

Probing the Shadows: A Search for Dark Photons (10-40 GeV) Using Scouting Data

Analysis review

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1) Minimal Dark Photon Model

- **Minimal Dark Photon Model:**

- Extends the Standard Model (SM) with a new U(1) gauge symmetry in the dark sector.
- Introduces a dark photon (Z') that interacts with SM particles via kinetic mixing.

- **Kinetic Mixing:**

- Lagrangian term: $\mathcal{L} \supset -\frac{\epsilon}{2} F_{\mu\nu} F'^{\mu\nu}$
- ϵ is the kinetic mixing parameter allowing dark photon to couple to SM particles.

- **Free Parameters:**

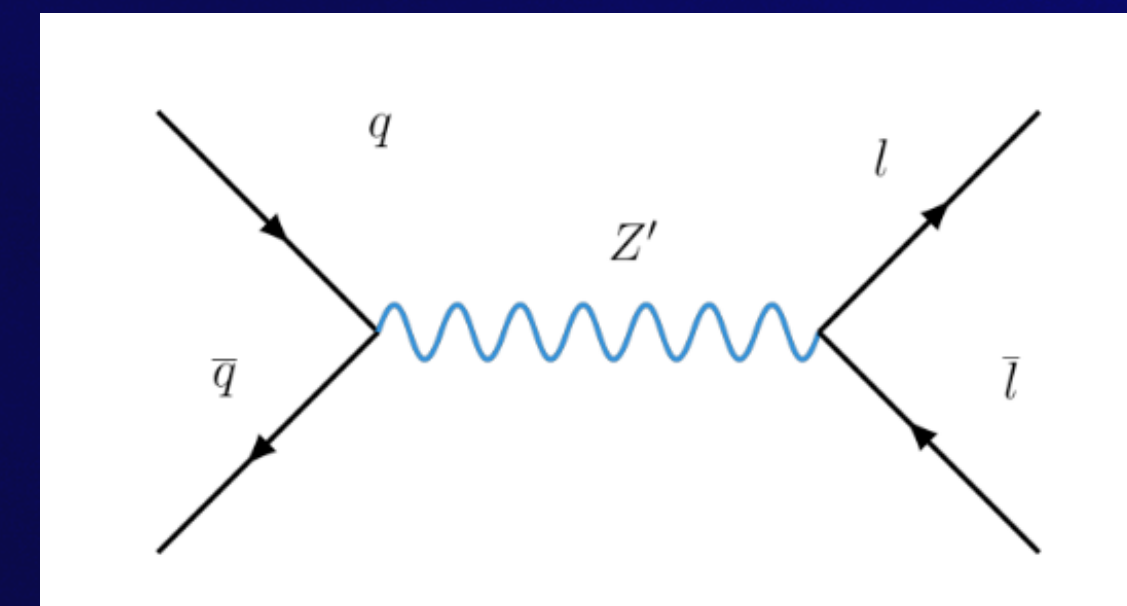
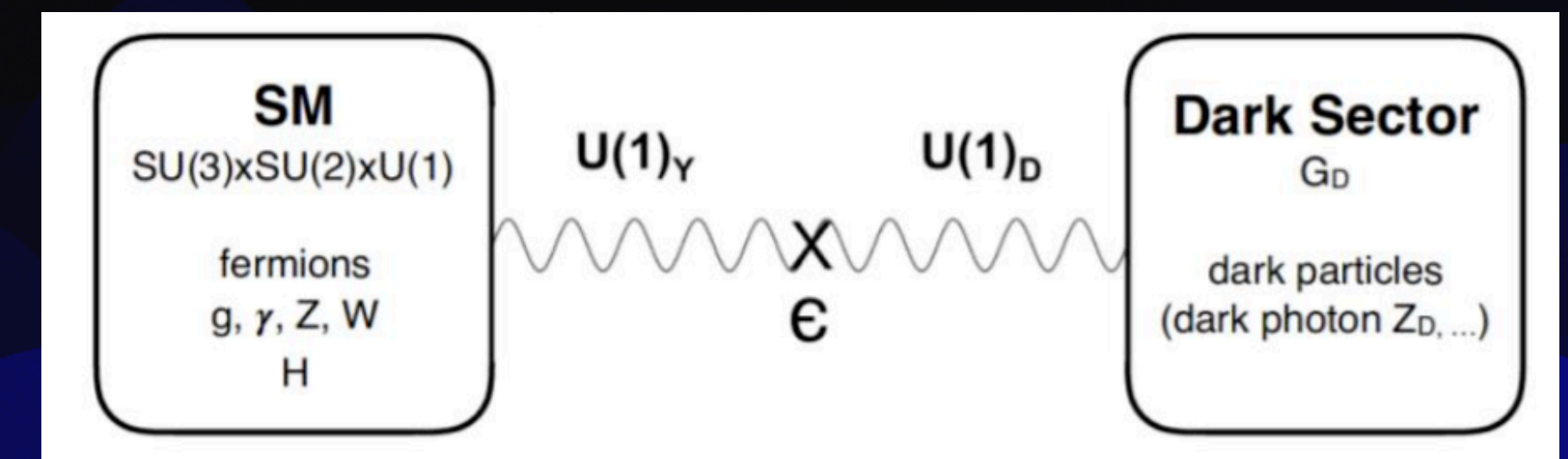
- Kinetic mixing parameter ϵ .
- Dark photon mass $m_{Z'}$
- The decay branching fraction of the dark photon into invisible dark-sector final states, typically assumed to be either unity or zero (corresponding to whether any invisible dark-sector final states are kinematically allowed or not)[†]

- **Experimental Signatures:**

- Dark photons produced in collisions, decay into SM particles (e.g., $\mu^+\mu^-$).
- Feynman diagram: dark photon (Z') mediating between quarks and leptons.

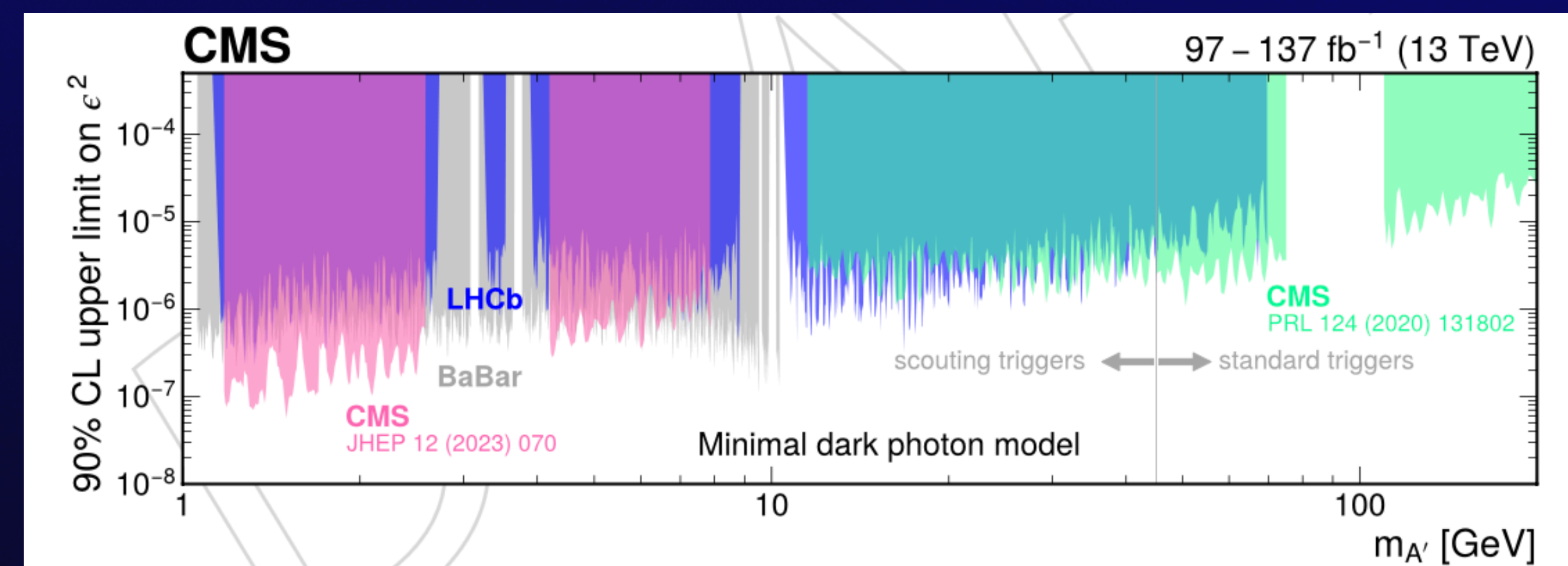
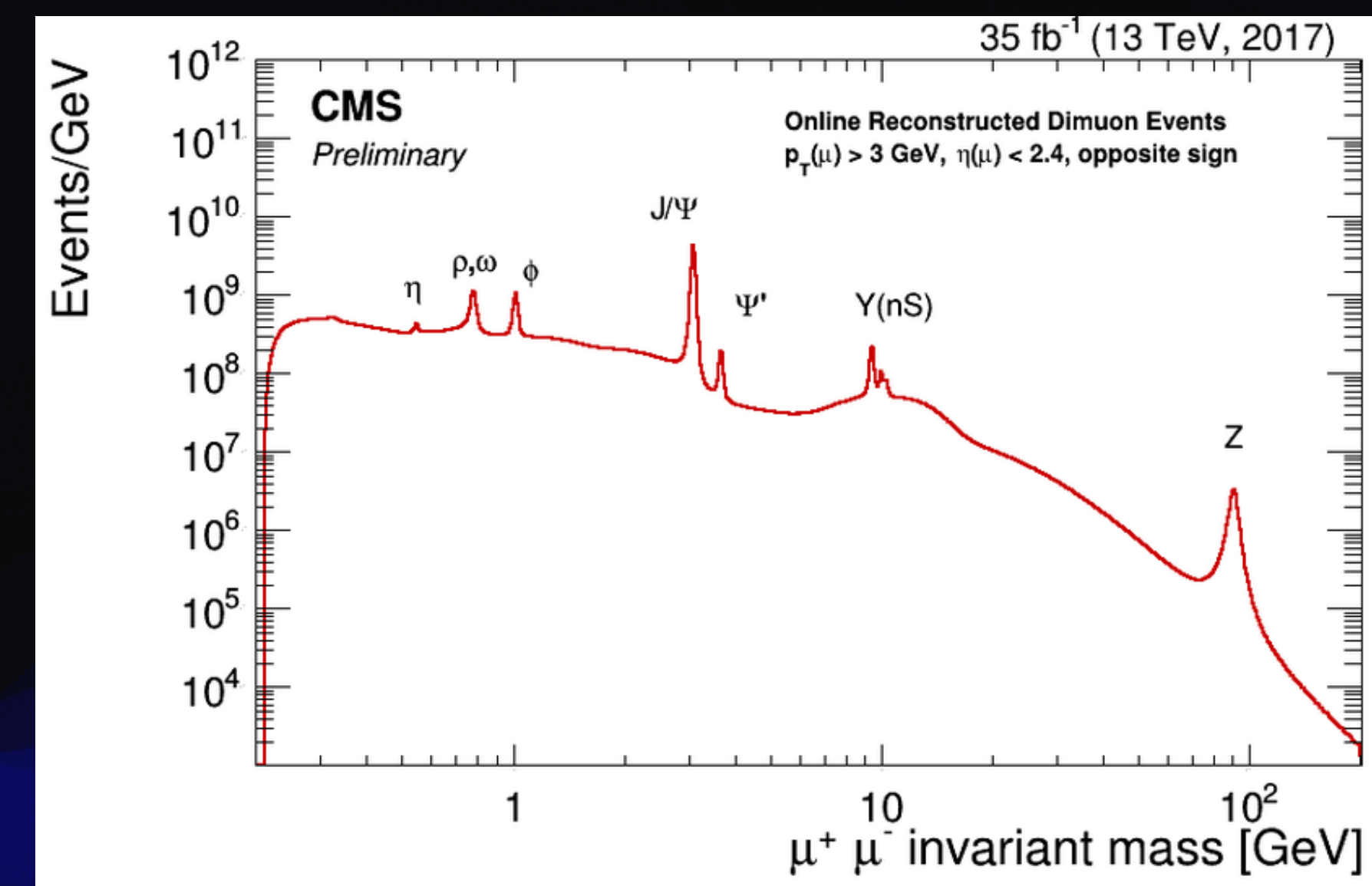
- **Motivation:**

- Explains astrophysical and cosmological observations suggesting dark matter.
- Dark photons as candidates for mediating interactions between dark matter and SM.



Dimuon Resonances

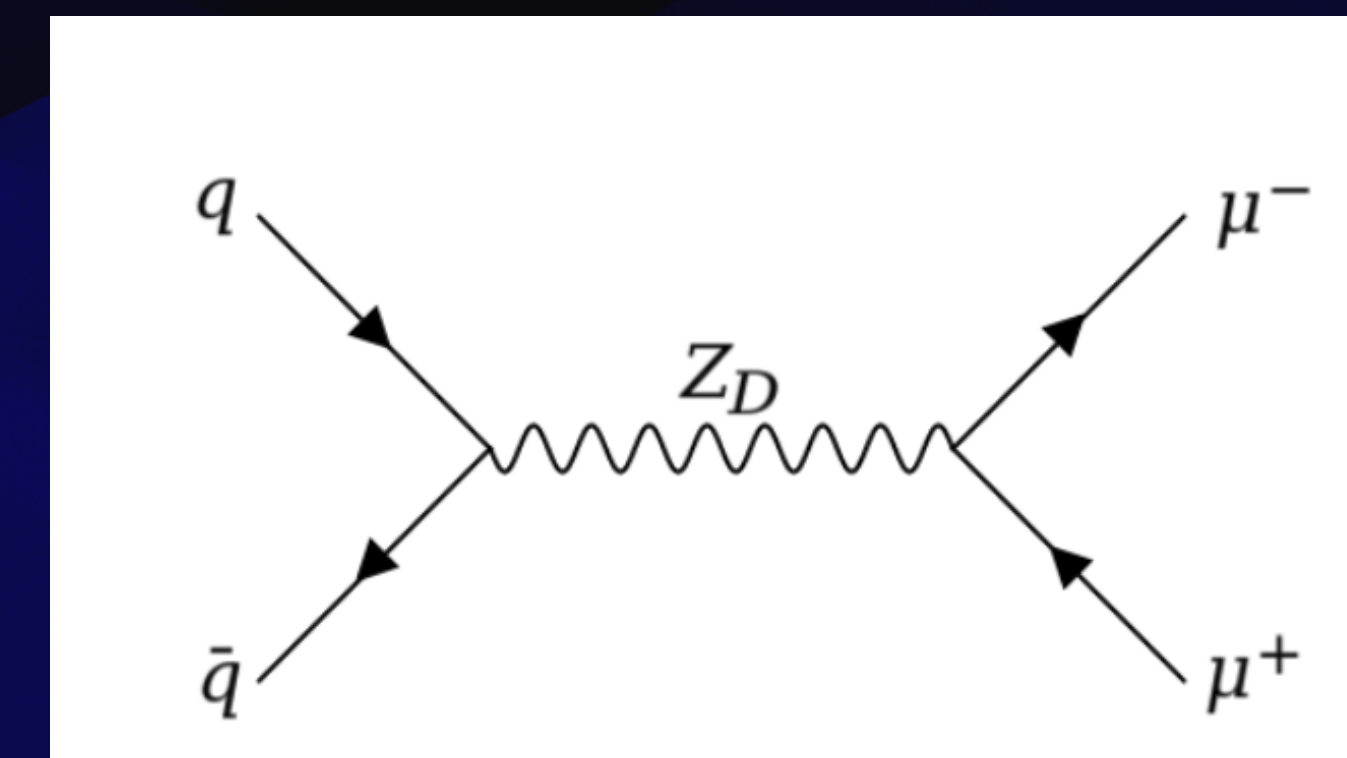
- Discovery of many new particles through the resonant particle pair production in dimuon channel
- Search for a narrow dimuon resonance at low mass using scouting data recorded by the CMS
- Study the dimuon final states to test the minimal Dark Photon model
- Most recent results for the observed upper limits on the square of the kinetic mixing coefficient ϵ



Observed upper limits
 CMS PAPER EXO-23-005

Analysis Strategy

- Investigating the existence of Dark Photons, a potential BSM mediator, within the 10-40 GeV range, through the decay into oppositely charged muon pairs, utilising scouting data from the CMS experiment
- Optimising event selection for (prompt) dimuon resonance signals and efficiency calculations
- Searching for a bump in the dimuon mass spectrum using analytical signal and background Pdfs
- Study systematic uncertainties
- Establish model-independent limits for the cross-section of low-mass dimuon resonant states



An expected production channel of a Dark Photon
<https://arxiv.org/pdf/2309.16003.pdf>

Employed Data

The datasets that are used in this analysis:

- **Observed Data**

- LHC Run 3, CMS Scouting Data: `/ScoutingPFMonitor/Run202*/RAW`

- **MC: Upsilon Samples**

- Upsilononto2Mu_UpsilonFilter_2MuFilter_TuneCP5_13p6TeV_pythia8 dataset
- Used to compare efficiencies around the Y region

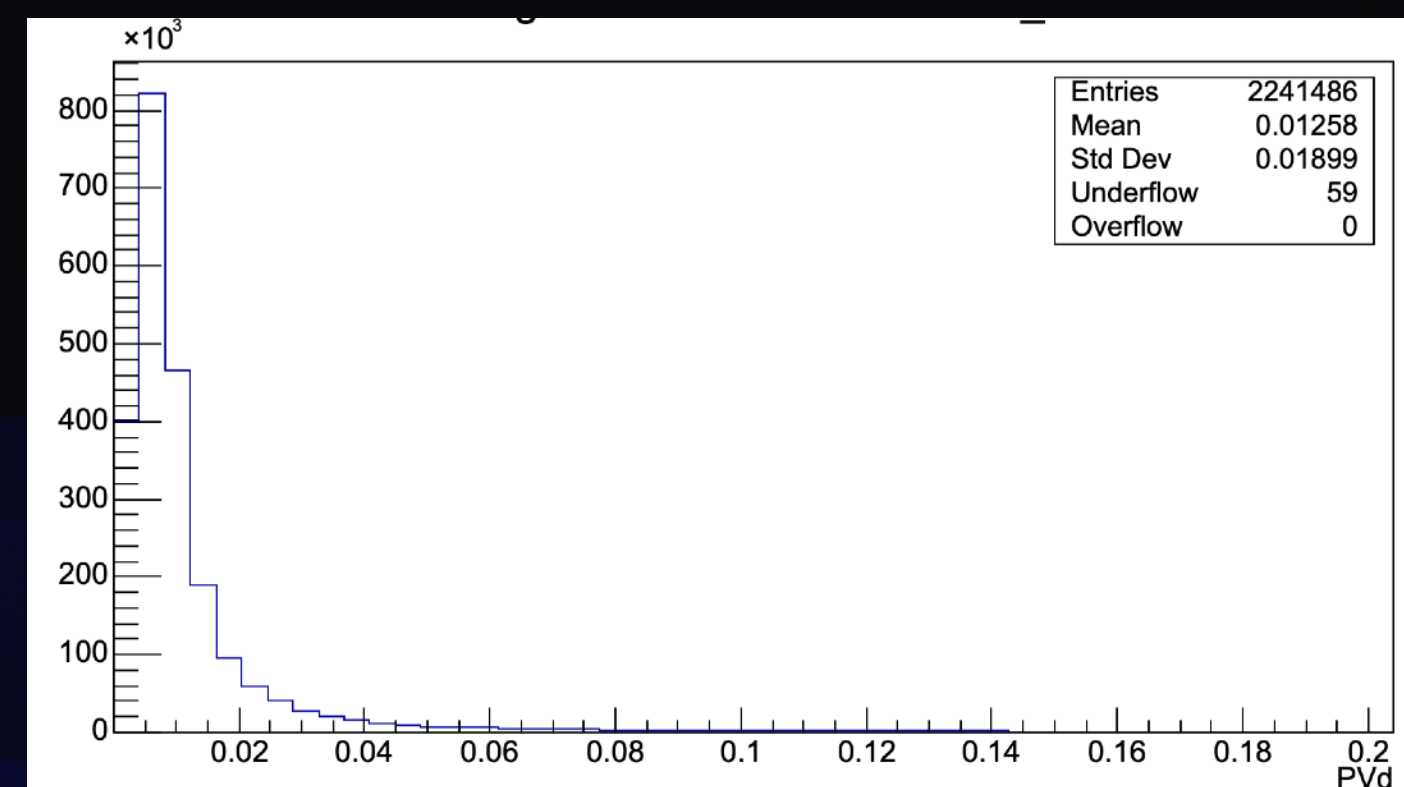
- **MC: DY Samples**

- Privately produced samples
- Used for efficiency calculations

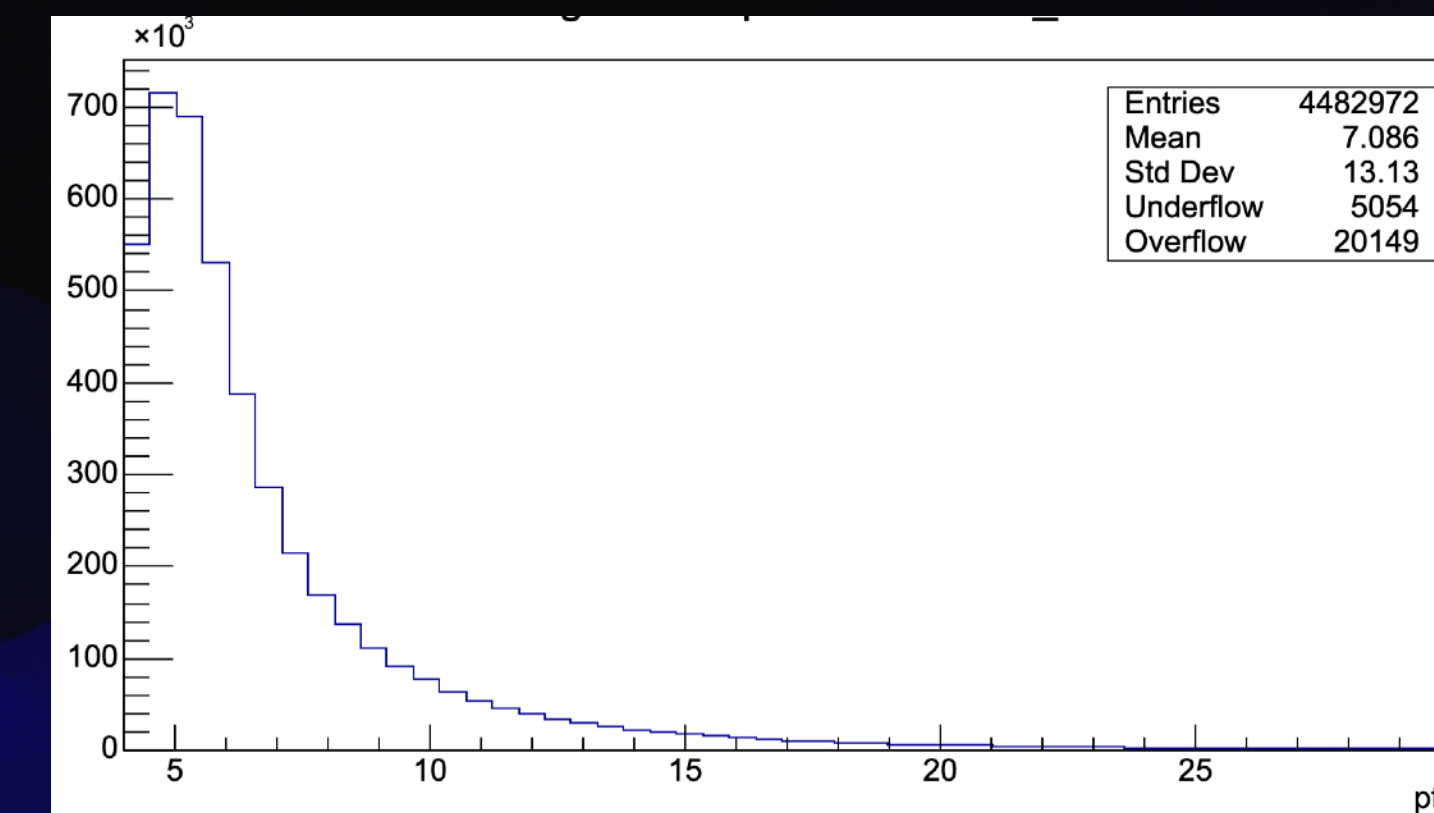
2) Event Selection and Efficiency Calculations

Event Preselection

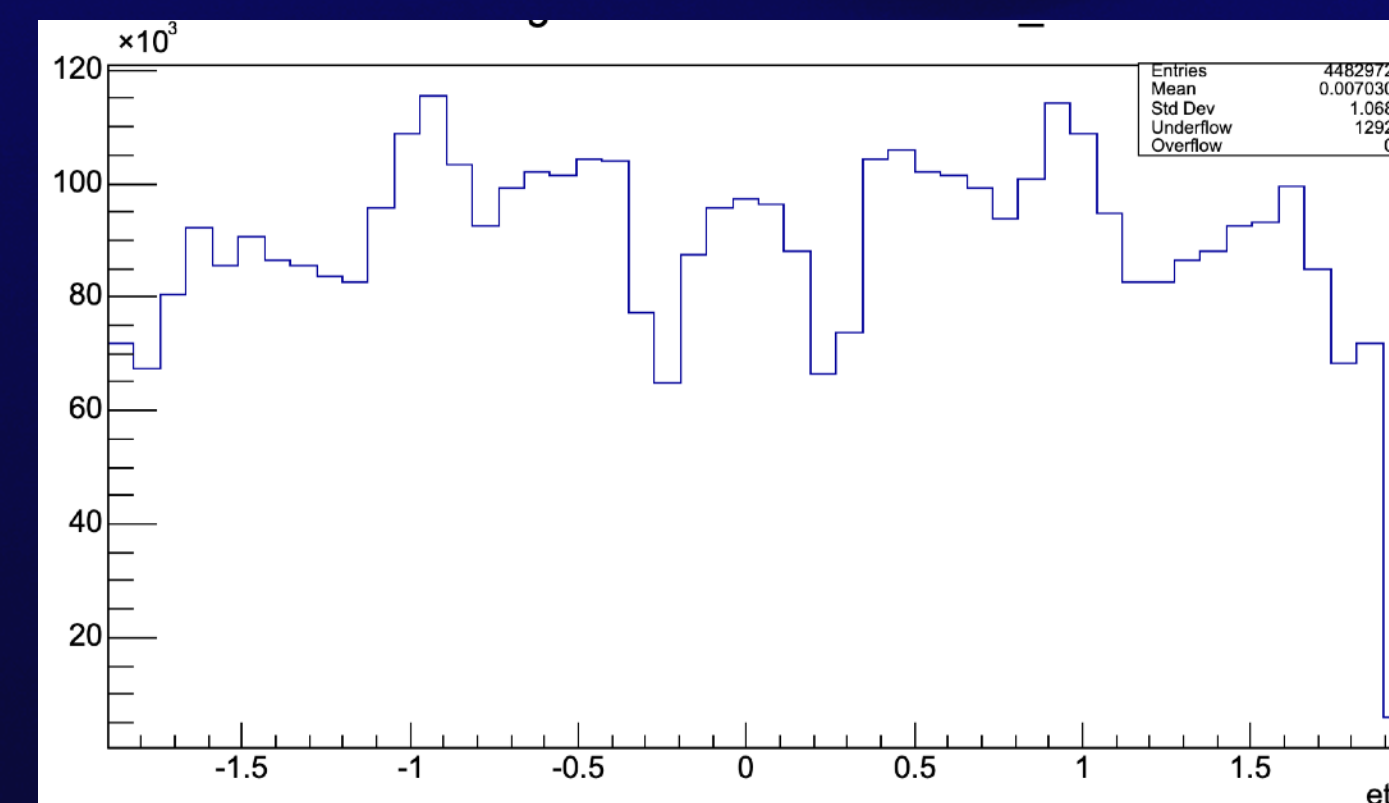
- Muon pair with opposite charges as the final state
- Prompt Production
 Transverse Displacement:
 $L < 0.2 \text{ cm}$
- Transverse Momentum:
 $p_t > 4 \text{ GeV}$
- Pseudorapidity: $|\eta| < 1.9$



PVd Distribution



p_t Distribution



η Distribution

Optimising the Event Selection

- Focus: Optimisation of event selection within the 10-40 GeV mass range.
- Strategy: Employing the Upsilon (Υ) resonance as benchmark for optimising event selection through neural network training and efficiency analysis of simple cuts.
- Goal: Enhancing signal detection sensitivity for prompt dimuon events.
- Approach: Comparison of neural network performance against traditional cut-based methods and refining the parameters for each case.

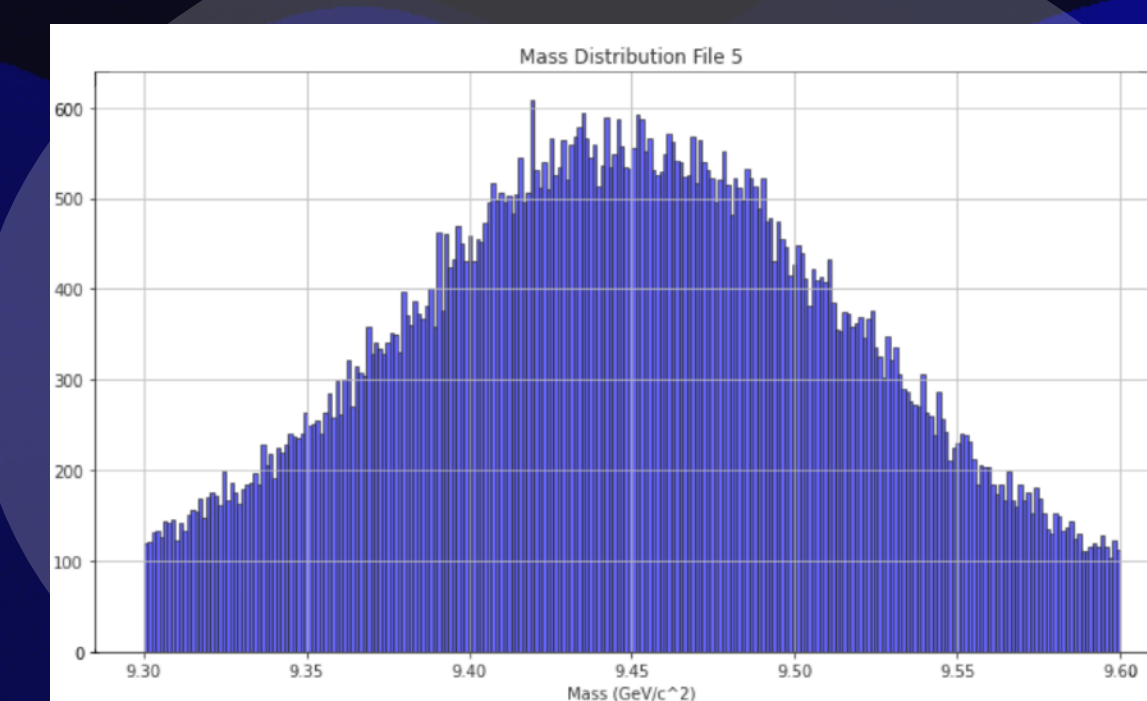
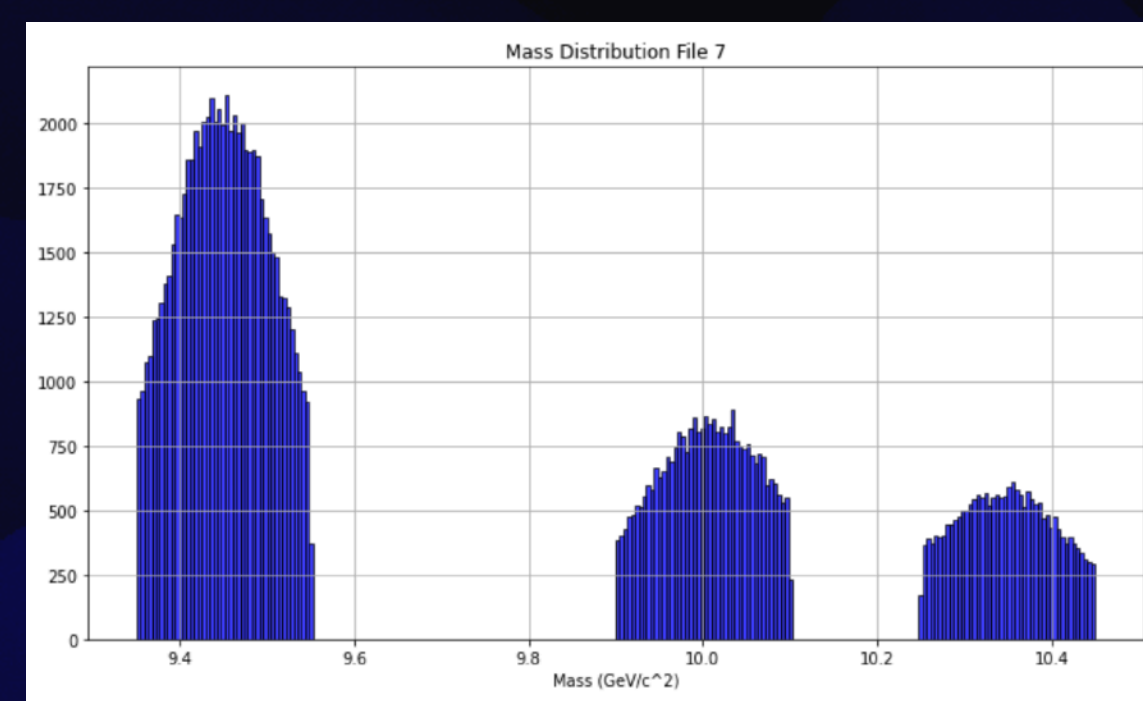
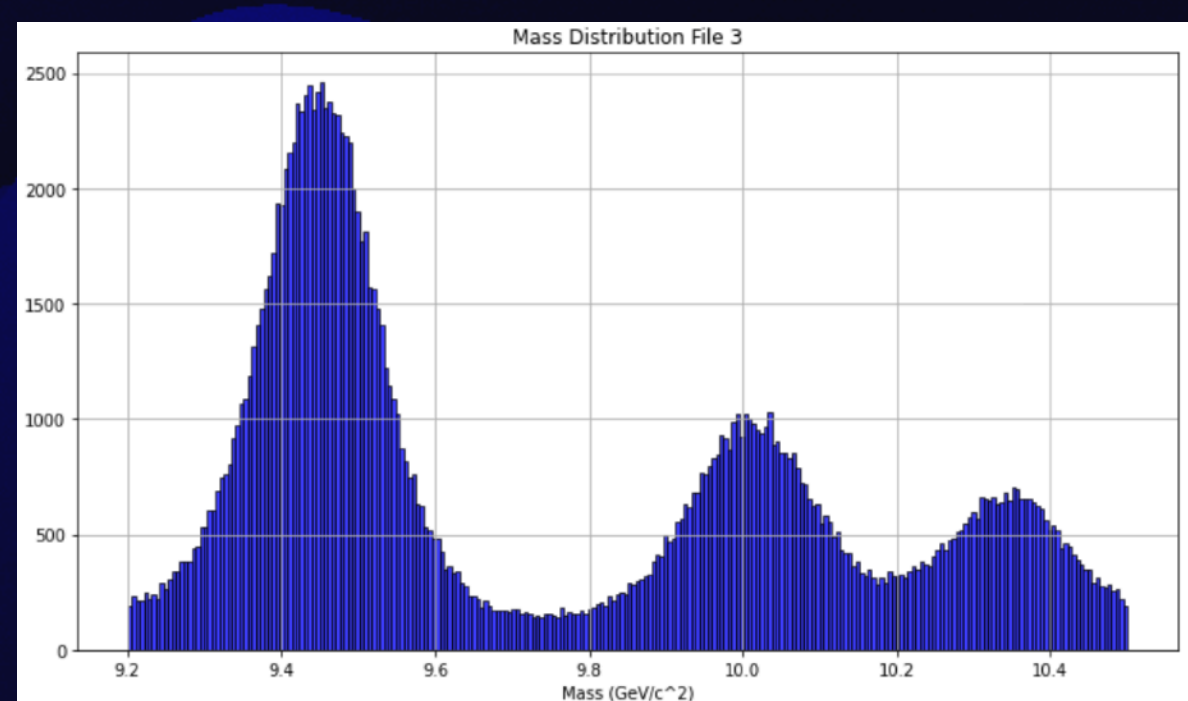
Approach 1. Employing Neural Networks

- i. Optimising the signal mass window
- ii. Choosing the variables for NN training
- iii. Choosing the optimiser algorithm
- iv. MVA analysis and MVA cut

Only for the MuonID

i) Optimising the Signal Mass Window

- Use all the candidate variables (will be optimised in the next slide)
- Deploy several mass windows as signal region for the training and compare the ROC curves
- Use AUC for comparison



The best response:

Mass Region 9.3 - 9.6 GeV

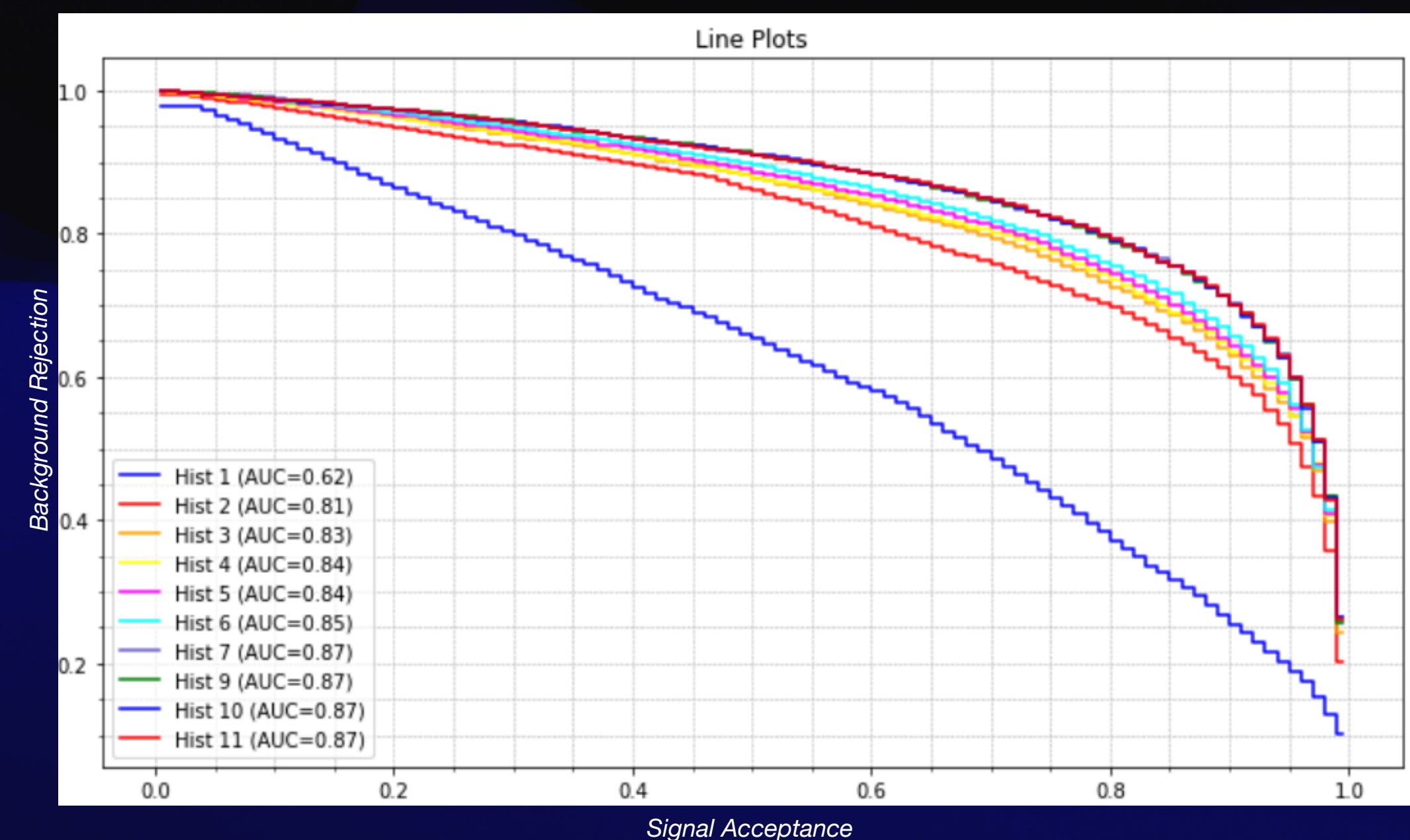
AUC 0.87

Signal Contamination 84%

ii) Choosing the Training Variables

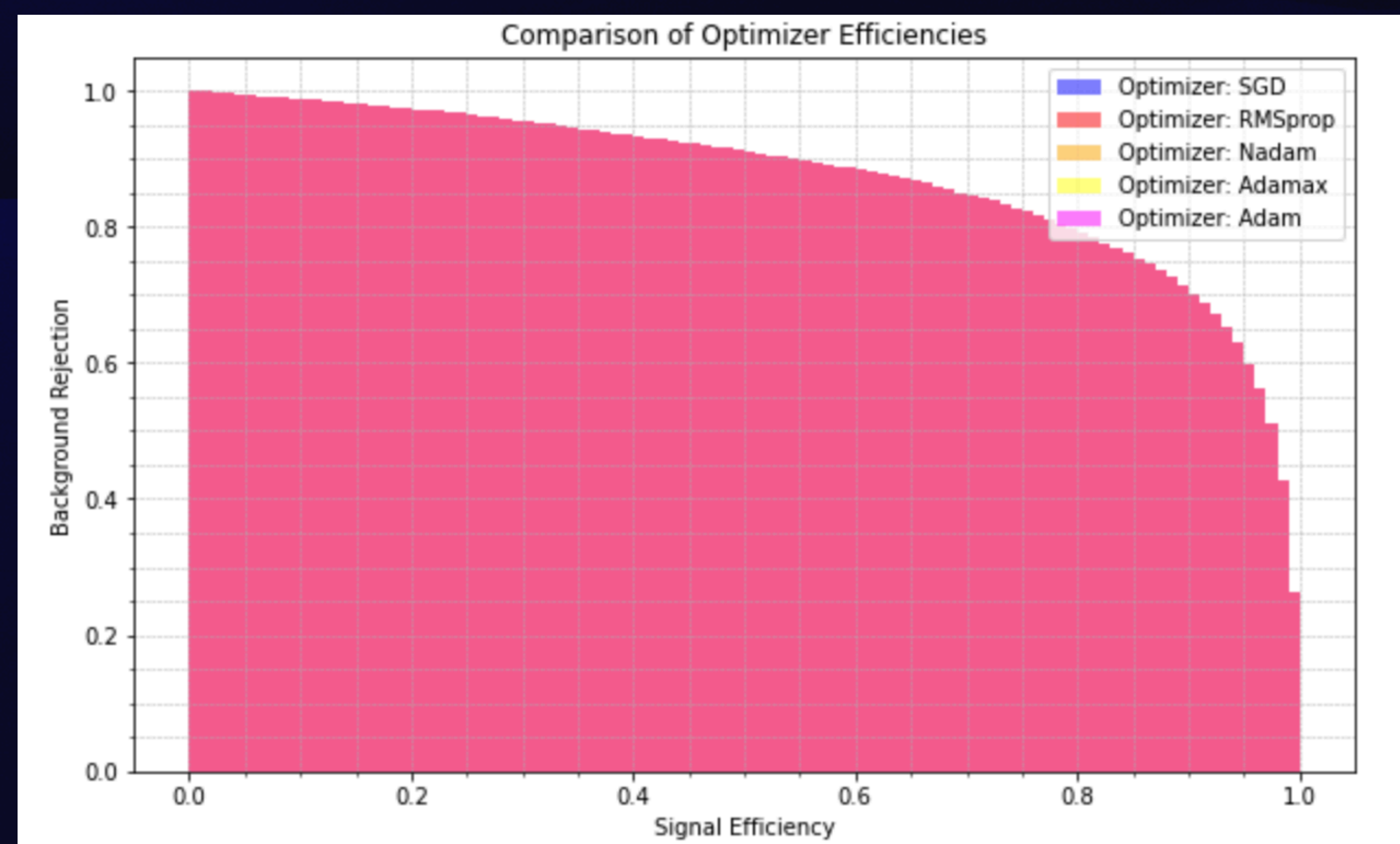
- Train with all the variables
- Modify the variable order according to importance
- Train with 1 variable and add the next one, repeat cumulatively
- Decide which variables to use

Used variables: "nmhits", "trkiso", "trkqoverp", "trklambda", "dxy",
 "ntklayers", "eta", "chi", "nphits"



iii) Choosing the Optimiser Algorithm

- Deploy several optimiser algorithms (SGD, Adam, Nadam...)
- Compare the ROC curves

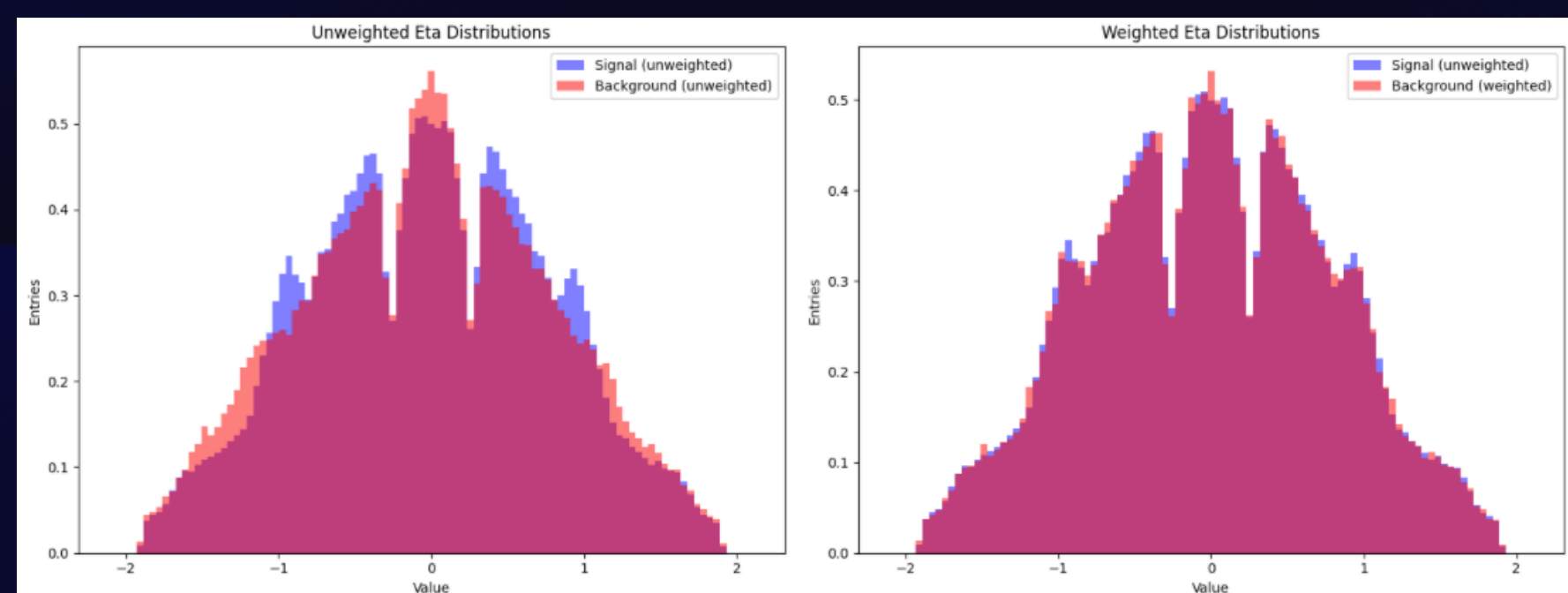


Almost same AUCs,
no visible improvement

Stick with SGD

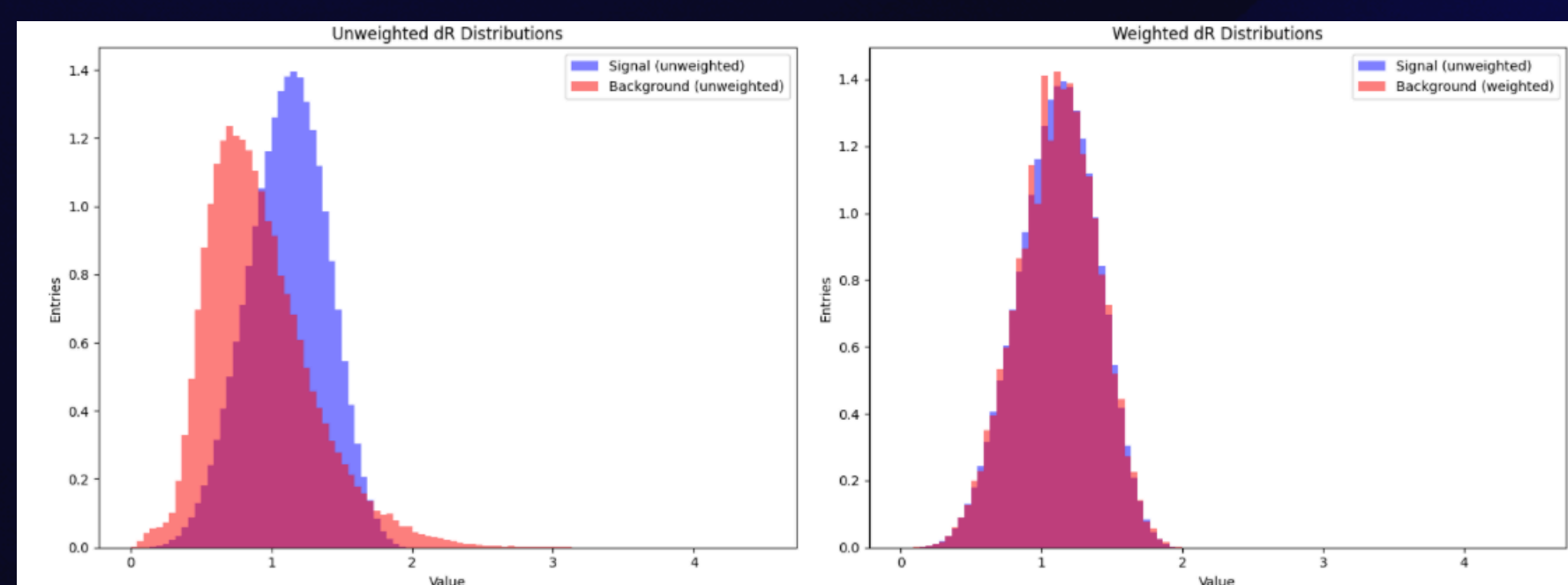
iv) MVA Analysis and MVA Cut

- Re-weigh data to balance background and signal to prevent model bias in distinguishing features



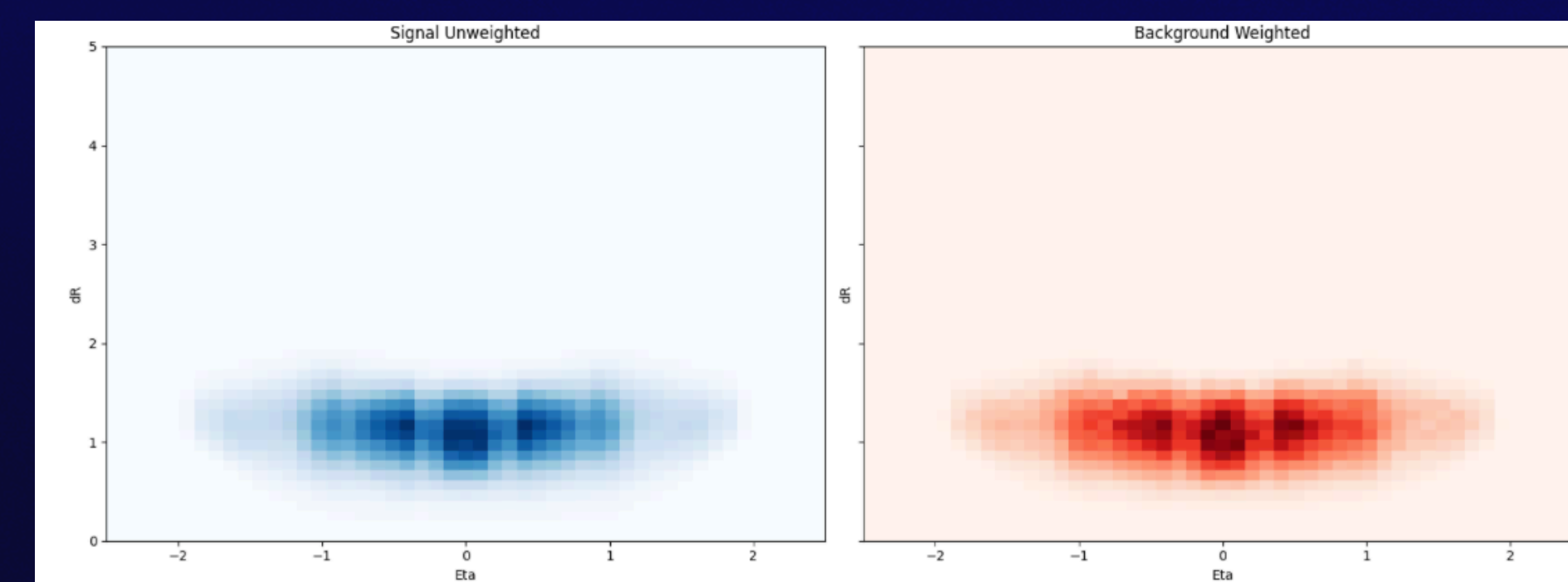
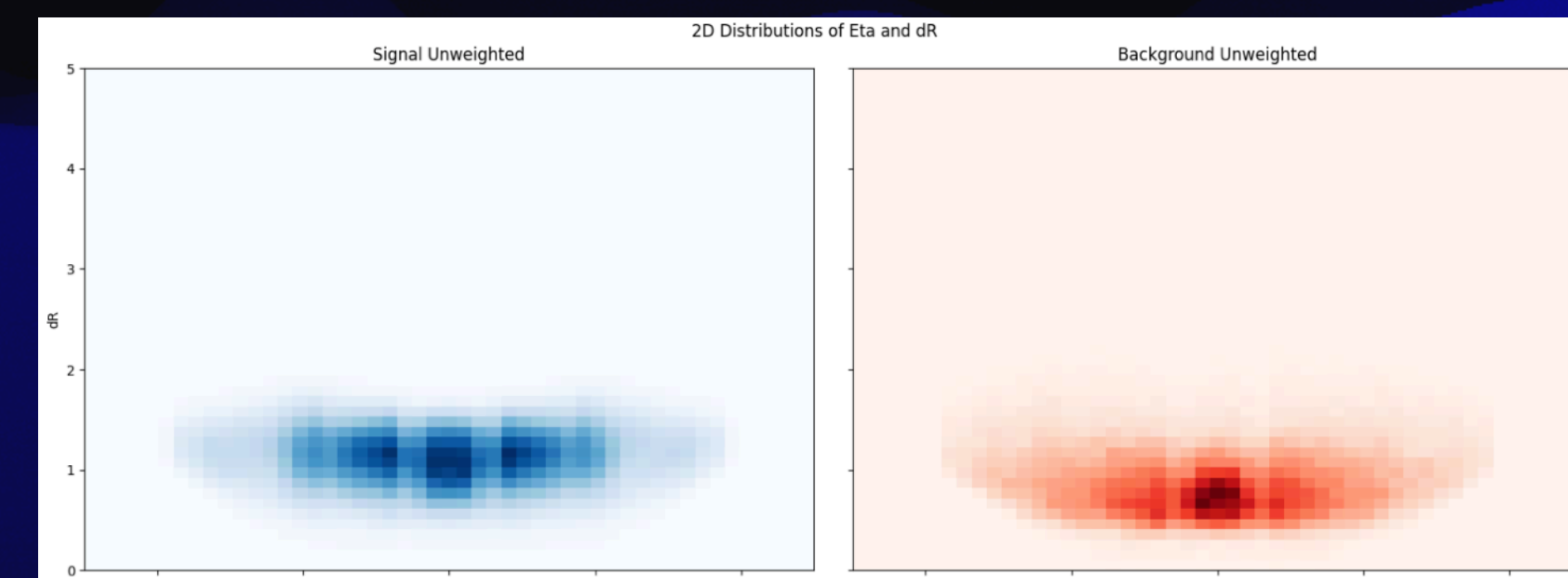
Eta

Eta

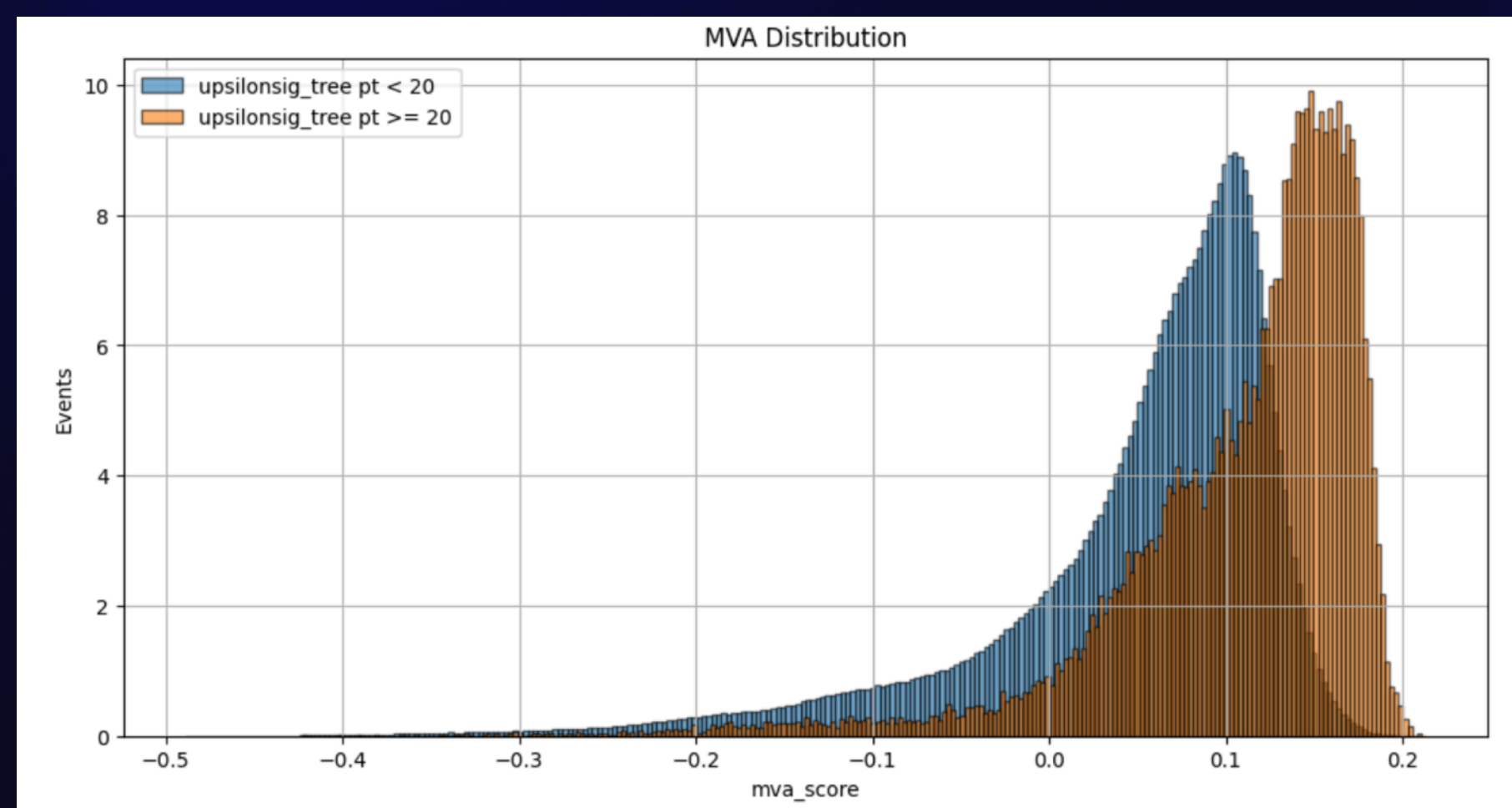
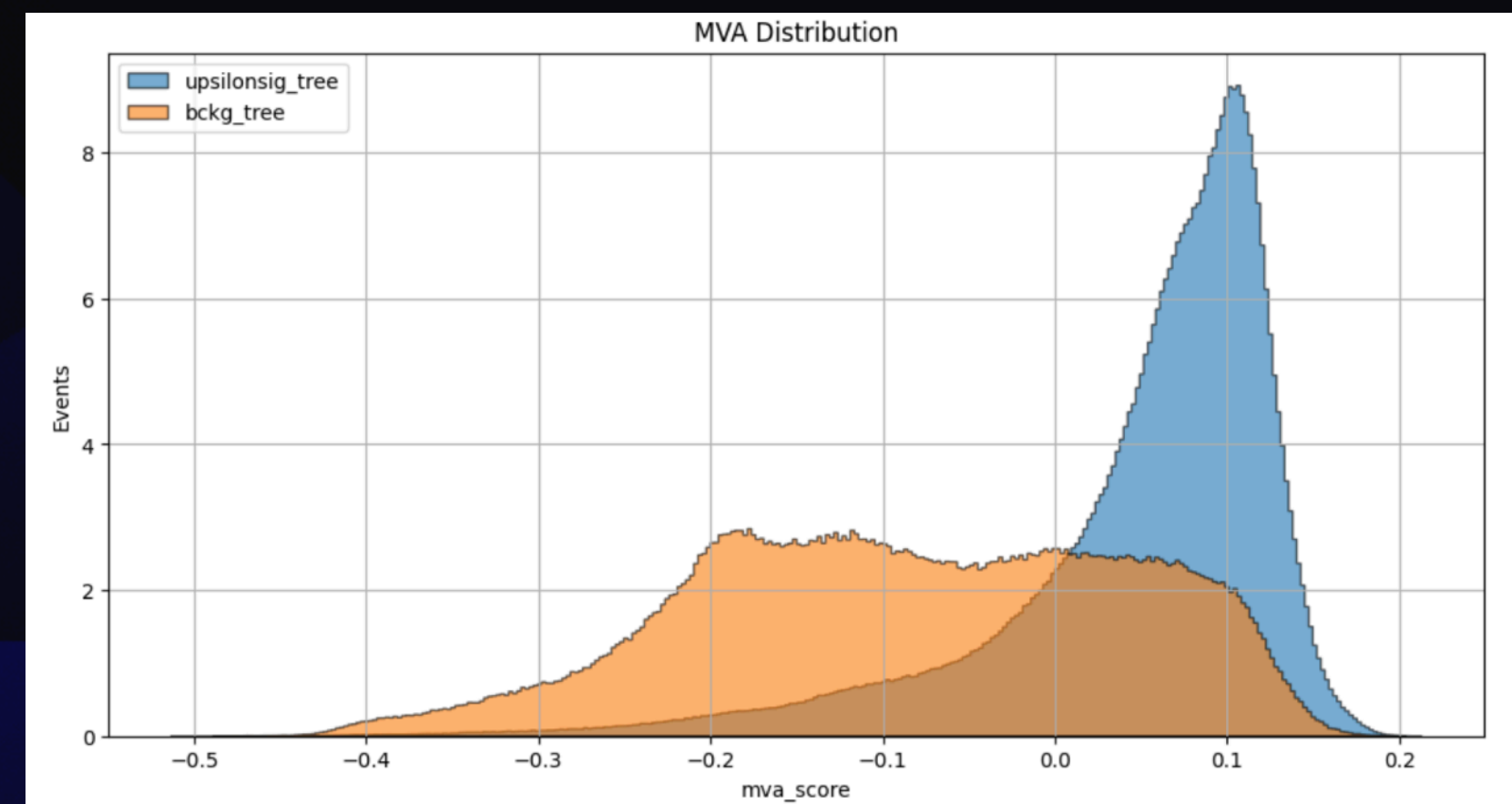


dR

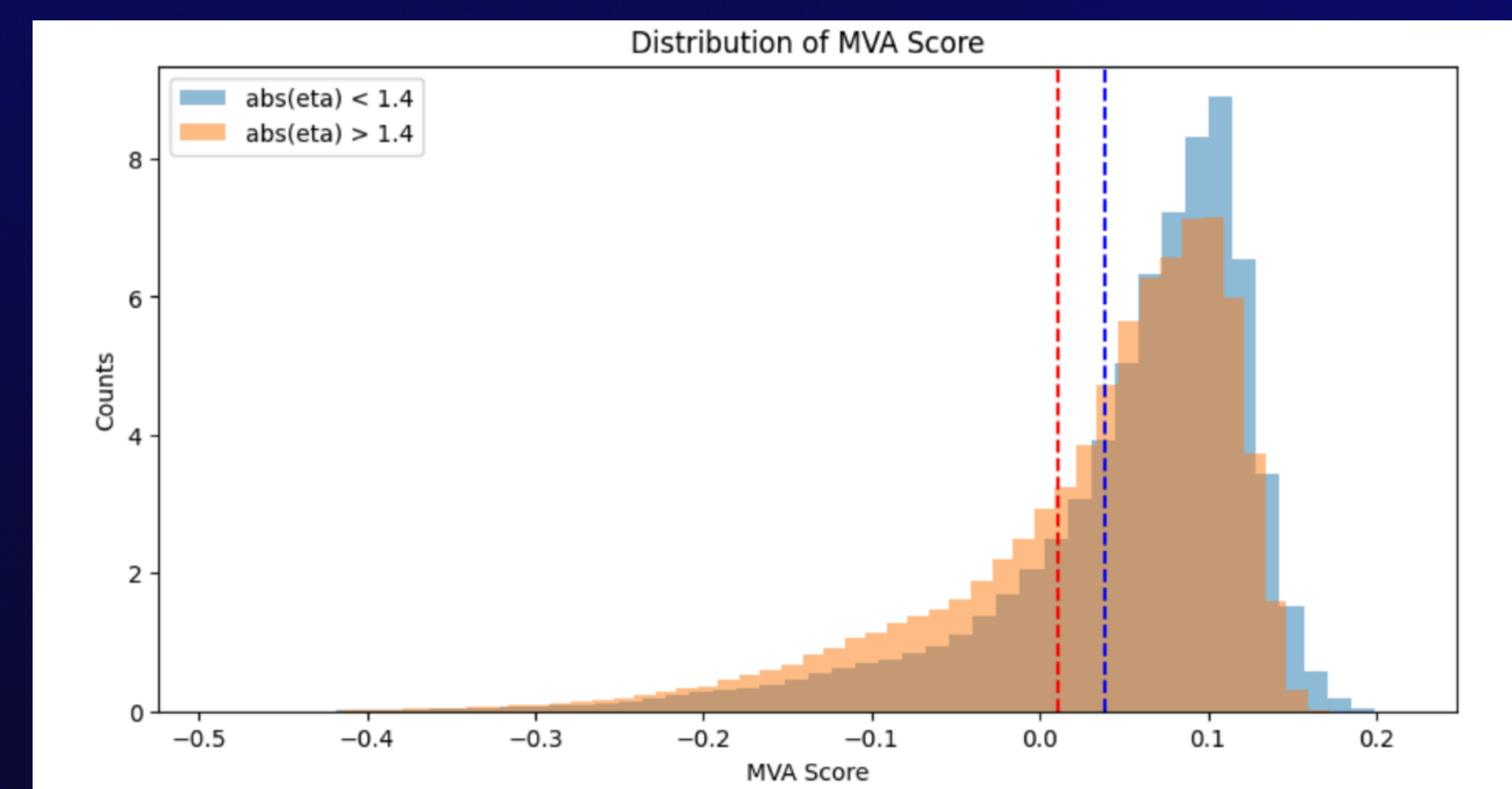
dR



- Add MVA using the best model's weights

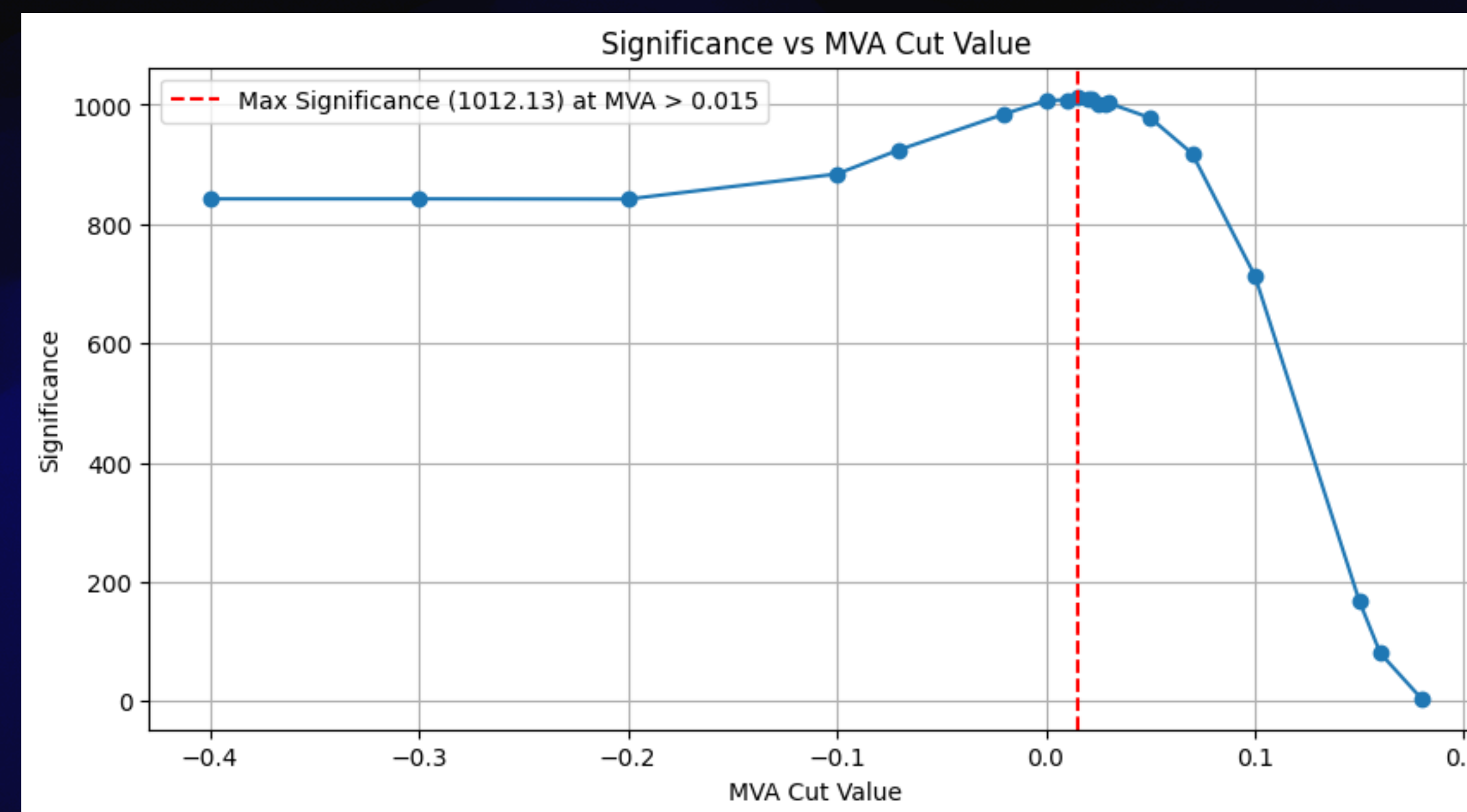
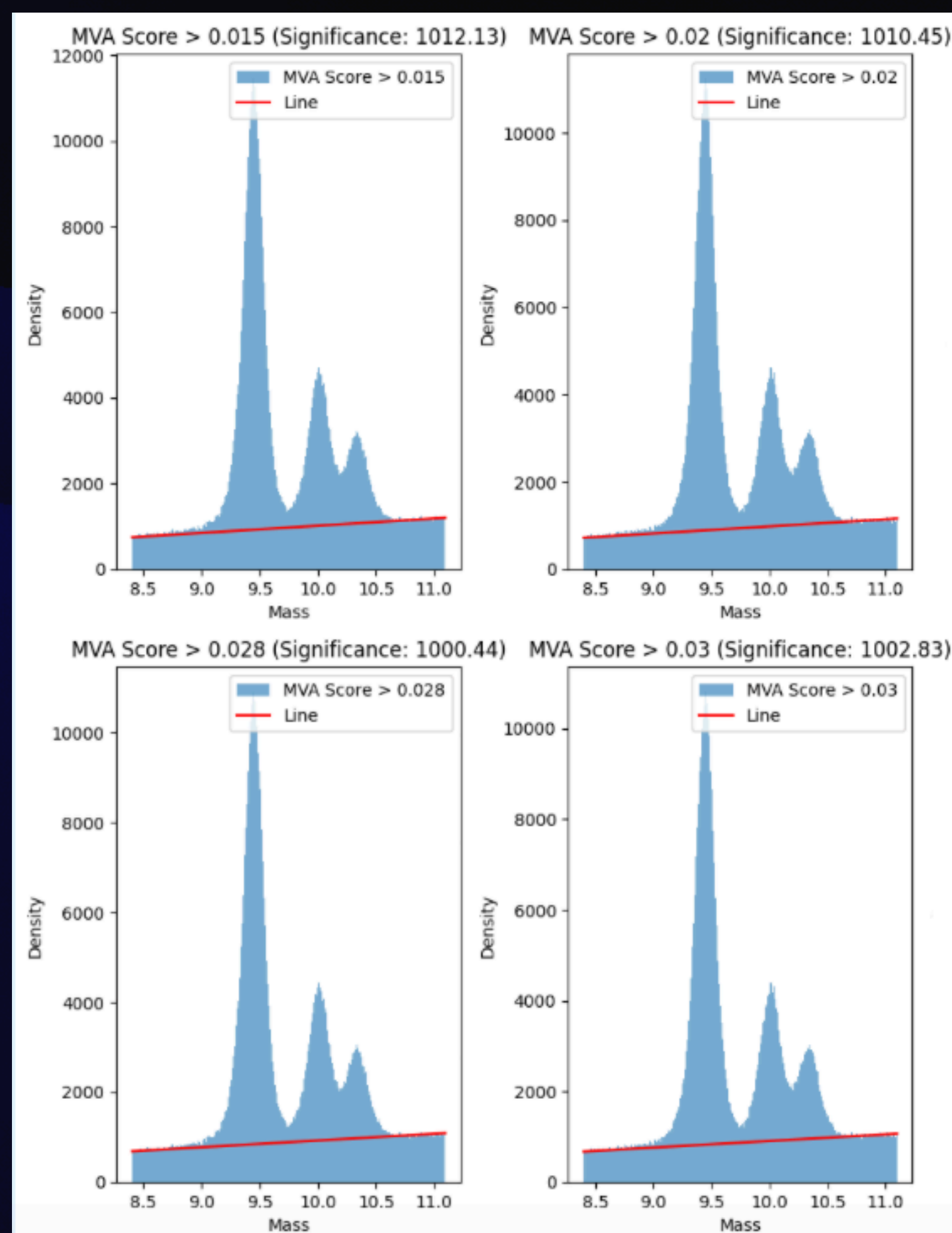


As expected



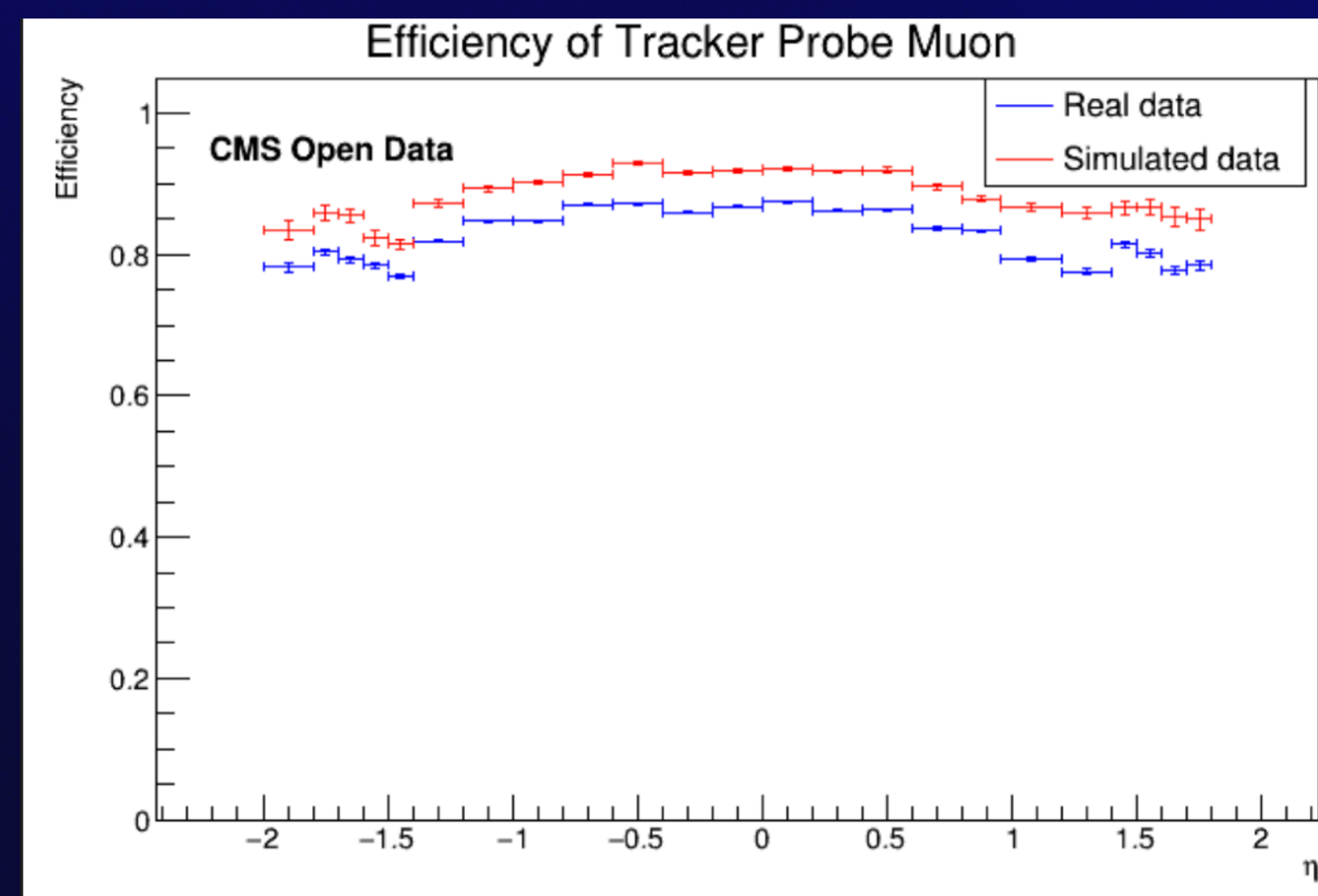
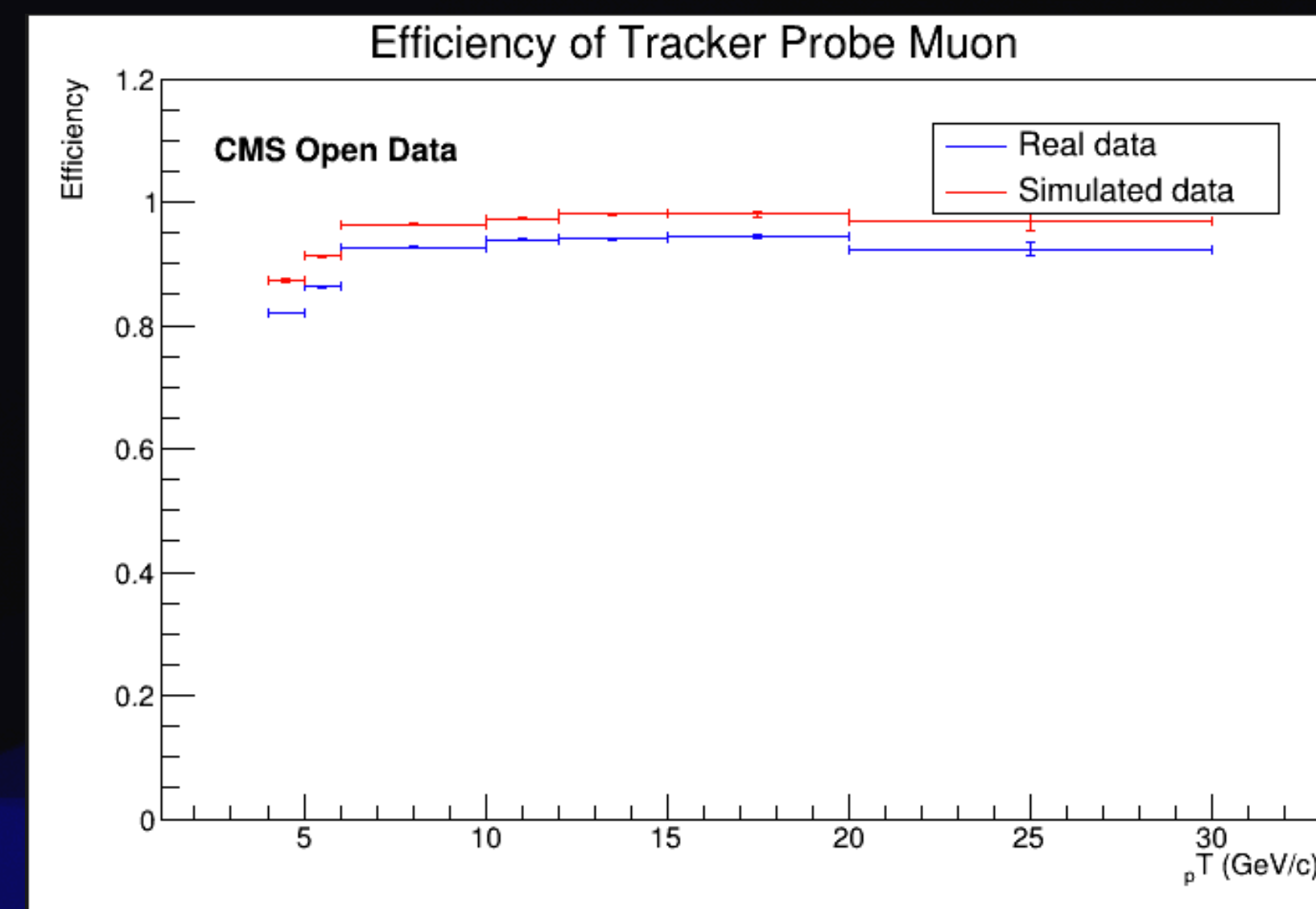
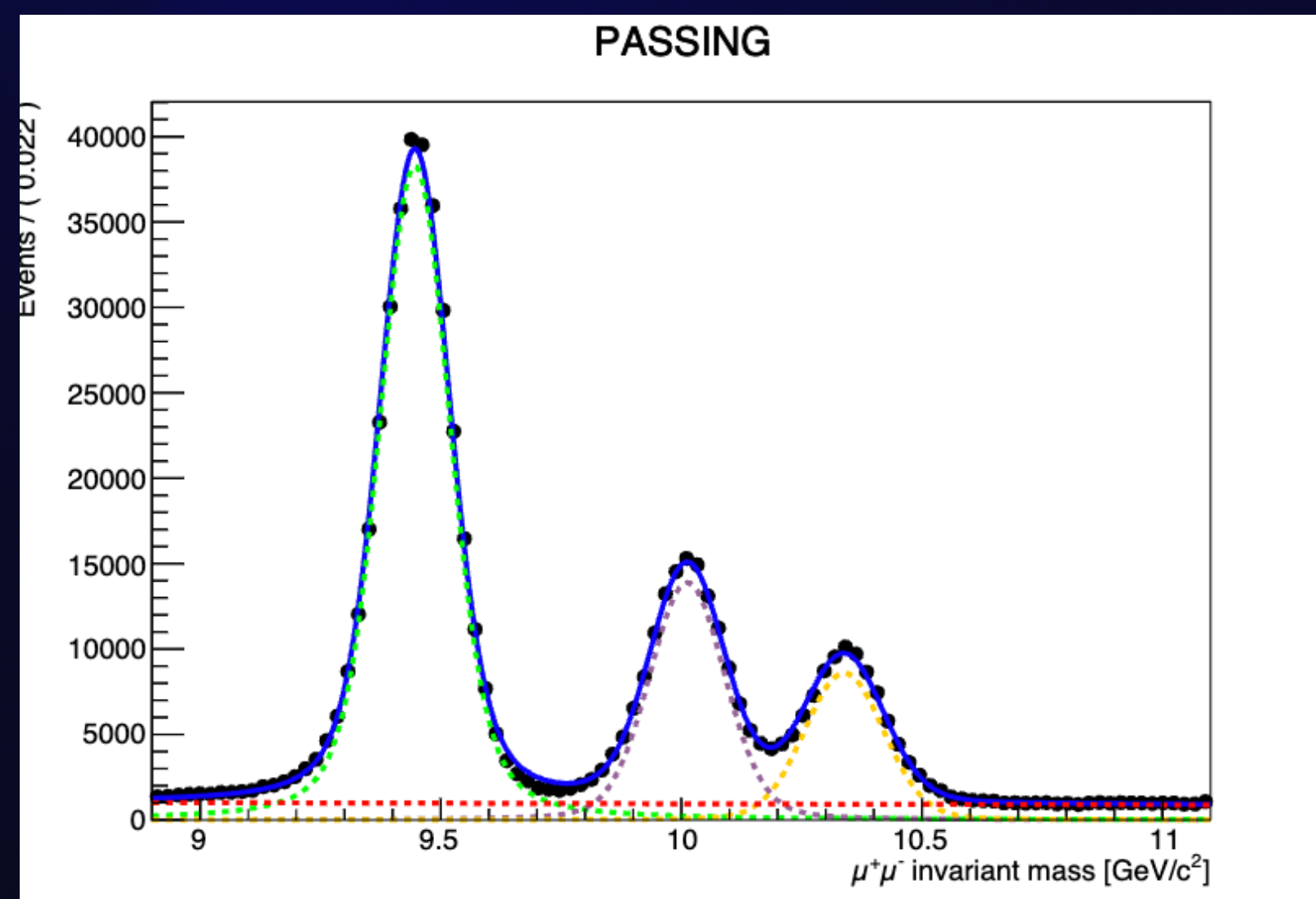
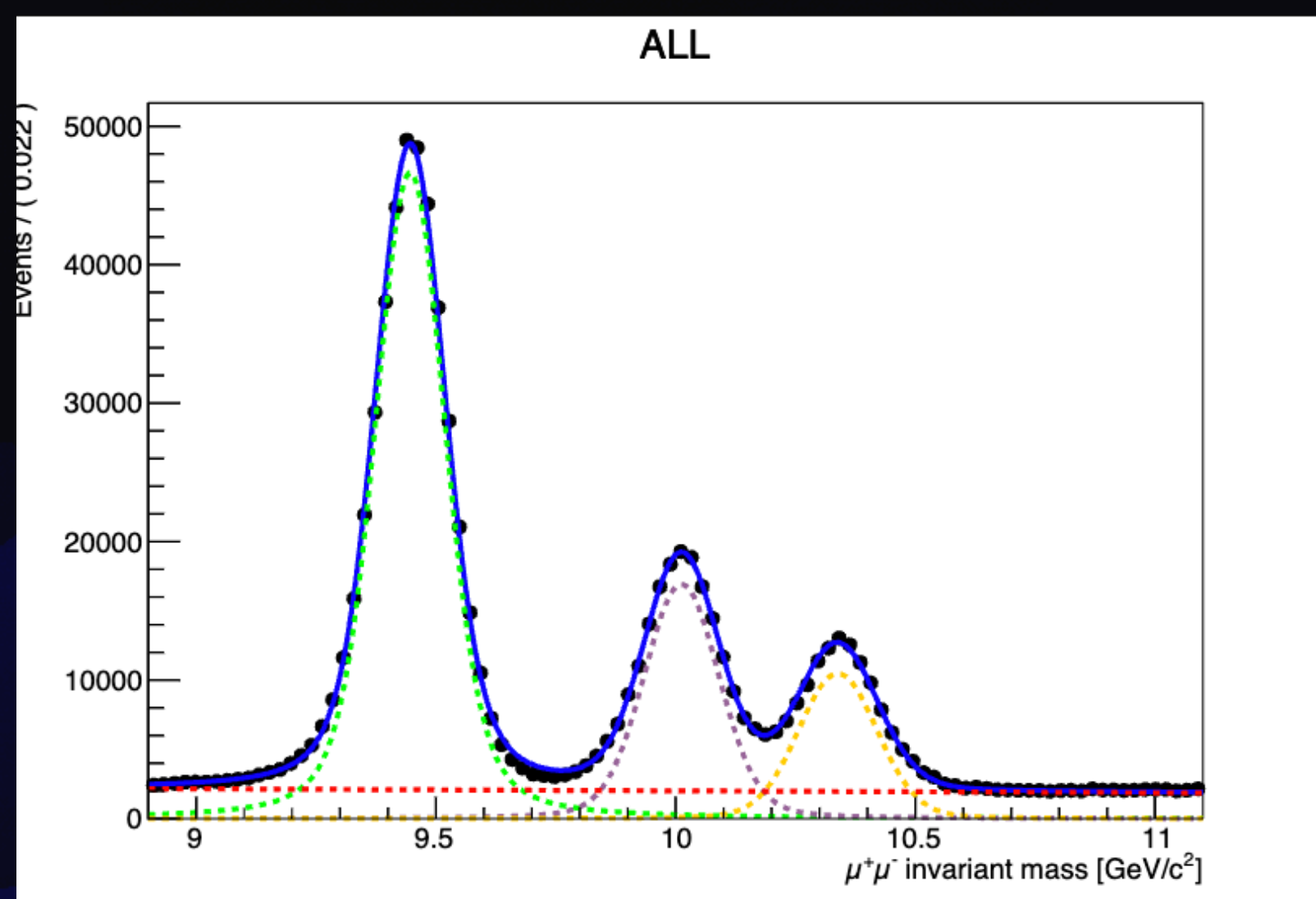
Previously this was a problem

• Optimise the significance



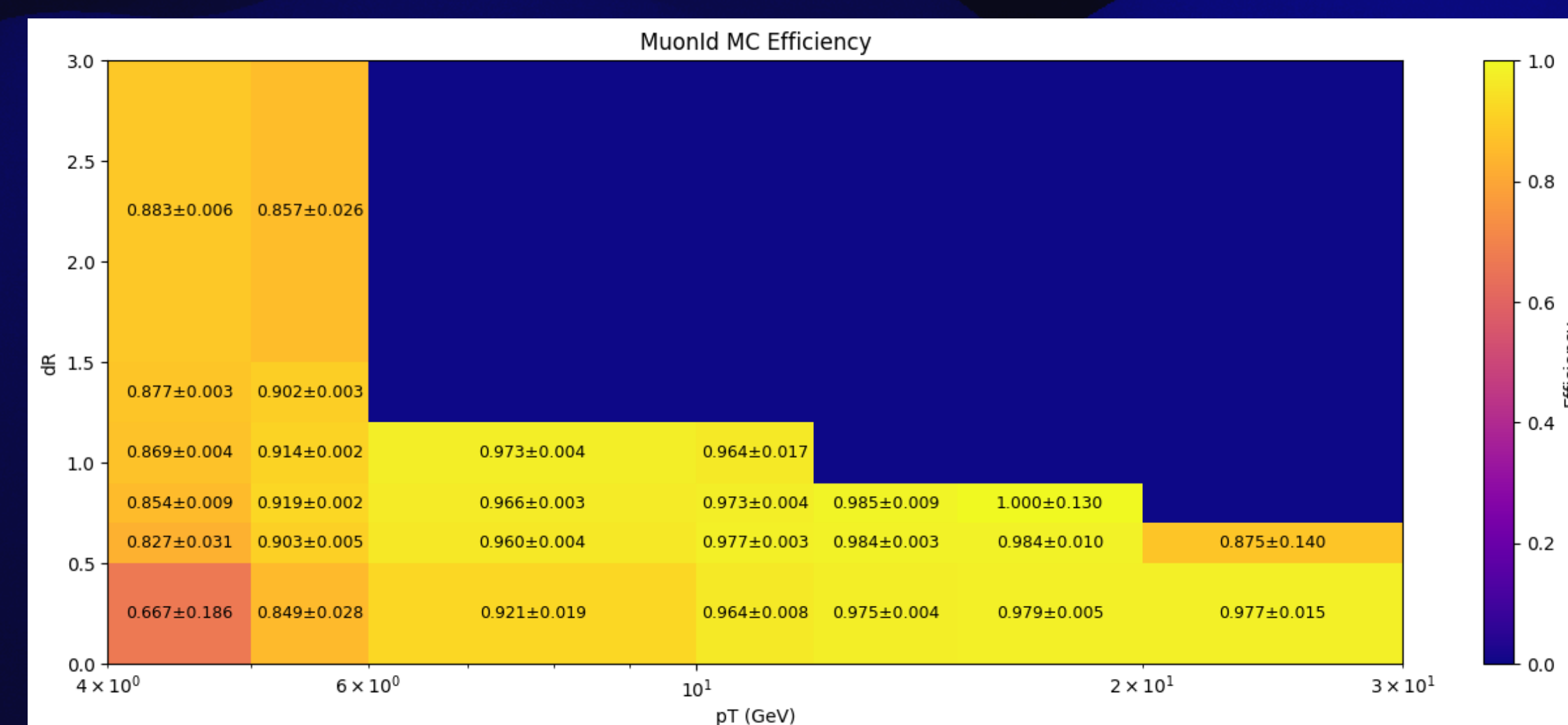
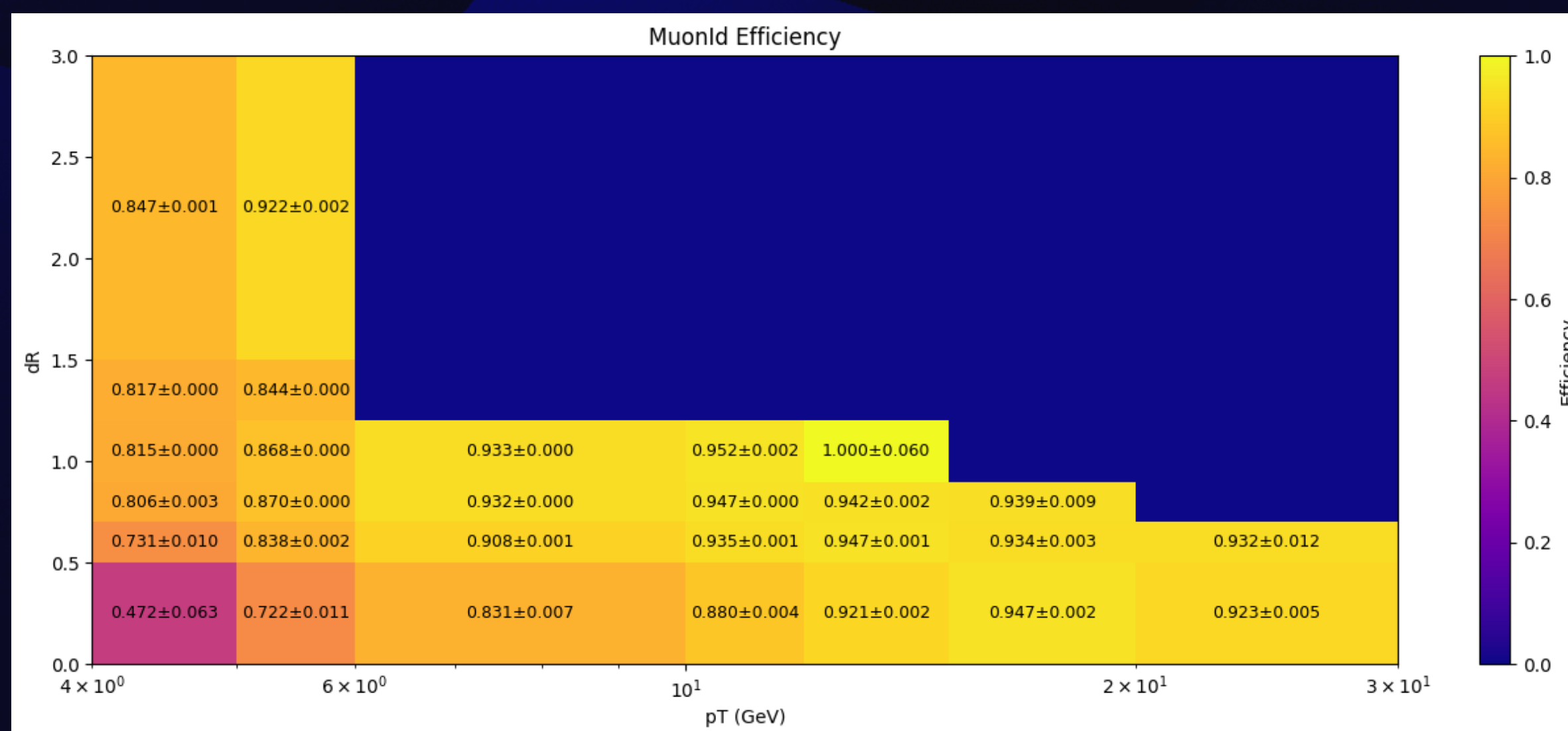
The maximum significance is 1012.13 for an MVA cut of > 0.015.

• MuonID (1: pass, 0: fail) Efficiency with Tag'n Probe Fitting Method



MuonId Efficiency

- Efficiencies for 2D (pT & dR) binnings
- The total efficiencies are calculated by integrating these values
- MC efficiencies for systematic uncertainties
- Efficiency does not change significantly wrt mass and taken constant as ~ 0.892



Approach 2. Traditional Cuts vs Training

- i. Compare the results for track isolation (trkiso)
- ii. Compare the results for vertex displacement (L_{xy})

Only for the vertex variables

i) Simple Cut vs Training: trkiso

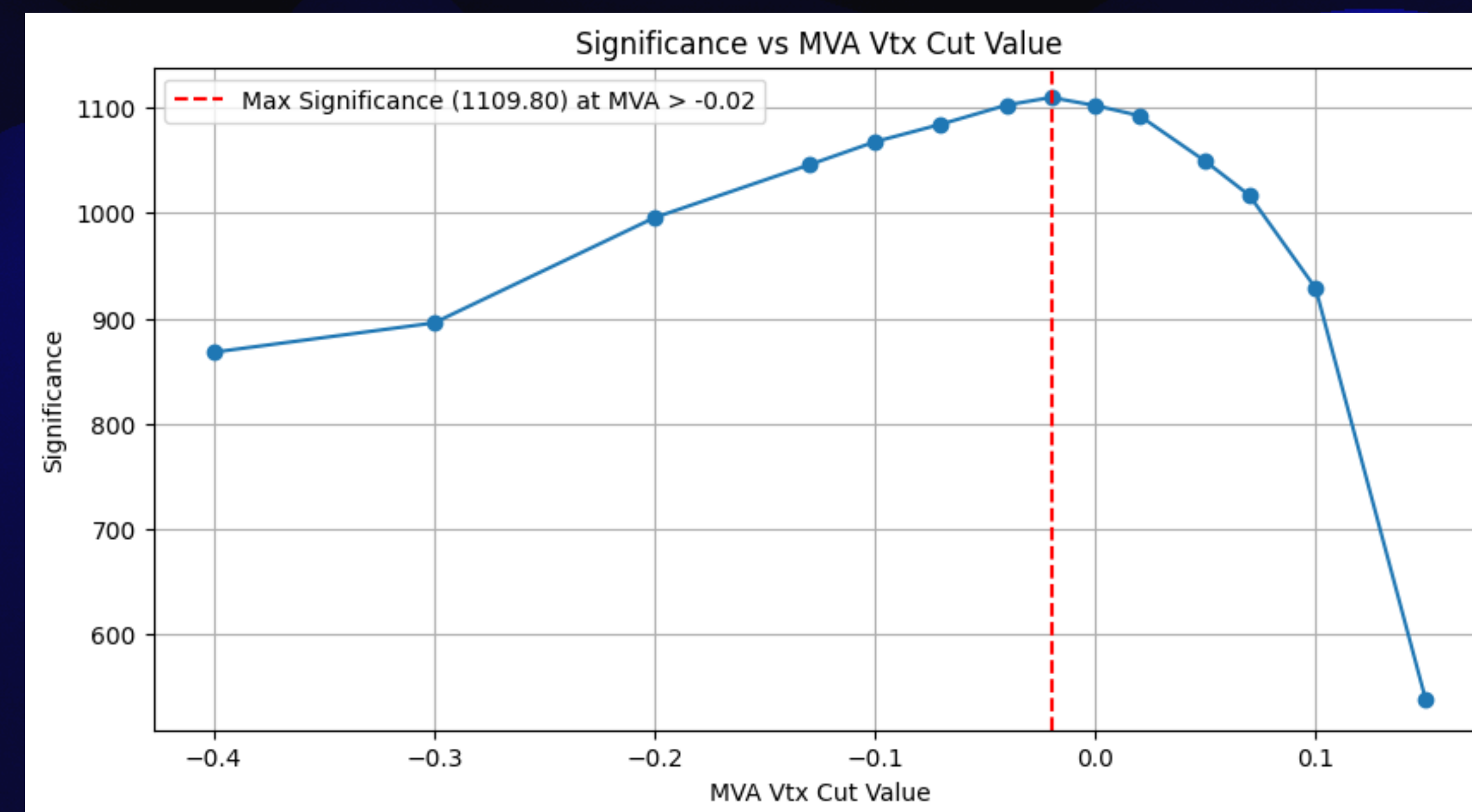
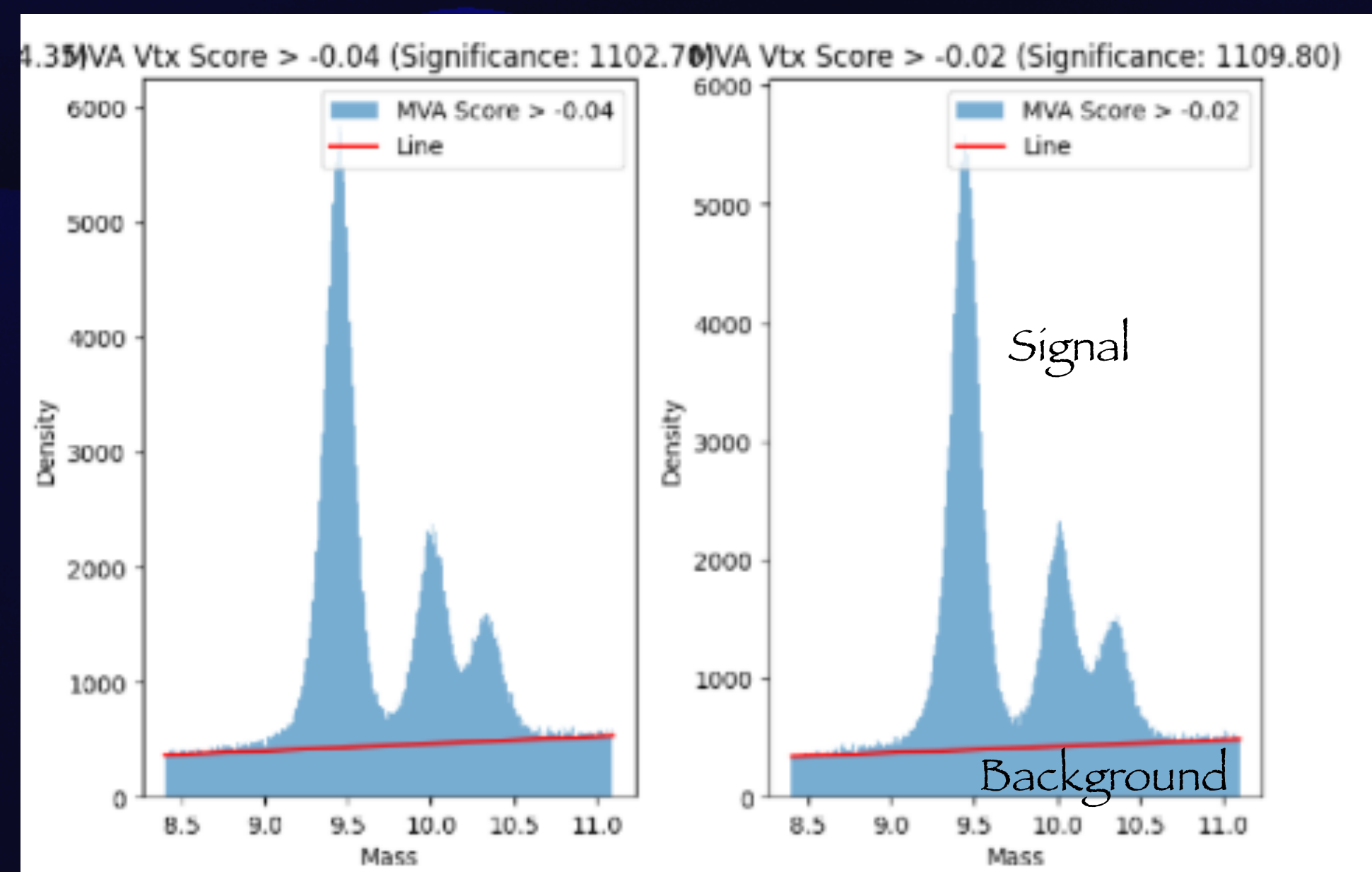
- Trkiso was used in the training for Run II data
- Use the weights of the training using track isolation, optimise the MVA cut, observe the significance
- Exclude trkiso from the training, optimise the MVA cut, optimise the trkiso cut
- Compare the results

	Trkiso included	Trkiso excluded + trkiso cut
Optimal significance	6210.69	6009.85
Signal efficiency at optimal cut:	0.6764	0.8278
Background rejection at optimal cut:	0.8398	0.6293

-> Including trkiso in the training is more effective!

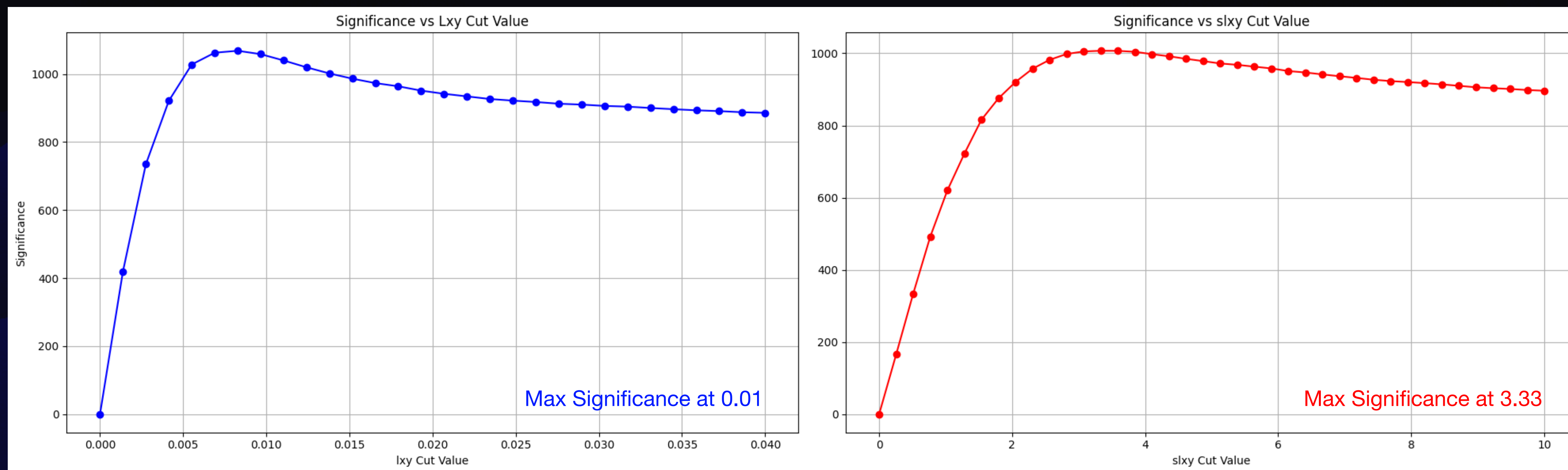
ii) Simple Cut vs Training: L_{xy}

- Train a model using vertex parameters[†], add MVA, optimise the MVA cut, observe the significance



The maximum significance is 1109.80 for an MVA cut of > -0.02.

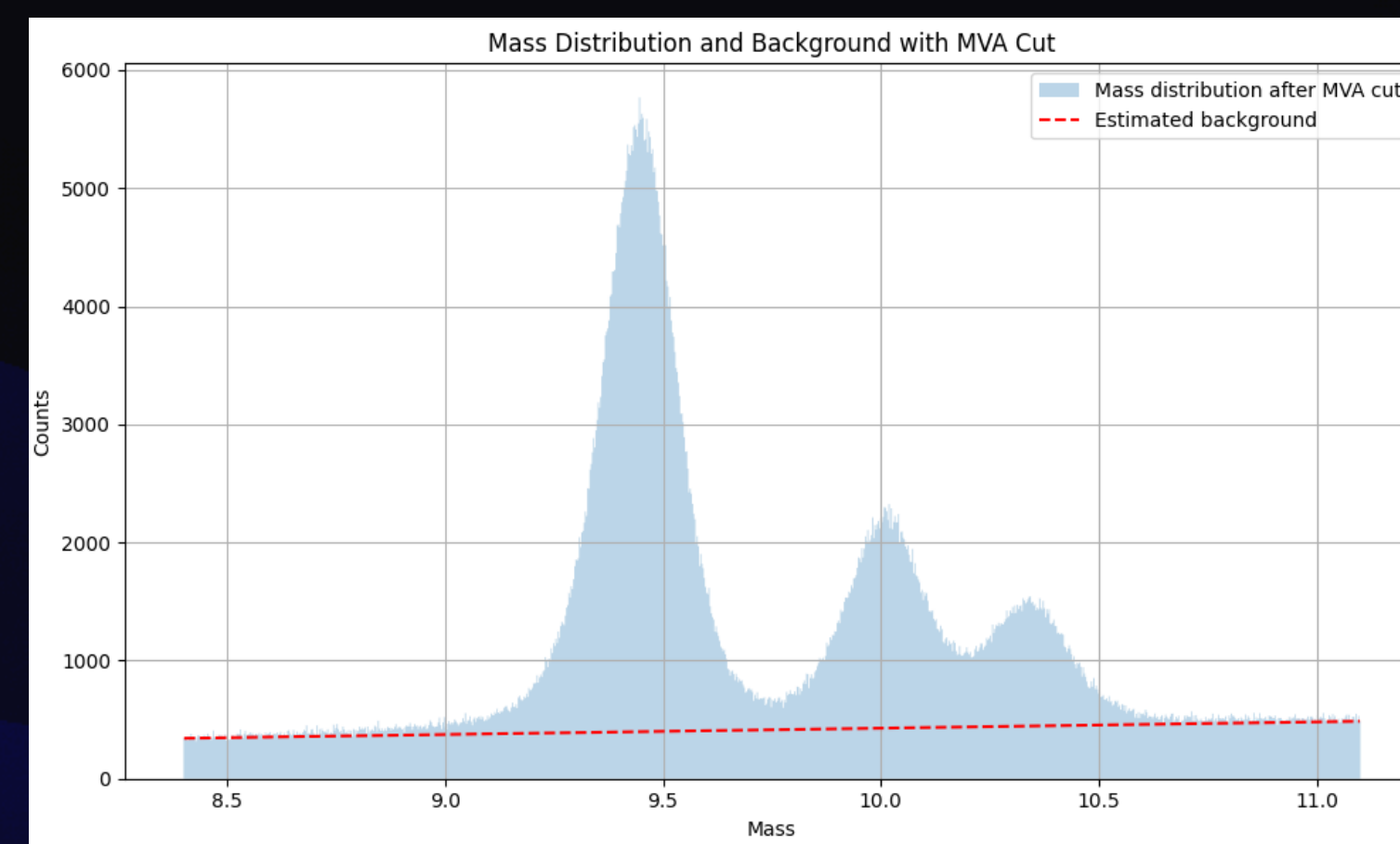
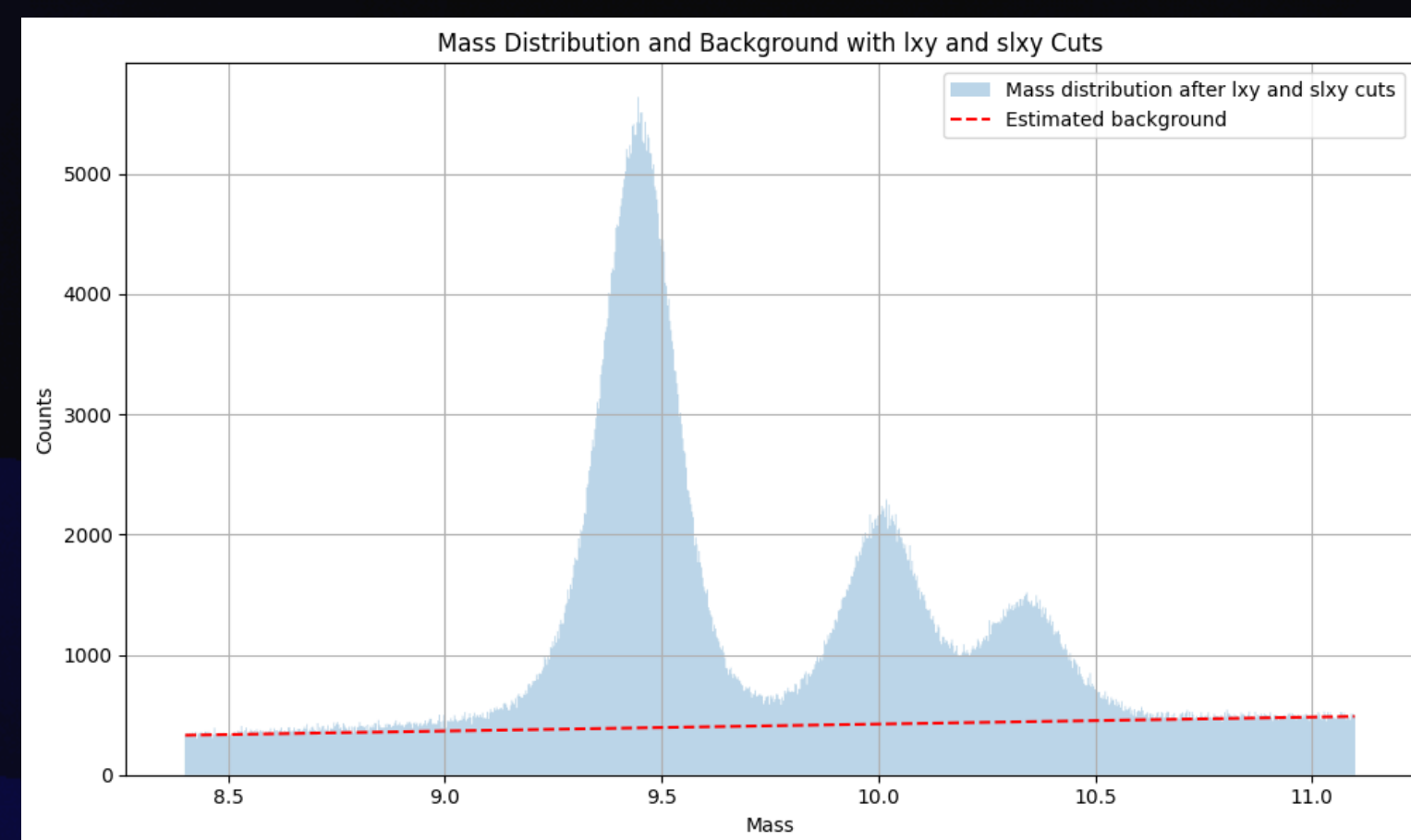
- Ignore the MVA, optimise the L_{xy} and Significance (L_{xy}/σ_{xy}) cuts



Max Significance after both cuts: 1072

Significance is higher with the MVA cut !

Comparison of the Results

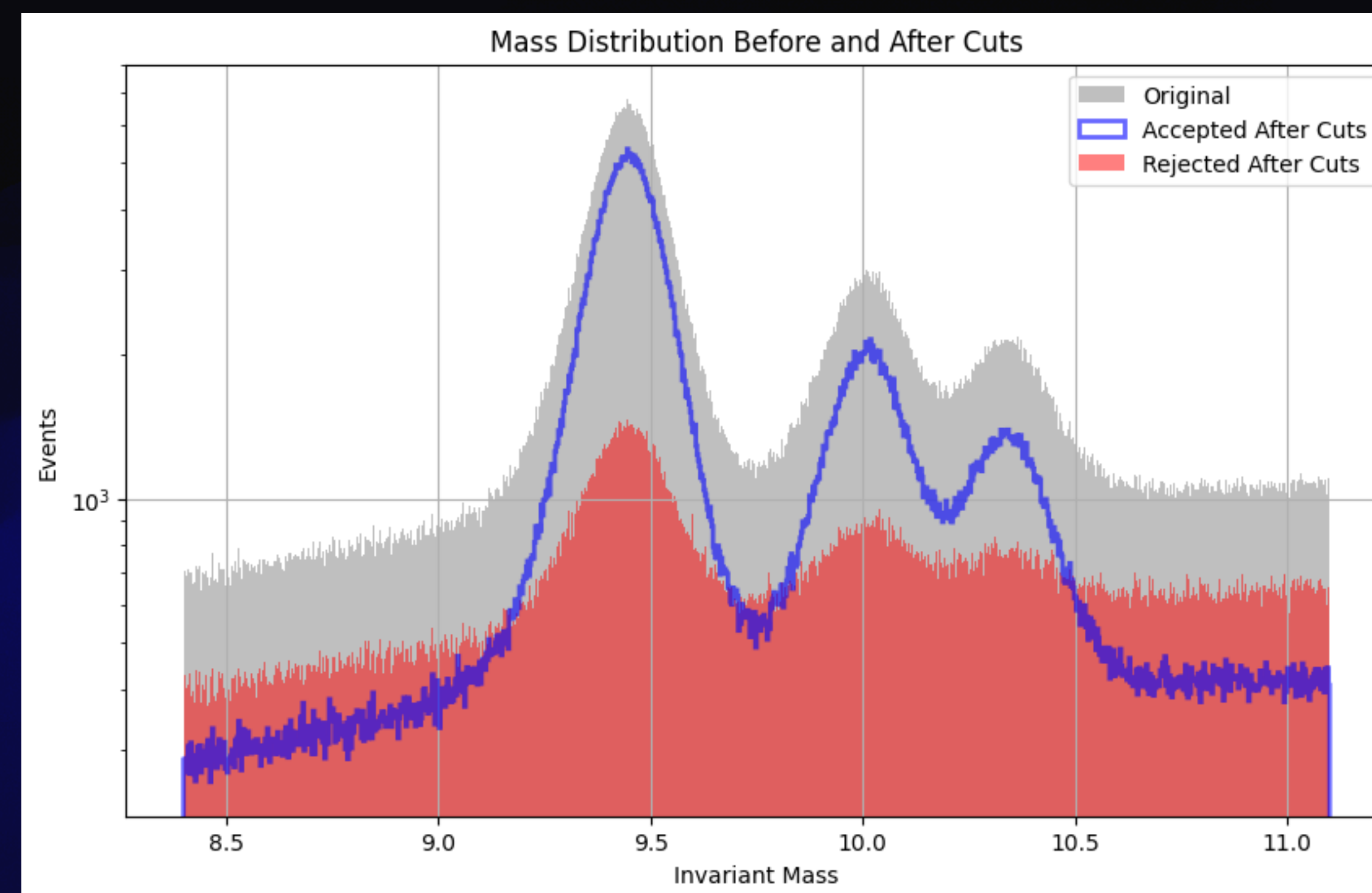


	Lxy + Slxy Cut	MVA Cut
Optimal significance	1072.85	1109.80
Signal efficiency at optimal cut:	0.8500	0.8846
Background rejection at optimal cut:	0.5320	0.5263

Vertexing Efficiency	
Data	0.850
MC	0.925
Ratio	0.919

Summary: Event Selection

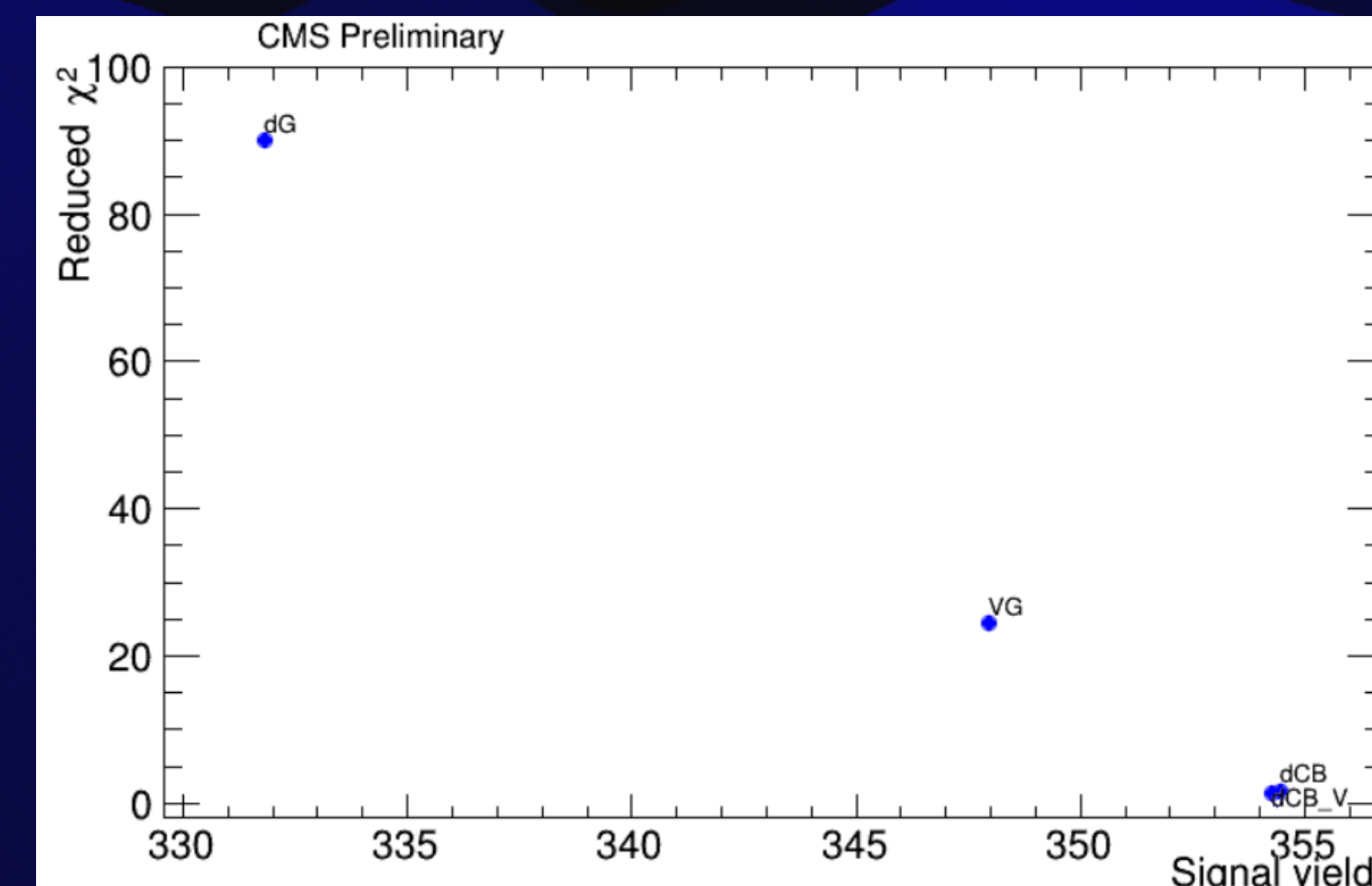
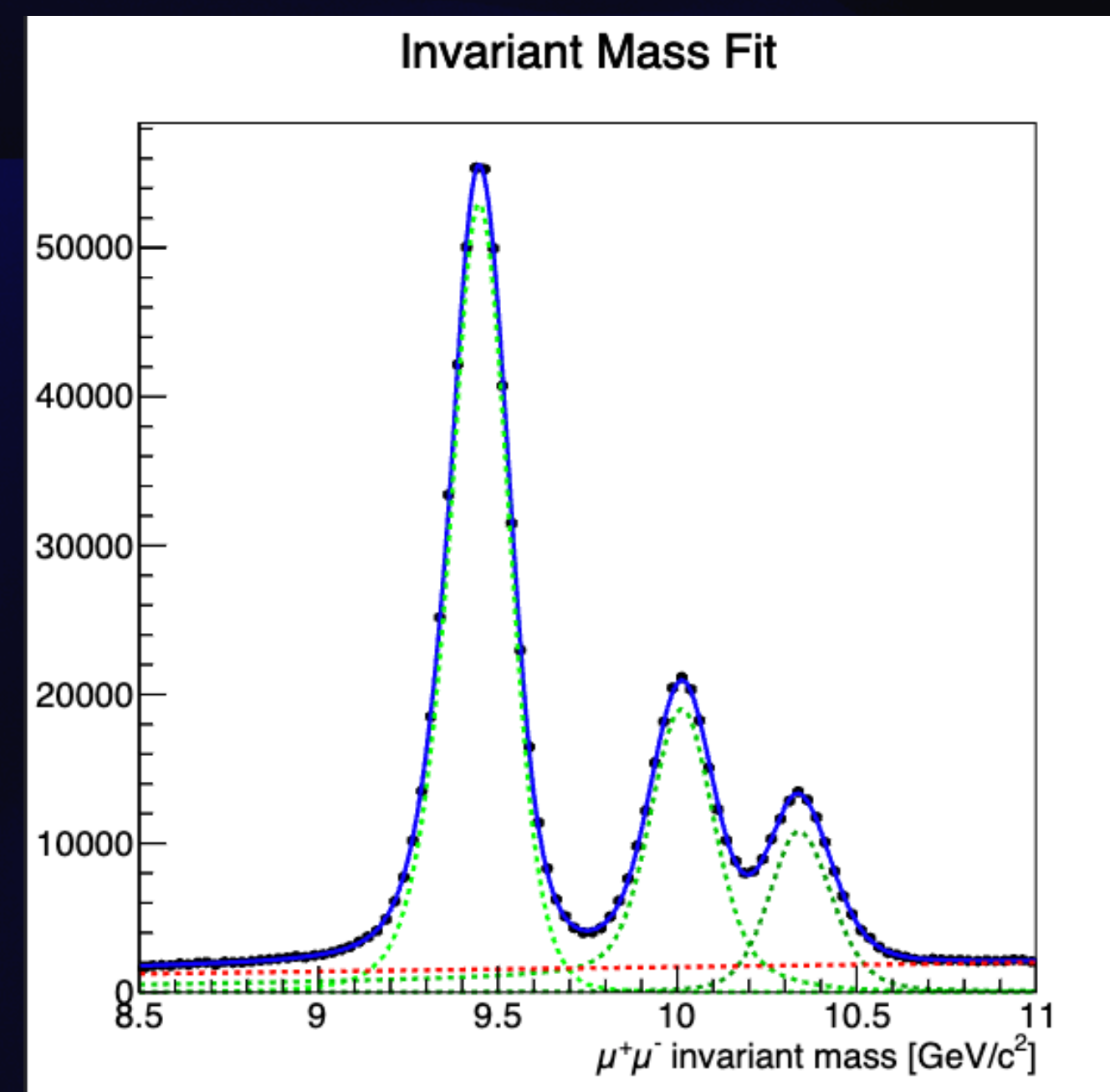
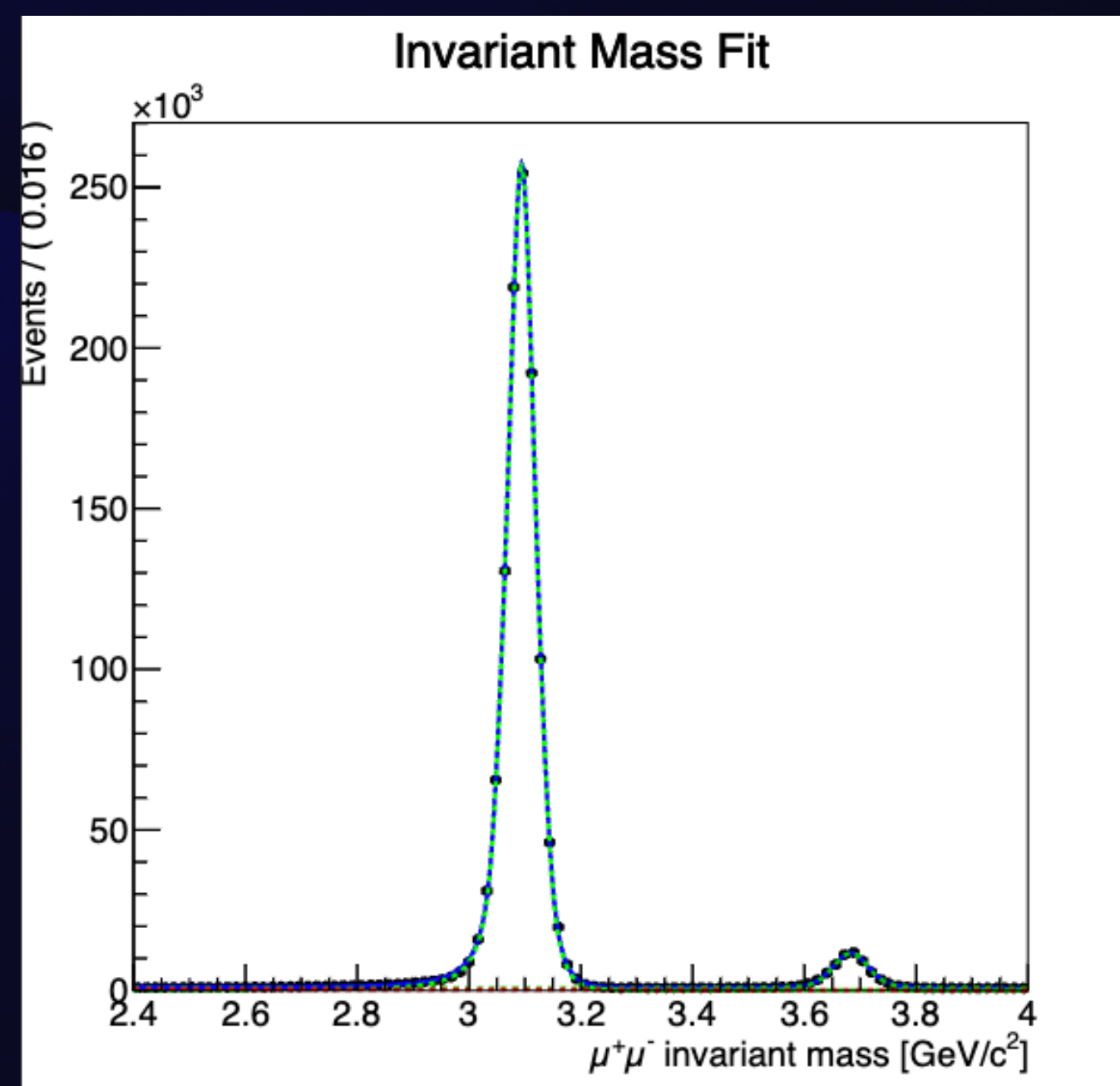
- Muon pair with opposite charges
- Transverse Momentum: $p_t > 4 \text{ GeV}$
- Pseudorapidity: $|\eta| < 1.9$
- MVA Cut: MVA Score > 0.015
- MVA_VTX Cut : MVA_VTX > -0.02



3) Signal and Background Modelling

i) Signal Modelling

- Tested Models:
- Gaussian
 - Voigtian
 - dCB + Gaussian
 - dCB + Voigtian (best model, also see Ludo's study)



From Ludo

ii) Background Modelling

Discrete profiling method to vary the function choice.

The families of functions up for investigation using RooMultiPdf:

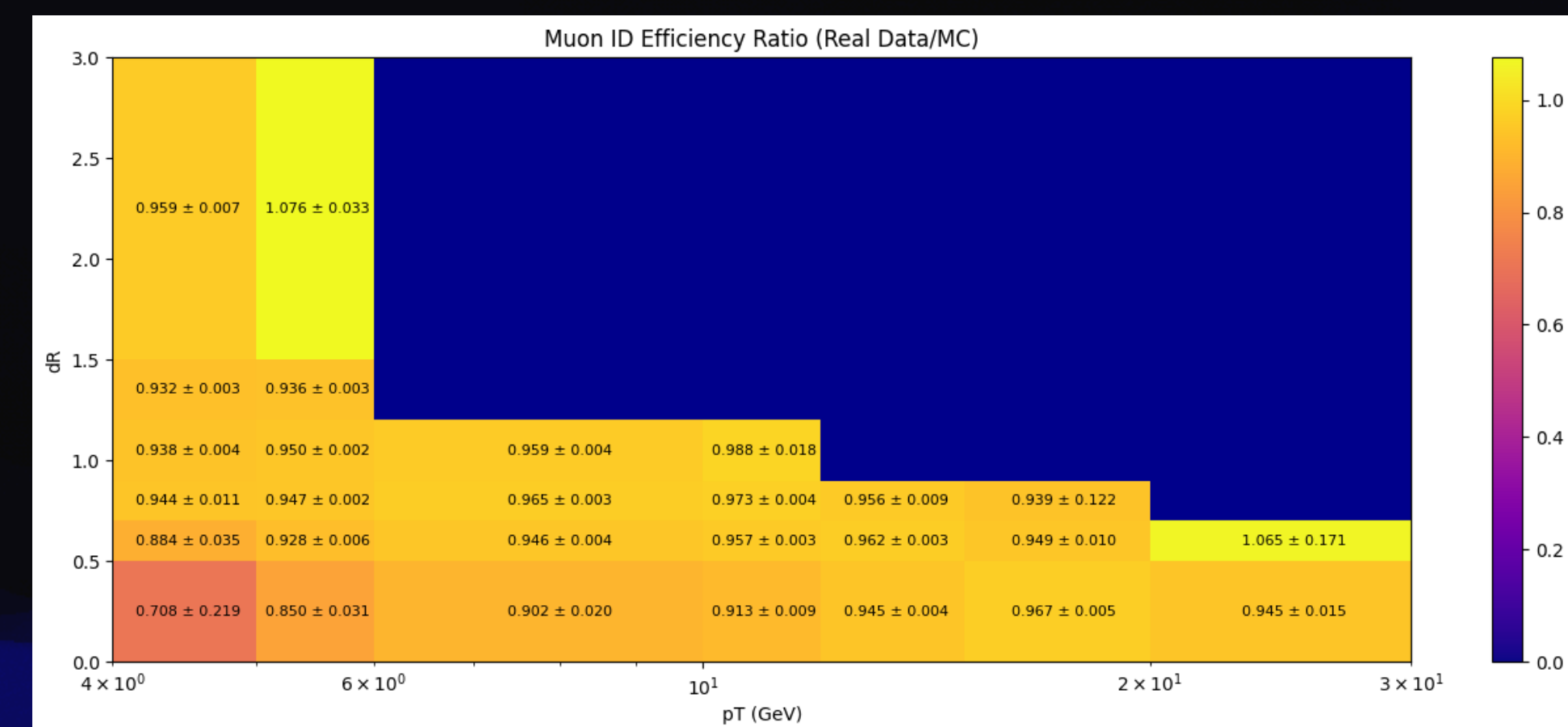
- **Bernstein Polynomial:** $B_n(x) = \sum_{v=0}^n \beta_v b_{v,n}(x)$, where $b_{v,n} = \binom{n}{v} x^v (1-x)^{n-v}$
- **Polynomial times exponential:** $P_n(x) = e^{c \cdot x} \sum_{n=1}^n \beta_n x^n$
- **Sum of exponentials:** $E_n(x) = \sum_{n=1}^n a_n e^{c_n \cdot x}$
- **Bernstein polynomial plus power law:** $B_{Pn}(x) = f B_n(x) + (1-f)x^a$

4) Systematic Uncertainties

Systematic Uncertainties

(i) On efficiency of data-driven selection

-Compare the efficiencies of MVA cut in Y region in data and MC (~6.9%)

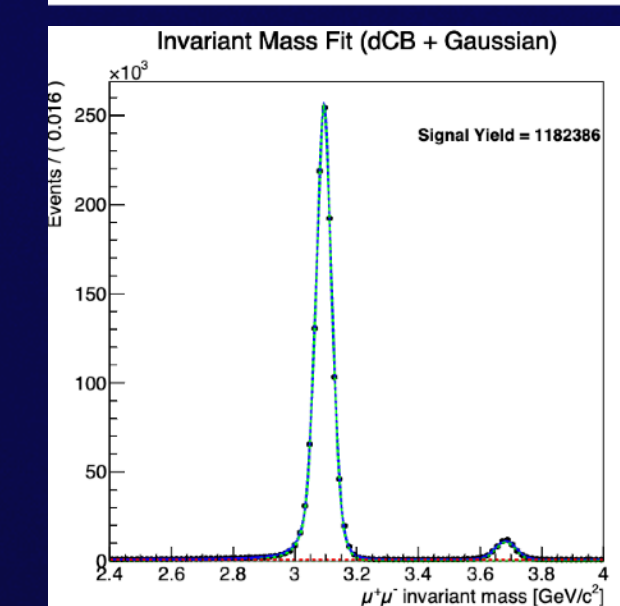
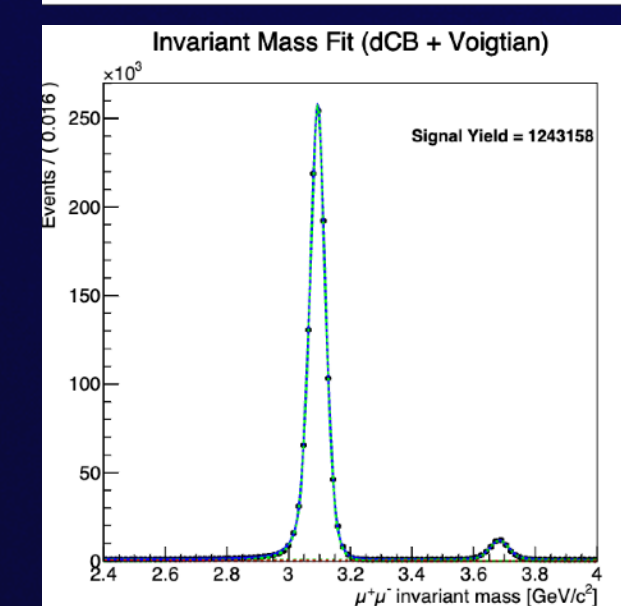
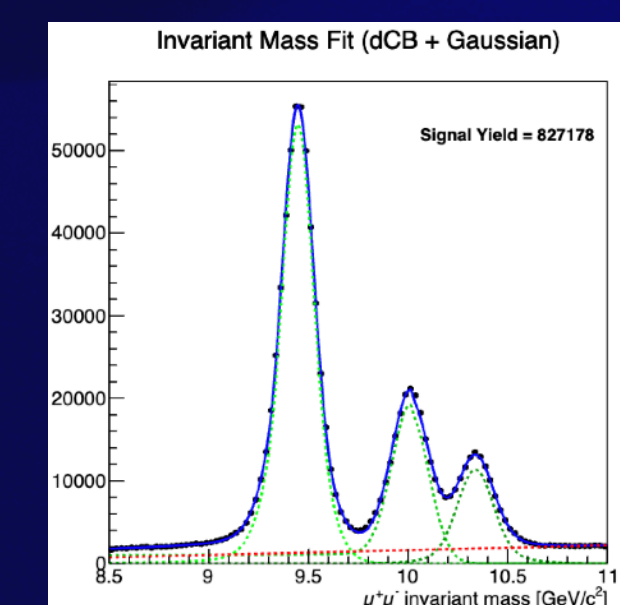
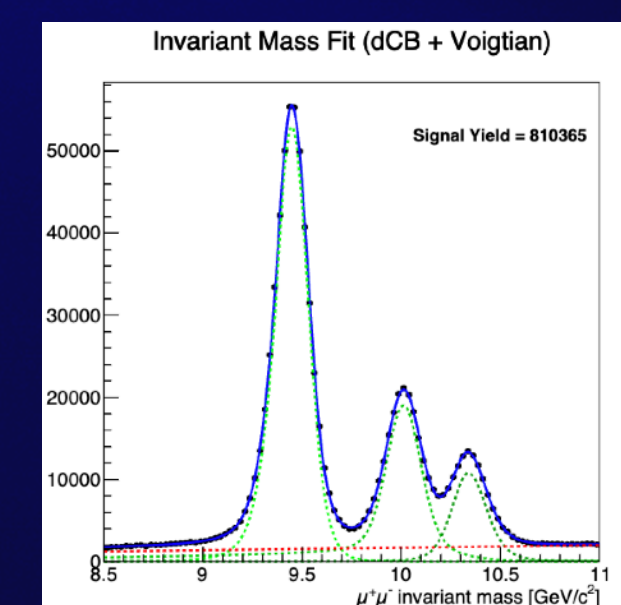


(ii) On signal modelling

-Compare yields of model candidates
-Negligible yield gap (~0.002)

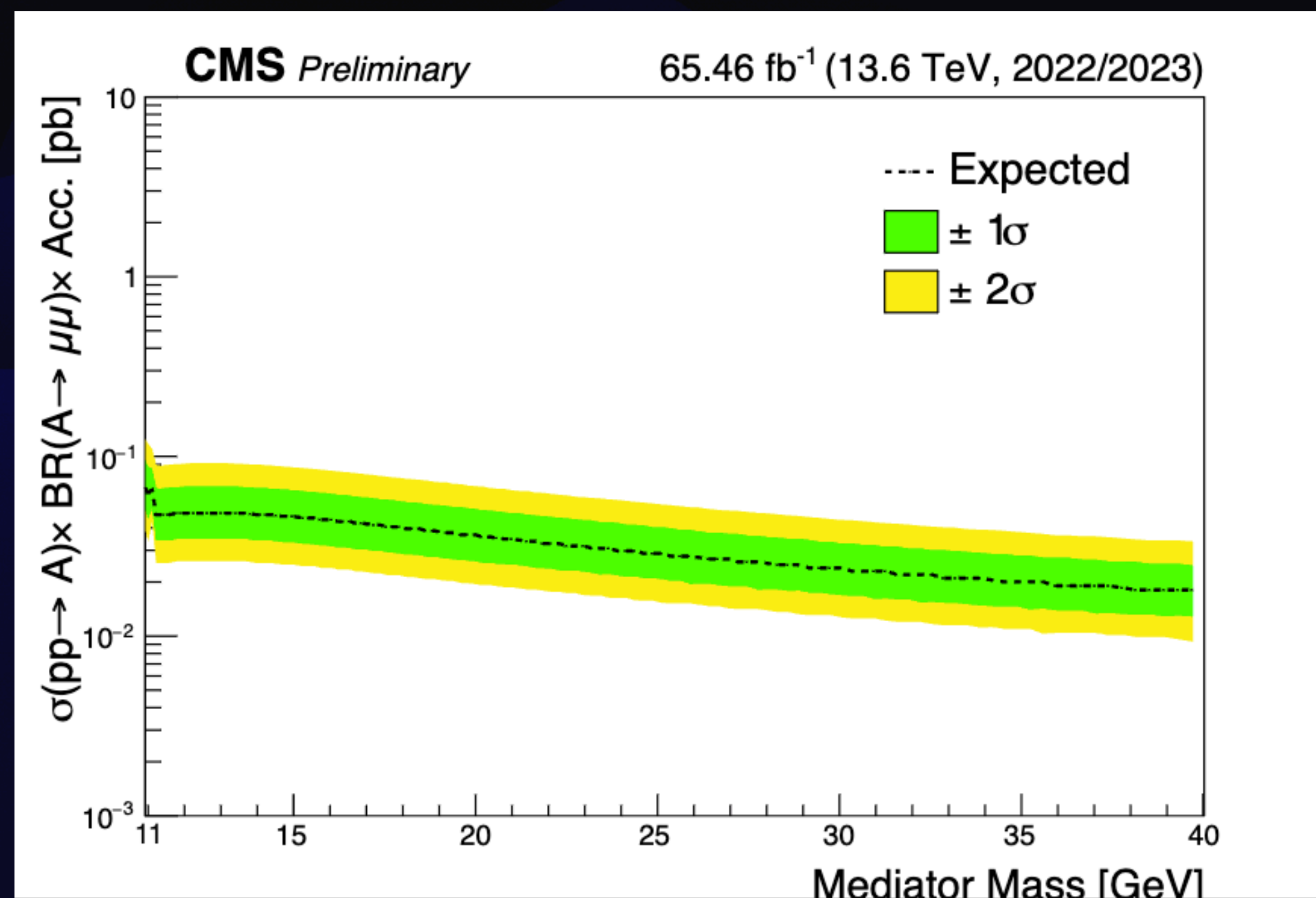
(iii) On Luminosity (65.46 fb⁻¹)

-Uncertainty ~ 2.3%



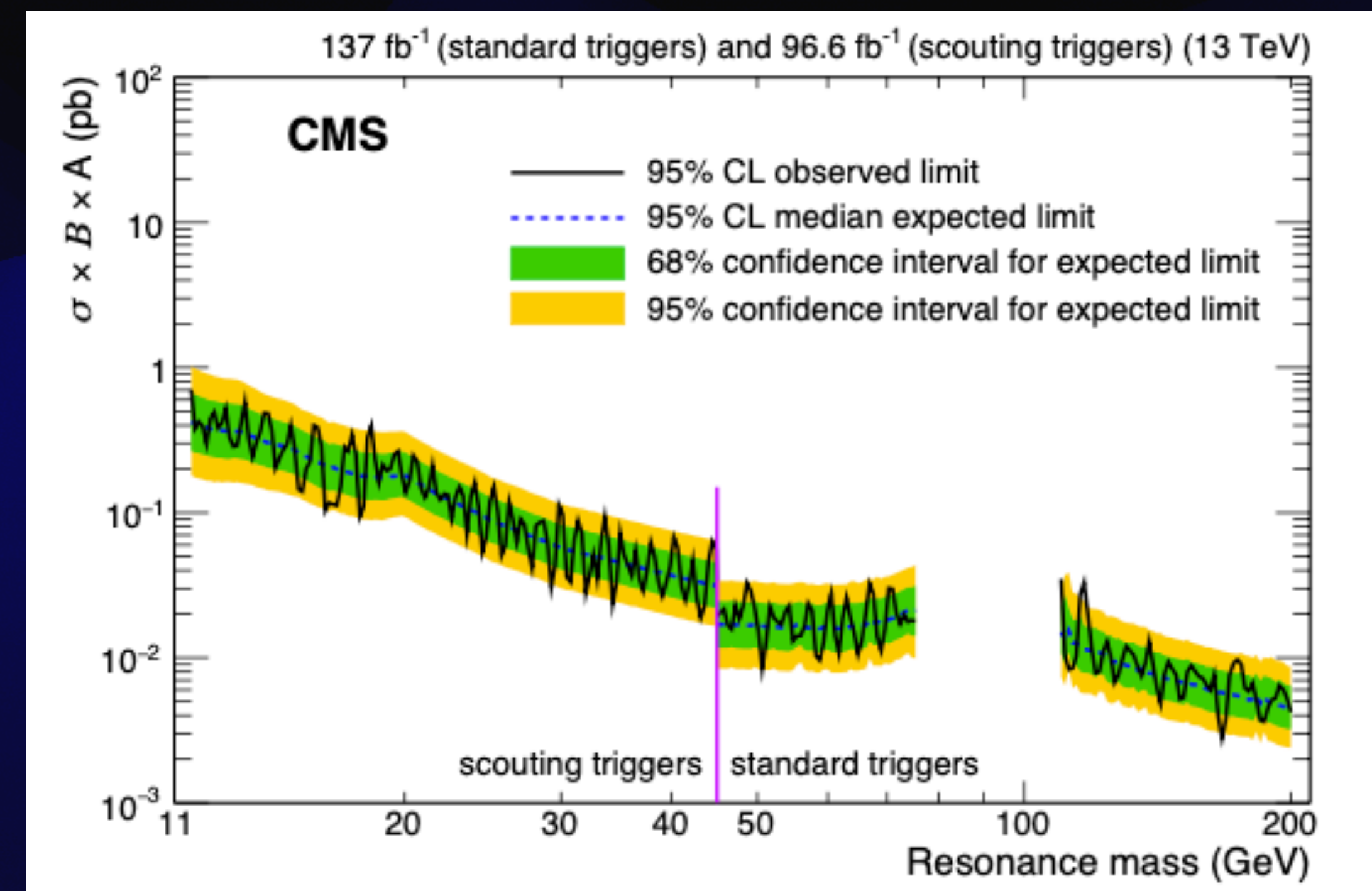
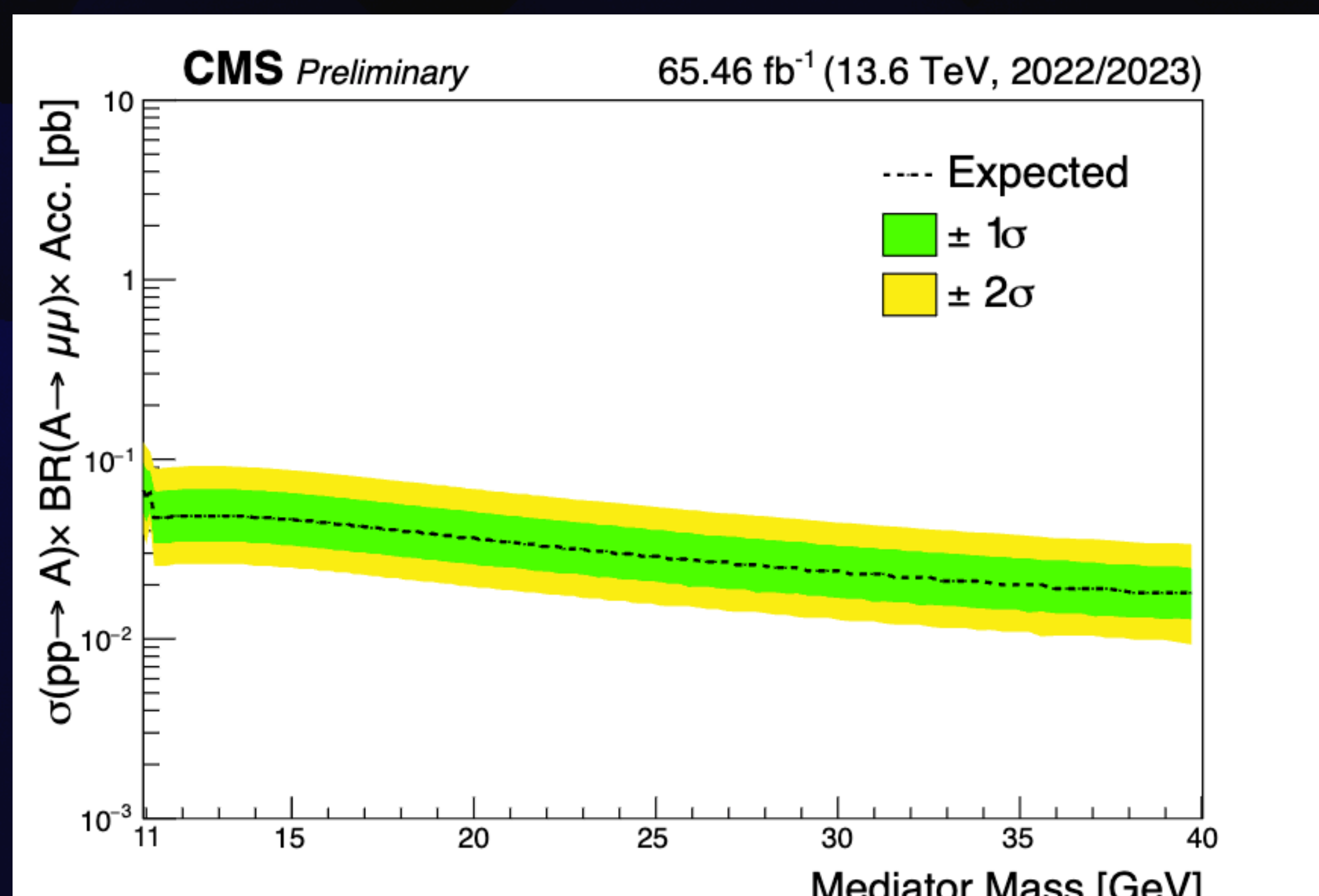
5) Exclusion Limits

Expected limits with discrete profiling



- Based on 65.46 fb⁻¹ of data collected at 13.6 TeV in 2022/2023, these limits constrain the parameter space of the minimal dark photon model.
- The trigger efficiency is taken from Run 2 and set to be constant at 0.80.

Comparison with Run 2 results



<https://arxiv.org/abs/1912.04776>

Summary

• Minimal Dark Photon Model

- Extends Standard Model with new $U(1)$ gauge symmetry
- Introduces dark photon (Z') interacting with SM particles via kinetic mixing
- Free parameters: kinetic mixing parameter (ϵ), dark photon mass ($m_{Z'}$)
- Experimental signatures: dark photons decay into SM particles ($\mu^+\mu^-$)

• CMS Run 3 Data and Scouting Trigger

- Utilized scouting trigger for data collection
- Analysis focused on 10-40 GeV mass range

• Event Selection and Efficiency Calculations

- Muon pair with opposite charges, transverse momentum ($p_T > 4$ GeV), pseudorapidity ($|\eta| < 1.9$)
- Neural network methods for MVA cut optimization
- Key variables: nmhits, trkiso, trkqoverp, trklambda, dxy, ntklayers, chi, nphits
- Efficiency calculated with Tag'n Probe fitting method

• Signal and Background Modelling

- Signal models: Gaussian, Voigtian, double Crystal Ball (dCB) + Gaussian, dCB + Voigtian (best) on multiple resonances within mass range
- Background models: Bernstein Polynomial, Polynomial times exponential, Sum of exponentials, Bernstein polynomial plus power law
- Discrete profiling method for background

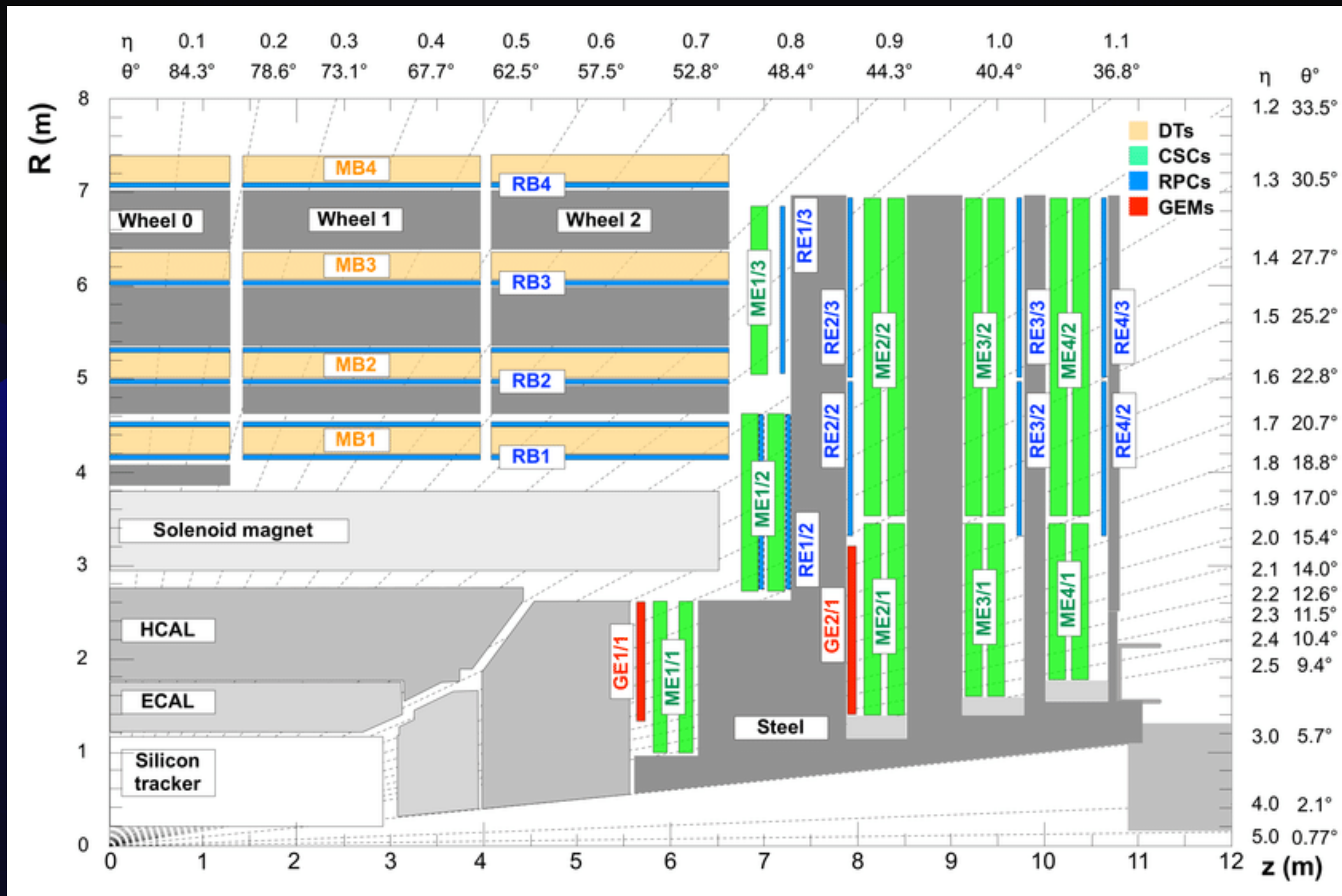
• Systematic Uncertainties

- Data-driven selection efficiency: $\sim 6.9\%$
- Signal modelling: Insignificant yield gap (~ 0.002)
- Luminosity uncertainty: $\sim 2.3\%$ (65.46 fb^{-1})

• Exclusion Limits

- Based on 65.46 fb^{-1} of data at 13.6 TeV (2022/2023)
- Limits constrain minimal dark photon model parameter space

Thank You!



CMS Muon Detectors