

# Performance and future Evolution of SIBYLL

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with R. Engel, A. Fedynitch, T.K. Gaisser and T. Stanev

HAP composition workshop 2015

Institut für Kernphysik

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S I B Y L L  2.1
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HADRONIC INTERACTION MONTE CARLO
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BY
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Ralph ENGEL
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R.S. FLETCHER, T.K. GAISSER
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P. LIPARI, T. STANEV
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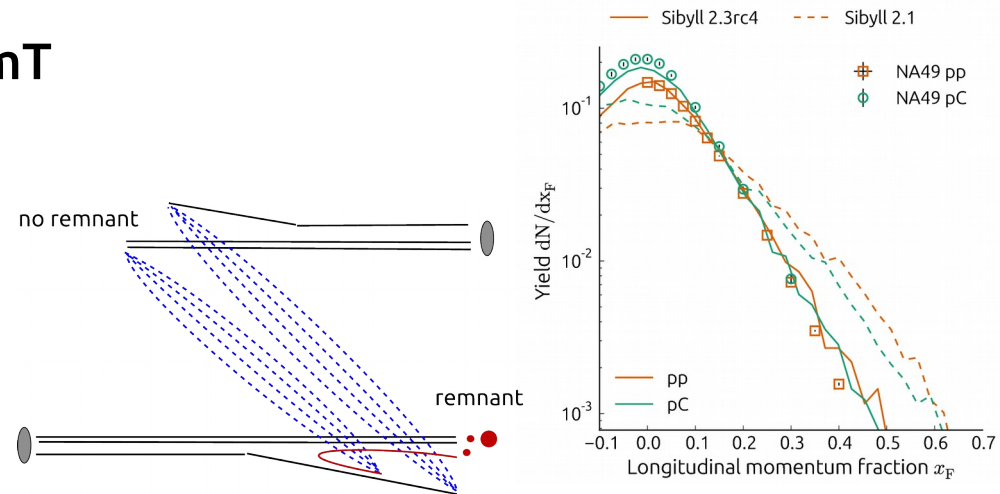
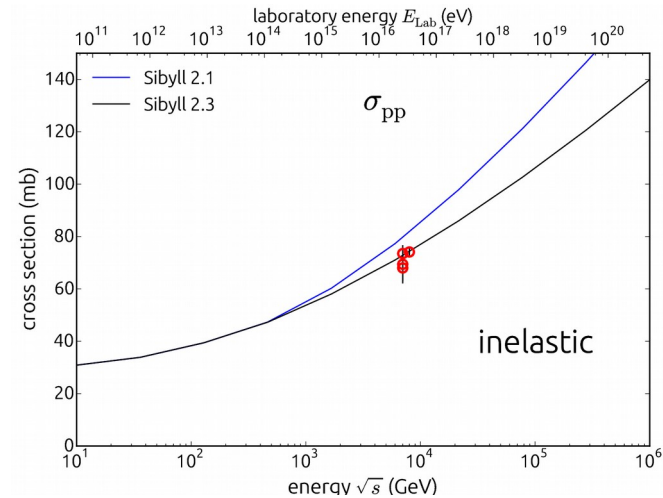
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Publication to be cited when using this program:  
R. Engel et al., Proc. 26th ICRC, 1 (1999) 415
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last modified: 28. Sept. 2001 by R. Engel
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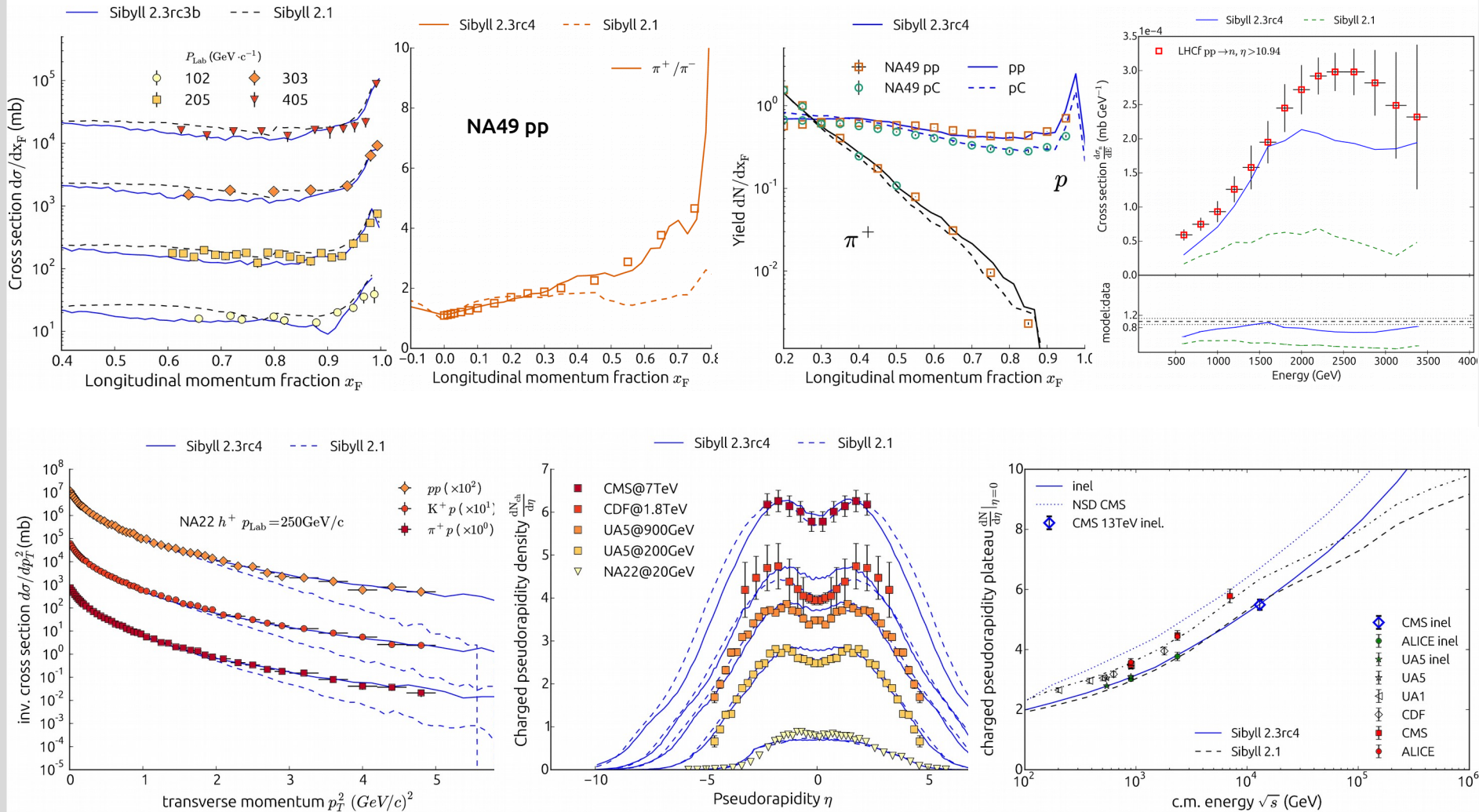
# Sibyll 2.3 – Release notes

Updates, changes to Sibyll 2.1  
(PRD 2009, last mod. 2001 !!):

- pp cross section
- Increased baryon pair production
- Added charm production
- Leading vector mesons
- Updated PDFs
- Changed pT – distributions to mT
- Revised leading particle model 'remnant treatment'
- Inel. Screening in nuclear collisions
- Bugfixes (off-shell particles, energy conservation)



# Sibyll 2.3 - Performance



# Muon number prediction

# Muon production in extensive air showers

Baryons:

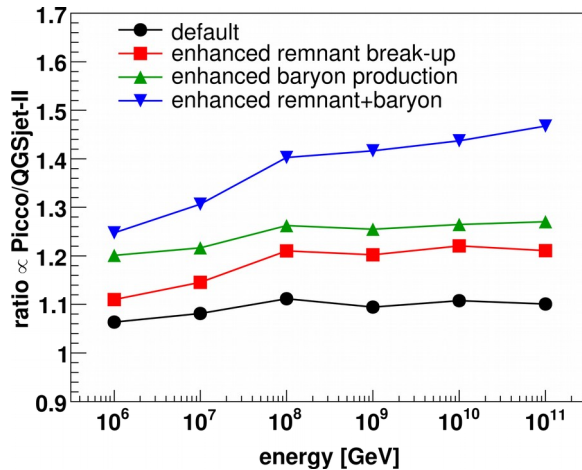
- More energy in hadrons
- Low energy muons

Neutral pions vs. rho:

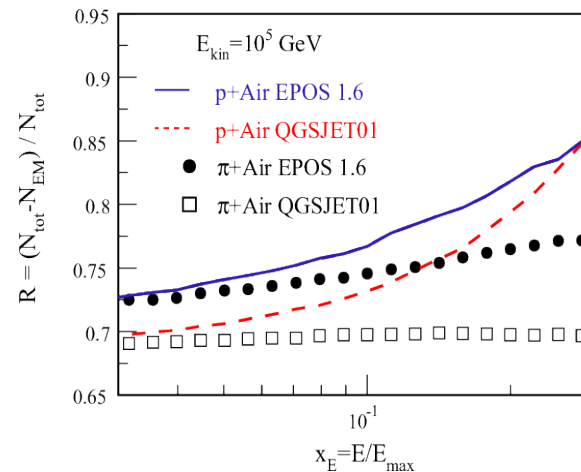
- Keep energy in hadrons
- Anti-correlation

Short-lived hadrons (charm):

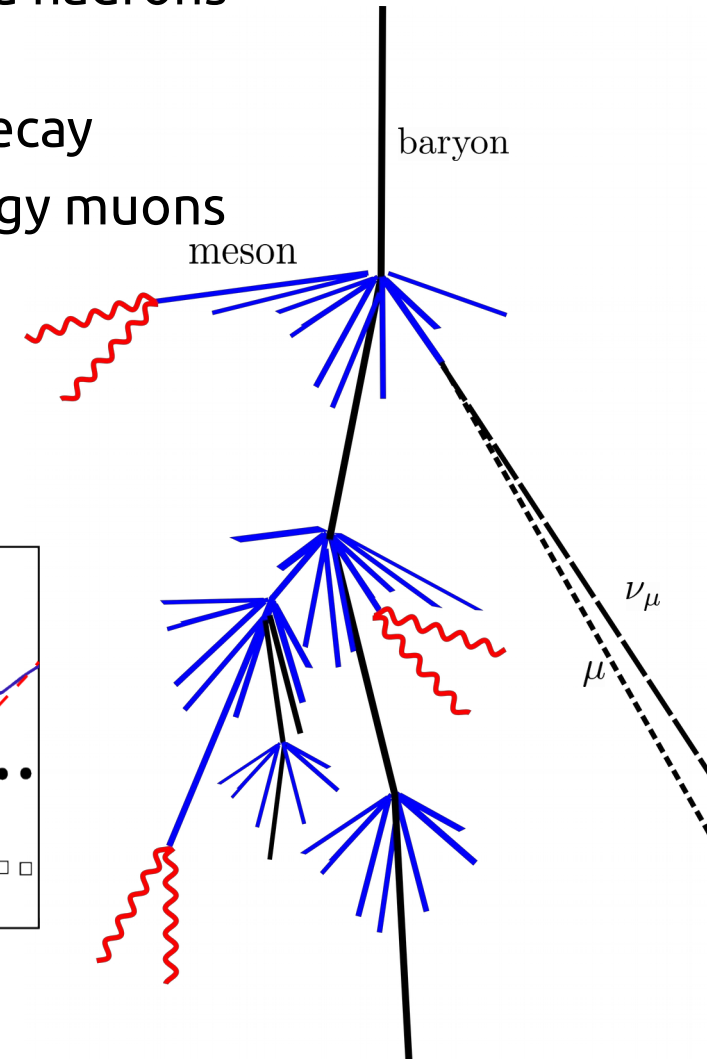
- Prompt decay
- High energy muons



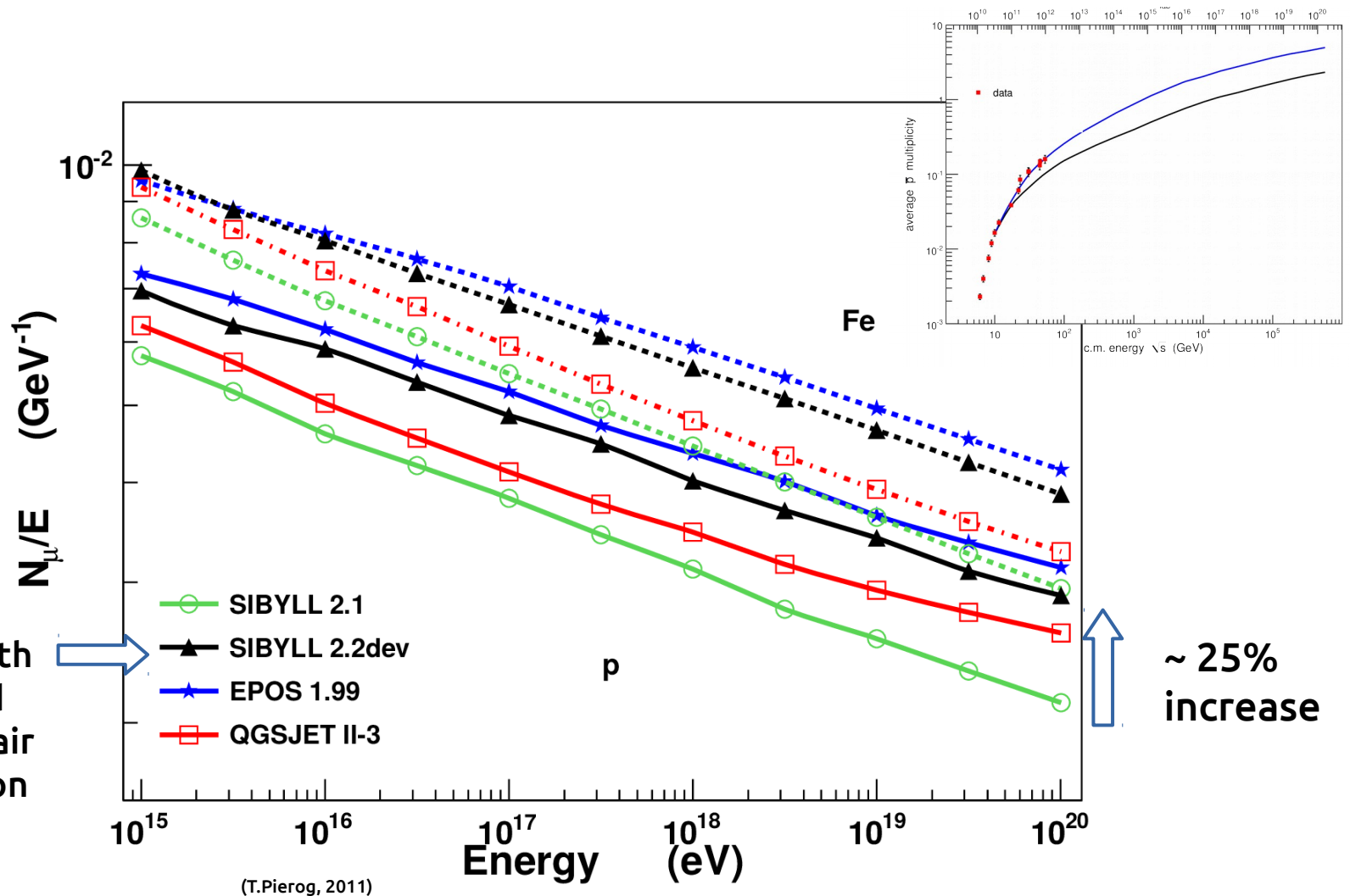
(Drescher, PRD 77 2008)



(Pierog, PRL 101 2008)



# Muon number with increased baryon prod.



Model with increased baryon pair production

(T.Pierog, 2011)

# Rho0 in air showers

- In itself not very important

- Negligible contribution to prompt muons

- New production mechanism!

- Data suggest few leading neutral pions in meson interactions

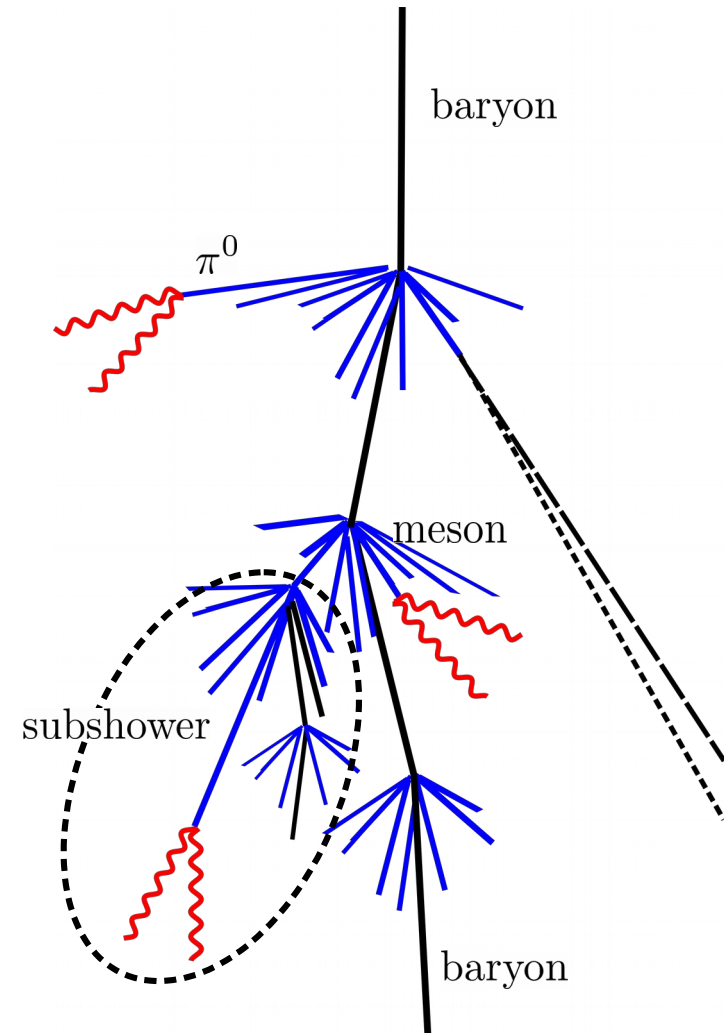
$$\pi^\pm + p \not\rightarrow \pi_{\text{lead}}^0 + X$$

Instead

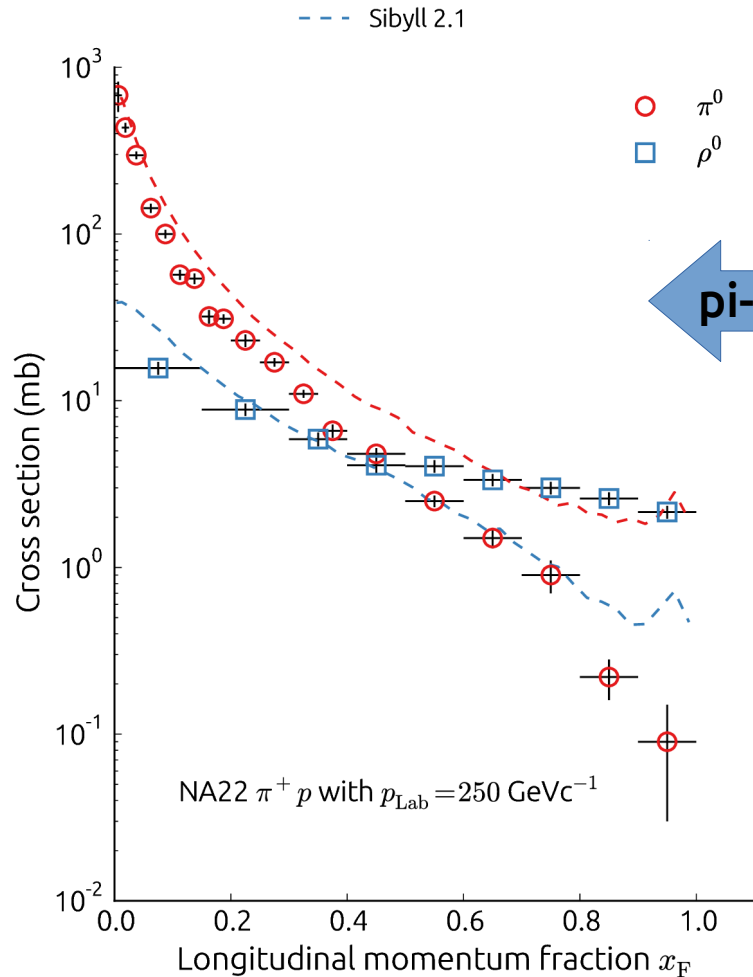
$$\pi^\pm + p \rightarrow \rho_{\text{lead}}^0 + X$$

- More hadronic subshowers!

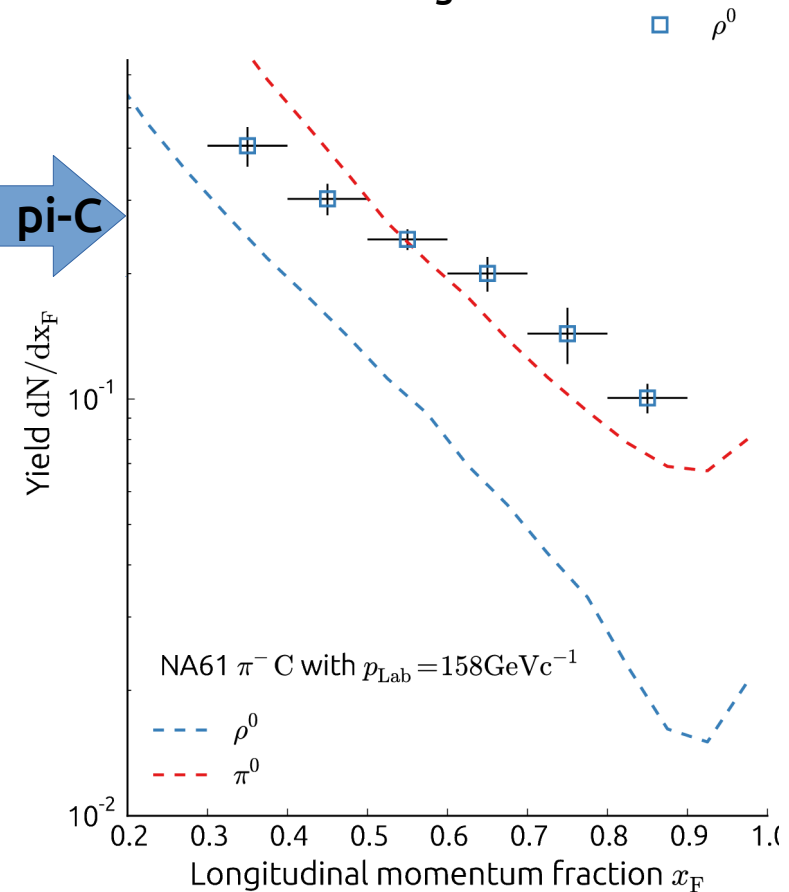
→ more muons



# Pi0 vs. Rho0 in pi-p and pi-C



NA61 (preliminarily) confirms rho enhancement for nuclear target!

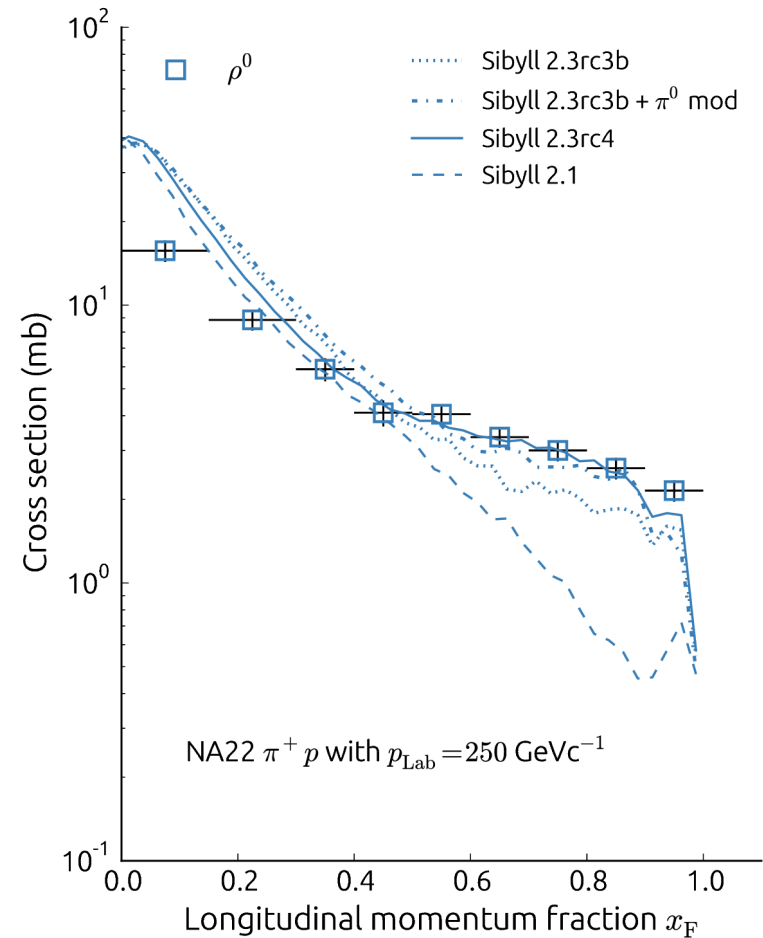
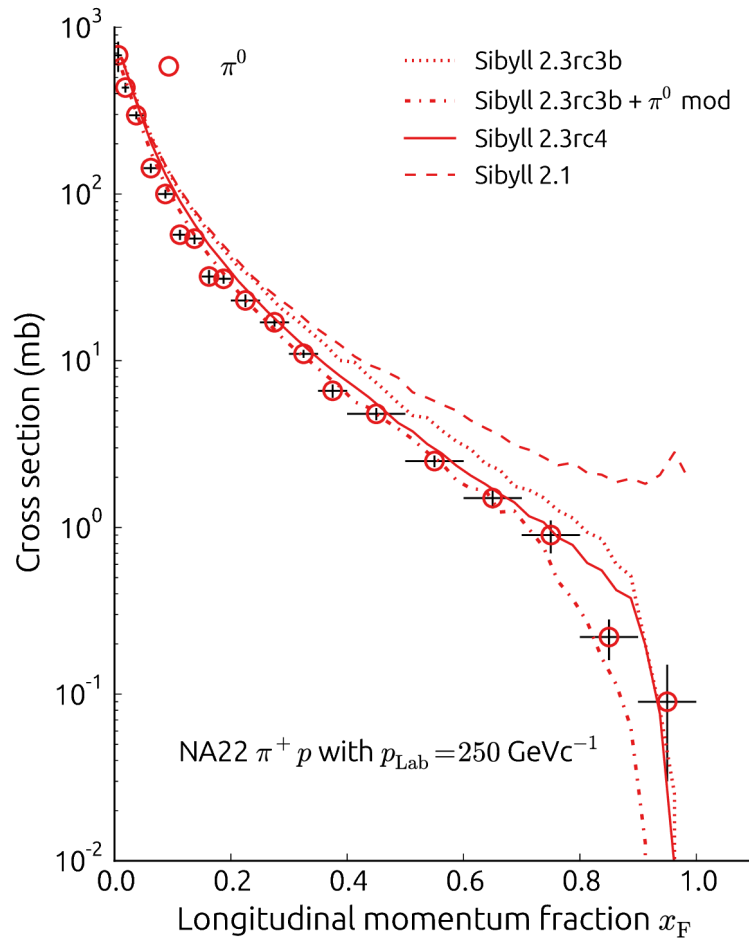


Presumably pi0 suppressed as in pi-p

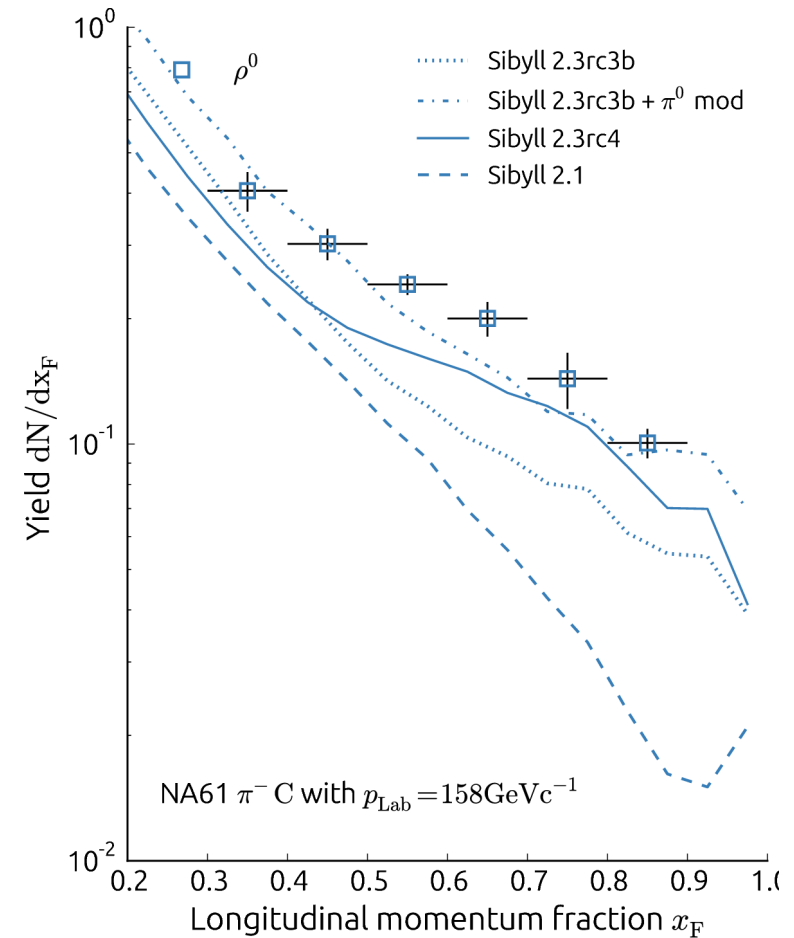
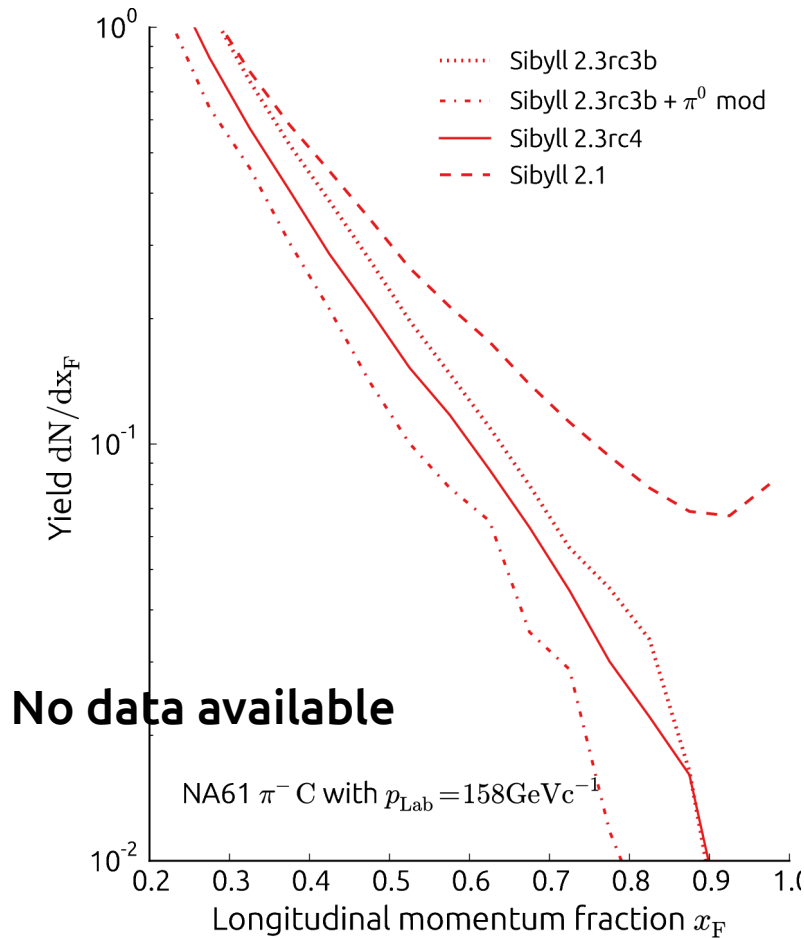
(NA61 data: Herve, ICRC 2015)



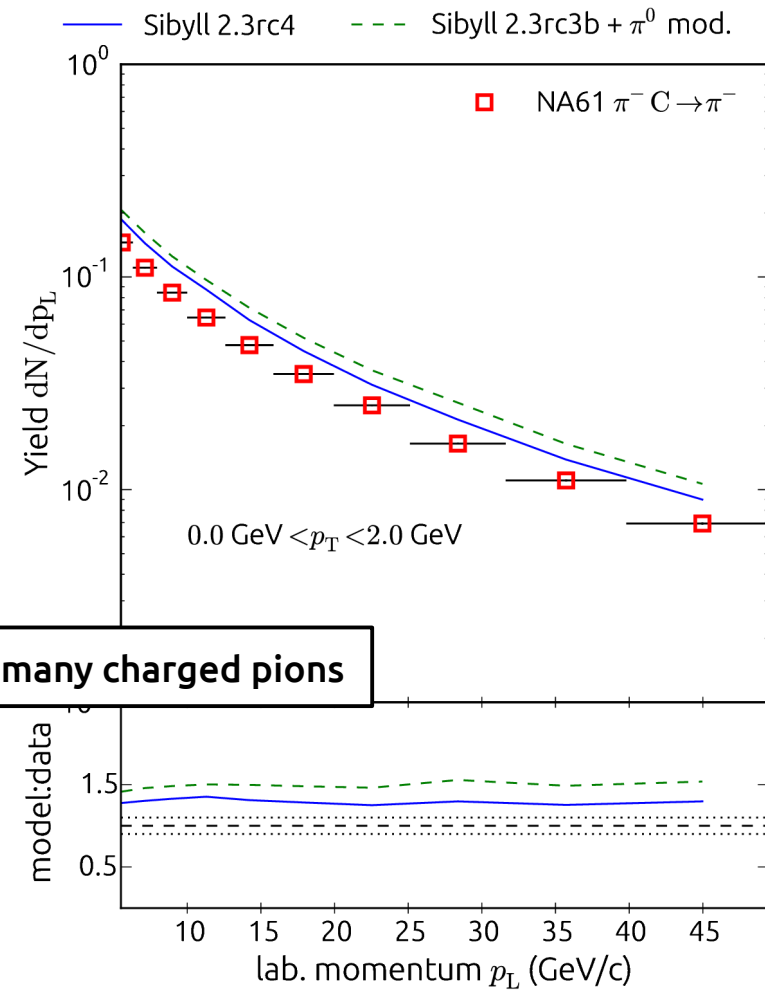
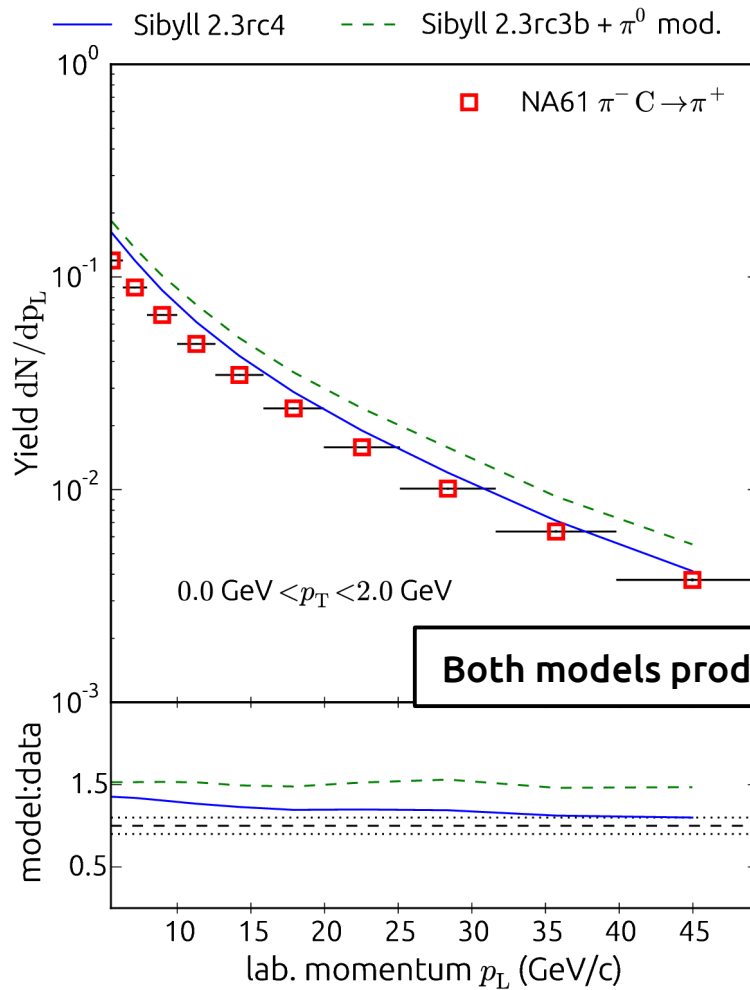
# NA22 $\pi^0$ / $\rho^0$ : four models



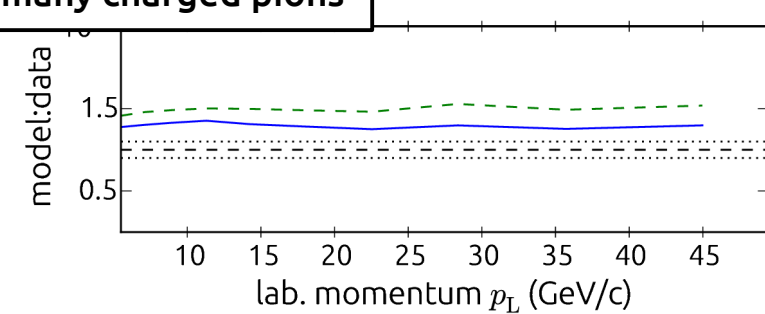
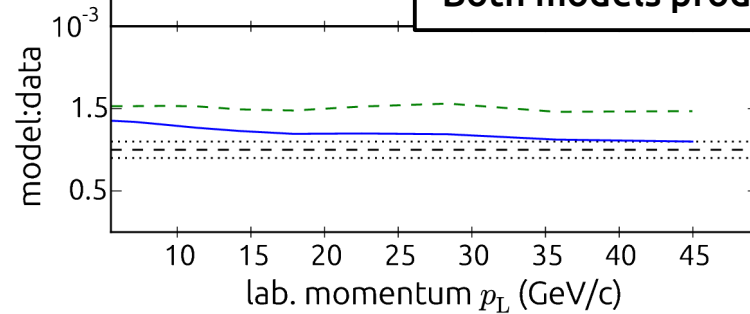
# NA61 pi0 / rho0: four models



# NA61: pion + carbon $\rightarrow$ charged pions

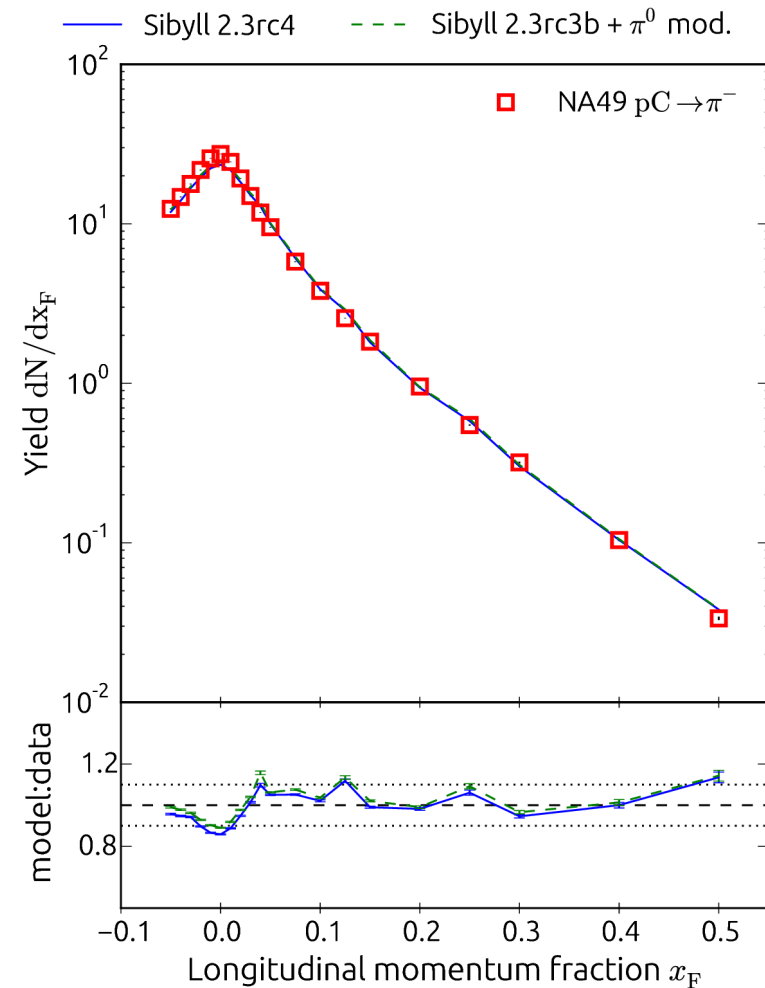
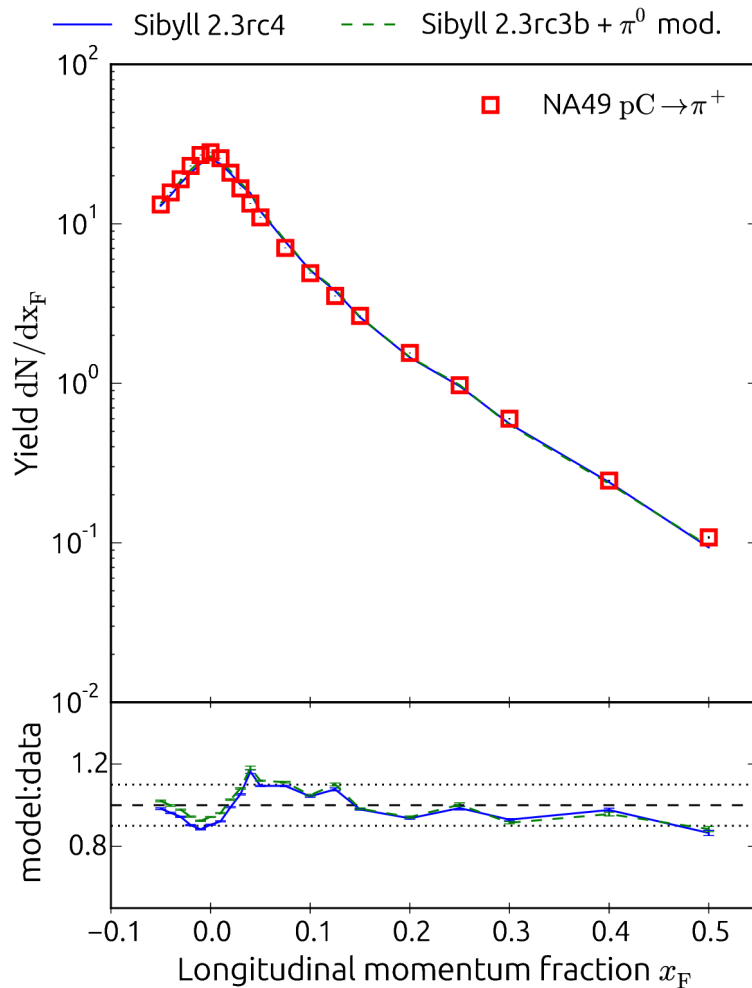


**Both models produce too many charged pions**



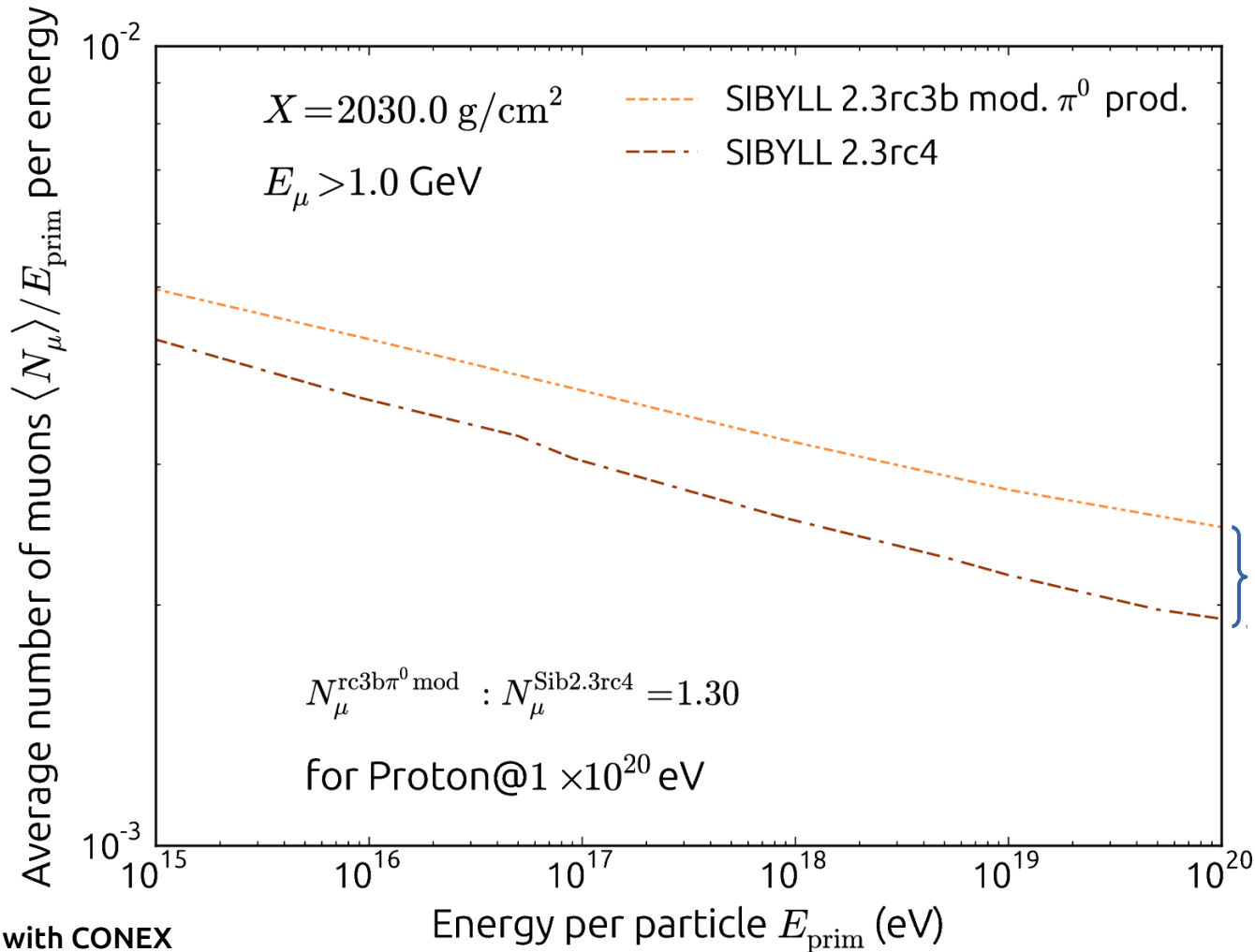
**Not much data available for Meson projectiles  
BUT central region expected to be ~universal (hadrons look alike on small scales)**

# NA49: proton + carbon $\rightarrow$ charged pions

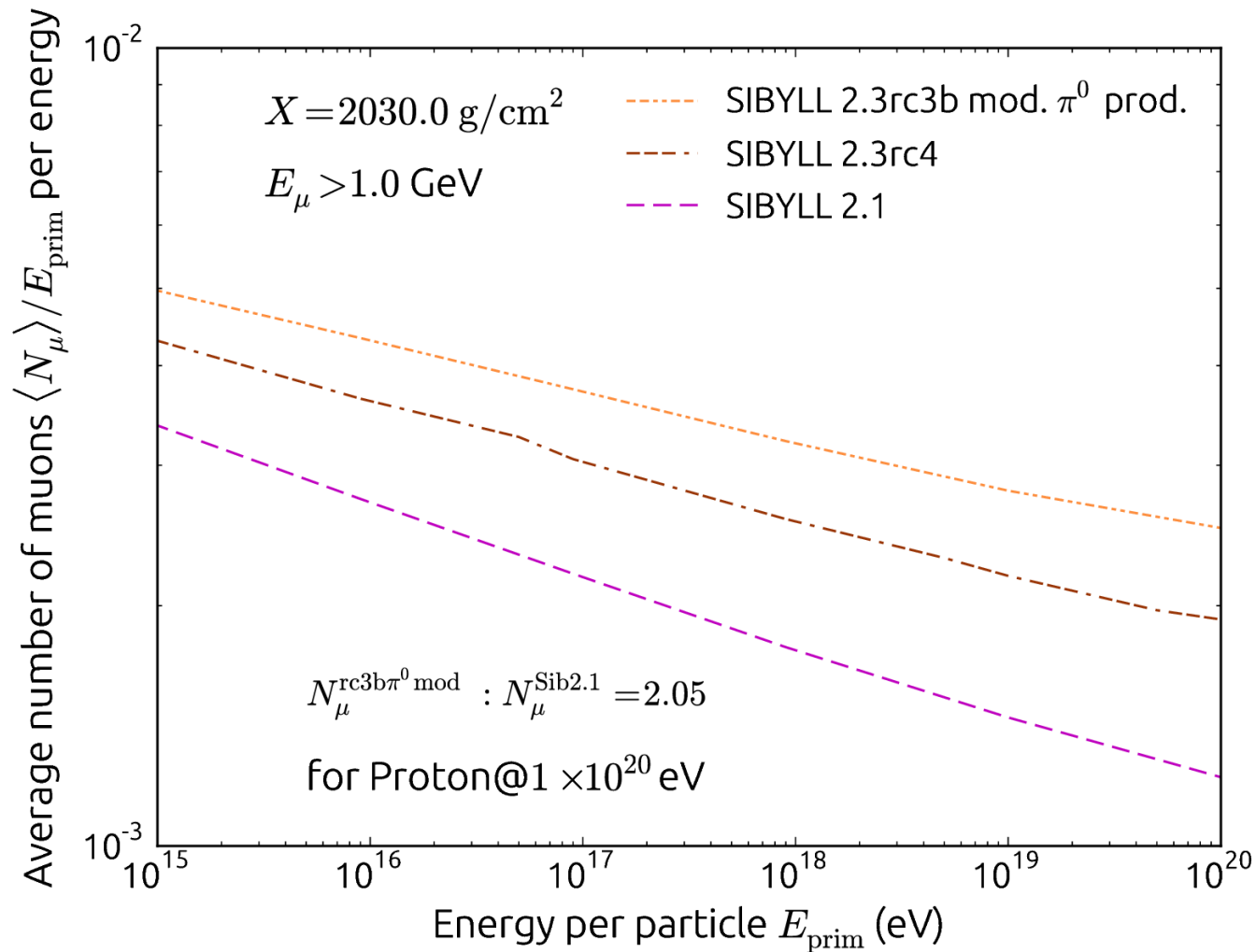


**Not much data available for Meson projectiles  
 BUT central region expected to be ~universal (hadrons look alike on small scales)**

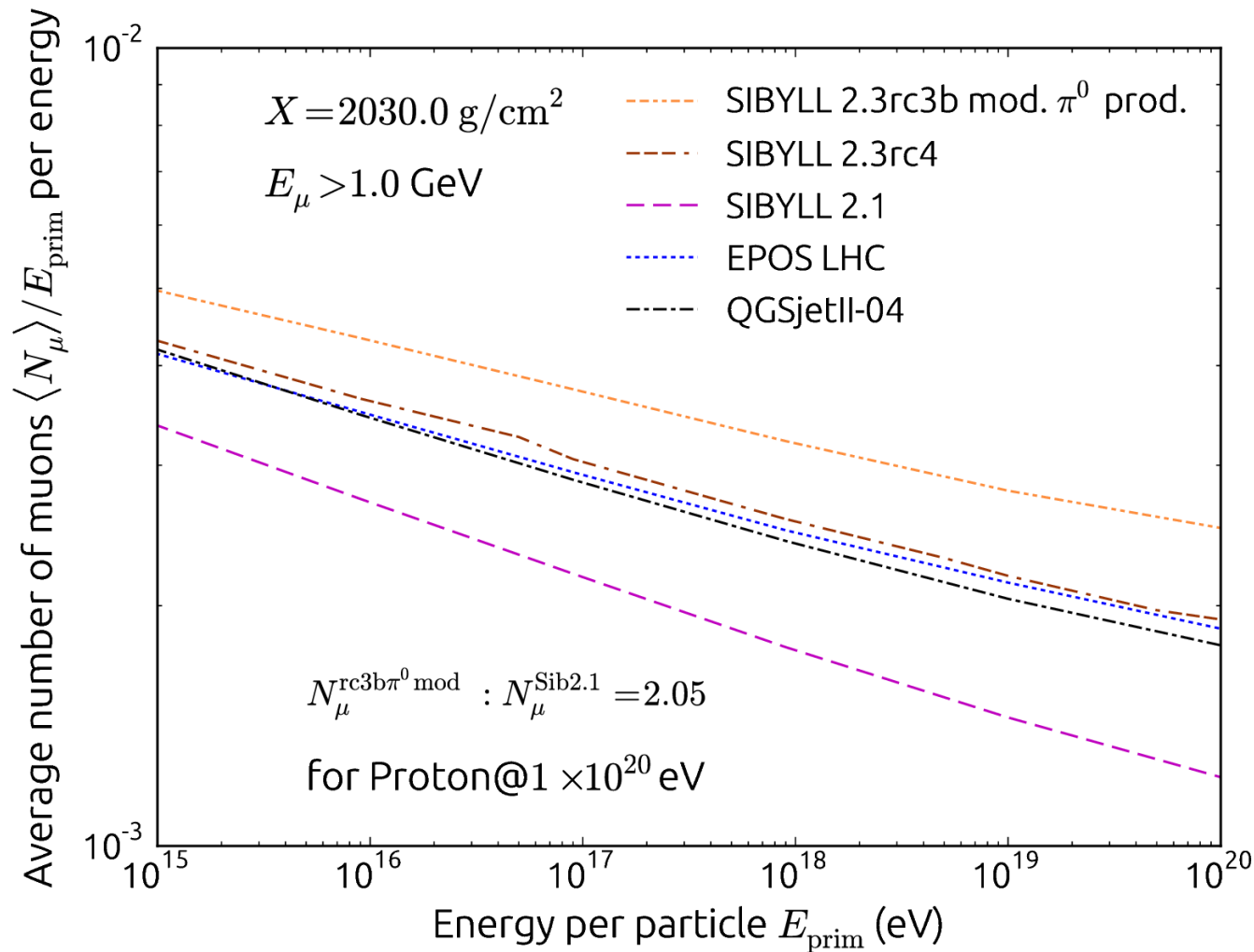
# Sibyll evolution: Number of muons



# Sibyll evolution: Number of muons



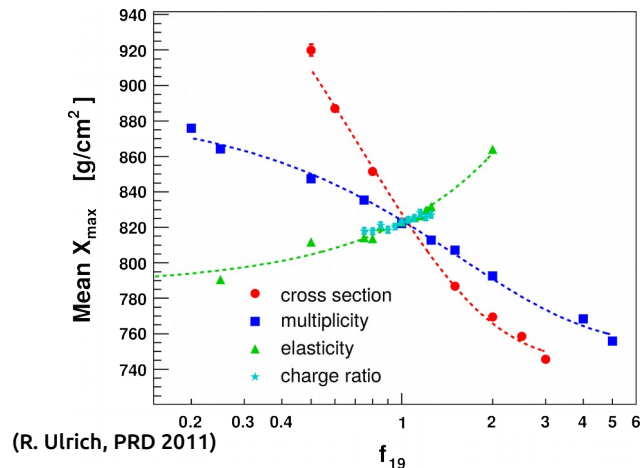
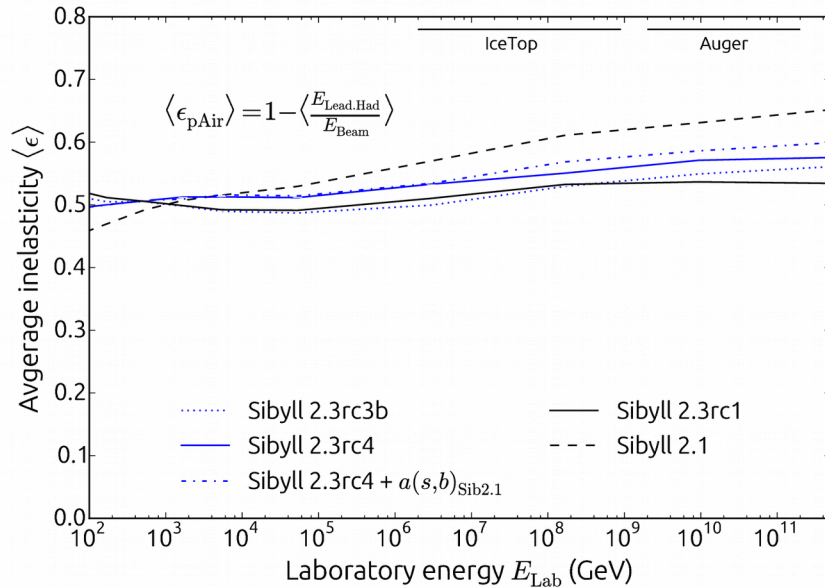
# Sibyll evolution: Number of muons



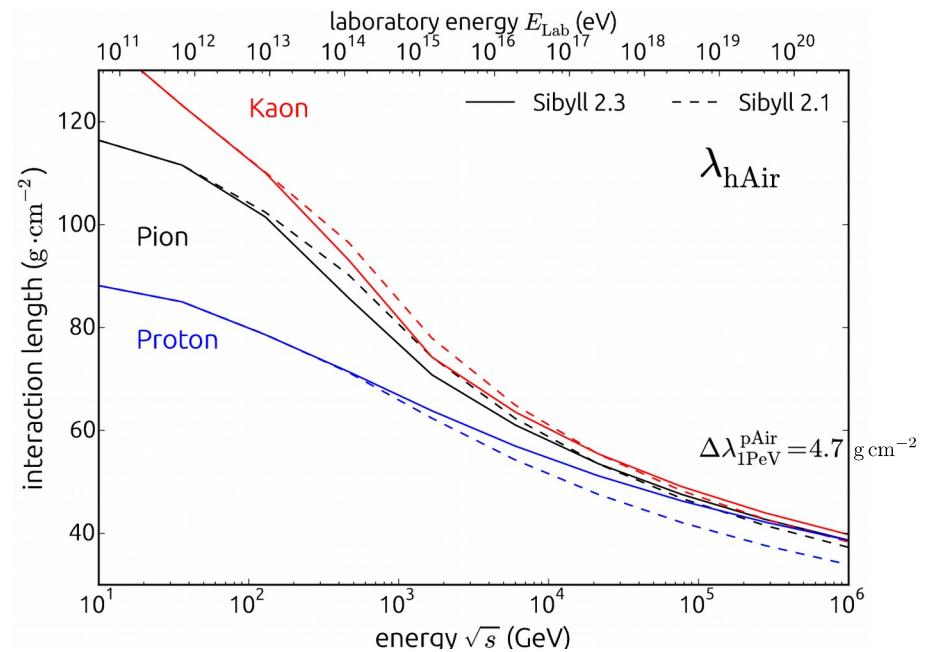
# Xmax prediction



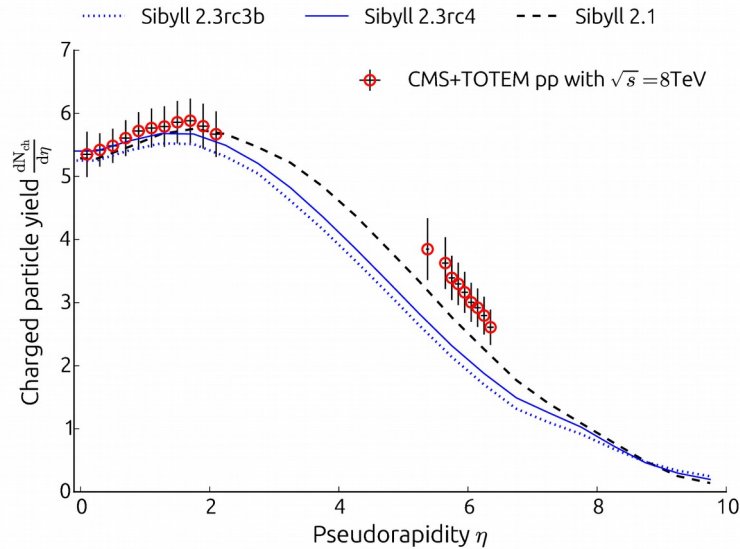
# Influences on Xmax



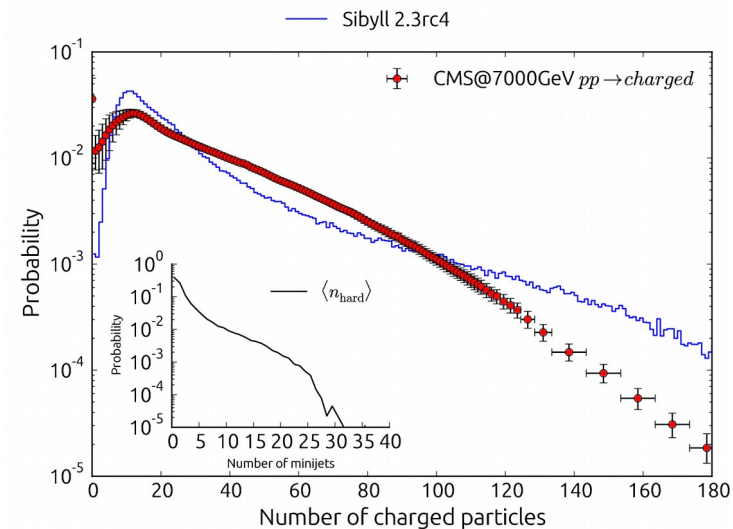
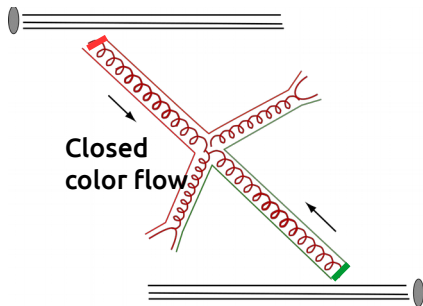
- Cross section / interaction length
- Inelasticity
- Multiplicity



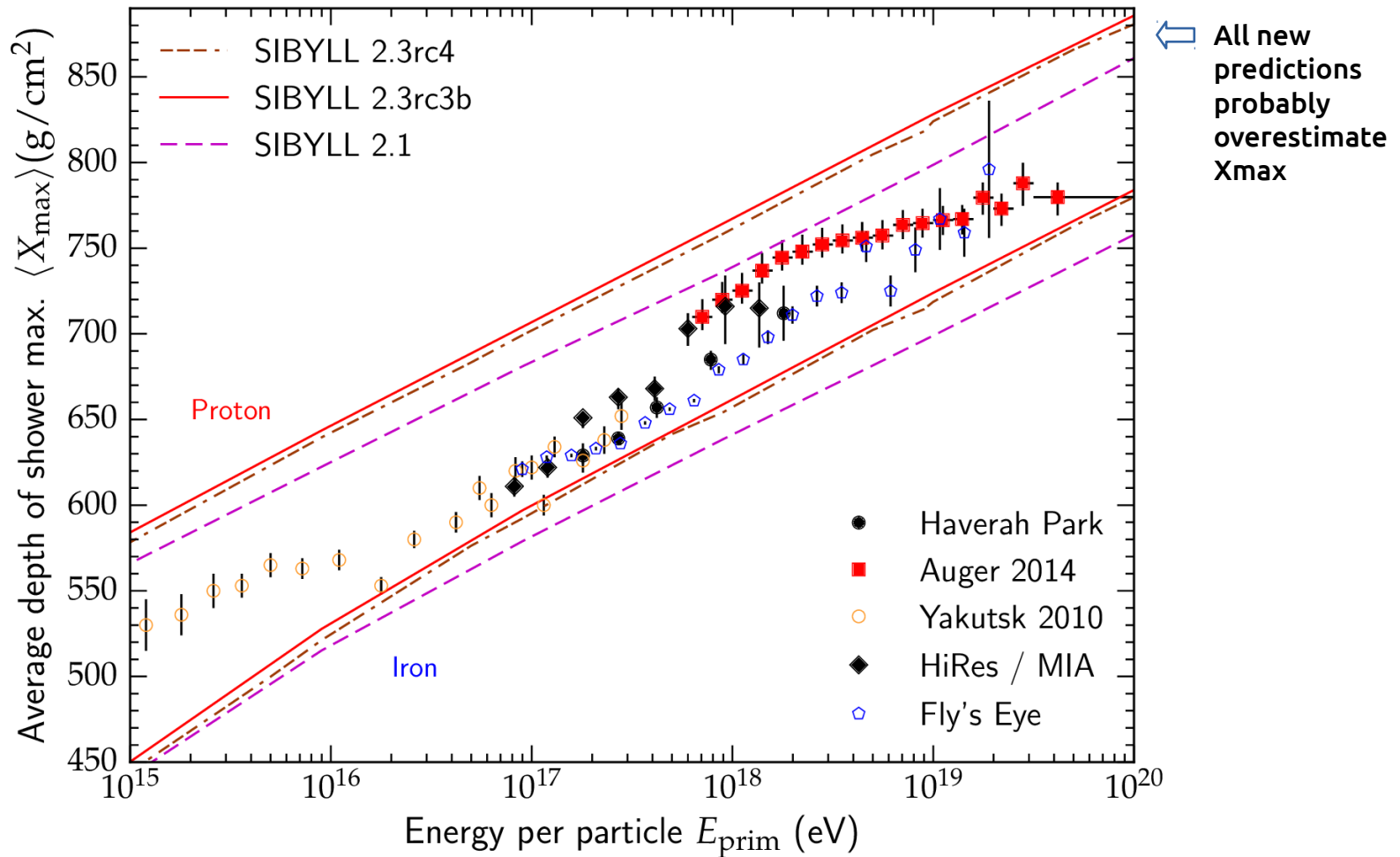
# Open problem: minijets



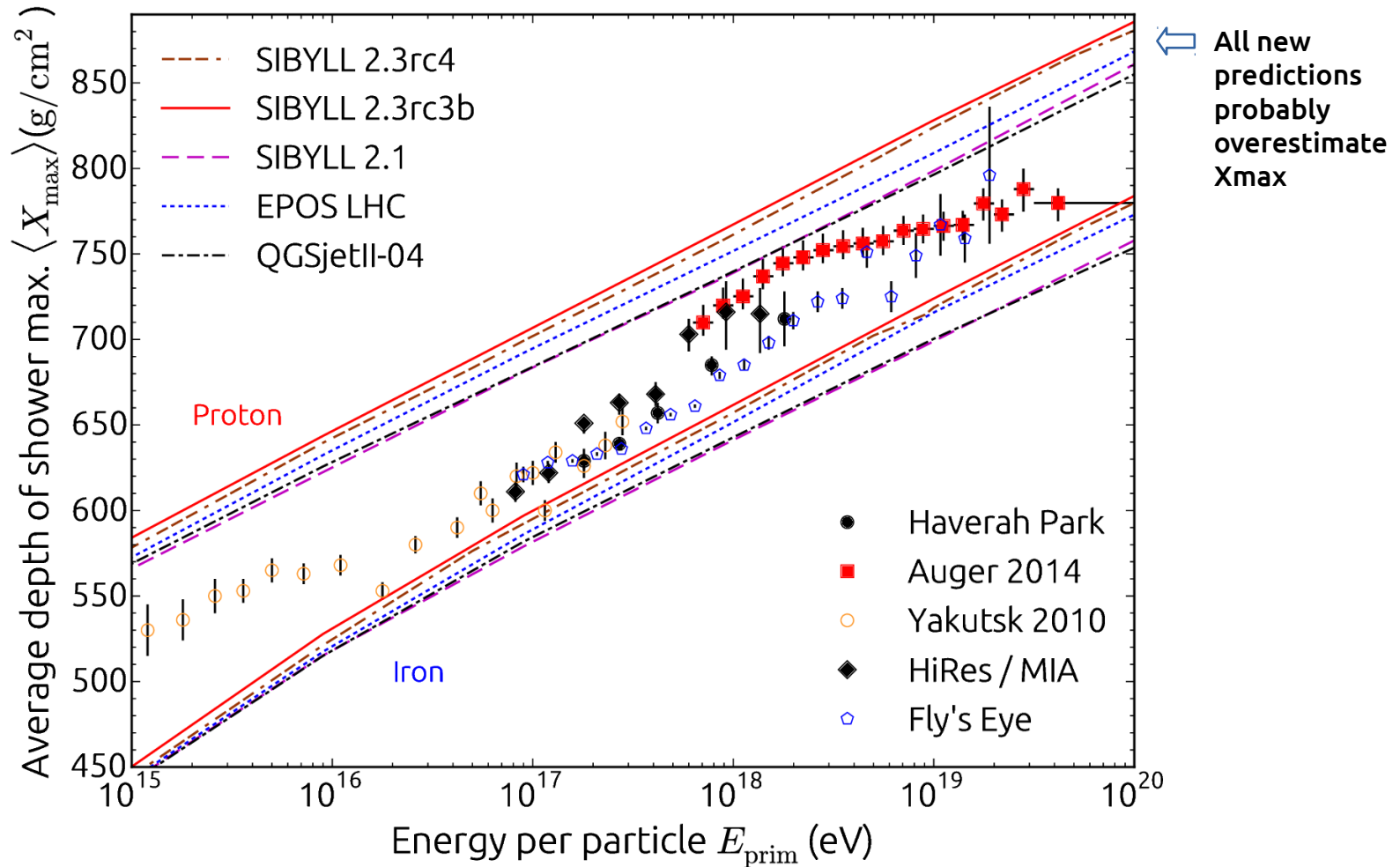
- Broad  $dN/d\eta$  in Sibyll 2.1 by accident
- Minijet color flow disconnected from rest of hadron
- Large tail in multiplicity distribution  
Number of minijets very high  
→ saturation effects missing



# Xmax prediction, Sibyll



# Xmax prediction, latest models

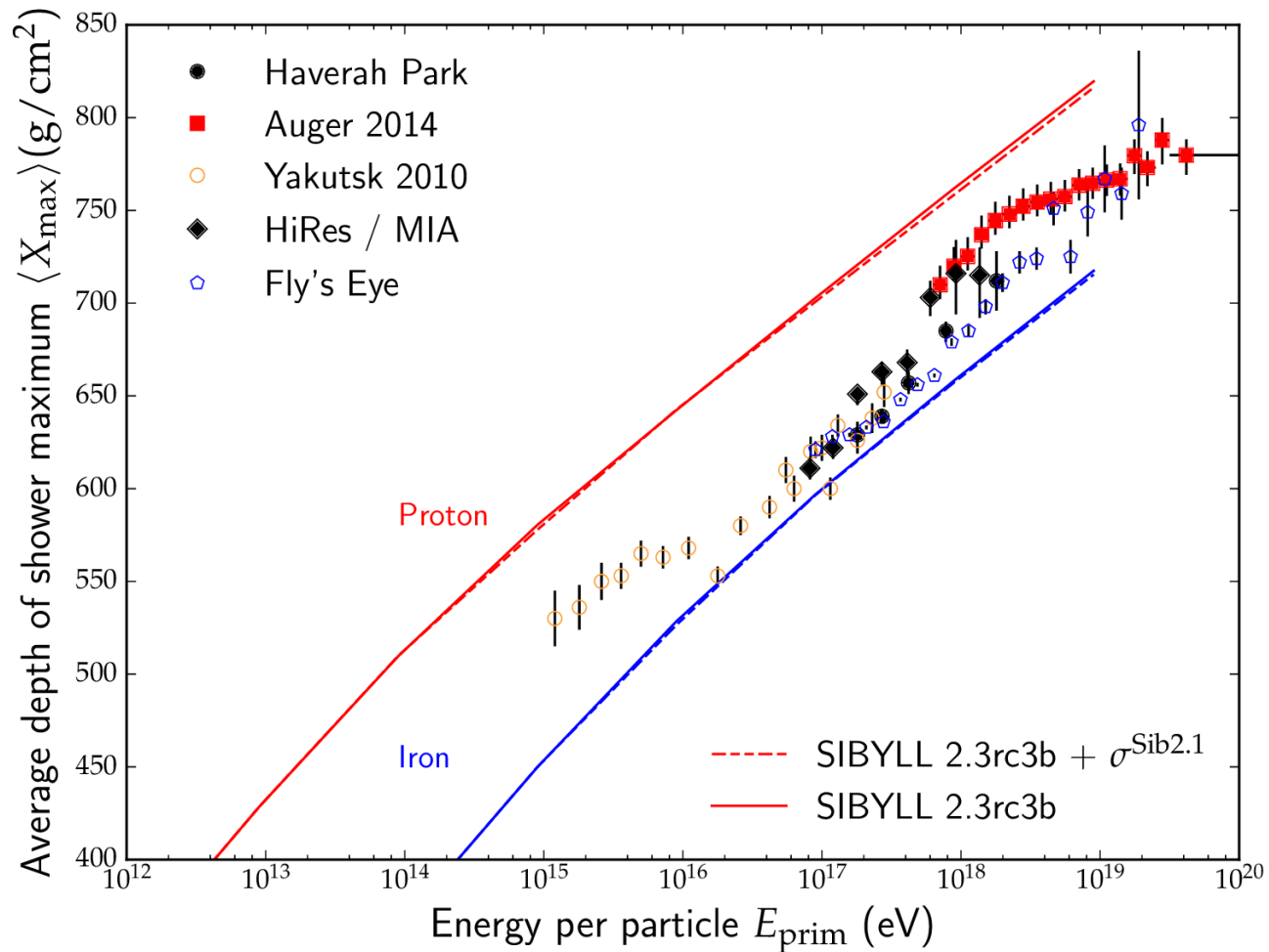


# Conclusions

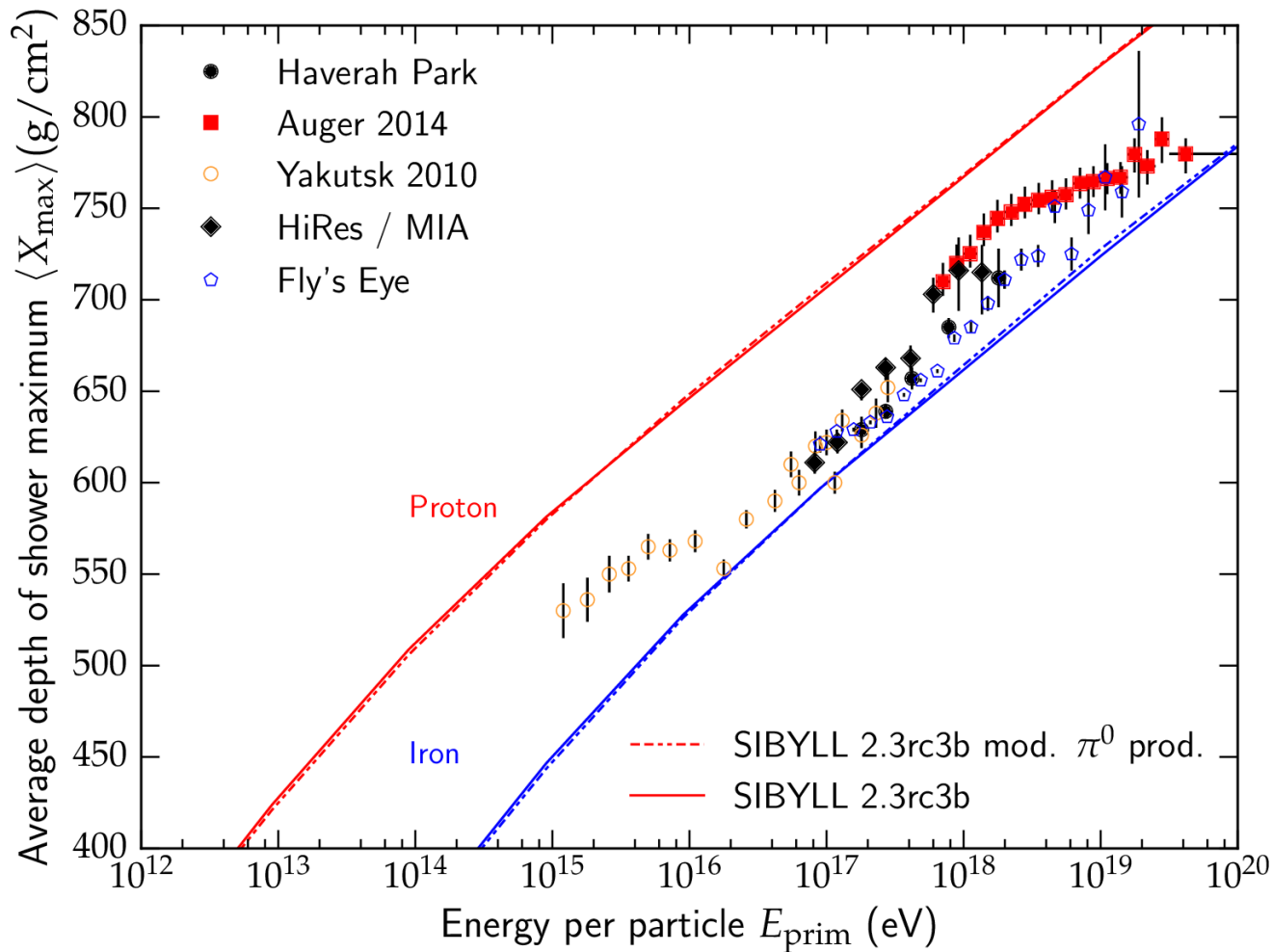
- **Model improved in many aspects**
  - Low & high energy particle production
  - Leading particles (LHCf neutrons & pions)
  - Charm production included
  
- **Open problems:**
  - Large tail in multiplicity distribution, incomplete saturation (?)
  - Narrow pseudorapidity distribution, disconnected minijets (?)
  
- **Predictions:**
  - Increased number of muons at ground (low and very high energy)
  - Much deeper  $X_{\max}$ , but inelasticity suggests  $X_{\max}$  probably lower
  - Consistency ( $X_{\max}$ ,  $\text{RMS}(X_{\max})$ ) probably improved



# Role of the p-p cross section



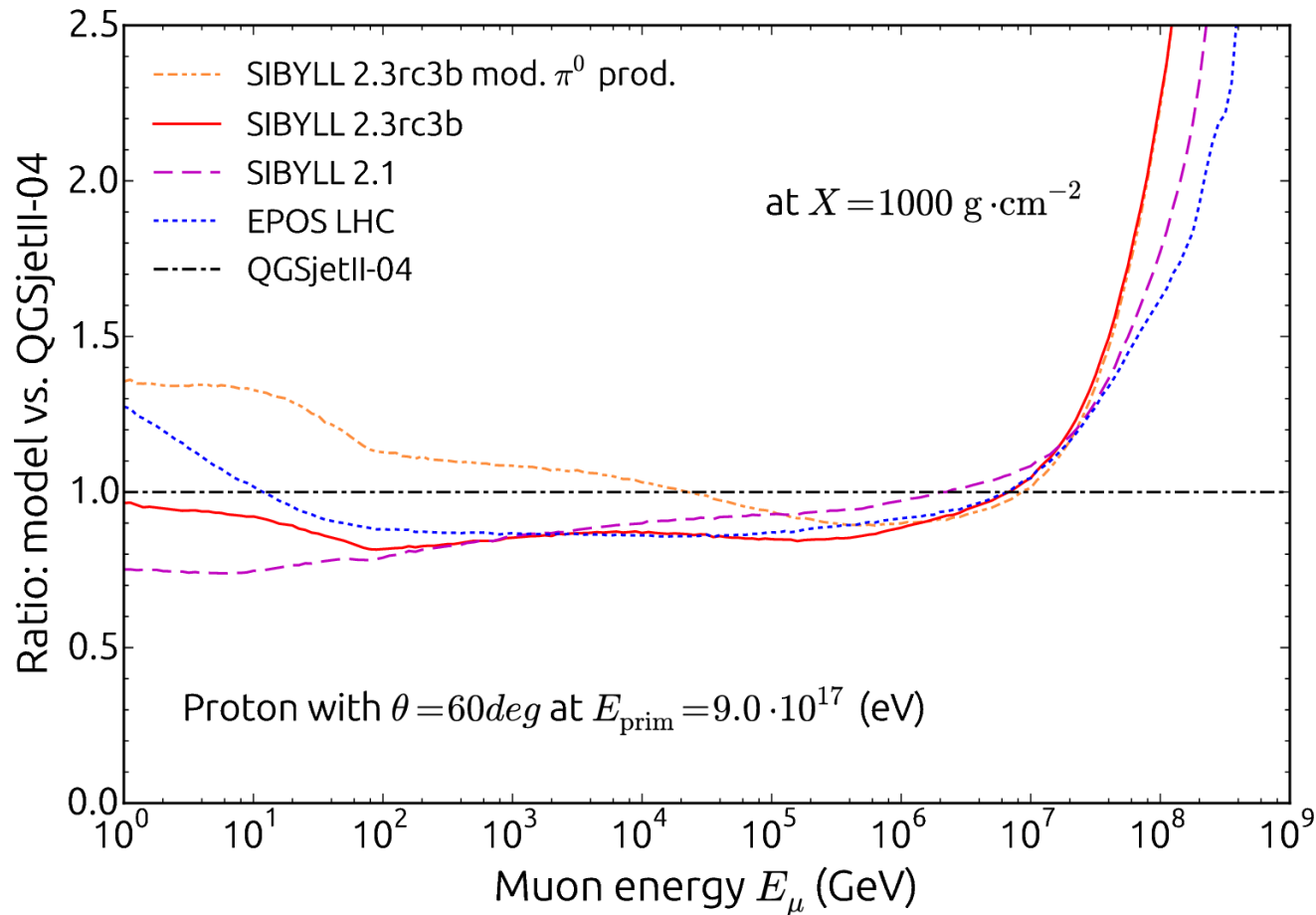
# Role of neutral pion production: $X_{\max}$



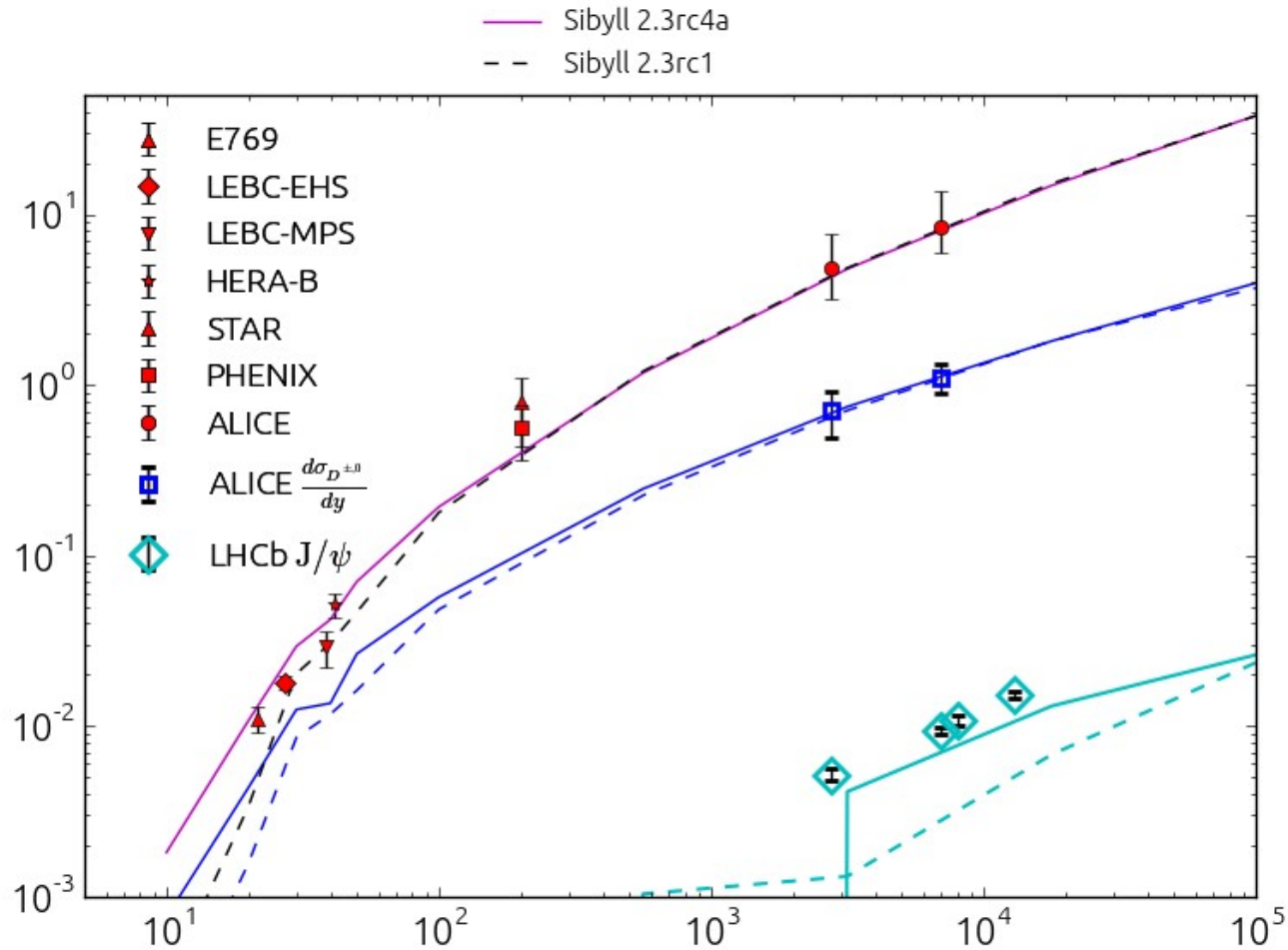
No influence on  $X_{\max}$ , probably because meson interactions are important in late stages when the bulk em cascade was formed



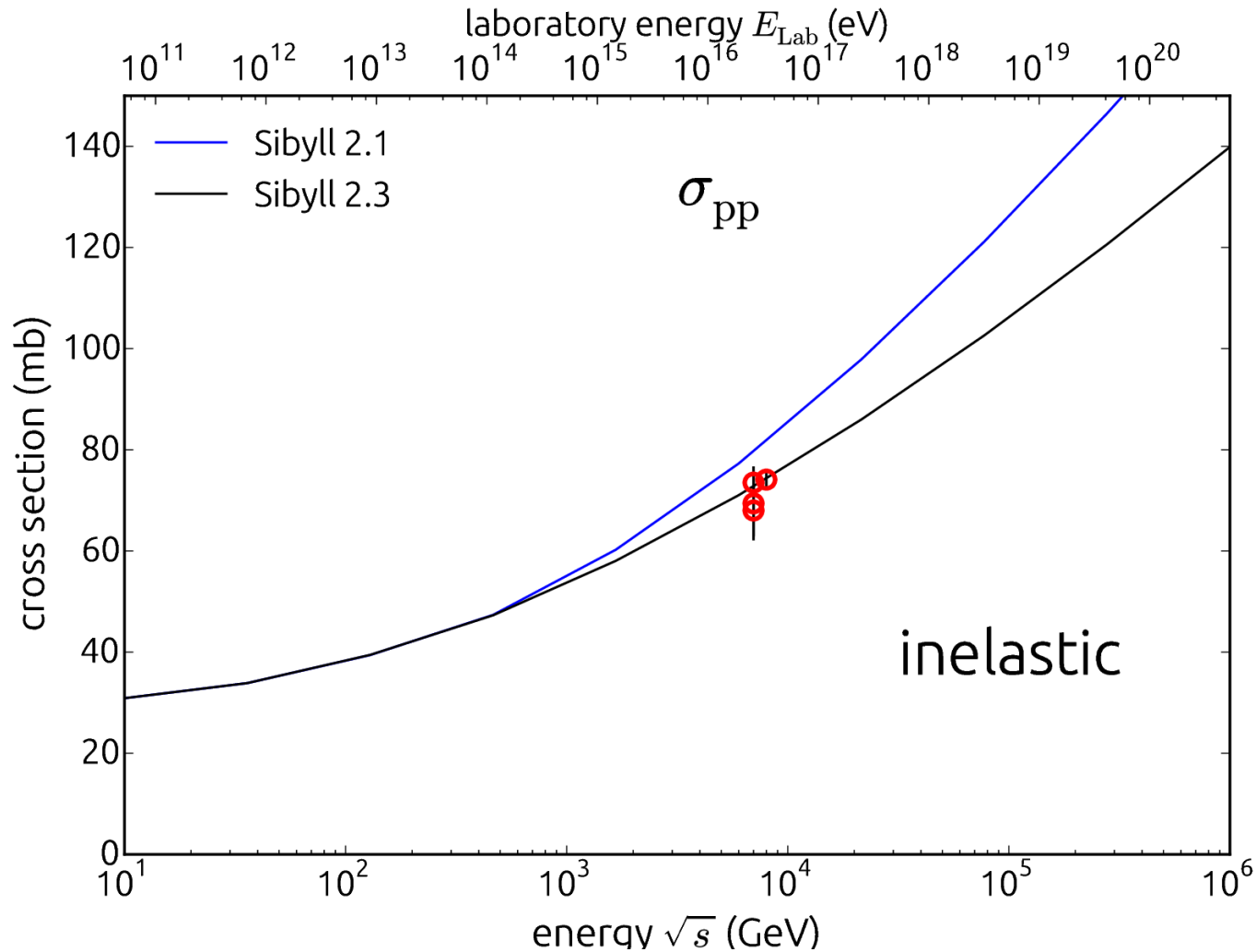
# Sibyll evolution: muon energy spectrum



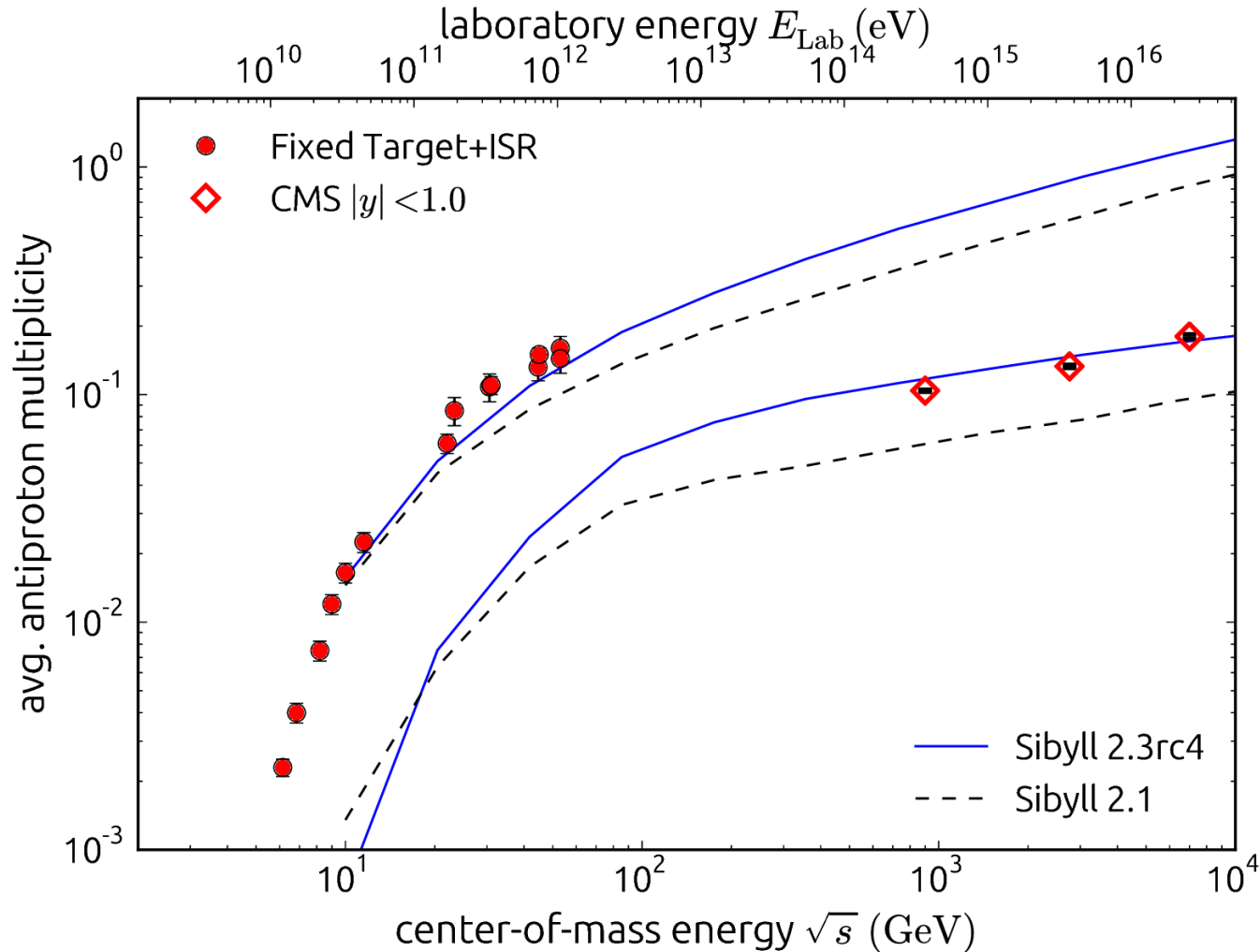
# Sibyll evolution: charm production



# Sibyll evolution: cross section



# Sibyll evolution: baryon production



# Sibyll evolution: remnant

# Inclusive atmospheric flux predictions

# RMS( $X_{\max}$ ) prediction

