





**Alliance for Astroparticle Physics** 

# Production of CaWO<sub>4</sub> Crystals for the CRESST Experiment & Development of Neganov-Luke Amplified Cryogenic Light Detectors

HAP Dark Matter 2015

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22.09.2015

# Production of CaWO<sub>4</sub> Crystals for the CRESST Experiment

## **CRESST III**

- Already introduced by H. Kluck
- Focus on low-mass WIMPs

(see also: Probing low WIMP masses with the next generation of CRESST detectors, arXiv 1503.08065)

- → important factors: Performance (Threshold) & Radiopurity
- New, fully scintillating, detector modules (crystal mass ~ 25g)





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- Before CRESST-II Phase 2: only commercial CaWO<sub>4</sub> crystals (Russia & Ukraine)
  - Difficult: influence on selection of raw materials
  - No influence on crystal growth parameters
- Since ~ 2011: in-house production of CaWO<sub>4</sub> crystals at TUM for CRESST & EURECA\*
  - Full control over raw materials (screening & selection)
  - Full control of growth / post-growth treatment parameters!
- Dedicated Czochralski furnace for CaWO<sub>4</sub> crystals





\* A. Erb and J.C. Lanfranchi, Growth of high-purity scintillating CaWO4 single crystals for the lowtemperature direct dark matter search experiments CRESST-II and EURECA, CrystEngComm, 2013

#### HAP Dark Matter 2015



- Steady development towards complete in-house production of CaWO<sub>4</sub> crystals
  - Production of raw powder:  $CaCO_3 + WO_3 \rightarrow CaWO_4 + CO_2$
  - Crystal growth (Czochralski method) in Rh crucible (99%Ar / 1% O<sub>2</sub>)
  - Post growth treatment (annealing) to improve optical quality (100% O<sub>2</sub>)
  - Cutting & polishing of crystals
- Dedicated oven & powder grinding machine to produce CaWO<sub>4</sub> powder
- Dedicated annealing oven
- CRESST-II Phase 2 : 4 TUM-grown crystals
   CRESST-III : majority of crystals foreseen to be TUM-grown



 $CaCO_3 + WO_3$ 

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### Improved Radiopurity: TUM40

Total  $\alpha$  - Activities (1.5 - 7 MeV)

#### $e^{-/\gamma}$ Background events below 50 keV



TUM40: 3.07 ± 0.11 mBq/kg commercial crystals: 3 - 107 mBq/kg (Münster et al., JCAP05(2014)018) TUM40: 3.51 / (keV kg day) commercial crystals: 6 - 30 / (keV kg day) (Strauss et al., JCAP06(2015)030)

→ Significant improvement in radiopurity!

For TUM40 results see: EPJC (2015) 75:352

### Further Improvement of Radiopurity

#### recrystallisation of CaWO<sub>4</sub> crystals

- growth of multiple crystals
   (segregation of impurities during crystal growth)
- cleaning of crucible / removal of remaining melt
- recrystallisation of produced crystals
- → growth of crystal with improved radiopurity
- chemical cleaning of raw materials
  - CaWO<sub>4</sub> powder cannot be cleaned (due to chemical / physical properties)
    - $\rightarrow$  cleaning of raw materials (currently CaCO<sub>3</sub>)
  - Preliminary results: U extraction factor ~ 5000, Th extraction factor ~ 1000 (detailed ICP-MS / HPGe measurements ongoing, check that no new impurities are introduced!)
- → growth of crystal with improved radiopurity
- both methods currently under investigation (PhD thesis A. Münster (in preparation))

   → operation of crystal with further improved radiopurity planned for CRESST III

#### Further activities:

- Study influence of crystal growth / annealing on optical properties of CaWO<sub>4</sub> crystals (PhD thesis, A. Münster, in preparation)
- Investigate radiopurity of CaWO<sub>4</sub> via scintillation spectroscopy (PhD thesis, M. v. Sivers, TUM 2014 (now University of Bern))
- Investigate scintillation mechanisms in CaWO<sub>4</sub>

(PhD thesis S. Roth, TUM 2014 (now Queens University, Kingston), in collaboration with A. Ulrich, TUM E12)

Investigate scintillator non-proportionality

(Master thesis C. Bruhn, TUM 2015)

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#### **Summary:**

- Complete in-house production of CaWO4 crystals for CRESST / EURECA ✔
- Significant improvement in radiopurity ✓ (+ further potential for improvement!)
- Ongoing activities to study scintillation mechanism & optical properties

# Development of Neganov-Luke Amplified Cryogenic Light Detectors

→ Goal: Improve sensitivity (threshold / resolution) of cryogenic light detectors relevant for background suppression in both DM ( $e^{-\gamma}$ / nuclear-recoil separation) and  $0\nu\beta\beta$  ( $e^{-}/\alpha$  separation) experiments!



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→ Drifting of charge carriers in Electric field leads to increased phonon signal

$$G_{th} = 1 + rac{eV_{NL}}{E_{ph}/\eta}$$

Ideally: no increase in el. noise  $\rightarrow$  improvement in S/N ratio  $\rightarrow$  improved sensitivity

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• <u>A typical silicon-based NL light detector:</u>

*light absorber:* 20 x 20 x 0.5 mm<sup>3</sup>, p-type, ρ > 10kΩcm *aluminum electrodes* ~ 20 mm x 0.2 mm x 250 nm, d ~ 6-18 mm ( produced using a photolithographic shadow mask and Ar-Sputtering / EBE of Al )



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- Typical NL voltages: ~ 70 150 V
- NL amplification: ~ 10 @ 70V (~ 25 expected)
- additionally: reduction of amplification with time solved now  $\checkmark$
- Currently under Investigation: Why can we apply "only" up to 150V? Why is the amplification lower than expected?

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## Neganov-Luke: Effective Gain

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Stark et. al, NIM A 545 (2005)



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Reduction of signal amplitude:

- Accumulation of drifted charge carriers close to aluminum electrodes
   → Reduction of V<sub>eff</sub> (& G<sub>eff</sub>) over time
- Annealing of substrates ( PhD S. Roth TUM 2014 )  $\rightarrow$  Significant improvement



Isaila et al. Physics Letters B 716 (2012)

## Neganov-Luke Light Detectors - Planar Electrode Geometry



- Planar electrode geometry (shallow (implanted) ohmic contacts)
   V<sub>NL</sub> applied between top/bottom face of light detector
- → Drift of e/h pairs through bulk / *no drift across surface* of substrate!
- Very high electric fields possible (~ 6.5 kV/cm)
- First measurements promising: no increase in noise / linear amplification / no charge trapping
- Detailed measurements in preparation!

In collaboration with X. Defay (Excellence Cluster Universe & TUM) "Cryogenic silicon detectors with implanted contacts for the detection of visible photons using the Neganov-Luke Effect", arXiv 1509.06266

## Application: Background Suppression in TeO<sub>2</sub> Crystals

- <sup>130</sup>Te (TeO<sub>2</sub>) promising  $0\nu\beta\beta$  candidate
- α-events expected to be dominant background contribution in next-gen. experiments
- future-gen. experiments (CUPID arxiv 1504.03599)
   plan to employ active background suppression techniques
   → detection of Cherenkov photons for e<sup>-</sup> / α discrimination (expected signal at Q<sub>ββ</sub>(<sup>130</sup>Te): ~ 100eV)



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@TUM: first measurements showing feasibility of Cherenkov detection from TeO<sub>2</sub> using NL light detectors (2015 JINST 10 P03003)



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## Thank you for your attention!