV 2024

・ロト ・回ト ・ヨト・

16 / 29

∃ ► Ξ|= <00</p>

Two sets of couplings  $|\theta| \ll 1$ 

(LH): 
$$\theta_{\alpha I}^L$$
, (RH):  $\theta_{\alpha I}^R \sim \frac{m_W^2}{m_{W_R}^2}$ 

to be constrained by the seesaw relation:

$$m_{\nu} = - \underbrace{\theta^L M \theta^L}_{\text{type-I seesaw}} + \underbrace{\frac{v_L}{v_R} M}_{\text{type-II seesaw}}$$

- complicated  $3 \times 3$  matrix equation.

has closed analytic solution for  $\theta_L$ , if  $\theta_R$ ,  $m_N$  fixed [2403.07756]

0					
0	ек	SIL	ĸu	en	ко

ELE DOG

$$\begin{split} \mathcal{L} \supset \bar{L}_{\alpha}([Y_{e}]_{\alpha\beta}\Phi - [Y_{\nu}]_{\alpha\beta}\sigma_{2}\Phi^{*}\sigma_{2})R_{\beta} + \\ &+ \bar{L}_{\alpha}^{c}[Y_{1}]_{\alpha\beta}i\sigma_{2}\Delta_{L}L_{\beta} + \bar{R}_{\alpha}^{c}[Y_{2}]_{\alpha\beta}i\sigma_{2}\Delta_{R}R_{\beta} + \text{h.c.} \\ \Phi \rightarrow v \operatorname{diag}(\cos b, -\sin b \, e^{-ia}) \qquad \Delta_{L,R} \rightarrow \begin{pmatrix} 0 & 0 \\ v_{L,R} & 0 \end{pmatrix} \end{split}$$

Generalized parity:  $Y_e^\dagger = Y_e, \quad Y_\nu^\dagger = Y_\nu, \quad Y_1 = Y_2$ 

After spontaneous symmetry breaking and diagonalization of I, N masses:

$$U_{\text{PMNS}}^{\text{diag}} M_{\nu}^{\text{diag}} U_{\text{PMNS}}^{\dagger} = -(vV_RY - m_l^{\text{diag}}be^{ia}V_R)[m_N^{\text{diag}}]^{-1}(vY^TV_R^T - V_R^Tm_l^{\text{diag}}be^{ia}) + \frac{v_L}{v_R}V_R^*m_N^{\text{diag}}V_R^{\dagger}$$
with  $Y = V_P^{\dagger}Y_{\nu}V_R$ ,  $Y^{\dagger} = Y$ ,  $V_P^{\dagger} = V_R$ 

◆□▶ ◆□▶ ◆三▶ ◆三▶ 三三 のの()

$$U_{\rm PMNS}^{*} m_{\nu}^{\rm diag} U_{\rm PMNS}^{\dagger} = -(vV_RY - m_l^{\rm diag}be^{ia}V_R)[m_N^{\rm diag}]^{-1}(vY^TV_R^T - V_R^Tm_l^{\rm diag}be^{ia}) + \frac{v_L}{v_R}V_R^*m_N^{\rm diag}V_R^{\dagger}$$

$$\theta_{\alpha I}^{R} = \frac{m_{W_{L}}^{2}}{m_{W_{R}}^{2}} [V_{R}]_{\alpha I}, \qquad \theta_{\alpha I}^{L} = \frac{i}{m_{N_{I}}} \left[ v V_{R} Y - b e^{ia} m_{I}^{\text{diag}} V_{R} \right]_{\alpha I}$$

Oleksii Mikulenko

Quasi-Dirac HNLs in the LRSM

シック 正正 《田》《田》《田》 《日》 BLV 2024

$$U_{\rm PMNS}^{*} m_{\nu}^{\rm diag} U_{\rm PMNS}^{\dagger} = -(vV_RY - m_l^{\rm diag}be^{ia}V_R)[m_N^{\rm diag}]^{-1}(vY^TV_R^T - V_R^Tm_l^{\rm diag}be^{ia}) + \frac{v_L}{v_R}V_R^*m_N^{\rm diag}V_R^{\dagger}$$

$$\theta_{\alpha I}^{R} = \frac{m_{W_{L}}^{2}}{m_{W_{R}}^{2}} [V_{R}]_{\alpha I}, \qquad \theta_{\alpha I}^{L} = \frac{i}{m_{N_{I}}} \left[ v V_{R} Y - b e^{ia} m_{I}^{\text{diag}} V_{R} \right]_{\alpha I}$$

• Assume 2 quasi-Dirac pair N<sub>2</sub>, N<sub>3</sub> ( $|\theta_{L,2/3}|^2 \gg U_{\text{seesaw}}^2$ ) and a decoupled DM candidate ( $|\theta_{L,1}|^2 \ll U_{\text{seesaw}}^2$ )

$$m_N^{\text{diag}} = m_N \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}, \qquad Y = y \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & -i \\ 0 & i & 1 \end{pmatrix}, \quad y \gg 1$$

• and perturb this exact lepton symmetry

Oleksii Mikulenko

Quasi-Dirac HNLs in the LRSM

BLV 2024

・ロト ・四ト ・ヨト ・ヨヨ

# Type-I only: analytic solution

Neutrino masses: 0, 
$$m_2$$
,  $m_3$   
 $V^R = iU^*_{PMNS}O$ 

$$V_{\alpha 2}^{L} = -iV_{\alpha 3}^{L} = \frac{e^{i\beta}}{\sqrt{m_{2} + m_{3}}} \tilde{U}_{\mathsf{PMNS}}^{*} P \times \begin{pmatrix} 0\\ \mp e^{-i\eta}\sqrt{m_{2}}\\ \sqrt{m_{3}} \end{pmatrix}, \quad \frac{|\theta_{\alpha}^{L}|^{2}}{\sum_{\alpha} |\theta_{\alpha}^{L}|^{2}} = |V_{\alpha}^{L}|^{2}$$

with

$$O = \frac{1}{\sqrt{2(m_2 + m_3)}} \times \\ \times \begin{pmatrix} \sqrt{2(m_2 + m_3)} & 0 & 0 \\ 0 & -i(\sqrt{m_3}e^{-i\beta} \pm \sqrt{m_2}e^{i\beta}) & \sqrt{m_3}e^{-i\beta} \mp \sqrt{m_2}e^{i\beta} \\ 0 & -(\sqrt{m_3}e^{i\beta} \mp \sqrt{m_2}e^{-i\beta}) & i(\sqrt{m_3}e^{i\beta} \pm \sqrt{m_2}e^{-i\beta}) \end{pmatrix}$$

- two free parameters: Majorana phase  $\eta$  and angle  $\beta$ 





- $V^L_{lpha}$  remain **the same** as in the minimal case only depend on  $\eta$
- $V^R_{\alpha 2}$ ,  $V^R_{\alpha 3}$  depend on both  $\eta$ ,  $\beta$
- $|V_{\alpha 1}^R|^2$ ,  $|V_{\alpha 1}^R|^2 + |V_{\alpha 1}^L|^2$  are fixed
- HNL mass splitting  $|m_{N_2} m_{N_3}|$  is arbitrary

## Corrections

# Type-II corrections $\kappa = \frac{v_L m_N}{v_R (m_2 + m_3)}$





Quasi-Dirac HNLs in the LRSM

BLV 2024

## Interesting signatures?

▲ @ ▶ < ≥ ▶</p>

∃ ► Ξ|= <00</p>

• Decoherent pair:

number of ev.  $(X \to I_{\alpha}N \to I_{\alpha}I_{\beta}) \propto |V_{\alpha 2}|^2 |V_{\beta 3}|^2 + |V_{\alpha 3}|^2 |V_{\beta 2}|^2$ 

• Coherent pair:

number of ev. 
$$(X \to I_{\alpha}N \to I_{\alpha}I_{\beta}^{\pm}) \propto |V_{\alpha 2}V_{\beta 3}^{(*)} + V_{\alpha 3}V_{\beta 2}^{(*)}|^2$$

Summing up over one lepton flavor ( $\alpha$ ) — everything reduces to  $\propto |V_{\beta 2}|^2 + |V_{\beta 3}|^2$ 

0	leksii	Mikı	ılen	ko	

◆□▶ ◆□▶ ◆三▶ ◆三▶ 三三 のの()

### Probing decoherence at SHiP

• The initial lepton is lost in the target — no information?

-						
0	п	ksii	- Mit	ĸп	len	ko

#### Probing decoherence at SHiP

- The initial lepton is lost in the target no information?
- Not if kinematic constraints help us
- For example,  $D \rightarrow I_{\alpha} N \rightarrow I_{\alpha} I_{\beta}$  cannot have  $\tau$ -leptons for a GeV HNL

イロト (周) (ヨト (ヨト ) 三日 ののの

#### Probing decoherence at SHiP

- The initial lepton is lost in the target no information?
- Not if kinematic constraints help us
- For example,  $D \to l_{\alpha} N \to l_{\alpha} l_{\beta}$  cannot have  $\tau$ -leptons for a GeV HNL



#### Probing decoherence with Keung-Senjanović process

- Full reconstruction of event matrix  $X \to I_{\alpha}^{-} N \to I_{\alpha}^{-} I_{\beta}^{\pm}$
- In the coherent case, LNV can dominate LNC decays



< 🗇 🕨



-

## DM at SHiP



#### No $N_{2,3} \rightarrow N_1$ decay in the minimal LH case



## DM at SHiP

• Benchmark model:

$$\begin{split} |V_e^L|^2 &: |V_{\mu}^L|^2 : |V_{\tau}^L|^2 = 0.11 : 0.22 : 0.67 \\ |V_{e2}^R|^2 &: |V_{\mu2}^R|^2 : |V_{\tau2}^R|^2 = 0.16 : 0.46 : 0.38 \\ |V_{e3}^R|^2 &: |V_{\mu3}^R|^2 : |V_{\tau3}^R|^2 = 0.16 : 0.46 : 0.38 \\ |V_{e1}^R|^2 &: |V_{\mu3}^R|^2 : |V_{\tau3}^R|^2 = 0.49 : 0.22 : 0.30 \end{split}$$

Fraction of RH interactions:

$$\mathcal{R}\equiv rac{U_R^2}{U_L^2+U_R^2}$$

For a given R < 1, we want to distinguish</li>
no light N<sub>1</sub> versus LH-only HNL with arbitrary V<sup>L</sup>
with light N<sub>1</sub> versus LH-only HNL with arbitrary V<sup>L</sup>,
with light N<sub>1</sub> versus HNL with both LH, RH-interactions, arbitrary couplings V<sup>L</sup>, V<sup>R</sup>, R, but no N<sub>1</sub>.



▲ @ ▶ < ≥ ▶</p>

포네크



- Precision physics at Intensity Frontier is possible
- We need to know what results to expect and how to interpret them

Oleksii Mikulenko

Quasi-Dirac HNLs in the LRSM

BLV 2024

글 🖌 글 글

# Back-up


Quasi-Dirac HNLs in the LRSM

(本間) (本語) (本語)

▶ Ξ = ୬ ۹ ୧

FCC-ee



BLV 2024

< 17 ►

2/2

三日 のへの